A night sky filled with stars, with a faint, curved comet tail visible on the left side. In the foreground, the silhouette of a large telescope or observatory structure is visible against the dark sky. The horizon shows a faint orange glow, suggesting a sunset or sunrise.

Photometric studies of Solar System minor bodies

Colin Snodgrass

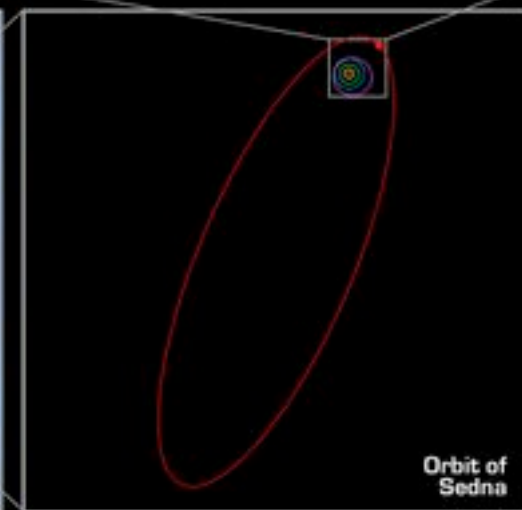
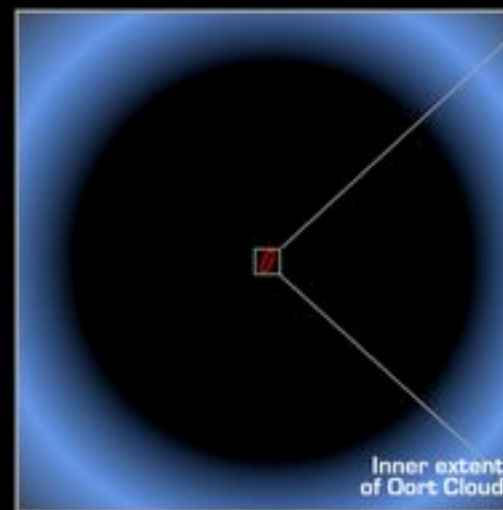
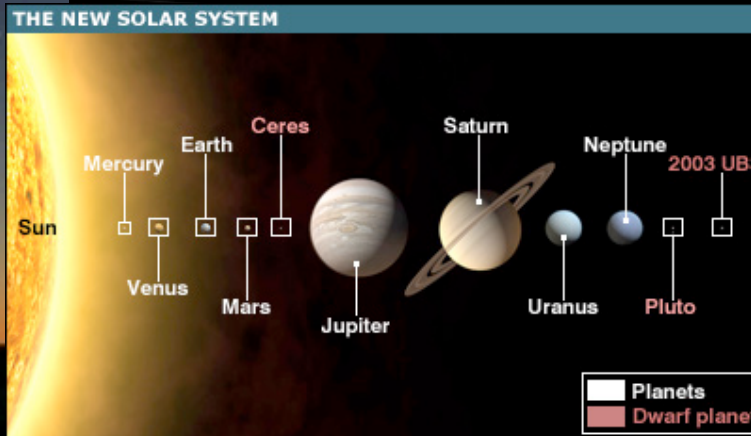
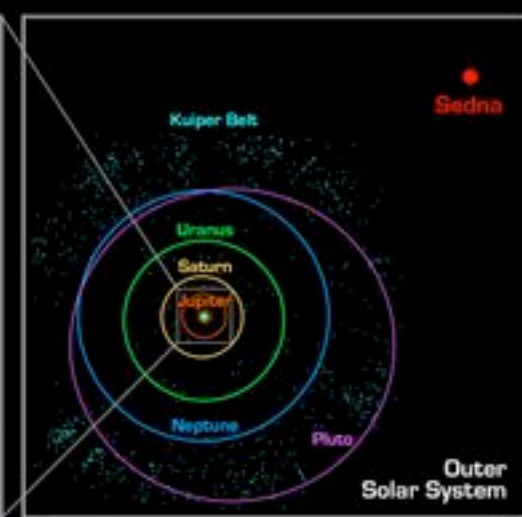
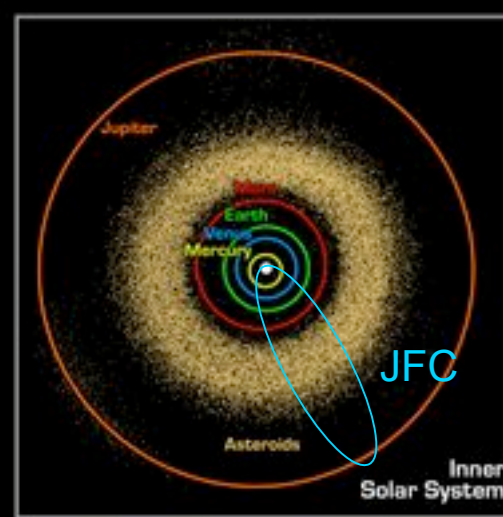
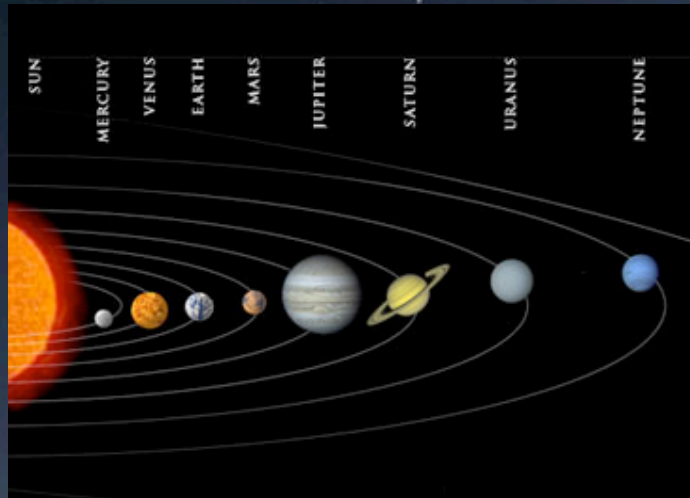
Outline

- Introduction
 - Who I am
 - Solar System, minor bodies
 - Comets
- Distant JFC work
 - Thesis work: Sizes, shapes, densities, colours
 - New projects: Phase functions, SEPPCoN, 67P
- Active comets
 - McNaught
 - Holmes

Who am I?

- I'm from Scotland.
- I studied at St Andrews in Scotland and at Queen's University Belfast in Northern Ireland.
- I've been an ESO fellow since October 2006.
- I'm based in Chile, with duty station at La Silla.
- I work on planetary science in general, with a focus on Solar System minor bodies.

Solar System minor bodies



Comets

- Small bodies of ice and rock.
- Remnants of Solar System formation.
- Activity due to sublimating ices when near the Sun.

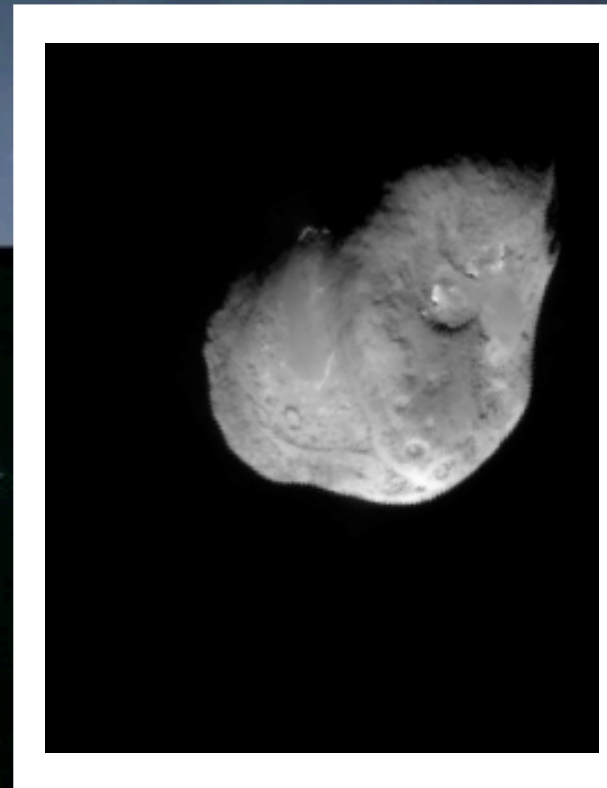


<http://antwrp.gsfc.nasa.gov/apod/ap071105.html>

Credit & Copyright: Vicent Peris and José Lúis Lamadrid (astrofoto.es)

Cometary nuclei

- Nuclei are small (radii a few kilometres) & dark (albedos typically only 4%)
- When active, the nucleus is obscured by coma unless imaged at very high spatial resolution.

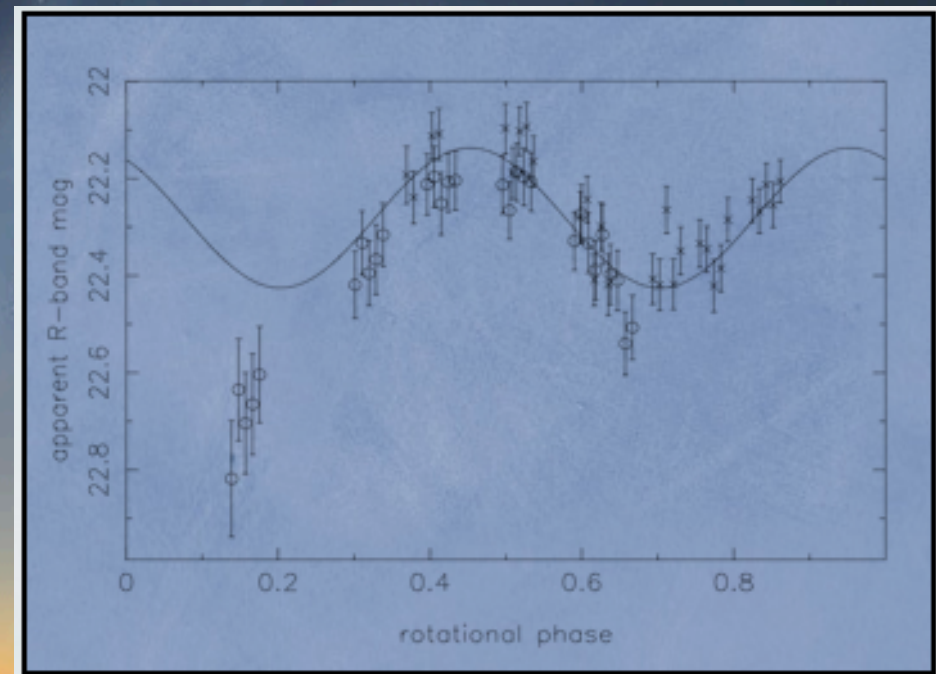


Observing the nucleus

- Canonical cut off for activity is $R_h = 3$ AU.
- We perform multi-filter, time-series photometry (CCD imaging). Requires 2-4m class telescope.
- From this we can measure surface properties (colours) and the bulk physical parameters of size, shape, rotation period and density.
- By studying a large number of comets in this way, we can constrain the general properties of the population.

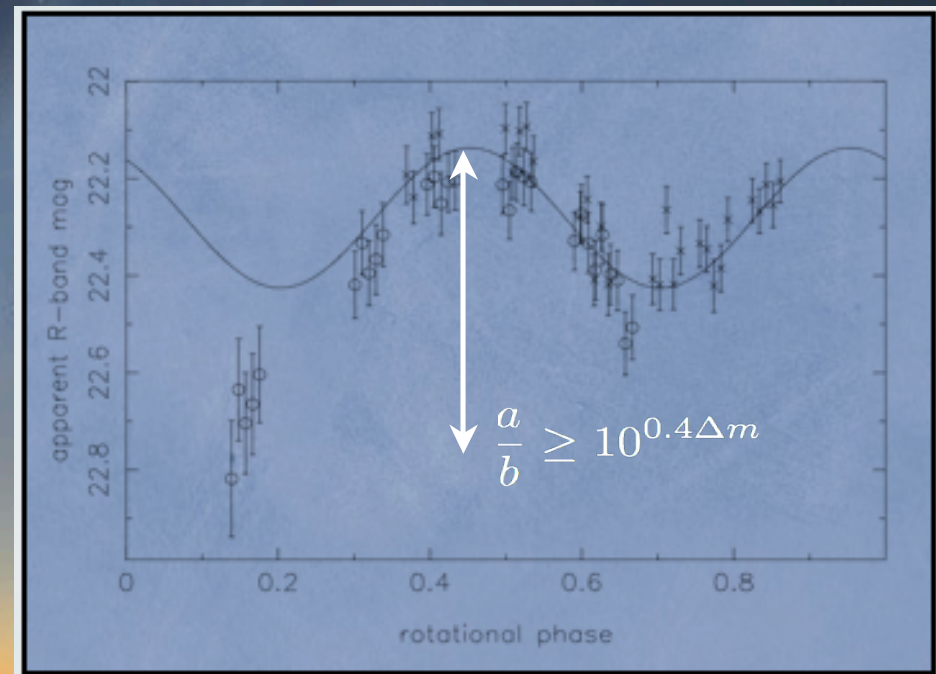
Time-series photometry

- Rotation of non-spherical nucleus causes variation in brightness.
- Searching for periodicity gives rotation rate.
- Mean brightness gives size.

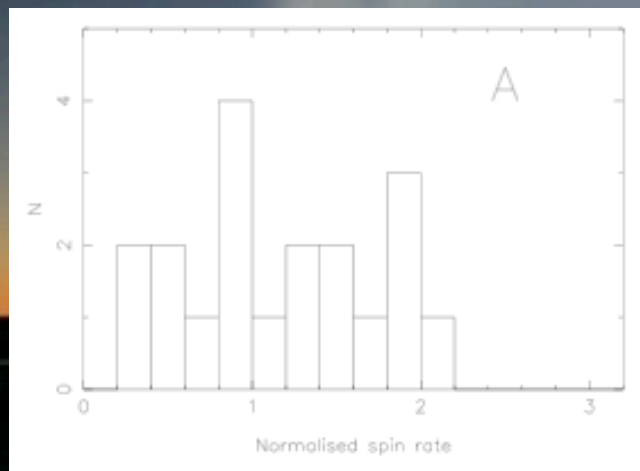
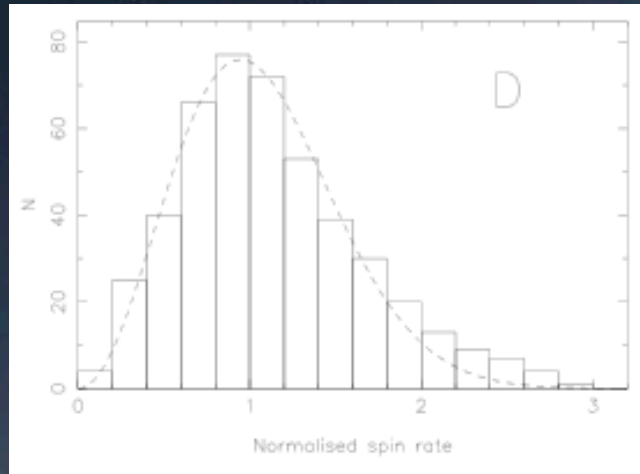


Time-series photometry

- Amplitude of light-curve gives elongation of nucleus, a/b .
- P_{rot} and a/b give minimum bulk density, by balancing self gravitation and centrifugal forces.

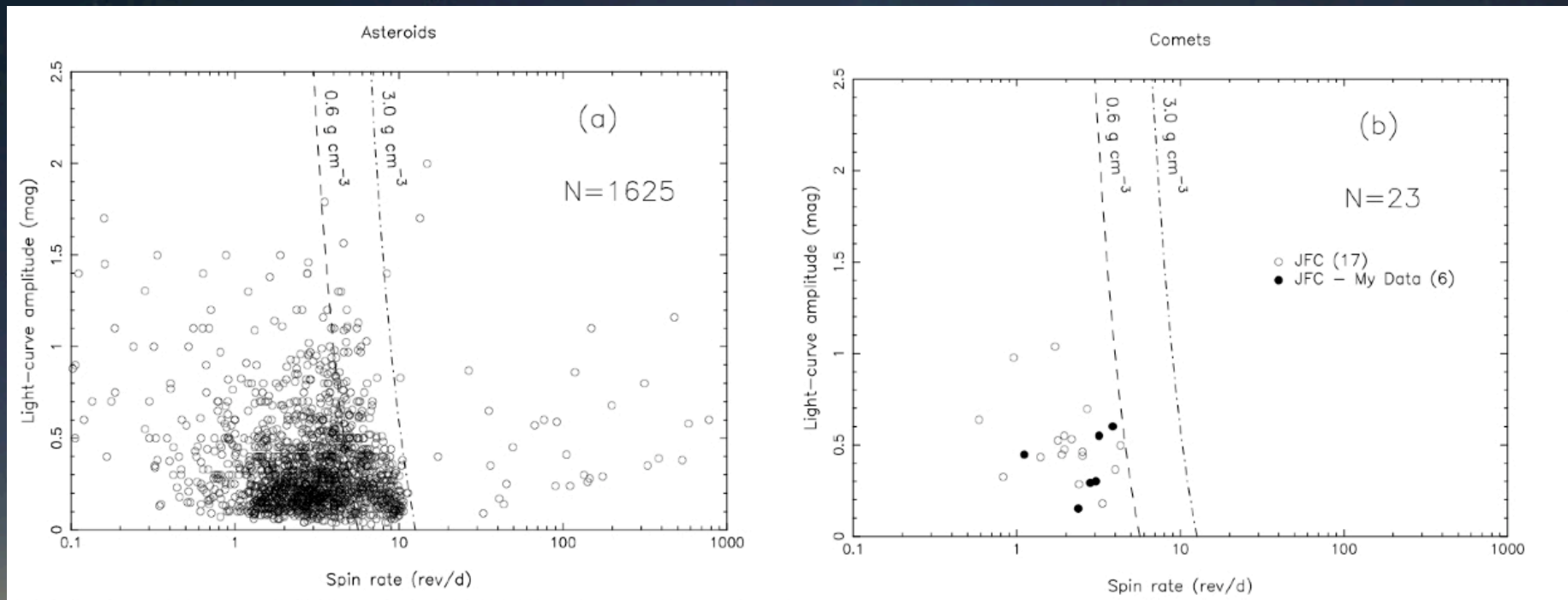


Rotational statistics



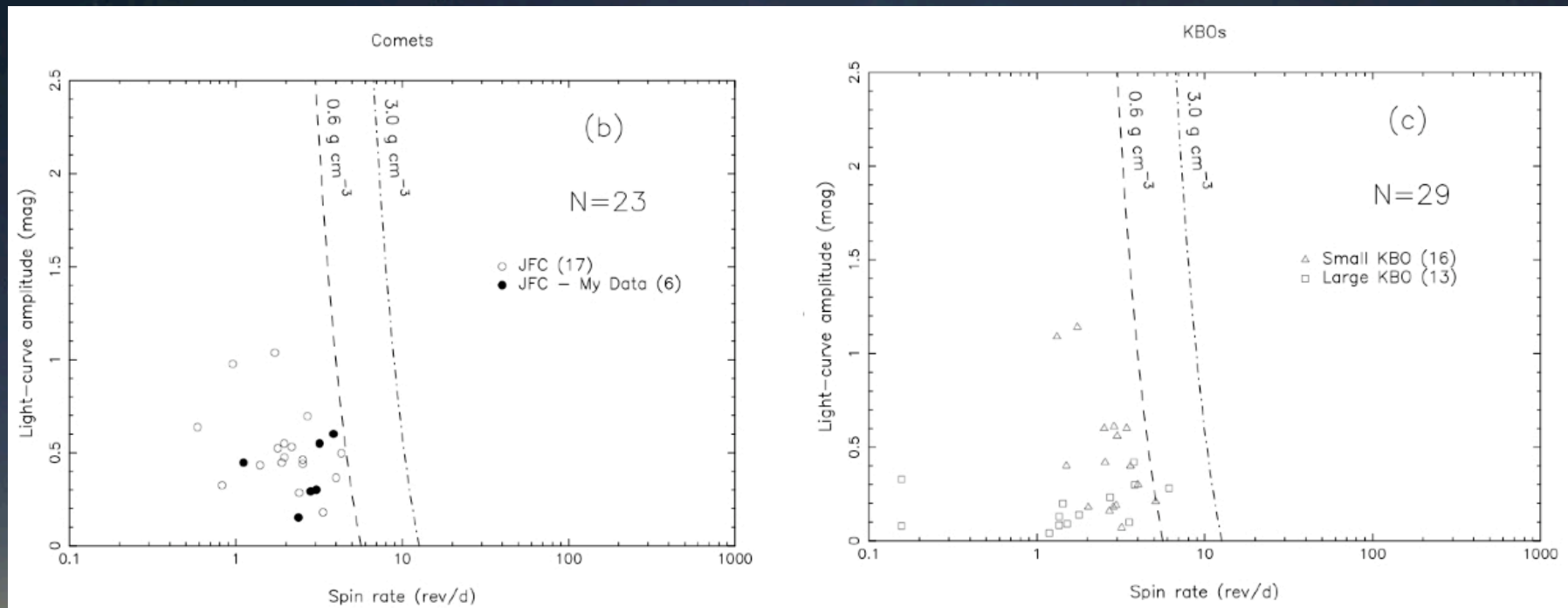
- Spin frequencies should have a Maxwellian distribution for a collisionally relaxed population.
- Comets appear to have a more flat distribution: torques from jets of cometary activity dominate the spin state of nuclei.

Bulk Densities



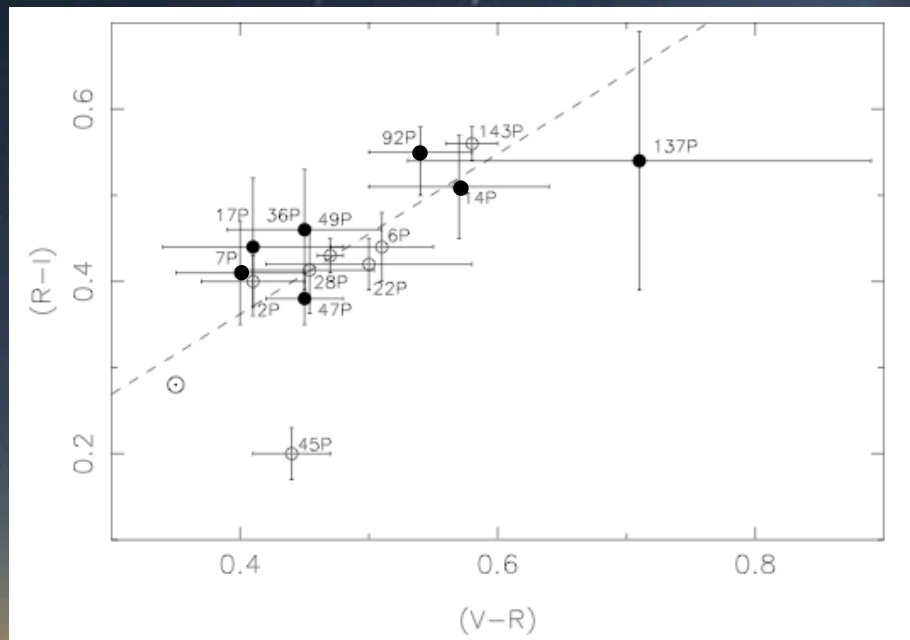
- Light-curve amplitude against spin-rate.
- Minimum bulk density for each minor body (rubble pile).
- Cut off gives average bulk density for population.

Bulk Densities (2)



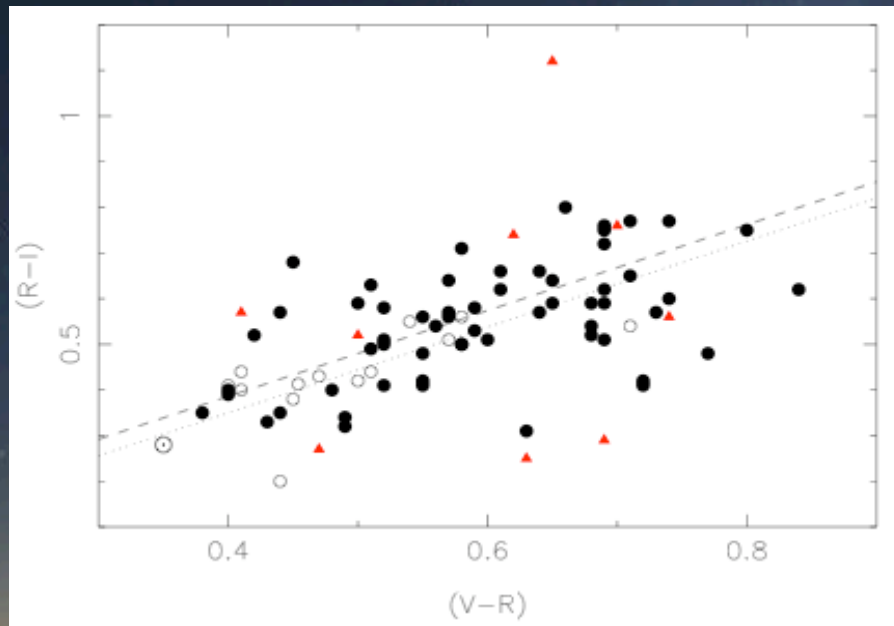
- Same plot for KBOs - split into small ($r < 200 \text{ km}$) and large.
- KBOs appear to have a similar cut off to JFCs, despite different sizes of observed bodies.

Surface Colours



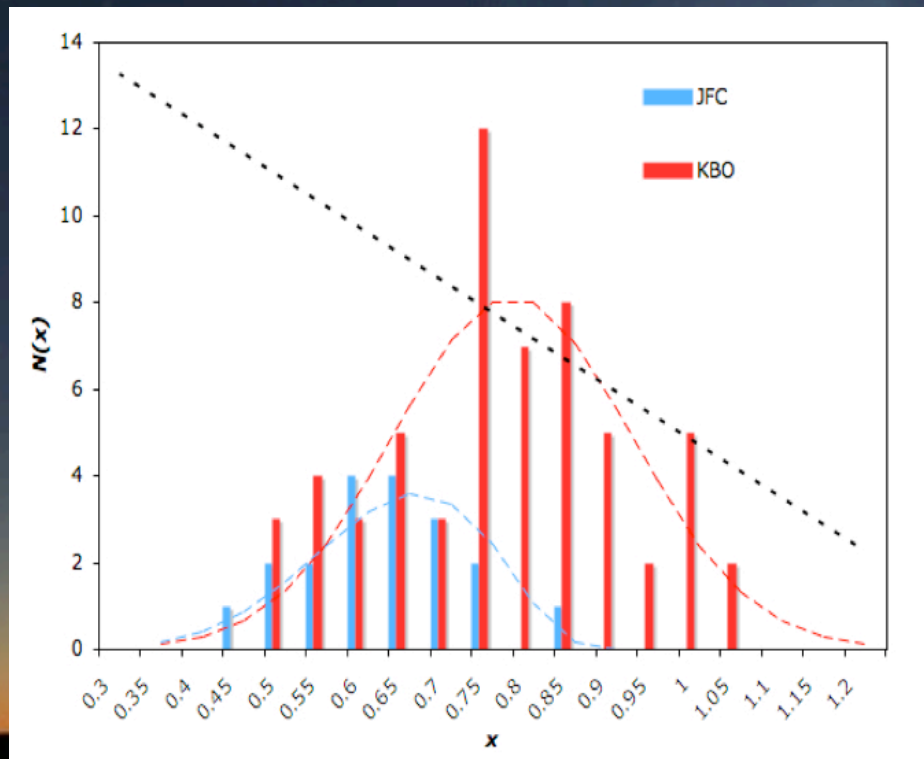
- There is a linear relationship between the measured colour indices $(V-R)$ and $(R-I)$.
- This implies increasing albedo through V, R & I bands, i.e. greater reflectance at greater wavelength.

Surface Colours (2)



- The same relationship between $(V-R)$ and $(R-I)$ holds for KBOs.
- Nuclei are found to be red, but not as red as KBOs.
- This is attributed to removal or covering of the 'ultra-red matter' seen in KBOs by cometary activity.

Surface Colours (3)



- Reducing the colours to a single parameter gives approximately normal distributions for both JFCs and KBOs.
- The JFC distribution can be reproduced by applying a linear 'de-reddening' function to the KBO one: i.e. the reddest surfaces are the most depleted.

Phase functions

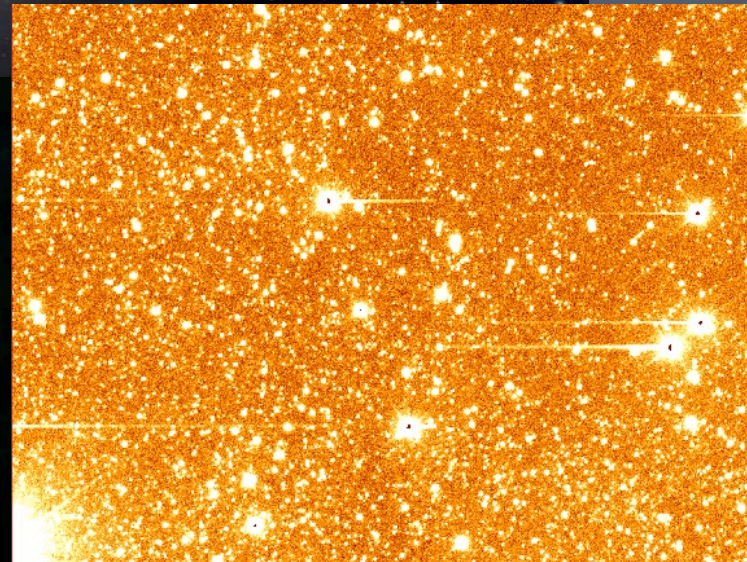
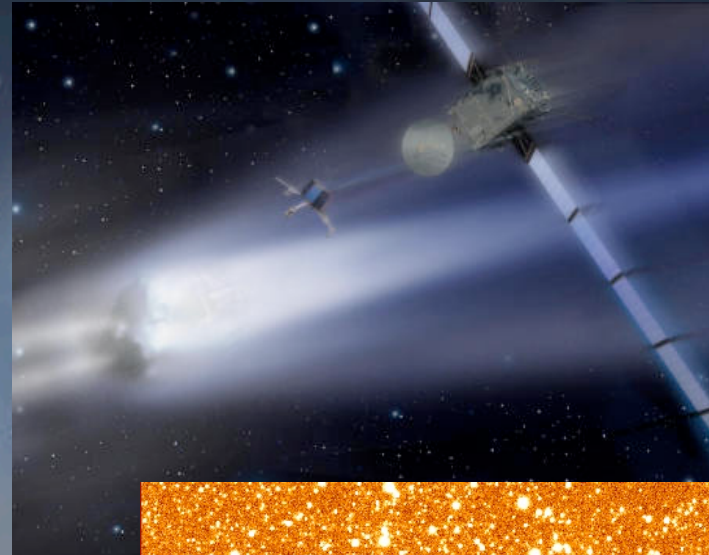
- Size distribution depends on assumed phase function.
- Phase functions measured for only 6 JFCs.
- Requires a light-curve to remove rotational variation at each phase angle:
 - Need time-series observations at many epochs.
- Possible to do with one good template curve, and then short segments.
- Regular monitoring with robotic 2m Liverpool Telescope.
 - Data taken this year.
 - Together with previous points, 11 epochs over 7 degrees in linear part of phase curve.

Albedos

- Size distribution also critically depends on assumed albedo.
- Spitzer program to get observations of 100 JFC nuclei - SEPPCoN, P.I. Y. Fernandez.
- Ground based semi-simultaneous optical at many facilities worldwide (inc 6 nights NTT so far).
- Combination of thermal IR + Optical gives albedo and size distributions.
- Also colours, light-curves, phase functions for a large subset of the comets.

67P/Churyumov-Gerasimenko

- Target of the Rosetta mission, due to arrive at the comet in 2014.
- Observations of the comet this year at the same orbital position it will be in 2014.
- Light-curves, colours, phase function etc.
- Tricky as comet is in front of galactic centre at the moment - crowded field.
- Activity may start this year.

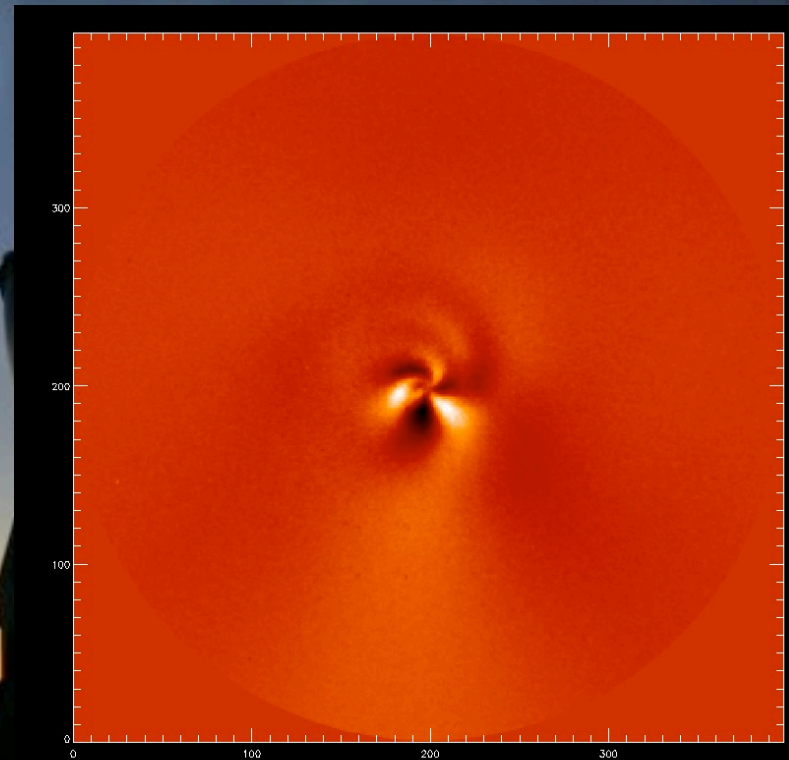


Active comets

- Two unexpected events this year produced fantastic comet displays visible to the naked eye.
 - Comet C/2006 P1 (McNaught) is a dynamically new Oort cloud comet which passed close to the Sun in January and became the brightest comet in 40 years.
 - Comet 17P/Holmes is a JFC which underwent a spectacular outburst and brightened by 14 magnitudes in late October.
- We got data on both comets...

Comet McNaught

- Perihelion on 12th January.
- Observed with the NTT starting on the 27th, at airmass 5 in twilight...
- Strong CN gas jets seen, and exotic species identified in gas coma (Sodium and other metals).



Comet Holmes

- 17P/Holmes was out bound at ~ 2.5 AU and fading normally at $V \sim 17$ in late October.
- At around Oct 24.0 UT a sudden outburst occurred, and the comet reached $V \sim 3$.
- Reported by amateurs - quick reaction meant that we have data starting at Oct 24.9.
- Data from numerous Northern sites. Unusually circular coma.
- We see an expanding shell of material, now being swept into a tail.

