

# VST science verification

## TEST DATA SET

Aniello Grado

&

Luca Limatola

---

# C1 – Test data set

***Which correction for seeing variations (field-to-field and across different bands) is applied in order to obtain consistent (point) source fluxes in the final catalogues?***

Presently no PSF homogenization is supported in VST-Tube. A user defined cut on FWHM can be applied to the exposures contributing to the final mosaic.

# NGC253 u

## Absolute photometry:

Zero point:  $22.717 \pm 0.006$  Color term:  $0.033 \pm 0.016$  Ext. coeff.:  $0.05$

## IC correction:

Statistics on:	Average	Median	STD	MAD
<i>Reference-us</i>	<i>0.0016</i>	<i>0.0031</i>	<i>0.10</i>	<i>0.09</i>
<i>Residuals</i>	<i>0.0014</i>	<i>0.0013</i>	<i>0.08</i>	<i>0.04</i>

## Relative photometry:

### All detections

Instrum mag RMS  
**0.025**

Ndets  
17952

### High S/N sources

Mag RMS  
**0.019**

Ndets  
13762

## Astrometric calibration:

High S/N Internal Astrometric Sigma	Sigma w.r.t. Astrometric Ref Catalog	Detections for statistics w.r.t. Astrometric Ref Catalog
<b>0.081''</b>	<b>0.16''</b>	<b>564</b>
<b>0.071''</b>	<b>0.13''</b>	

# NGC253 g

## Absolute photometry:

Zero point: **24.845 ±0.006** Color term: **0.021 ±0.016** Ext. coeff.: **0.040**

## IC correction:

Statistics on:	Average	Median	STD	MAD
Reference-us	0.041	0.046	<b>0.10</b>	<b>0.07</b>
Residuals	0.000130	-0.00028	<b>0.07</b>	<b>0.03</b>

## Relative photometry:

### All detections

Instrum mag RMS                      Ndets  
**0.045**                                      8880

### High S/N sources

Mag RMS                                  Ndets  
**0.02**                                      5967

## Astrometric calibration:

High S/N Internal Astrometric Sigma	Sigma w.r.t. Astrometric Ref Catalog	Detections for statistics w.r.t. Astrometric Ref Catalog
<b>0.0517''</b>	<b>0.296''</b>	<b>801</b>
<b>0.0451''</b>	<b>0.237''</b>	

# NGC253 r

## Absolute photometry:

Zero point:  $24.799 \pm 0.006$  Color term:  $0.049 \pm 0.016$  Ext. coeff.  $0.070$

## IC correction:

Statistics on:	Average	Median	STD	MAD
<i>Reference-us</i>	<i>0.037</i>	<i>0.044</i>	<i>0.09</i>	<i>0.06</i>
<i>Residuals</i>	<i>0.002</i>	<i>0.003</i>	<i>0.08</i>	<i>0.03</i>

## Relative photometry:

### All detections

Instrum mag RMS  
**0.048**

Ndets  
23542

### High S/N sources

Mag RMS  
**0.031**

Ndets  
19306

## Astrometric calibration:

High S/N Internal Astrometric Sigma	Sigma w.r.t. Astrometric Ref Catalog	Detections for statistics w.r.t. Astrometric Ref Catalog
<b>0.066''</b>	<b>0.23''</b>	1160
<b>0.062''</b>	<b>0.20''</b>	

# NGC253 i

## Absolute photometry:

Zero point:  $24.279 \pm 0.006$  Color term:  $-0.002 \pm 0.016$  Ext. coeff.:  $0.01$

## IC correction:

Statistics on:	Average	Median	STD	MAD
<i>Reference-us</i>	0.13	0.11	0.19	0.10
<i>Residuals</i>	0.02	-0.01	0.18	0.06

## Relative photometry:

### All detections

Instrum mag RMS  
**0.033**

Ndets  
**6479**

### High S/N sources

Mag RMS  
**0.021**

Ndets  
**4371**

## Astrometric calibration:

High S/N Internal Astrometric Sigma	Sigma w.r.t. Astrometric Ref Catalog	Detections for statistics w.r.t. Astrometric Ref Catalog
<b>0.056''</b>	<b>0.28''</b>	<b>827</b>
<b>0.047''</b>	<b>0.27''</b>	

# NGC253 NB\_659

## Absolute photometry:

Zero point: 21.98 Ext. coeff.: 0.09

## IC correction:

Used r band IC

## Relative photometry:

### All detections

Instrum mag RMS

**0.092**

Ndets

17564

### High S/N sources

Mag RMS

**0.092**

Ndets

17564

## Astrometric calibration:

High S/N Internal Astrometric Sigma	Sigma w.r.t. Astrometric Ref Catalog	Detections for statistics w.r.t. Astrometric Ref Catalog
<b>0.063''</b>	<b>0.22''</b>	<b>1348</b>
<b>0.058''</b>	<b>0.20''</b>	

# NGC253

## *Summary on mosaics*

Band	u'	g'	r'	i'	NB_659
exptime	29008	2100	3717	1500	5300
#exposures	28	7	21	5	13
Completeness 50%	25.8	25.2	24.9	23.8	23.6



---

# NGC253

Fractional signal variation between *OMEGA.2011-08-31T03:57:06.098.fits*  
and *OMEGA.2011-08-31T04:32:40.418.fits* (band *r'* night 2011-08-30)

$$Im = Filt[(Im_1 - Im_2 - off_{12})/Im_1]$$

*Filt* is a median average with kernel 7x7

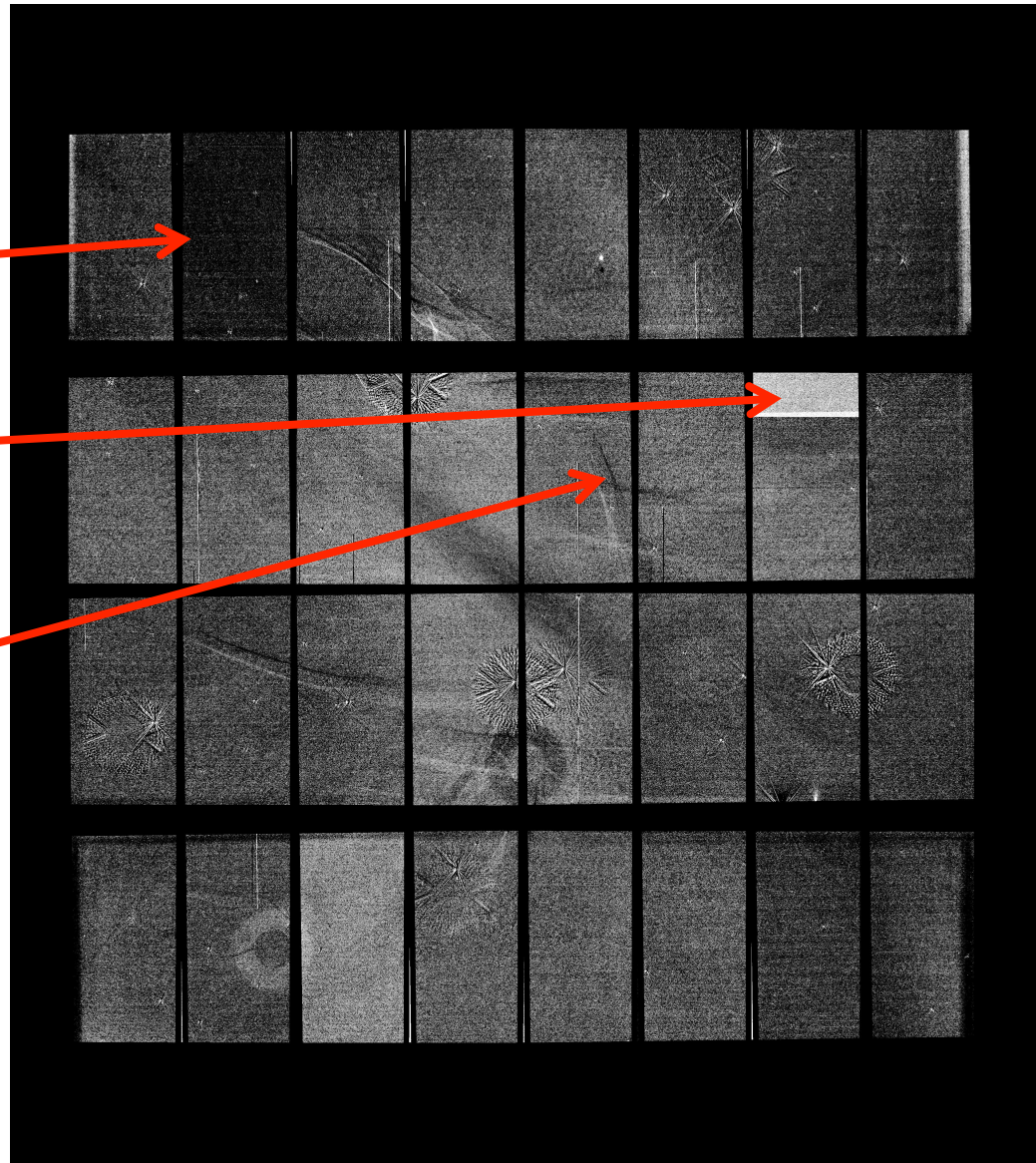
*Images* processed (flat using twilight and sky images), overscan correction row by row.)

# NGC253

*ccd95*

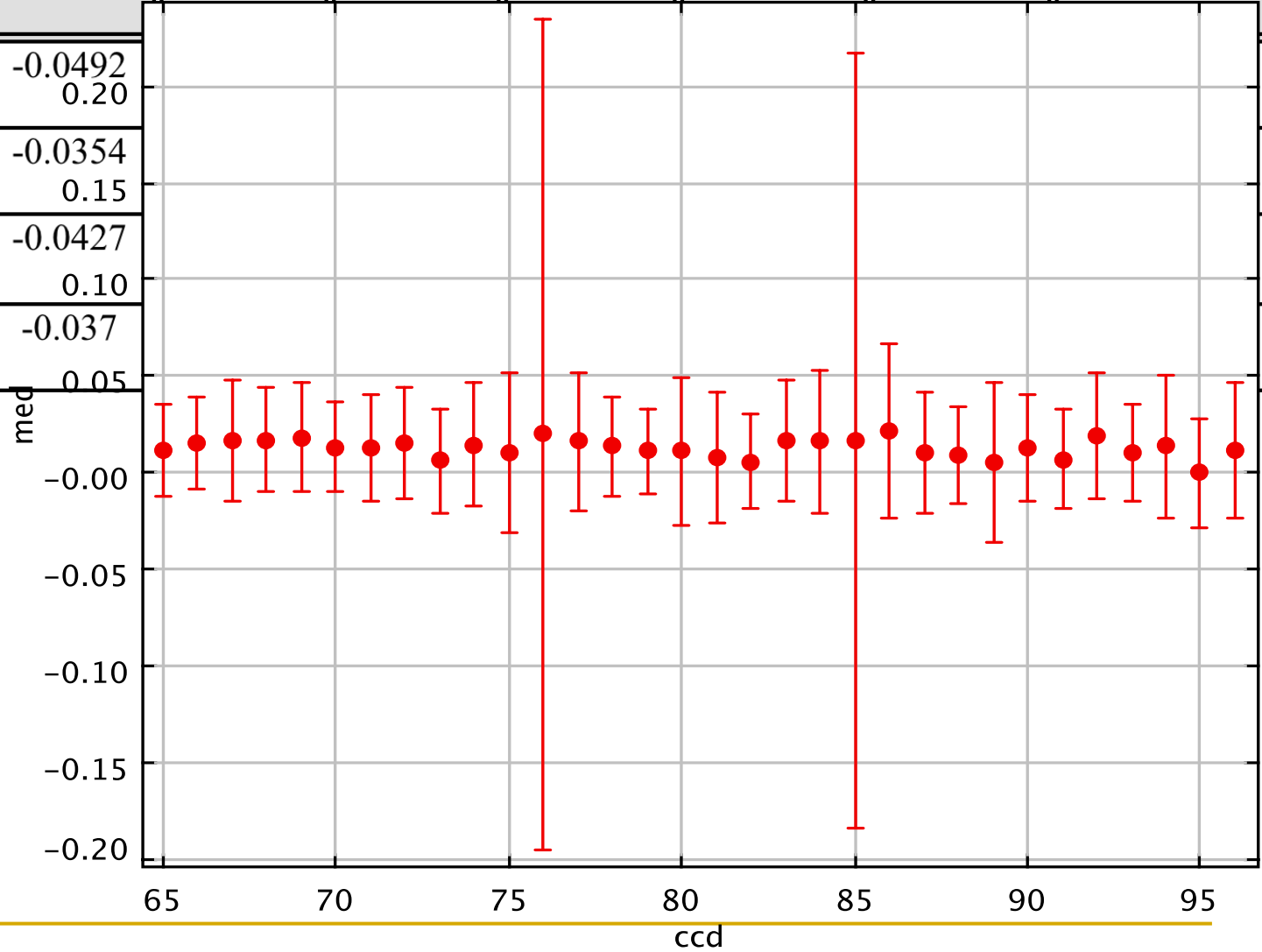
*ccd82*

*Light reflections*



# NGC253

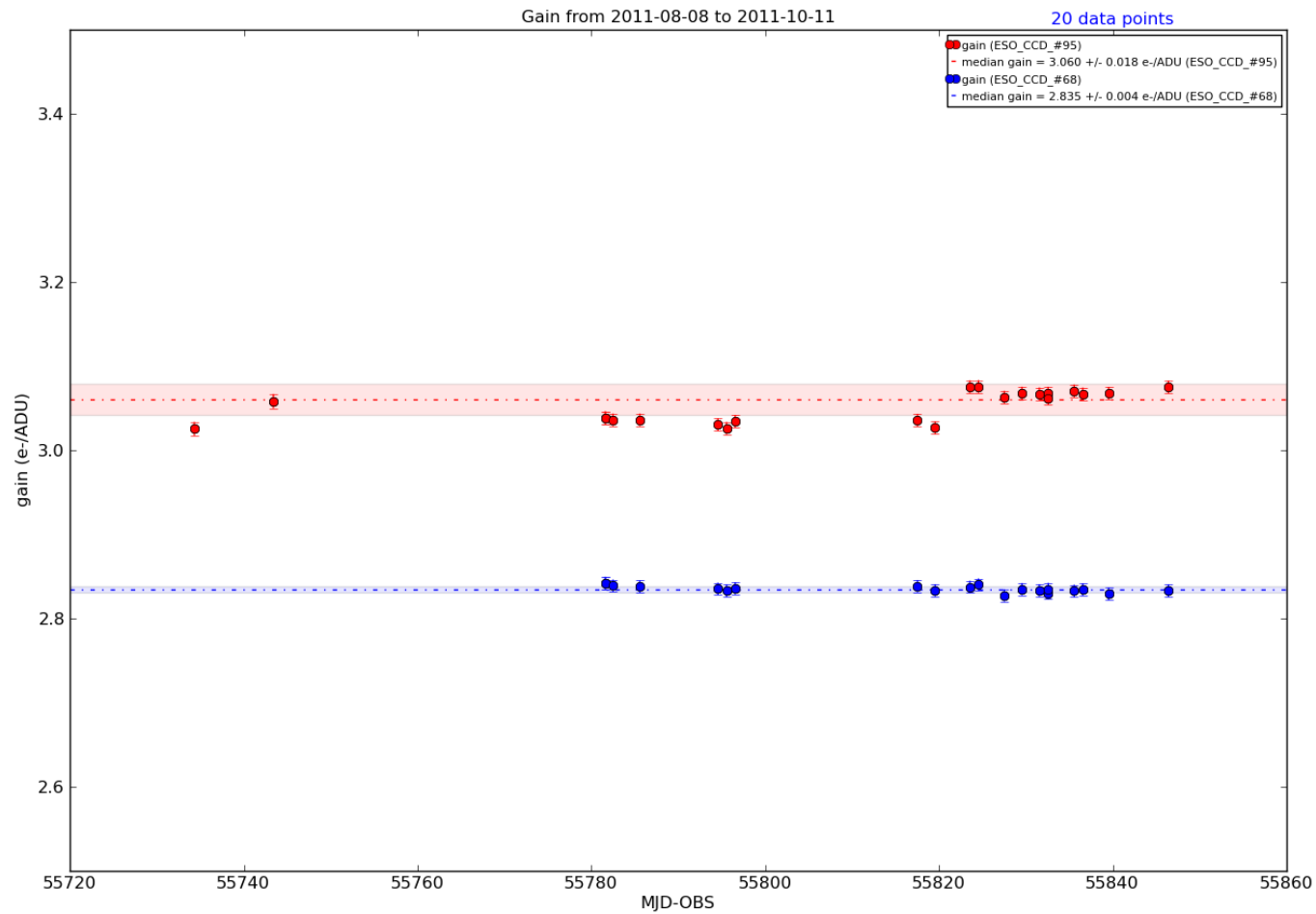
CCD	Position	Min	Max	Ave	Sigma	Skewness	Kurtosis	Median
95	x=3480 y=17760	-0.0492 0.20						
88	X=2800 Y=14600	-0.0354 0.15						
73	X=17800 Y=7000	-0.0427 0.10						
68	X=10600 Y=4500	-0.037 0.05						



# NGC253

From e-mail of Mark *Fri*, Oct 14, 2011

Time Dependency of OmegaCAM Gain





*Thanks*