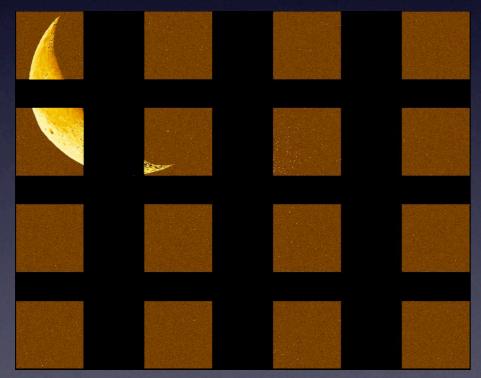
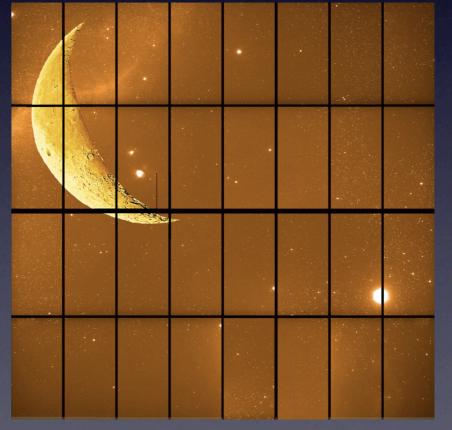
ESO Public Surveys Phase 2 Workshop



VISTA Public Surveys Simulating Phase 2



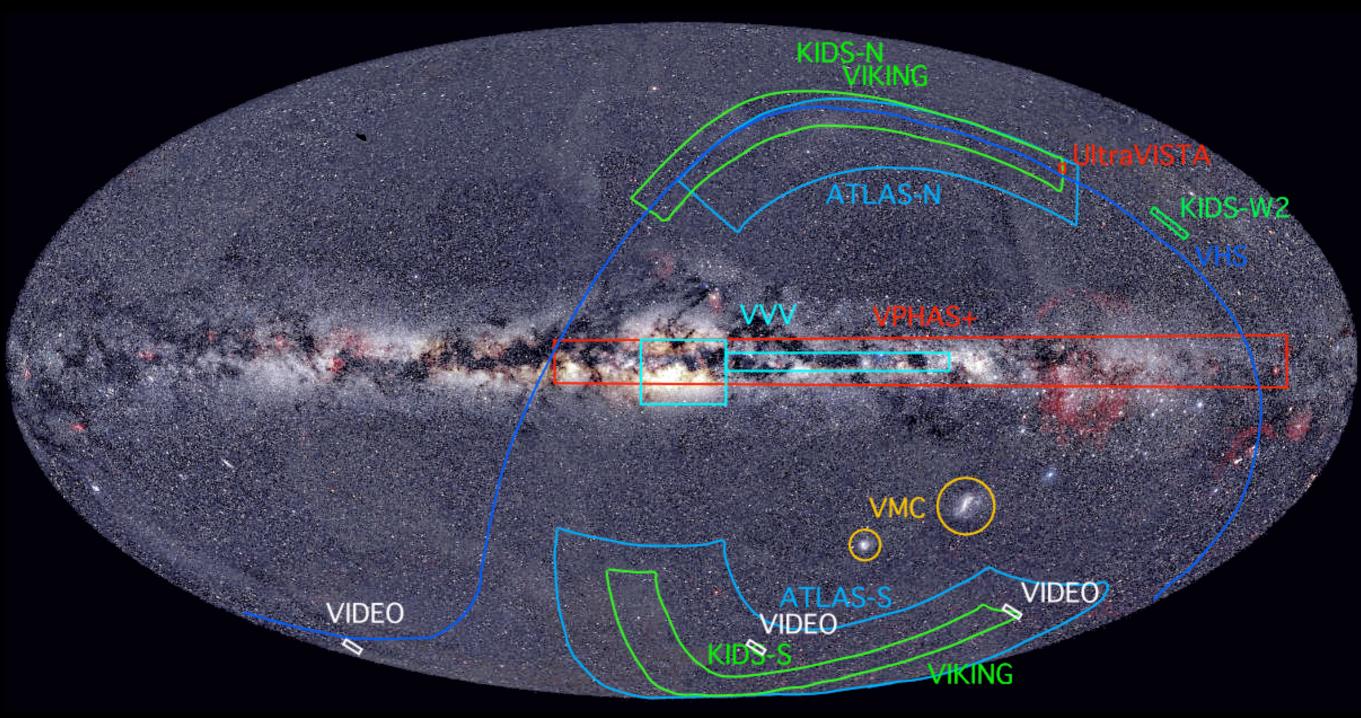


VIRCAM

Mark Neeser ESO Survey Team OmegaCAM

Summary of VISTA/VST Public Surveys





VST and VISTA will image the equivalent of more than $1^{1}/_{2} \times full$ southern sky each year

Survey Scheduling Project (mock phase 2)



<u>ESO Public Surveys</u> = 9 (6 VISTA / 3 VST) large Public Surveys covering a significant fraction of the southern sky in as many as 9 bands over 5 years.

- ~15 800 OB's/year
- => a significant challenge to scheduling logistics

Use EST familiarity with the SMP to model the first two semesters of VISTA operations. These OB's are the input for the VISTA scheduling simulation (afternoon talk by V. Ivanov)

Simulating phase two VISTA public surveys:

- VIKING
- VHS
- VVV

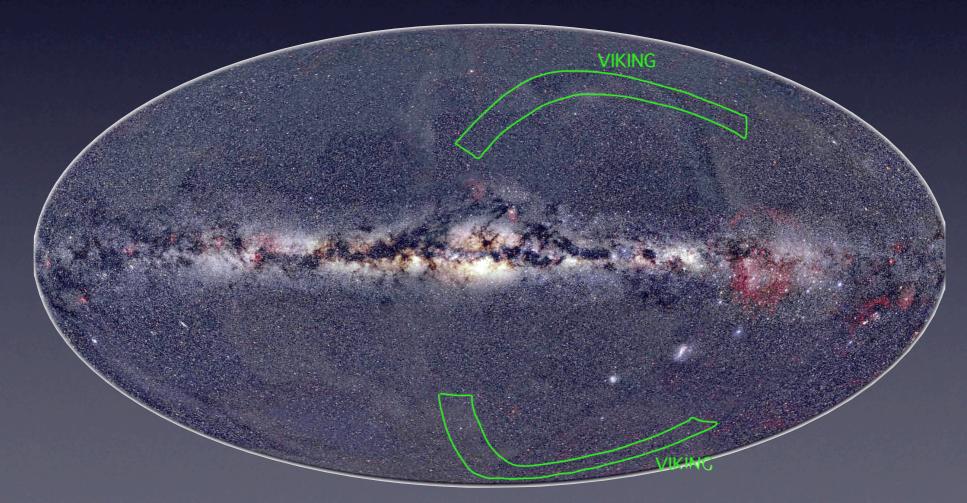
VIKING: VISTA Kilo-Degree INfrared Galaxy Survey (PI: W. Sutherland / Cambridge)



Science Case:

- key complement to VST/KIDS (9 bands => very accurate photo z's: $\Delta z/(1+z) \sim 0.03$ to z = 1)
- with KIDS: weak lensing and baryonic oscillations detection of very high redshift QSO's (z>7), ultra-cool brown dwarfs, galaxy evolution and morphology, and clustering

1490 deg² covering all KIDS fields in: Z, Y, J, H, Ks



Constraints:

- seeing <= 1.0 arcsec & THN
- | Tile => 2 OB's:

OB₁: Z,Y,J (dark or grey)

- OB₂: J₂, H, Ks (any moon)
- OB_1 and OB_2 should be observed within one month of each other (addressed using p2pp groups)

Strategy:

locate field centers along *tramlines* of constant declination (matching those of KIDS)
 => 2 arcmin overlaps (N/S); I arcmin overlaps (E/W)

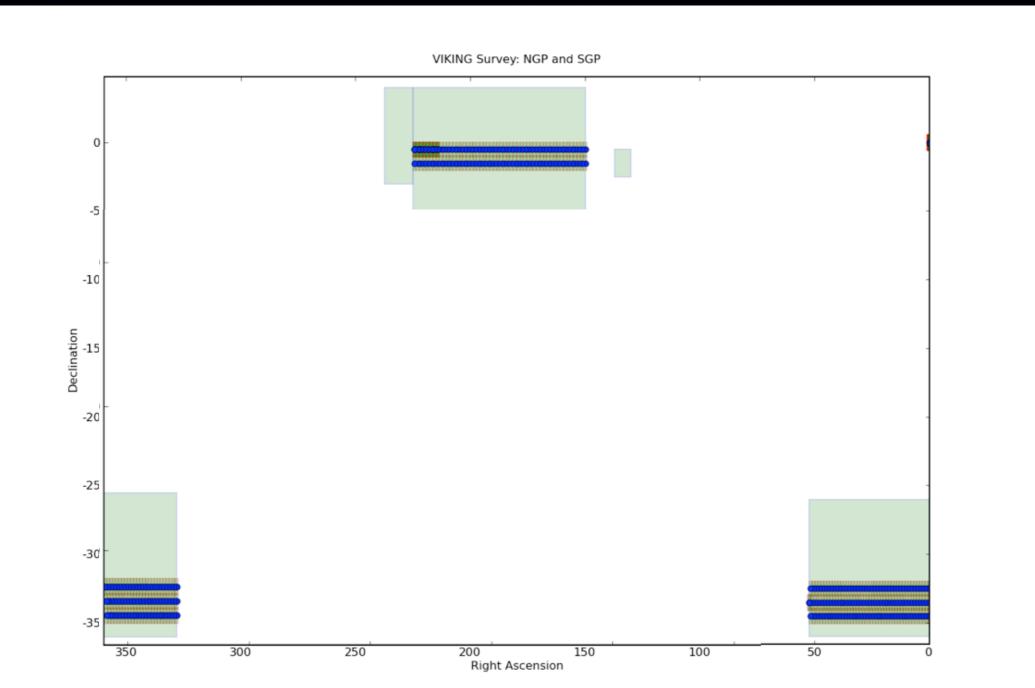
=> 49 x 10 tiles (SGP) and 61 x 9 tiles (NGP) x 1.435 deg² per tile = 1490 deg²

filter	Ζ	Y	J	Н	Ks	
depth (vega 5 sigma)	22.6	21.7 21.3		20.2	19.4	
DIT	50	50	25	10	10	
$nDIT \times nJitter \times nPAW$	Ix5x6	Ix4x6	2x2x6	5x3x6	6x4x6	
Total Exp (sec)	1500	1200	600	900	1440	



Priority Fields:

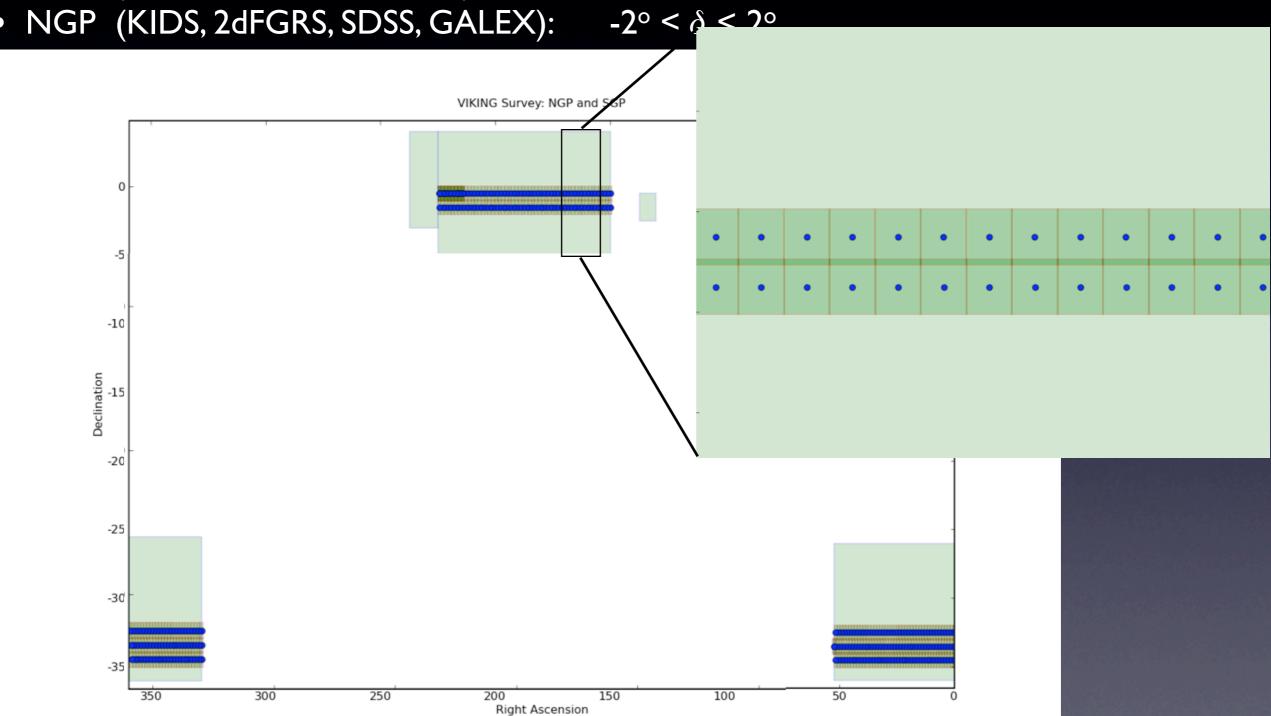
- 600 deg² subset with maximal overlap with ongoing surveys
- SGP (KIDS, 2dFGRS, SPT, DES): $-30^{\circ} < \delta < -36^{\circ}$
- NGP (KIDS, 2dFGRS, SDSS, GALEX): $-2^{\circ} < \delta < 2^{\circ}$



+ES+ 0 +

Priority Fields:

- 600 deg² subset with maximal overlap with ongoing surveys lacksquare
- SGP (KIDS, 2dFGRS, SPT, DES): -30° < *δ* < -36° ightarrow
- NGP (KIDS, 2dFGRS, SDSS, GALEX): ightarrow







Compromises:

- fixed declination tiling needed to be coded by hand and fed to SADT as field centers
- restricted to one OB per filter (workshop version of p2pp now supports this)
- $\bullet~N_{jitter}$ and N_{dit} as specified in the SMP not yet supported by p2pp
 - => (assumed DIT = total execution time)

Results:

• 1866 OB's covering 454 deg²

Observational Areas:

- VHS ATLAS (5000 deg²): divided ~evenly between N & S Galactic caps 60 sec. exposures in Y, J, H, and Ks
- VHS Dark Energy Survey (4500 deg²): SGC
 I 20 sec. exposures in J, H, and Ks
- VHS GPS (8200 deg²): Galactic Plane Survey (excl. VVV)
 60 sec. exposures in J, and Ks

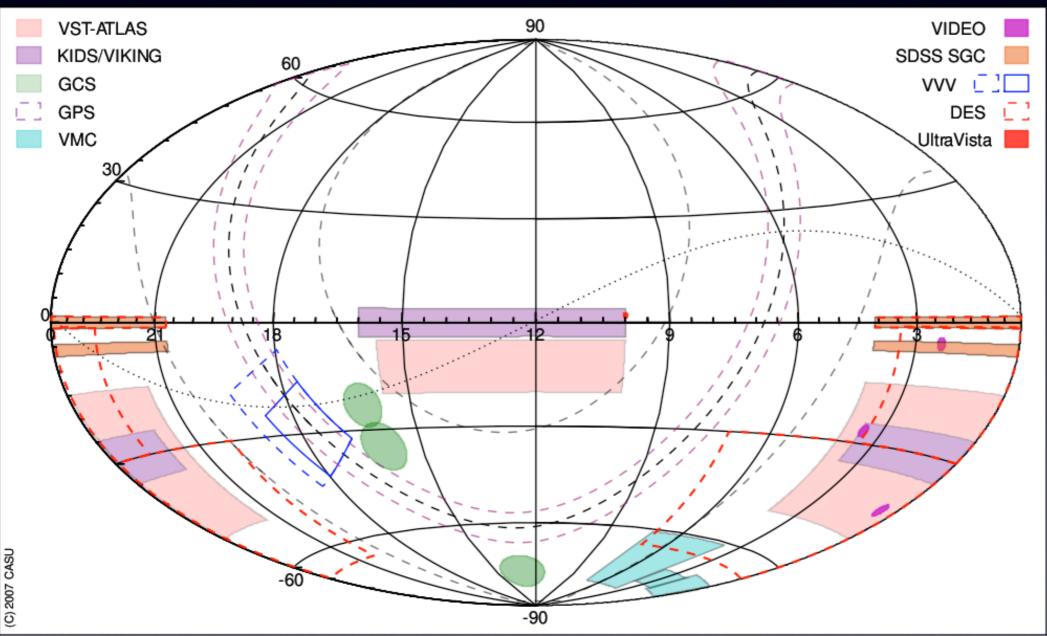




Observational Areas:

VHS - ATLAS (5000 deg²): divided ~evenly between N & S Galactic caps 60 sec. exposures in Y, J, H, and Ks

VHS - Dark Energy Survey (4500 deg²): SGC 120 sec. exposures in J, H, and Ks

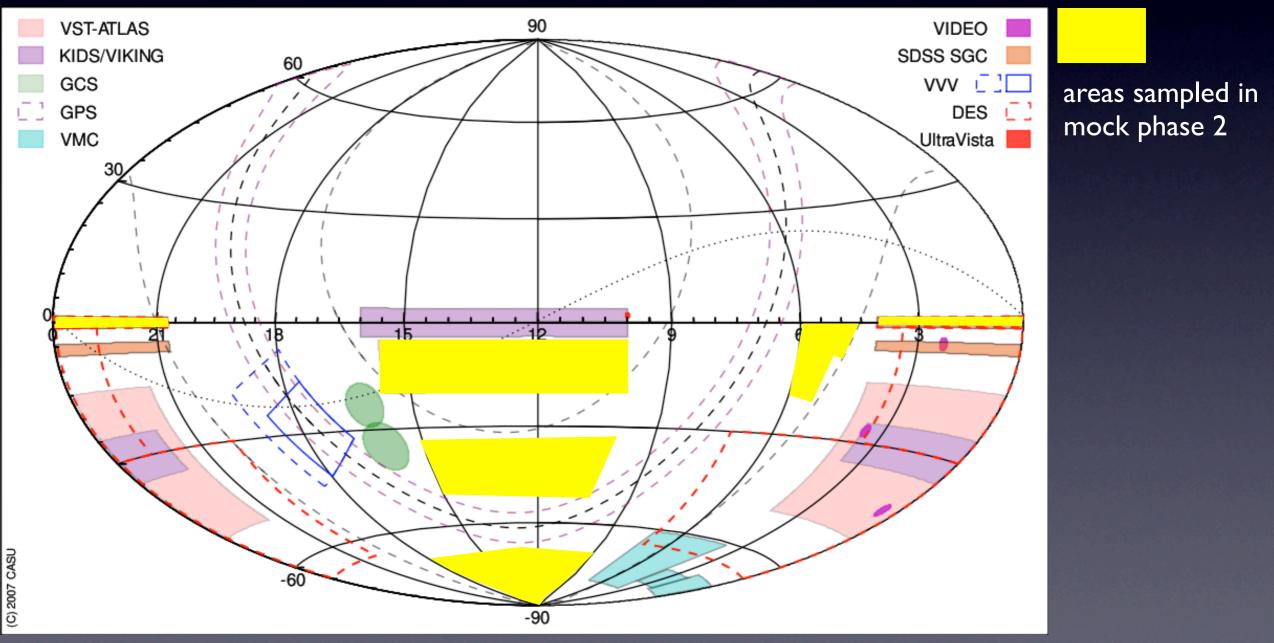




Observational Areas:

VHS - ATLAS (5000 deg²): divided ~evenly between N & S Galactic caps 60 sec. exposures in Y, J, H, and Ks

VHS - Dark Energy Survey (4500 deg²): SGC
 I20 sec. exposures in J, H, and Ks

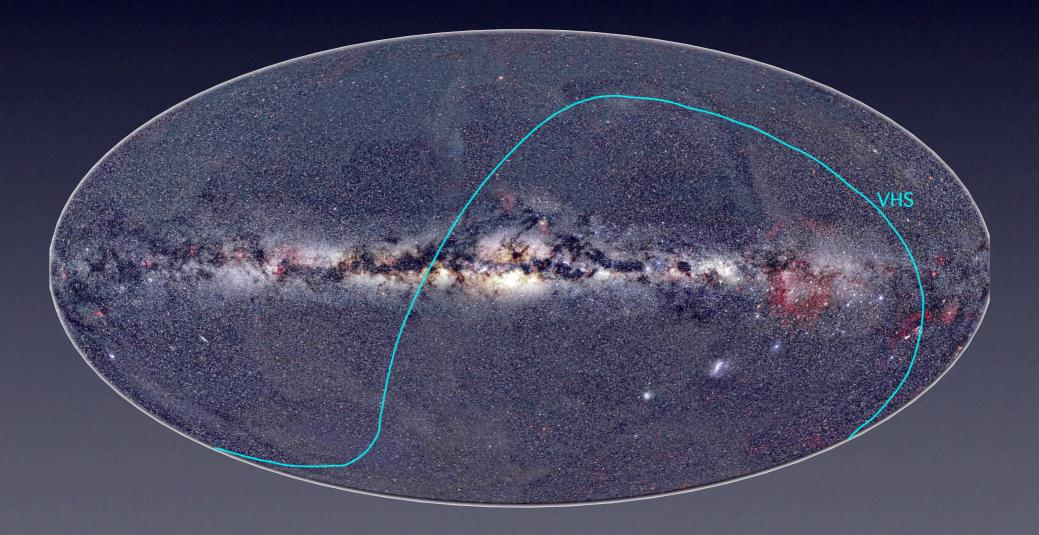


VHS: VISTA Hemisphere Survey (PI: R. McMahon / Cambridge)



Science Case:

- catalogue of low-mass and nearby stars and merger history of our Galaxy
- properties of dark energy via the large-scale structure to $z \sim I$
- search for extreme redshift quasars (z > 7)



Constraints:

- avoid observing in **both** worst seeing and in best conditions: THN: seeing <= 1.2 and CLR: seeing <= 1.4 arcsec
- J, H, and Ks (any moon); Y and J for ATLAS (grey)

Strategy:

- create OB's over 7000 deg² covering a wide range of RA and DEC
- Y (and/or J) and H and Ks tiles concatenated to ensure good sky subtraction
- time-linked so that all filters are observed within one month

Results:

• 4830 tiles covering: ATLAS (2800 deg²); DES (650 deg²); and GPS (3550 deg²)



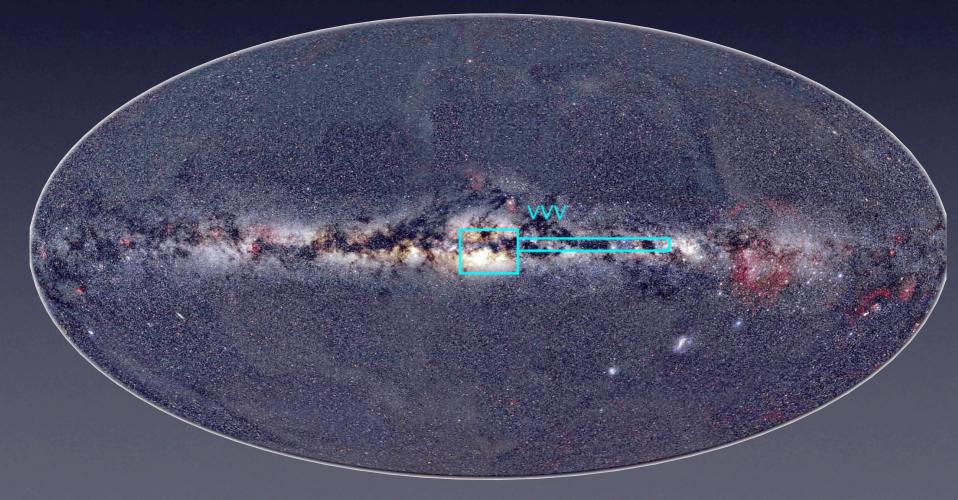
VVV: VISTA Variables in Via Lactea (PI: D. Minniti / Univ. Catolica)



Science Case:

- multi-epoch survey: Catalogue variable objects + high proper motion stars
- create a 3-d map of the Galactic bulge (RR Lyrae stars)
- cluster evolution (350 open and 33 globular clusters)
- ages of stellar populations and steller IMF

520 deg² of Galactic bulge and adjacent plane



VVV



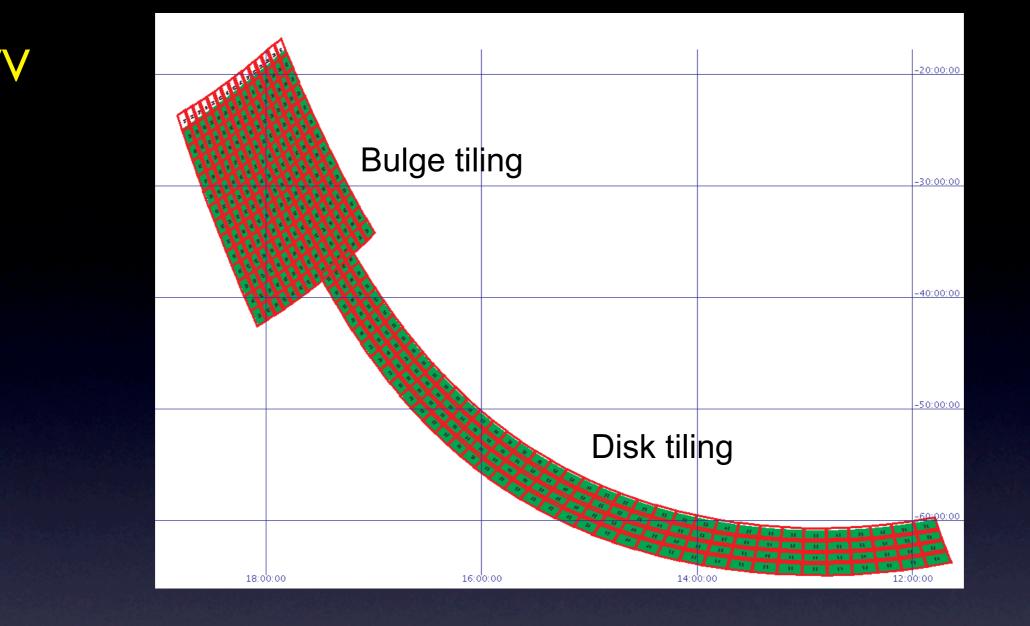
Constraints:

 seeing <= 0.8 arcsec / CLR / any moon applicable for both bulge and disk

Strategy:

 mock first year covers entire bulge/disk area (520 deg²) in: all 5 filters (Z, Y, J, H, and K_s) => concatenation then when complete repeat tiling in K_s at 5 further epochs with a 3 day time interval between successive epochs => time link

		Ζ	Y	J	Η	Ks
depth:	bulge	21.6	20.9	20.6	19.0	18.0
	disk	21.5	20.7	20.2	19.3	18.3
DIT:	bulge	10	10	6	4	4
	disk	20	20	10	10	10
$nDIT \times nJitter \times nPAW:$	bulge	Ix2x6	Ix2x6	2x2x6	Ix2x6	Ix2x6
	disk	Ix2x6	Ix2x6	2x2x6	2x2x6	2x2x6
Total Exp. (sec):	bulge	120	120	44	48	48 x 6 epochs
	disk	240	240	240	240	48 x 6 epochs





<u>Compromises</u>:

• needed to give *concatenation* a higher priority over *time-links* for constraints to work

Results:

• 5720 OB's covering 520 deg²

Summary of Public Surveys

Survey synergy	Area (deg ²)	Filters and limits (5 σ AB mag.)										
		u'	g	Hα	r'	;,	Ζ	Y	J	Н	Ks	NB
I. KIDS ^{2,8}	1500+	24.8	25.4	-	25.2	24.2				-	-	
2.ATLAS ^{1,5}	4500	22.7	22.9	-	22.9	22.0	21.2	-	_	-	-	-
3.VPHAS+ ⁷	1800	22.5	23.2	23.2	23.2	22.5	-	-	-	-	-	-
4. Ultra-VISTA	0.73	-	-	-	-	-	-	26.7	26.6	26.1	25.6	24.1
5.VHS 1,2,8	20000	-	-	-	-	-	-	21.2	21.1	20.6	20.0	-
6.VIDEO	12	-	-	-	-	-	25.7	24.6	24.5	24.0	23.5	-
7.VVV 3	520	-	-	-	-	-	22.5	21.9	21.1	19.6	20.0	-
8.VIKING ^I	1500	-	-	-	-	-	23.1	22.3	22.1	21.5	21.2	-
9.VMC	184	-	-	-	-	-	-	23.3	23.0	-	22.9	-

