



Commissioning Status

Jim Emerson Queen Mary, University of London

Telescope & Camera

- detector array modules (infrared and CCD)
- filter barrel
- Iens barrel
- baffle tube
- pressure window
- cryostat vessel
- electronics rack
- telescope structure and mirrors







Quick Summary

- VISTA is currently undergoing Engineering testing and Commissioning.
- All subsystems are functional and integrated, however their performance and interaction are not yet fully optimised or verified.
- The system has demonstrated fully "end-to-end integrated" functionality:
- Overall integrated control of the Telescope, Enclosure and Infra-Red Camera by the Telescope Control System
- Closed-loop control of the Primary and Secondary mirrors by the active optics system.
- Closed-loop auto-guiding of the Telescope axes.

Ongoing Activities

Ongoing Engineering and Commissioning activities include:

- Rigorous characterisation and calibration of the wavefront sensing and active optics system
- Refinement of the Telescope pointing model to improve pointing, acquisition and tracking.
- System-level shake-down of the Telescope/Enclosure/ Camera/control system.
- Refining the automatic location and acquisition of guidestars and aO reference stars.
- Optimisation of low-level Primary mirror support CANbus operation to improve performance.
- Investigation and characterisation of Secondary-mirrorrelated trefoil phenomenon
- Flat-field testing (sky-flat testing and Enclosure flat-field testing) to characterise and refine the system performance.
- Refinement of Template-driven observing and tiling.





Sub-Systems Status (31/08)

item	Problem s	Concerns	TBD
Ground Works	0	0	Cosmetic & Tidying up
Civil Works	0	0	0
Enclosure	0	Encoder wheel	Dome 'Flat' (linearity) screen
Plant & Services	0	0	0
Control hard & soft	0	0	Tuning software
Computing system	0	0	Full rate, Archive
Equipment	0	0	0
Safety Equipment	0	0	Alarm link to VLT
Spares	0	0	0
Documentation	0	0	Complete Verification
Coating Plant	0	Magnetron arcing, Washing	0
Telescope	0	Reflectivity, Trefoil	M1 Support characterisation
Camera	0	0	SIQ, Software (AG / aO), System
Data Flow System	0	TBD	beyond VISTA



Movement Overheads (so far)_

<u>Very</u> preliminary values

Overhead	Time (sec)	Comment
Preset	variable	Depends on $\triangle Az \ \Delta Alt \ \Delta Rot at that ST Acc'n 0.5 °/sec max Speed 2°/sec$
Guide Star Acquisition	~ 3 s	
Active Optics update	0 s / 30s	0 in parallel with observation 30 by itself
Pawprint movement	~ 5s	New guide and aO stars at each pawprint
Jitter movement	~2.5s	No change of guide/aO stars aO takes 30 sec
Microstep movement	tbd	No change of guide/aO stars aO takes 30sec



Filter change overheads (so

<u>Very</u> preliminary values

moved	seconds	From/to	atus	Reference po	sition (+ve clockwise)
1	20/21	J⇔Ks Ks⇔H H<=>Z' Z'⇔Y	blankG 8: DARK1	2:	blankA NB118
2	26/27	J⇔H Ks<=>Z' H⇔Y	blankF 7: Y	Filters Installed + (science filters framed)	blankB 3: J
3	TBD	J⇔Z' Ks⇔Y	blankE 6: DARK2	H0Jeyny	HOJcony HOJoany
4	40	J⇔Y	blankD		blankC
Digital I/O simula	ation Normal	F (ignored)		5: H	





Integration Overheads (so far)_

<u>Very</u> preliminary values

	Time (sec)		
T_res – Reset	1 sec		
Read	0		
T_red- Read (end)	1 sec		
T_pro + T_wri Process + Write	~12 sec variable - determined by IRACE system		
Total	T=[Ndit * (DIT + [T_res + T_red])+ [T_pro + T_wri]]		
Total	$T = (Ndit^{*}(DIT+2)) + 12$		
N.B. ETC assumed $T_res + T_red = 1$ not 2			

ETC assumed T_pro+T_wri ~4



Parallelisation of Overheads

	Overhead	 to	Comment
1	Preset	2	Acc'n 0.5deg/s Speed 2deg/ sec
2	Filter change	1	Depends on pair
3	Guide Star Acquisition	-	
4	Active Optics update	8, 9	0 in parallel with observation 30 by itself
5	Pawprint movement	4	If exceeds 30 sec
6	Jitter movement	4	If exceeds 30 sec
7	Microstep movement	4	If exceeds 30 sec
8	Integration	4	If exceeds 30 sec
9	Read out		If exceeds 30 sec

Camera data (mostly J)

- No surprises (so far)
 - No obvious crosstalk, even from very bright objects. Expect that correcting for crosstalk will be unnecessary.
 - Minimal persistence
 - Non-linearity generally 2-4% at the 10,000 ADU level.
 - astrometry: WCS well described by the ZPN projection with distortion coefficient is within ~ +/-1 of the predicted 42 radian/ radian³ distortion.
 - Systematic astrometric residuals already at <~50 mas over the whole detector array.</p>





END