

Structure of the Milky Way: Unusual Fossil Relics

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Between The Lines: a Stellar Spectroscopy Workshop
2-4 December 2024

Funding support by:



Comité Mixto
ESO-Gobierno
de Chile 2021

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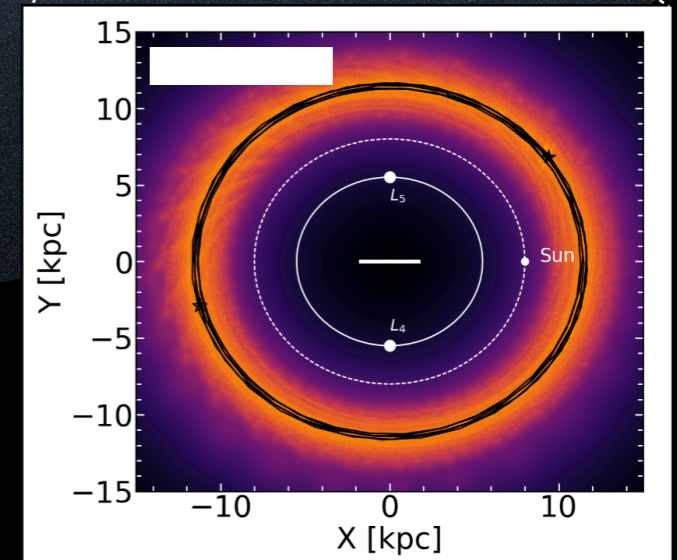
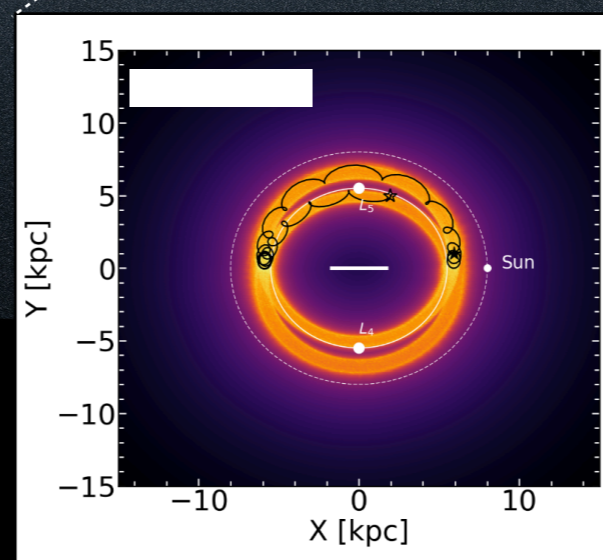
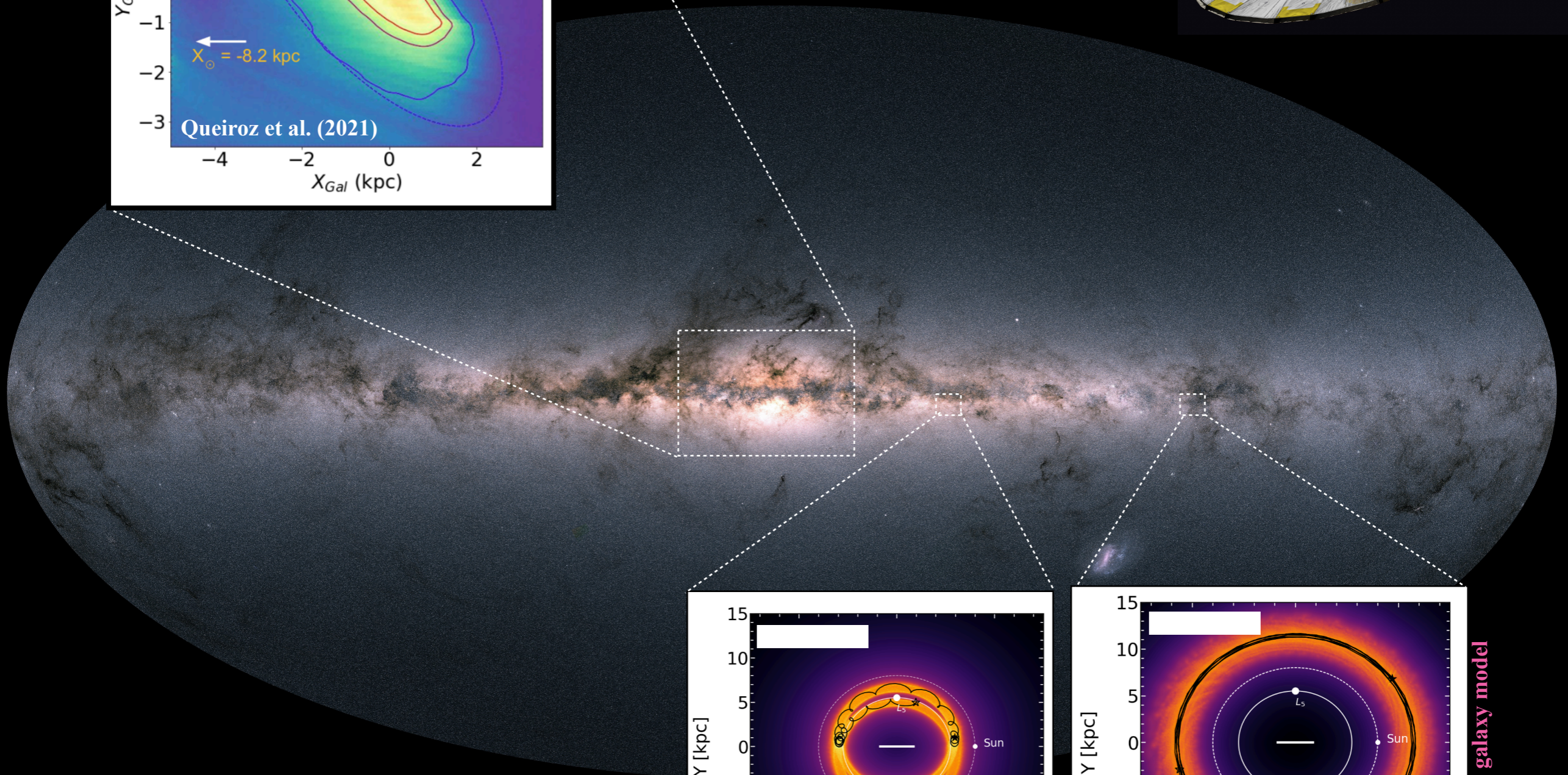
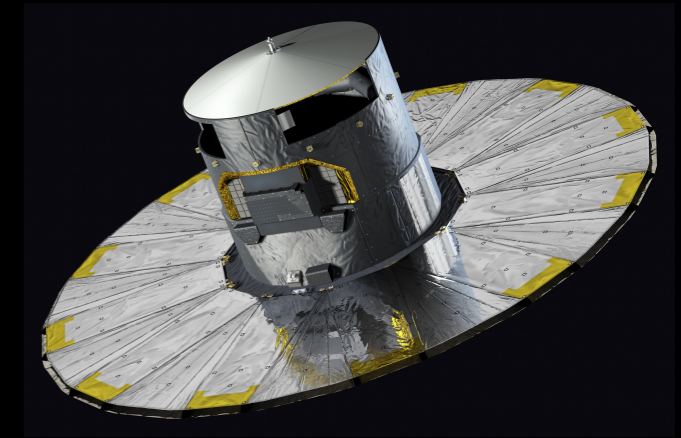
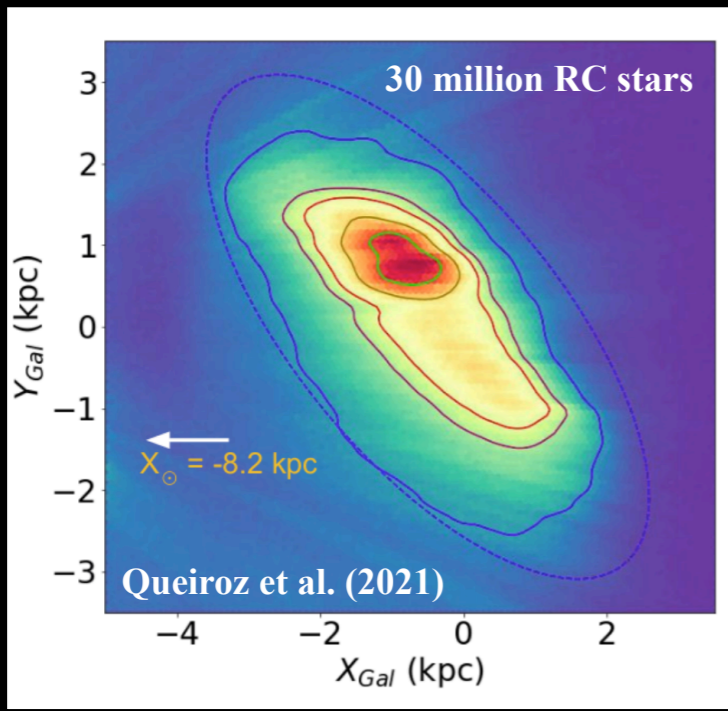


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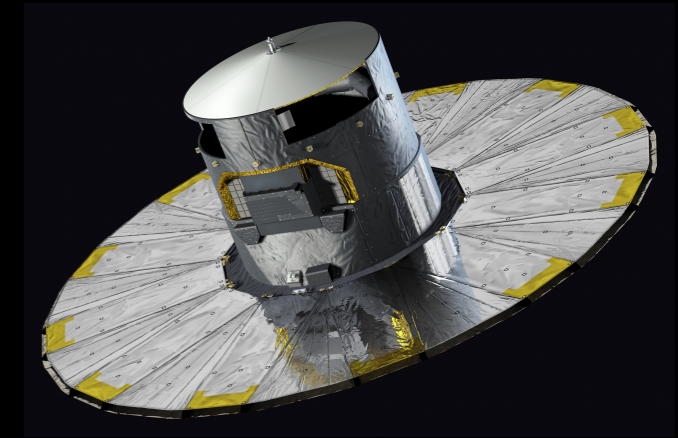




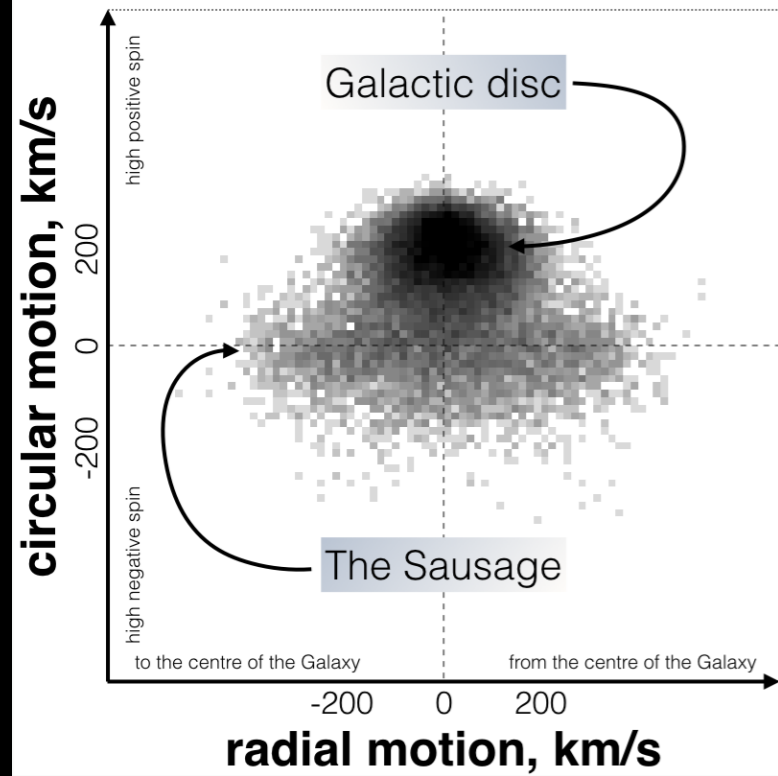
Credit: ESA/Gaia/DPAC, CC BY-SA 3.0 IGO

OC's: Guerço et al. in prep.

Violent past



Motions of 7,000,000 Gaia stars



Credit: V. Belokurov (Cambridge, UK and CCA, New York, US) and Gaia/ESA

Cause: *Gaia-Sausage/Enceladus*



Effect: *Likely the splash*

see Belokurov et al. 2020, MNRAS, 494, 3880 - AND-
Amarante et al. 2020, ApJL, 891, L30 (*different view*)

Credit: René van der Woude, Mixr.nl

Gaia-Sausage: Belokurov et al. 2018, MNRAS, 478, 611

Gaia-Enceladus: Helmi et al. 2018, Nature, 563, 85

Mix of stellar populations: Including Chemically unusual Stellar debris

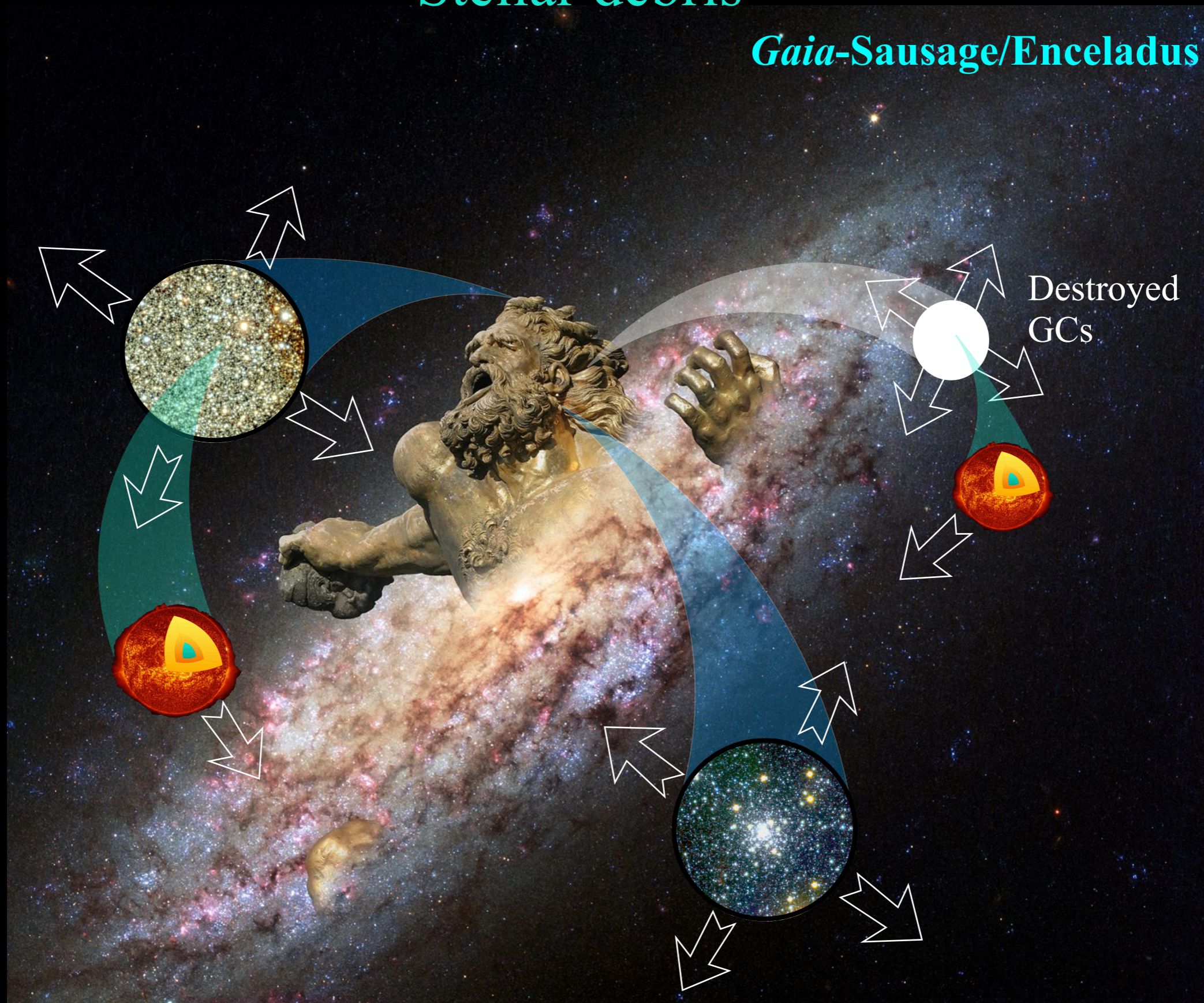
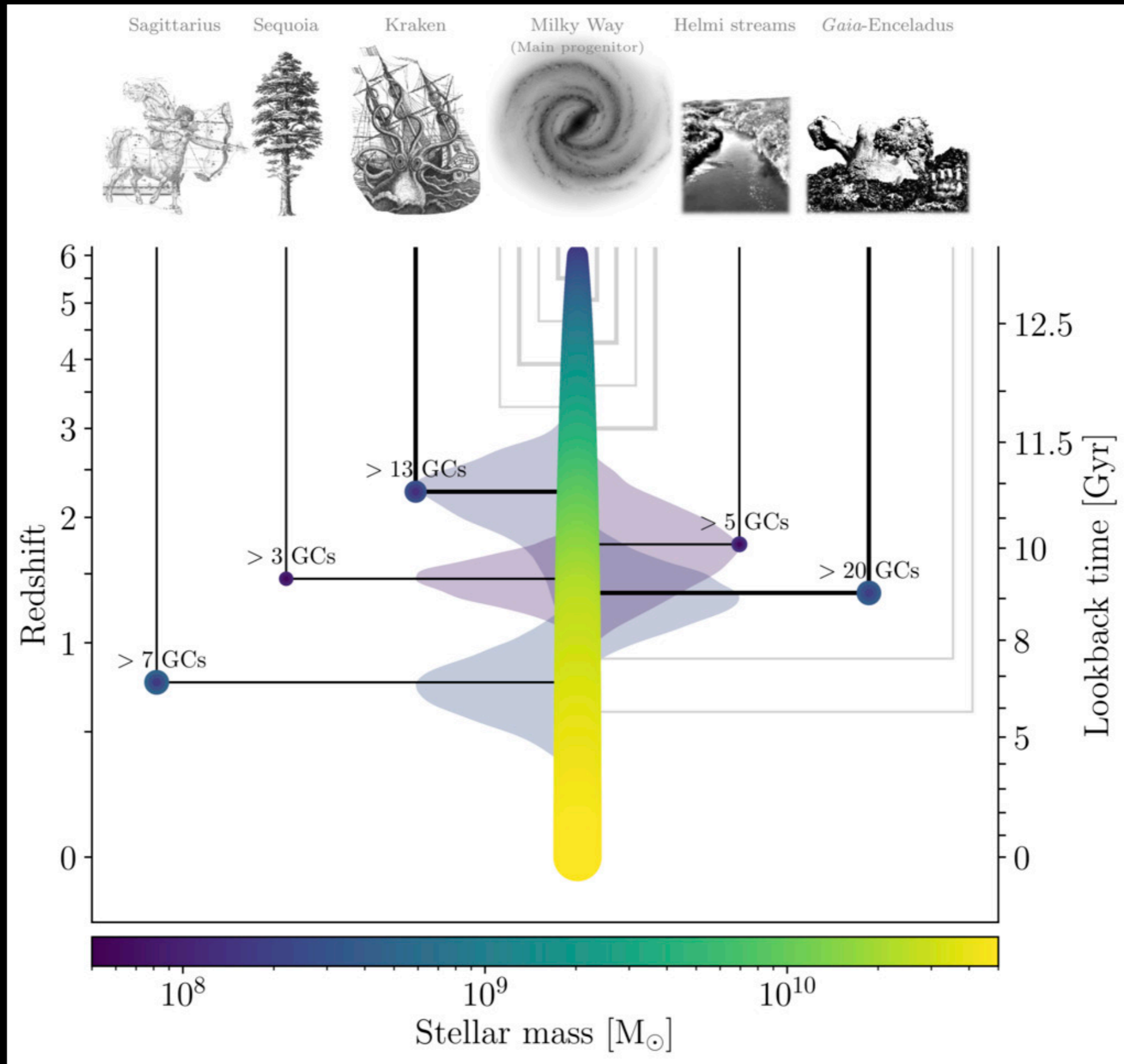


Figure modified from René van der Woude, mixr.nl (background image)

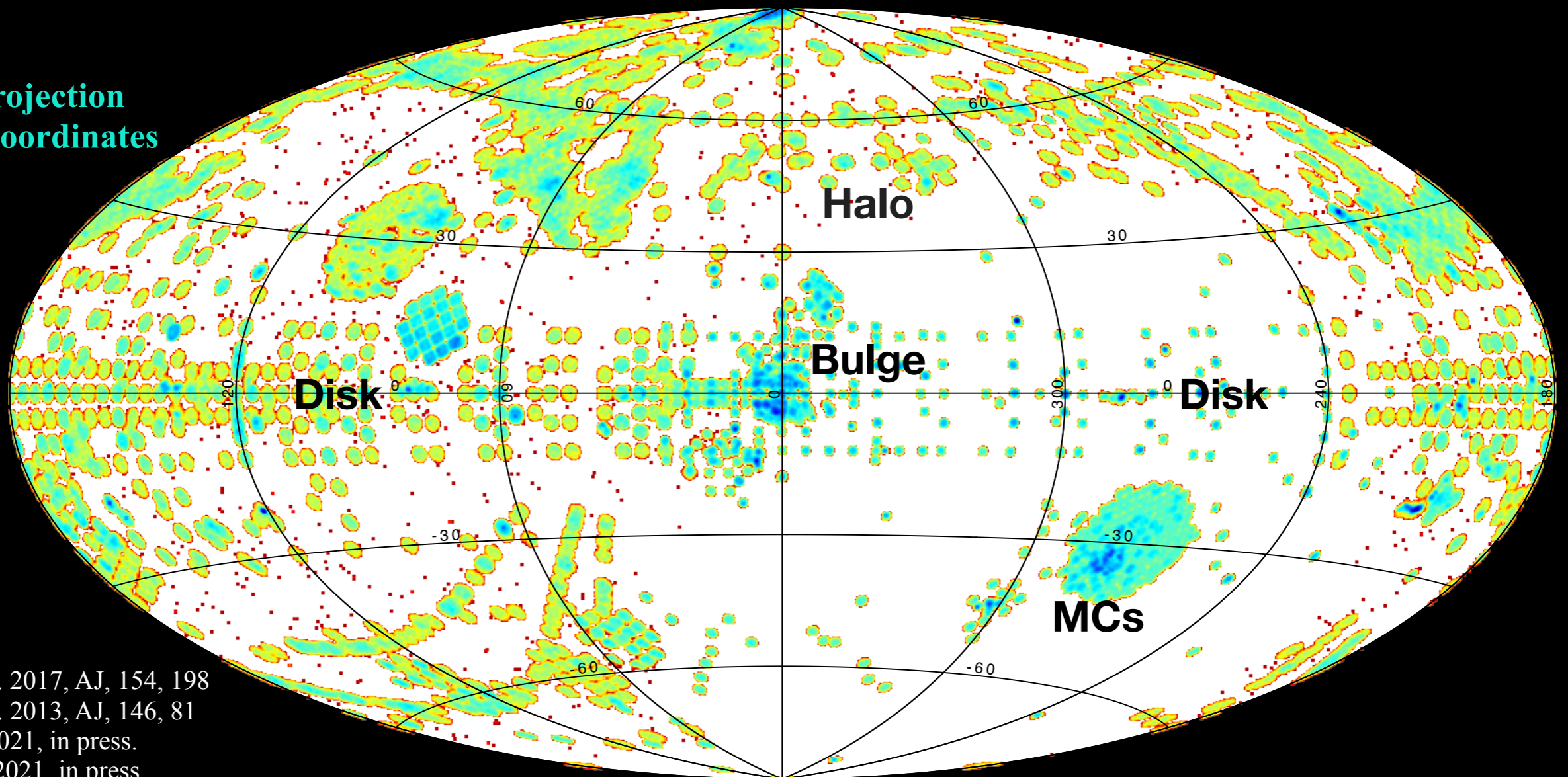
Past Accretion Events



APOGEE: The Apache Point Observatory Galactic Evolution Experiment survey

- ▶ Near-infrared survey / *H*-band (1.5 — 1.7) μm
- ▶ ~ 700 000 Stars (SDSS DR17)
- ▶ S/N > 70
- ▶ Radial Velocity error: 0.1 km s⁻¹
- ▶ $T_{\text{eff}} \pm 100$ K; $\log g \pm 0.1$ dex; [Fe/H] ± 0.04 dex; abundances ± 0.1 dex

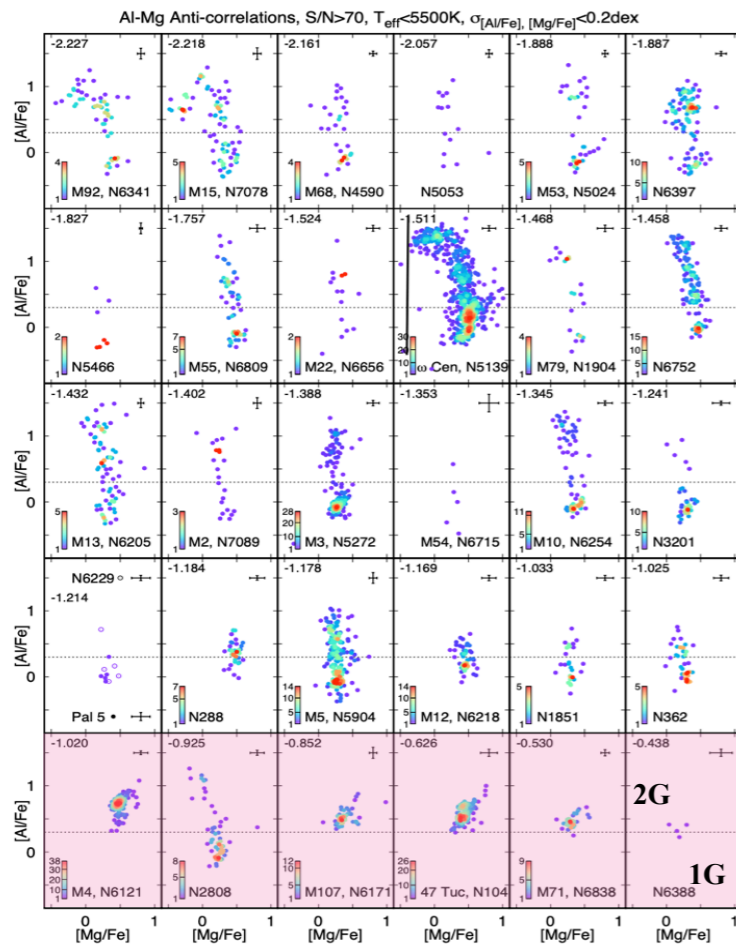
Aitoff projection
Galactic Coordinates



Key Chemical Species in GCs: Multiple-population Phenomenon

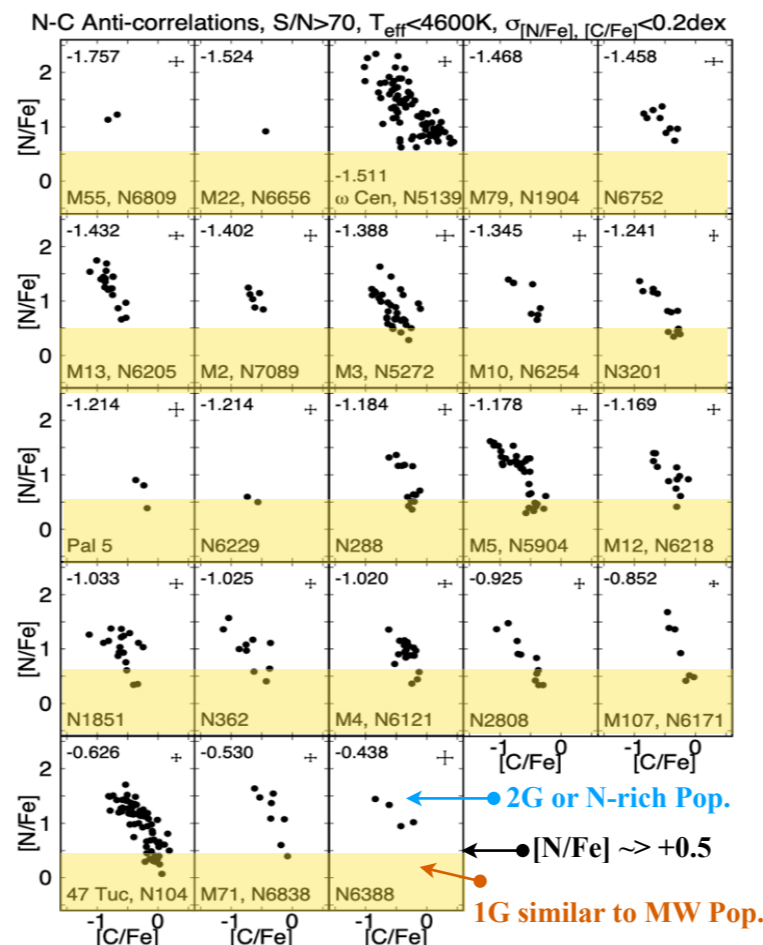


Al-Mg plane



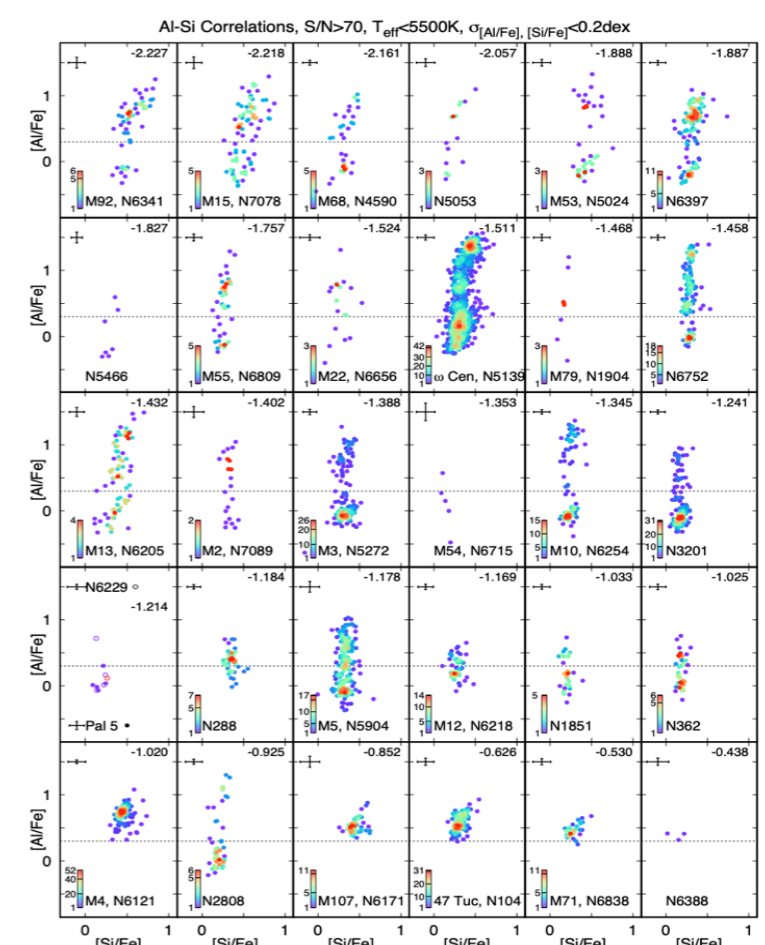
anticorrelation

N-C plane



anticorrelation

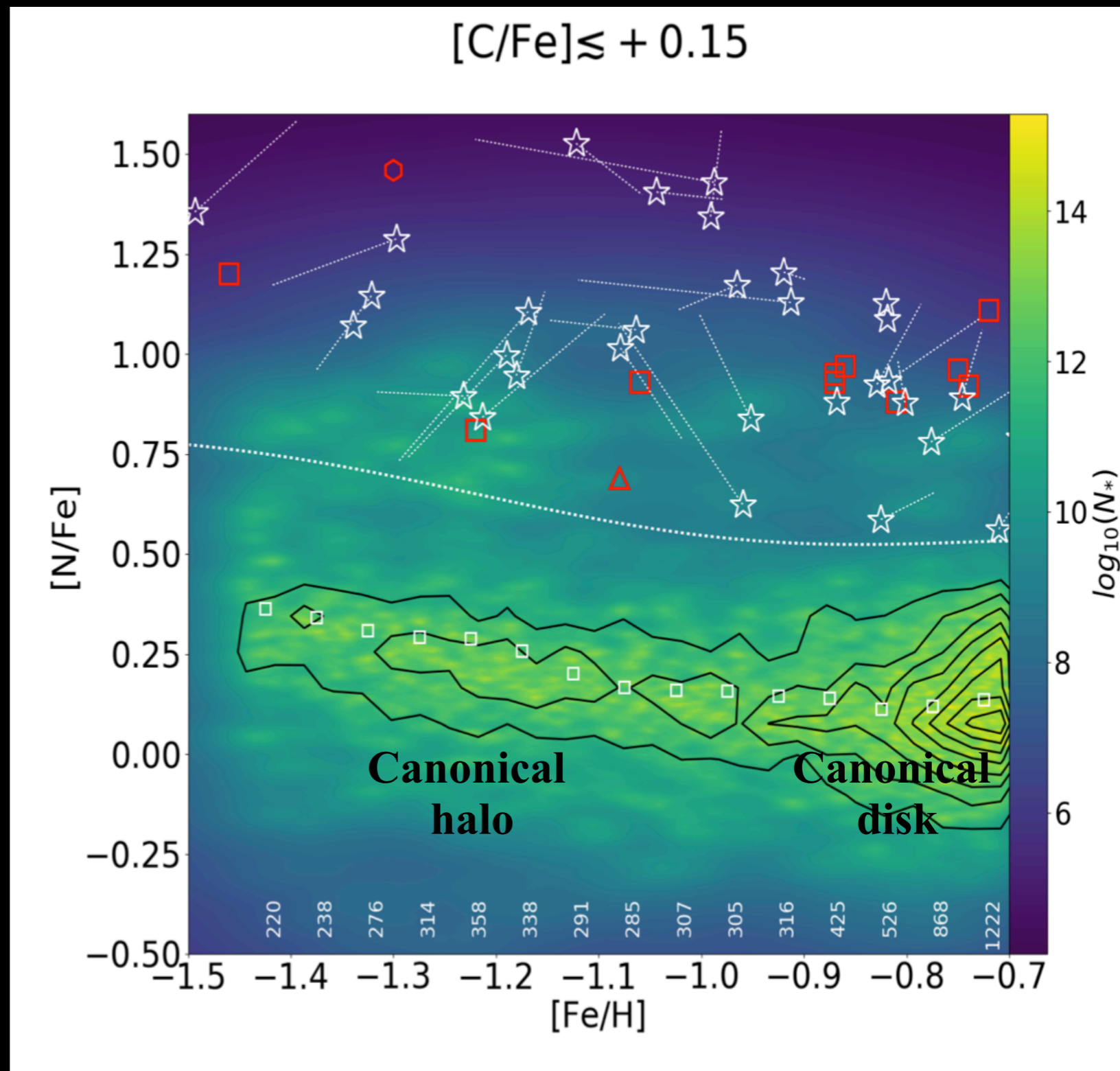
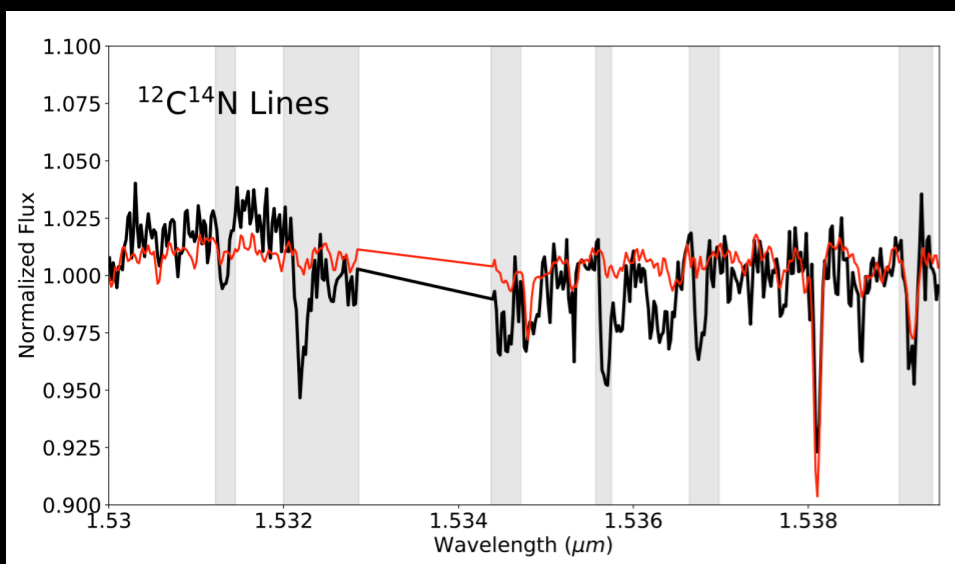
Al-Si plane



correlation

N-rich Stars: A new population in the Milky Way bulge, disk, and inner/outer halo

- Fernández-Trincado et al. 2016, ApJ, 833, 132
- Martell et al. 2016, ApJ, 825, 146
- Schiavon et al. 2017, MNRAS, 465, 501-524
- Fernández-Trincado et al. 2019, A&A, 631, A97
- Fernández-Trincado et al. 2019, MNRAS, 488, 2864
- Fernández-Trincado et al. 2020, MNRAS, 495, 4113



N-rich Stars with Extra-galactic Origins

12
Mg
Magnesium
24.305

13
Al
Aluminum
26.982

14
Si
Silicon
28.086

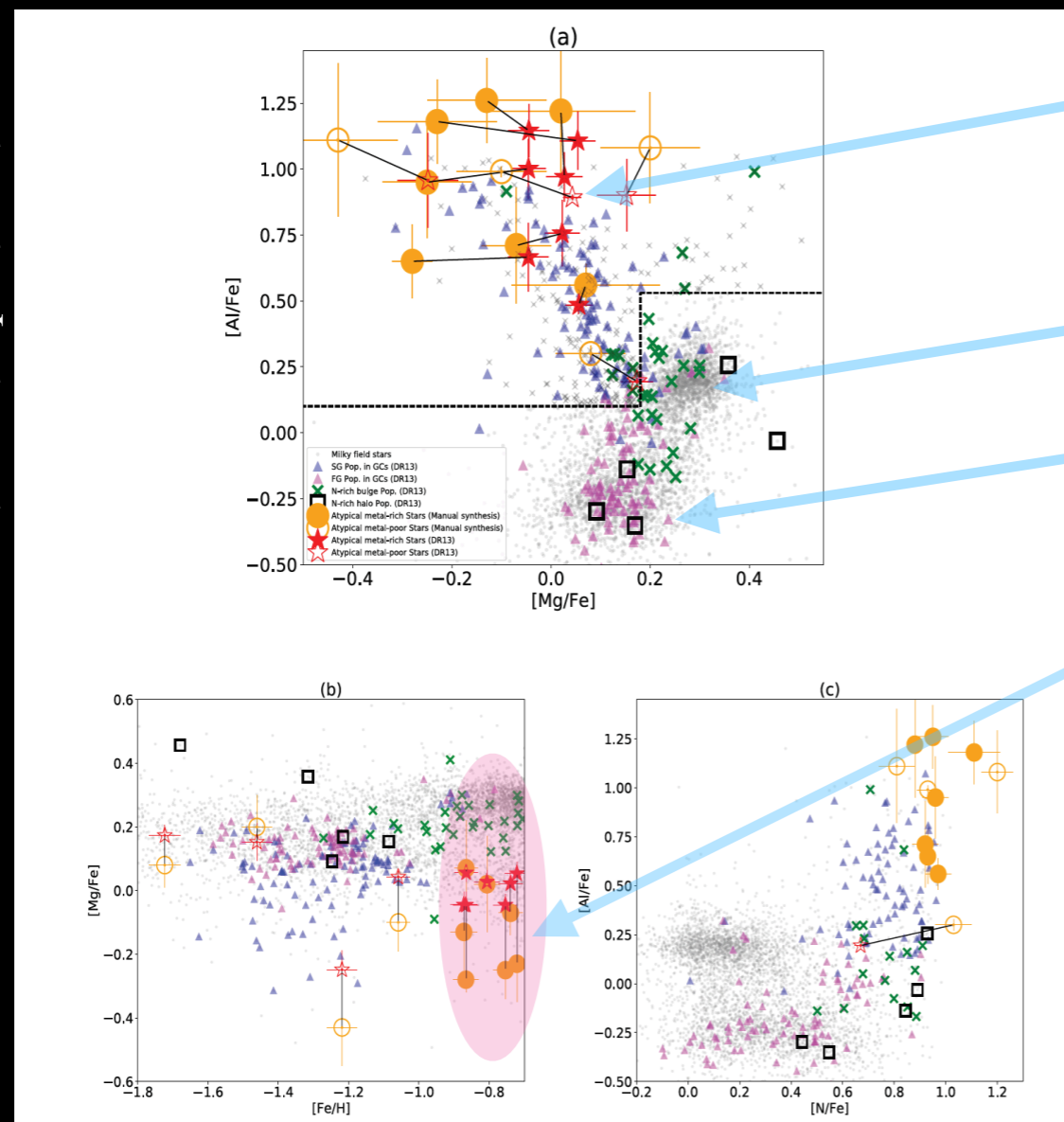
6
C
Carbon
12.011

7
N
Nitrogen
14.007

- ◆ Low Magnesium makes them **unlikely linked to Galactic Globular Clusters**
- ◆ High Aluminum enrichment makes them **unlikely linked to Dwarf Galaxy debris**
- ◆ Retrograde orbits suggest a probable link with **disrupted Globular Clusters**
- ◆ **Absence of radial velocity variation:** Difficult to support pollution by stellar winds from a binary companion

2G-like Abundance Patterns with Possible Extragalactic Origins

Fernández-Trincado et al., 2017, ApJL, 846, L2



GCs populations/debris
and unexplained objects

Canonical (thick) disk

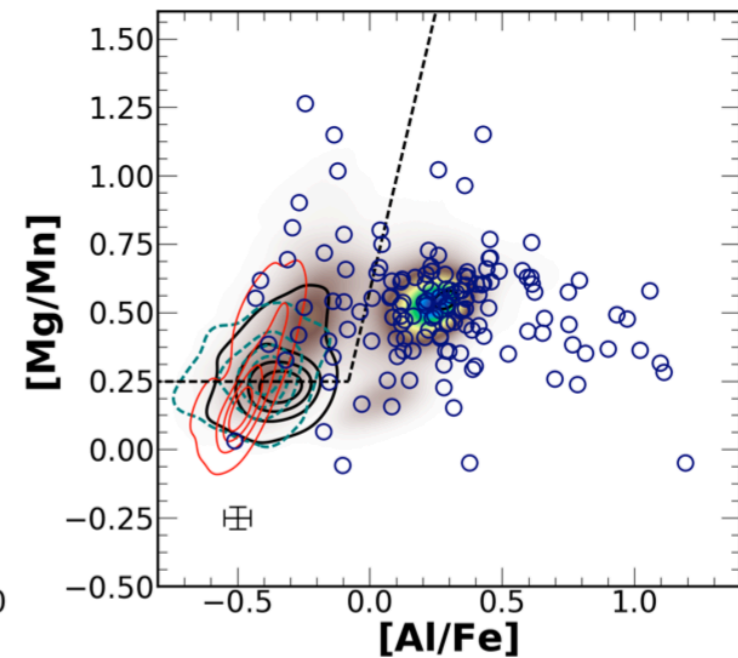
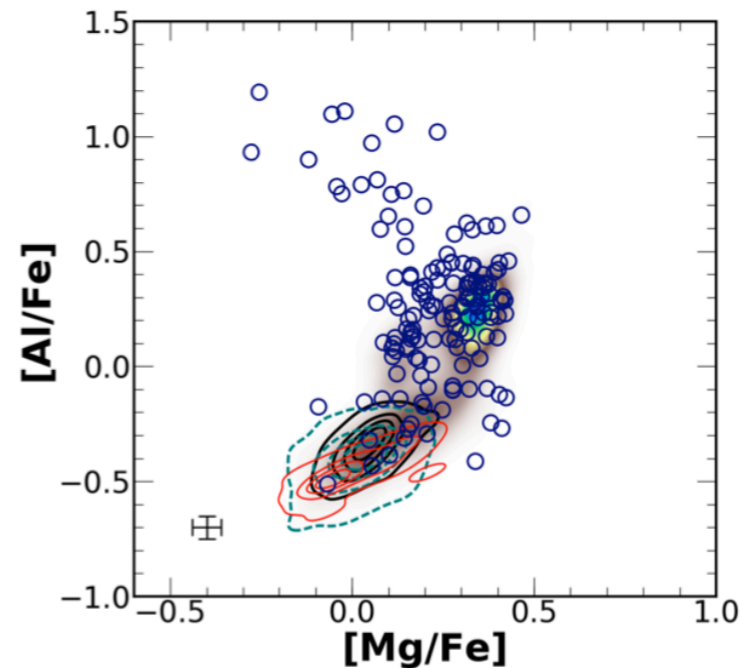
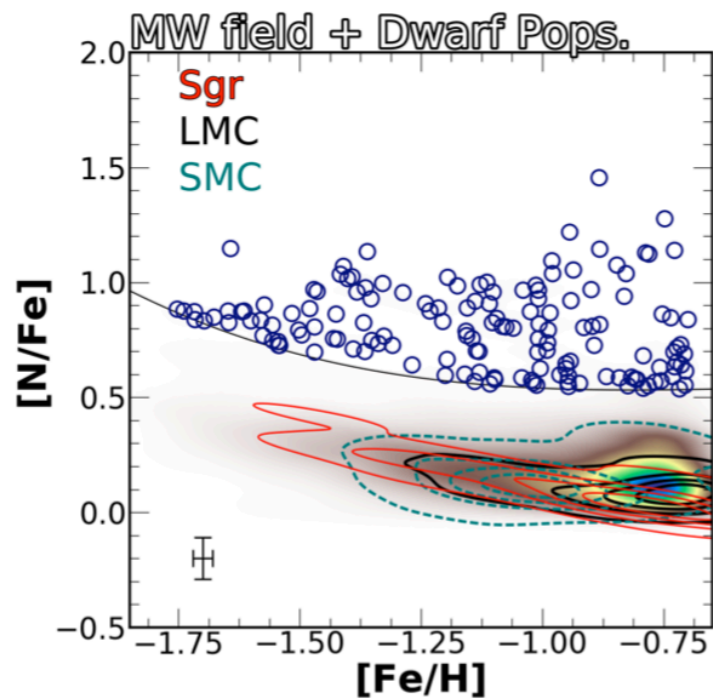
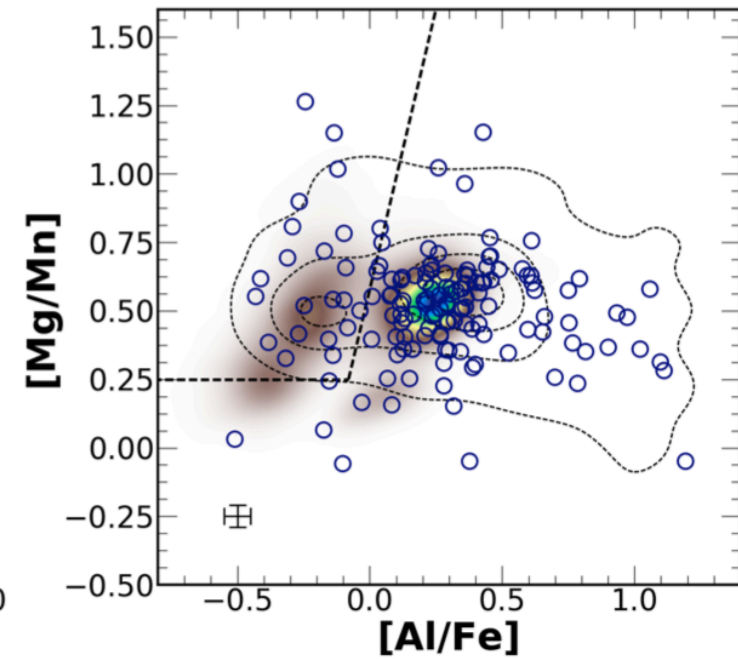
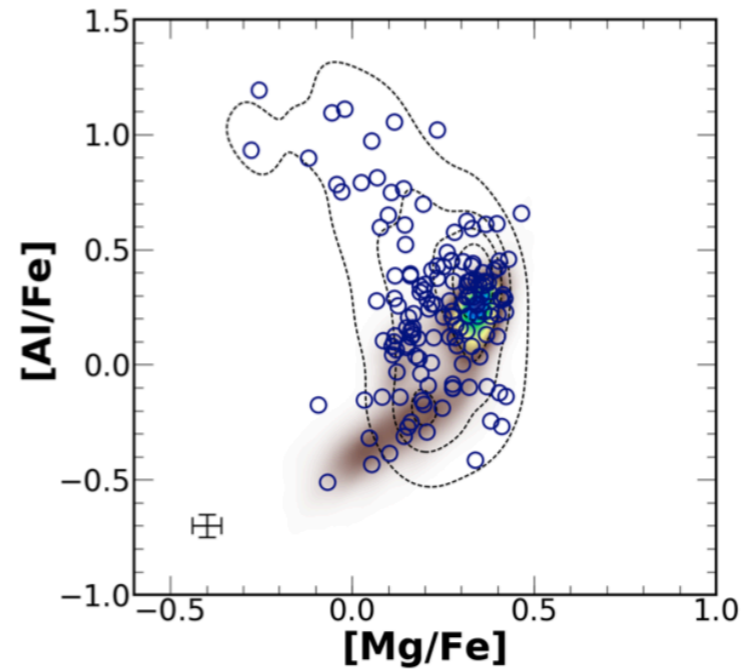
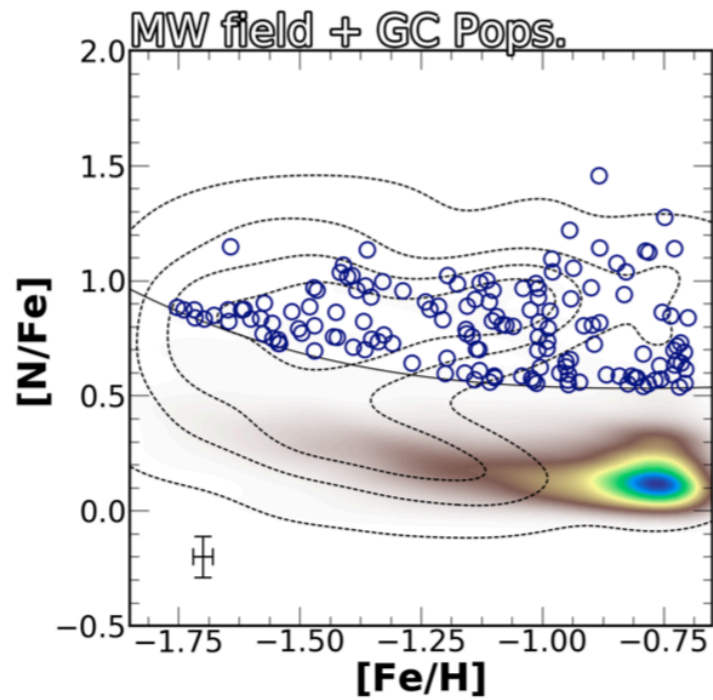
Canonical/Accreted Halo

NOT Galactic GCs with
 $[Mg/Fe] < 0$ in $[Fe/H] > -1.0$

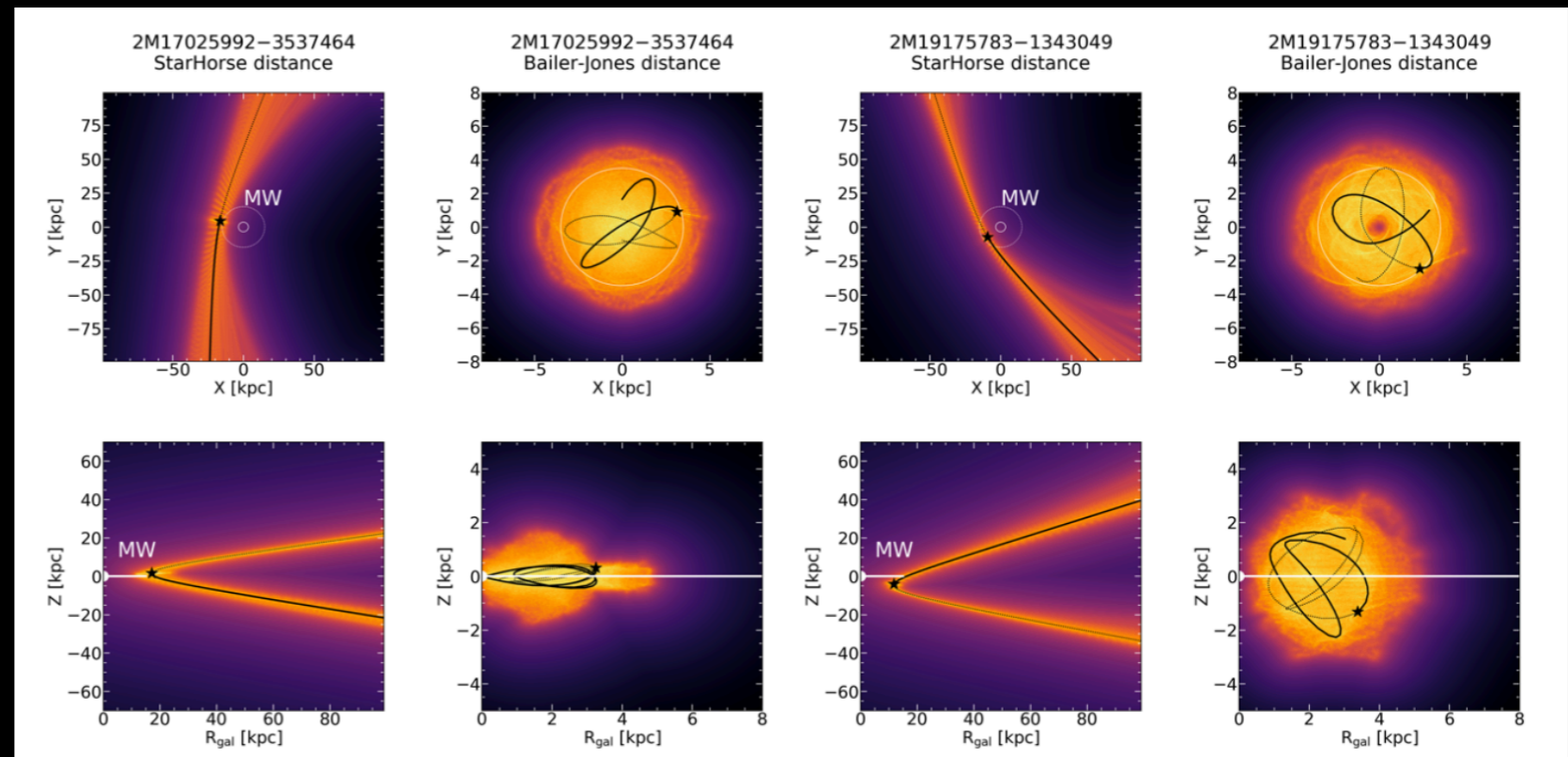
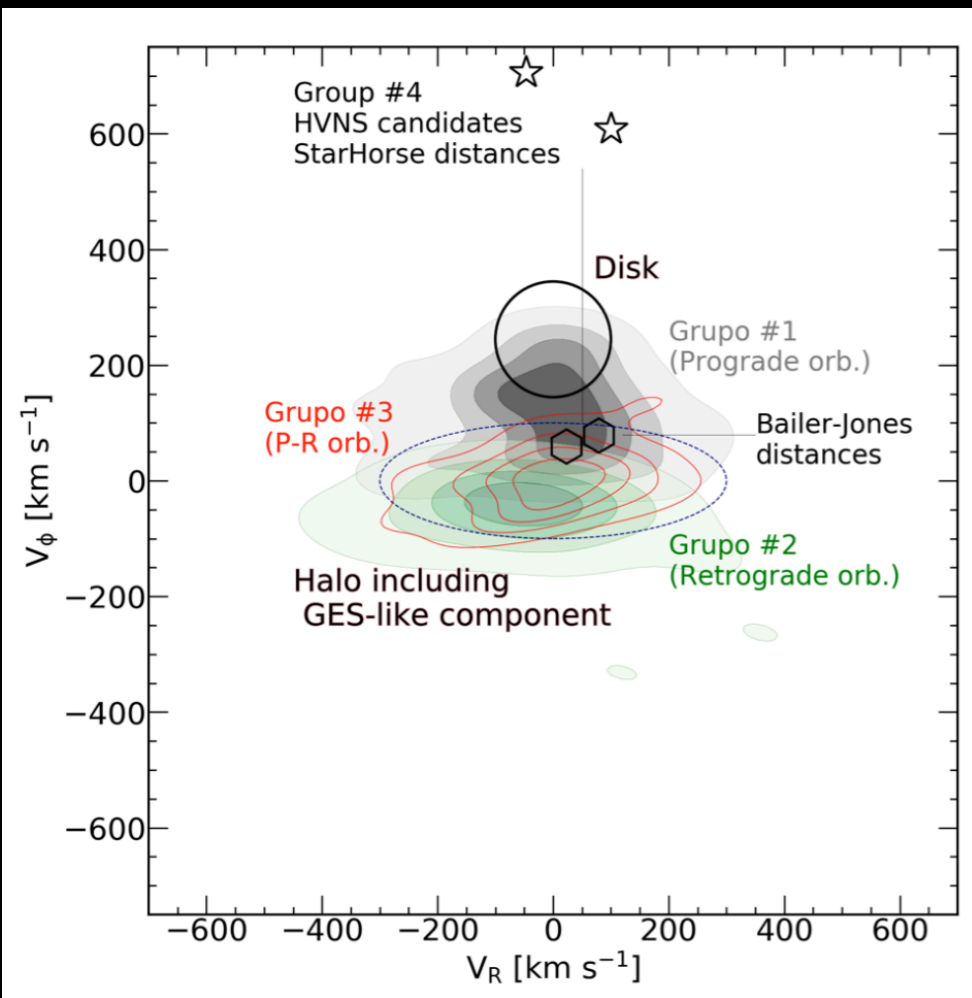
*~11 N-rich Stars
toward the inner Stellar Halo*

$[C/Fe] < +0.15$ dex

Large census of N-rich stars (Bulge+Disk+Halo): 412 unusual giant stars

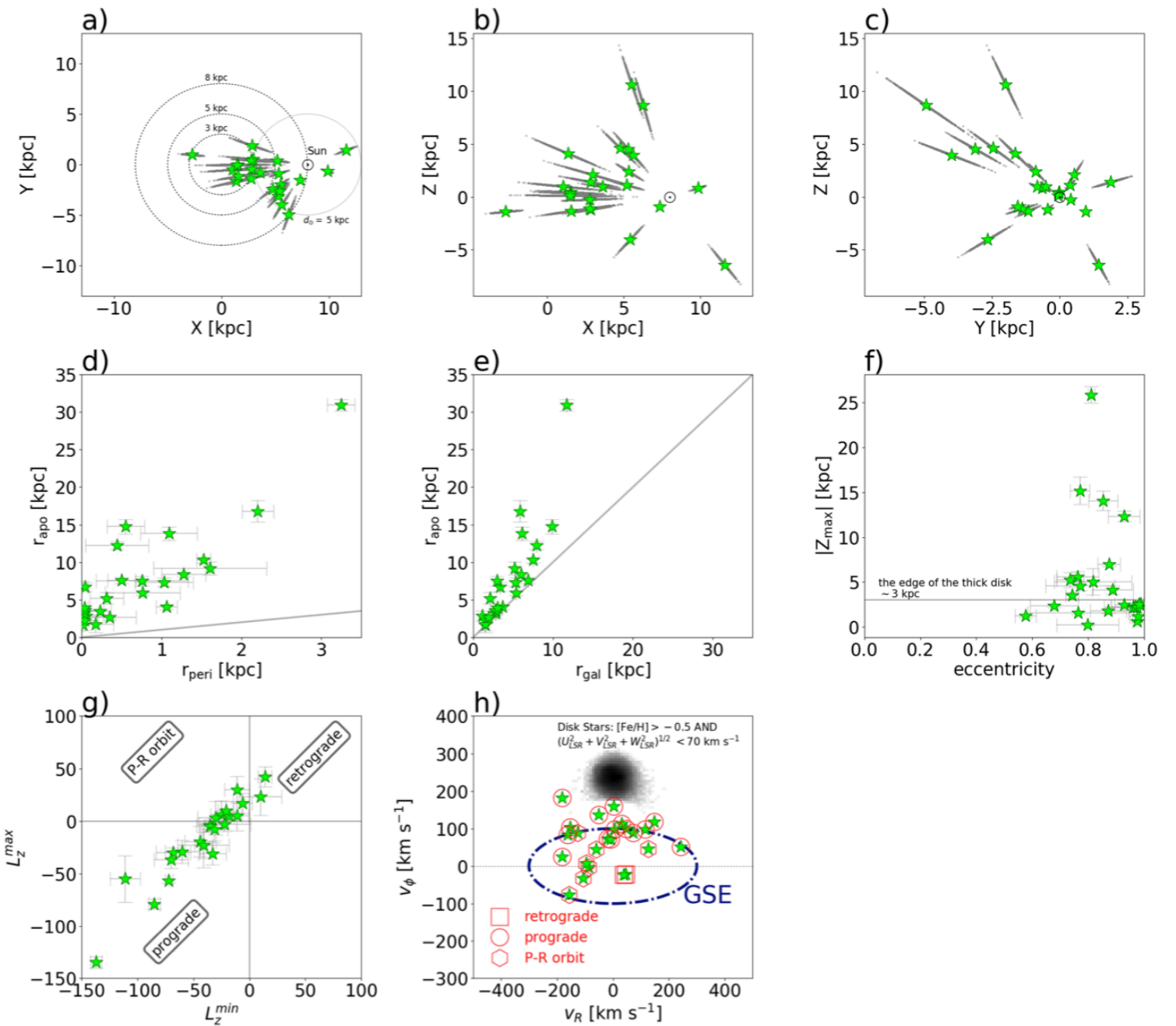
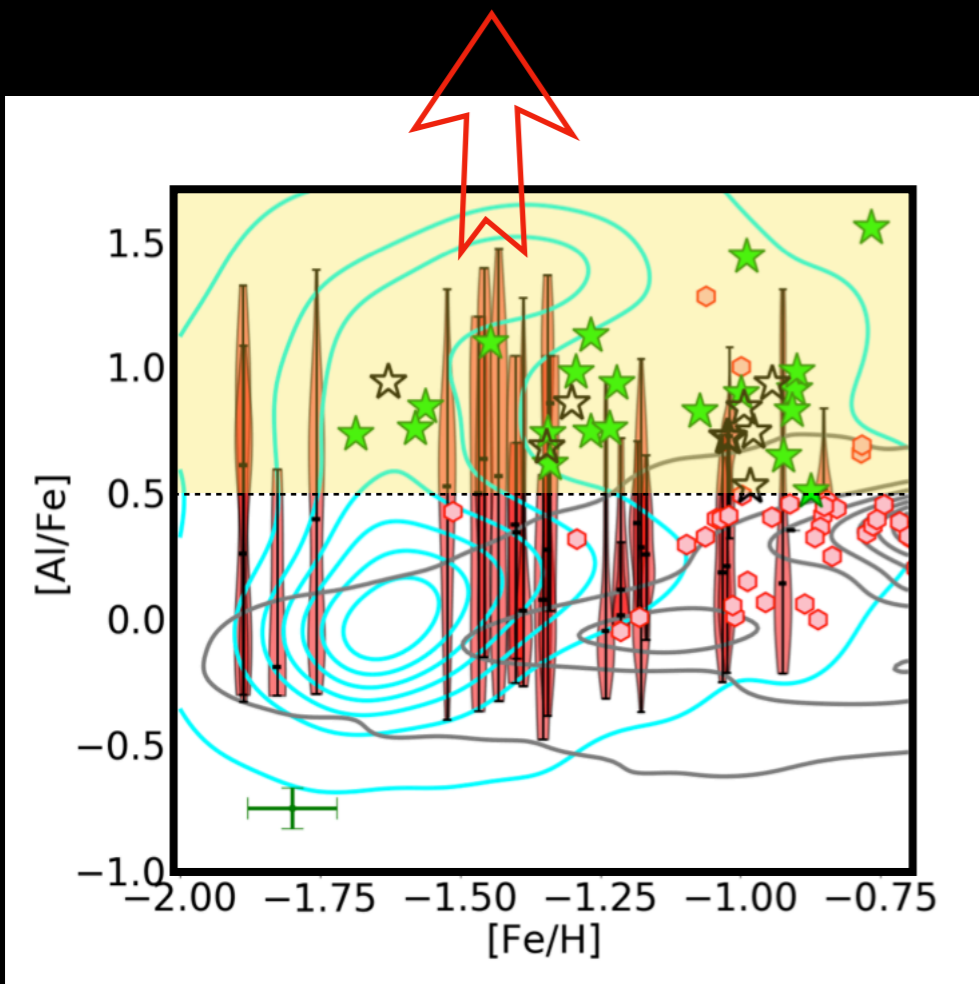


Large census of N-rich stars (Bulge+Disk+Halo): 412 unusual giant stars

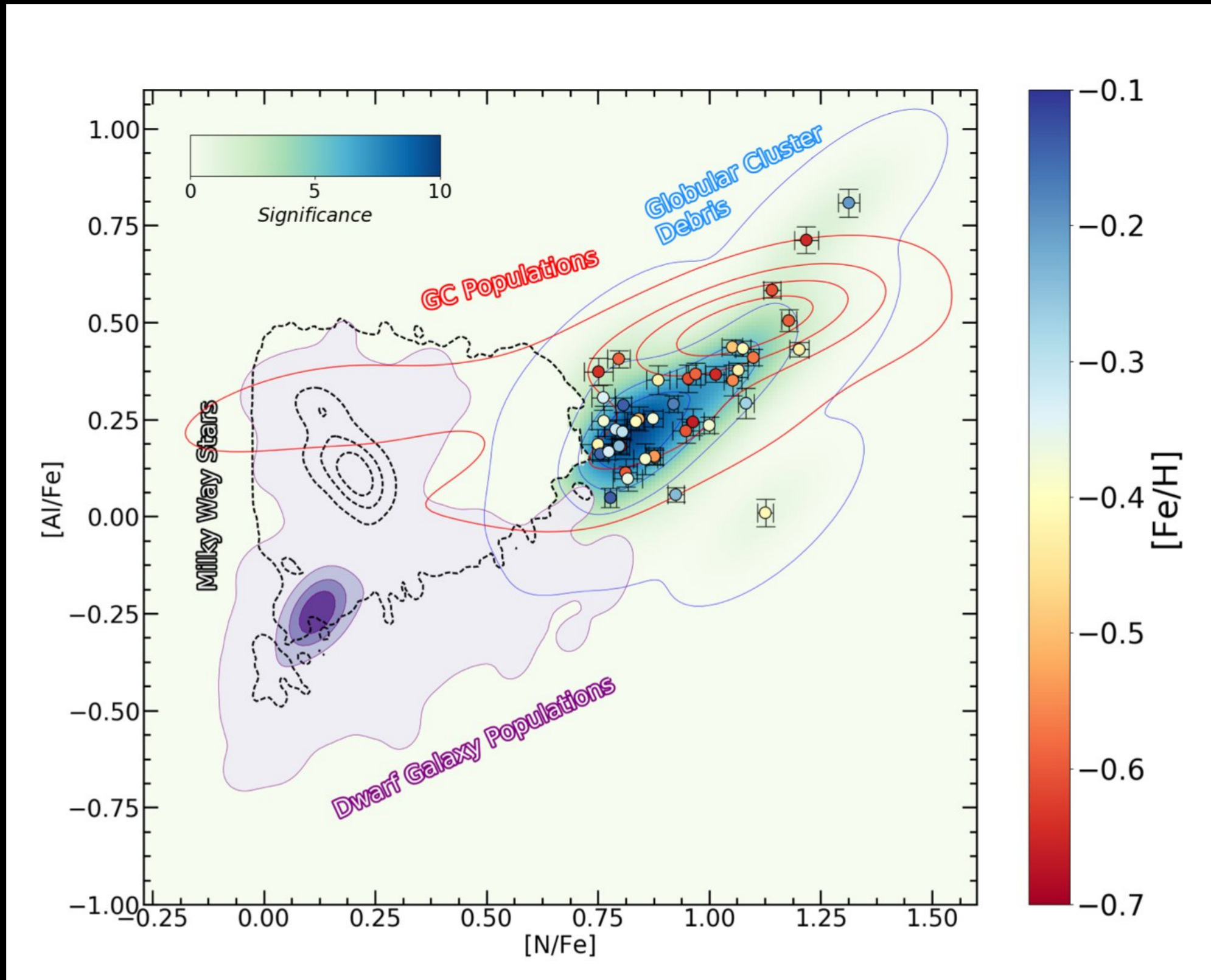


Discovery of a Population of Aluminium-enriched metal-poor stars buried in the Galactic Bulge

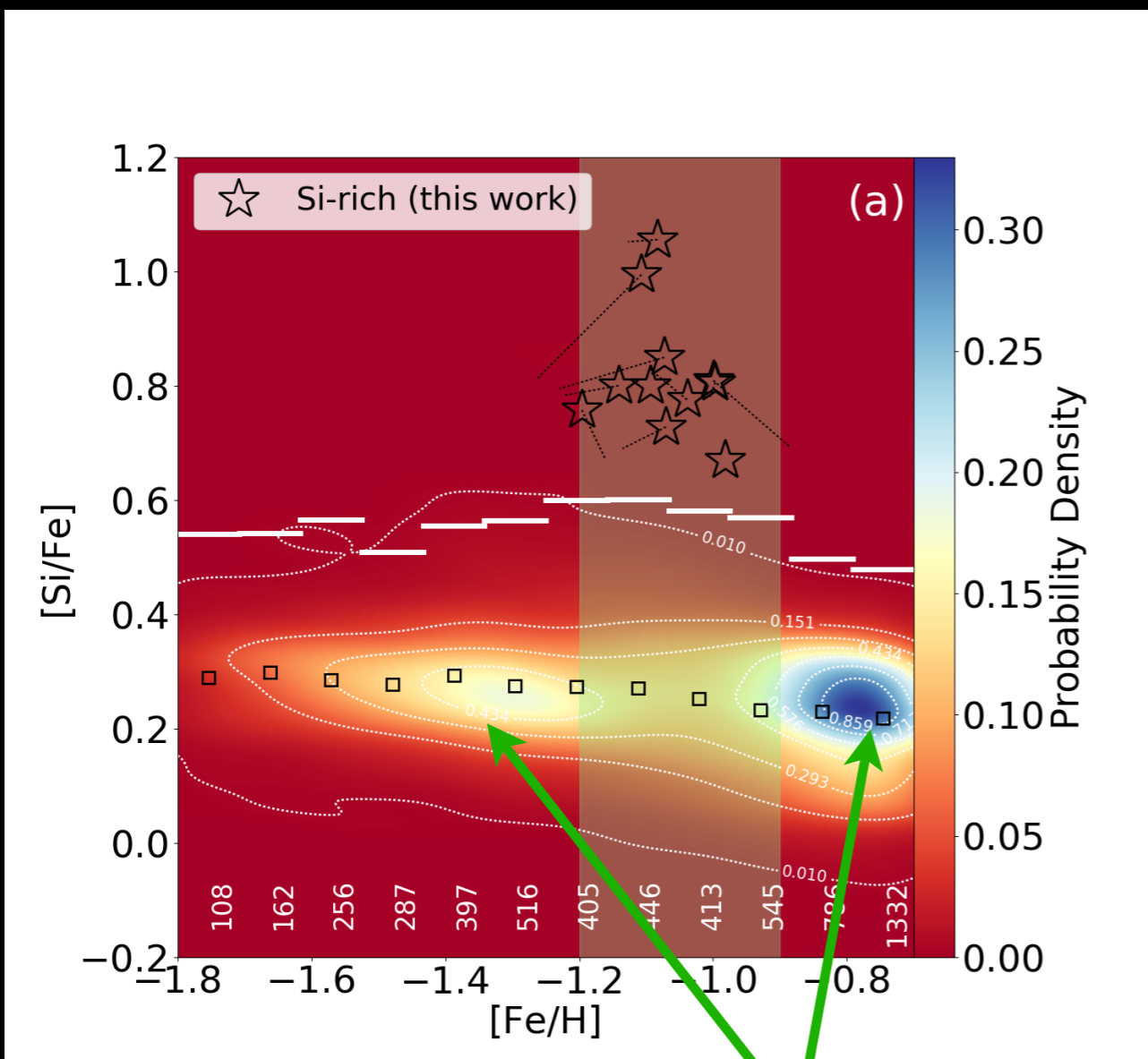
No evidence of Dwarf galaxies with $[Al/Fe] > +0.5$



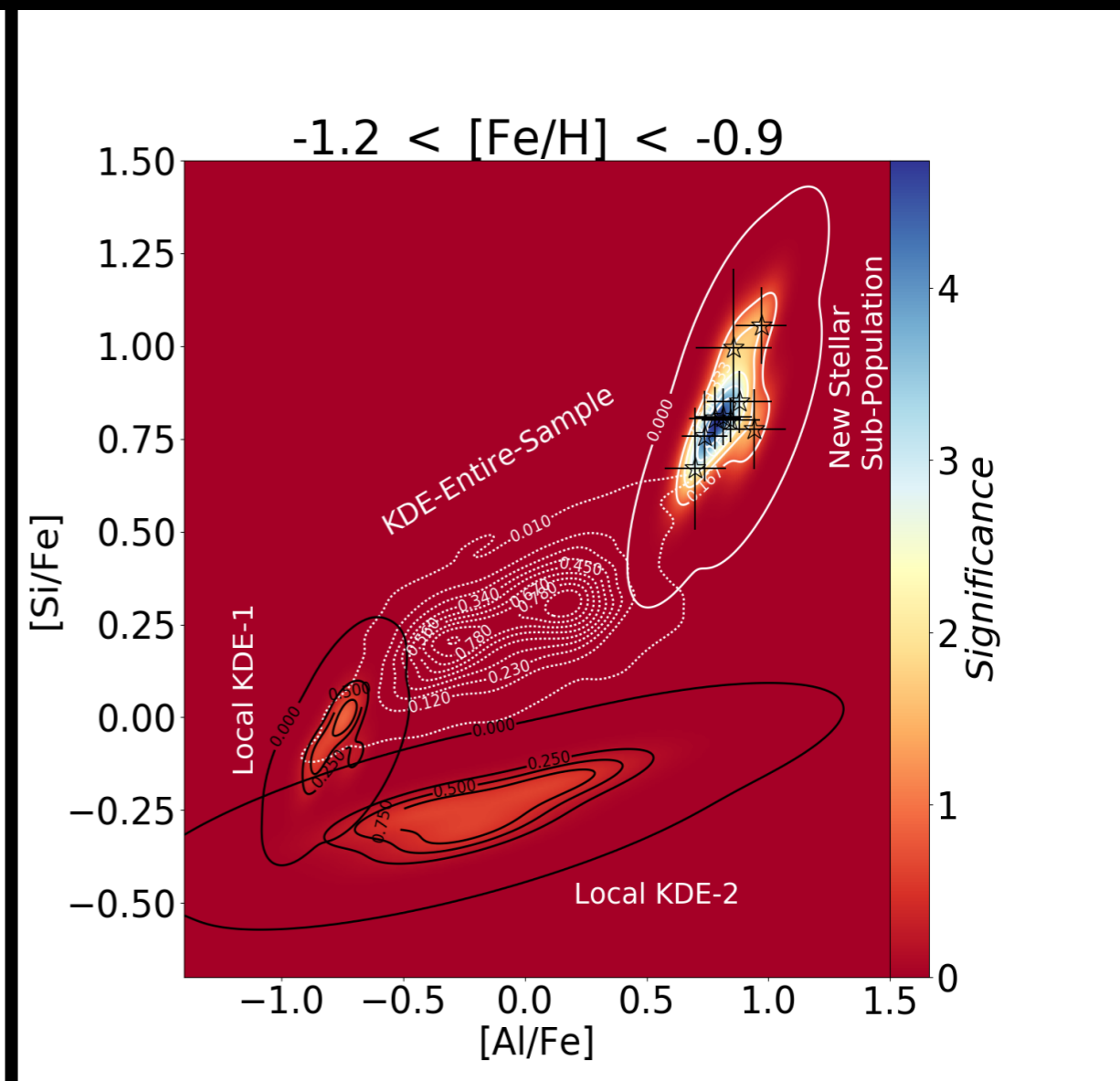
Discovery of a Large Population High-metallicity Aluminium+Nitrogen rich stars in the Inner Halo



Discovery of a New Sub-stellar population residing in the (inner) halo: Jurassic structure



Canonical (thick) disk & Halo



Fernández-Trincado et al., 2019, ApJL, 886, L8
 Fernández-Trincado et al. 2020, A&A, 644, A83

NGC 6723



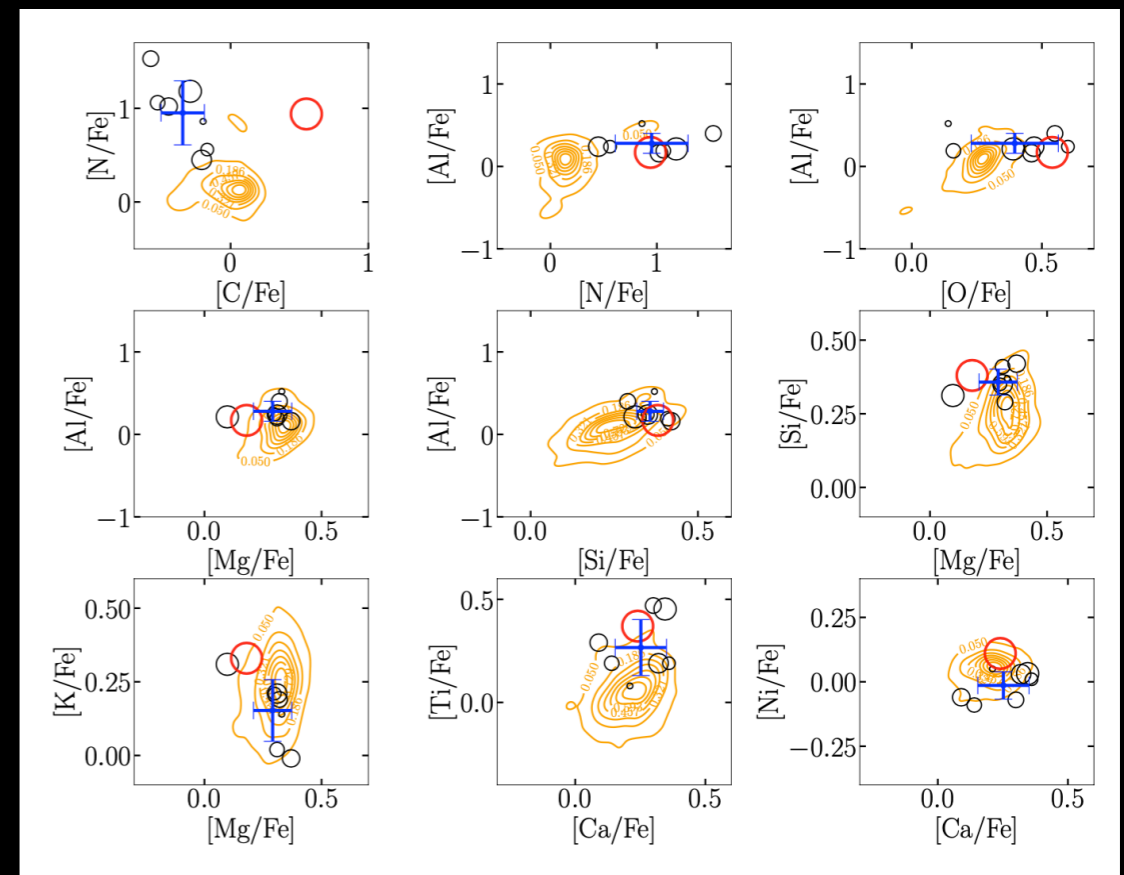
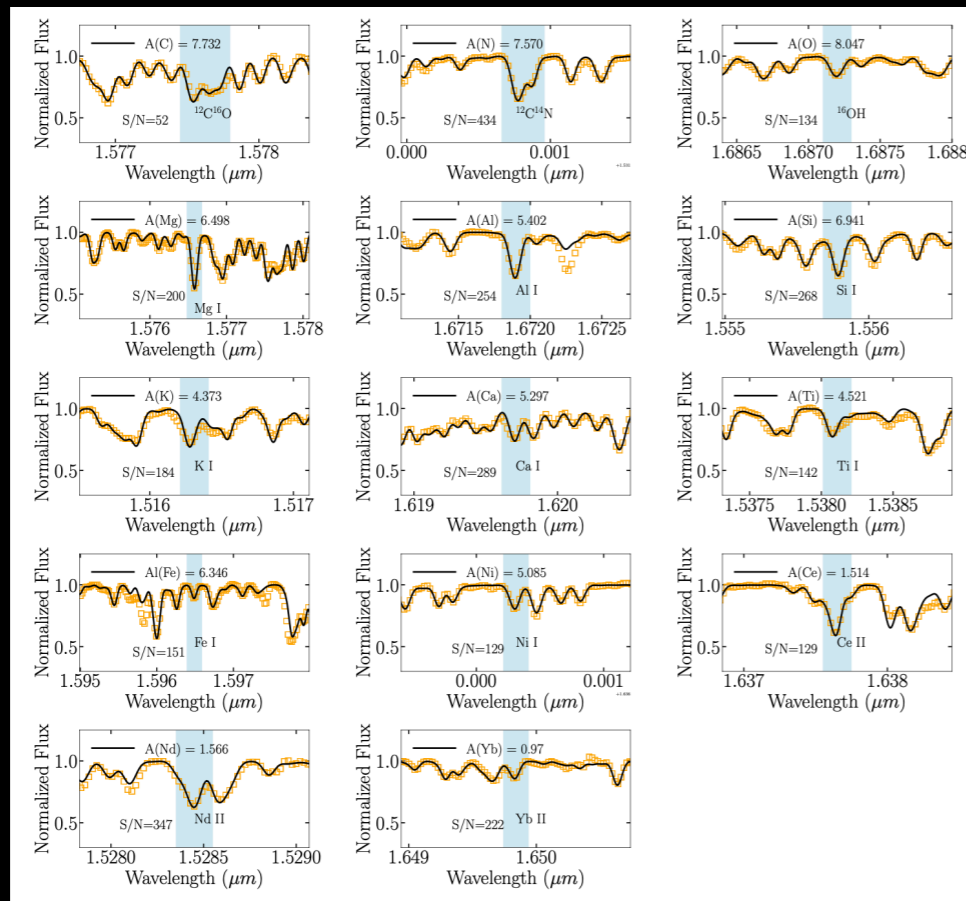
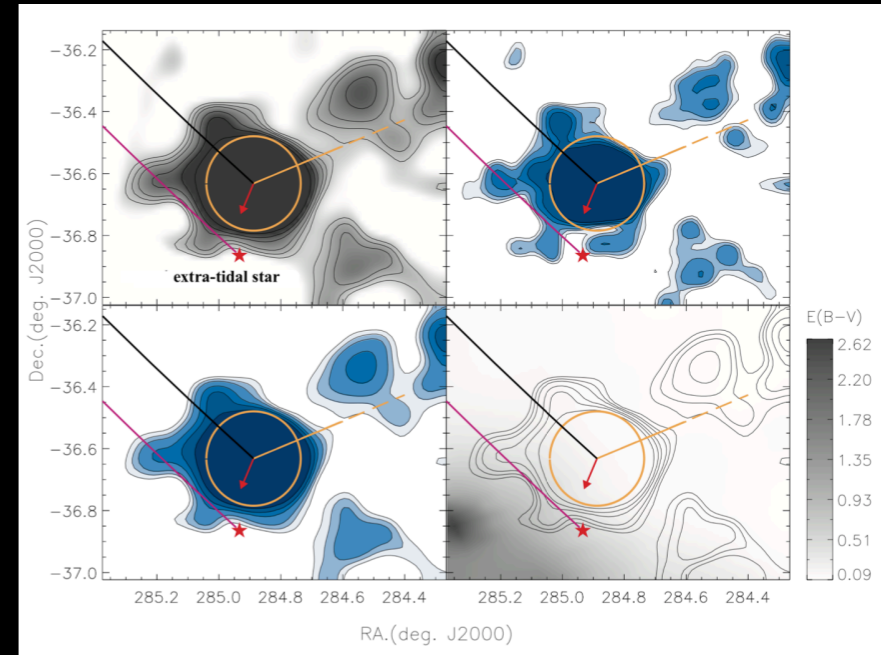
Image Credit: NASA, STScI, WikiSky

APOGEE DR16 Spectra

$[\text{Fe}/\text{H}] = -1.0 \pm 0.06$ dex

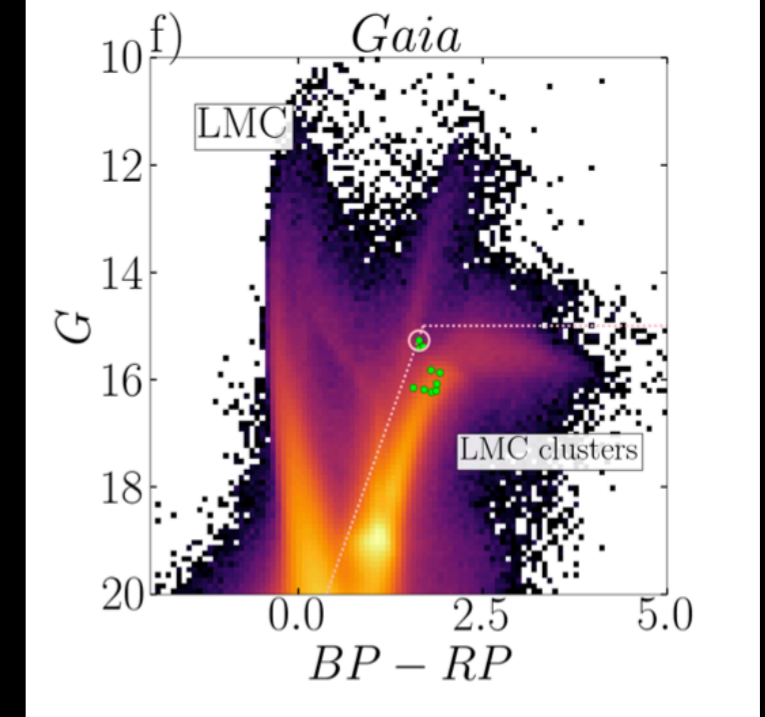
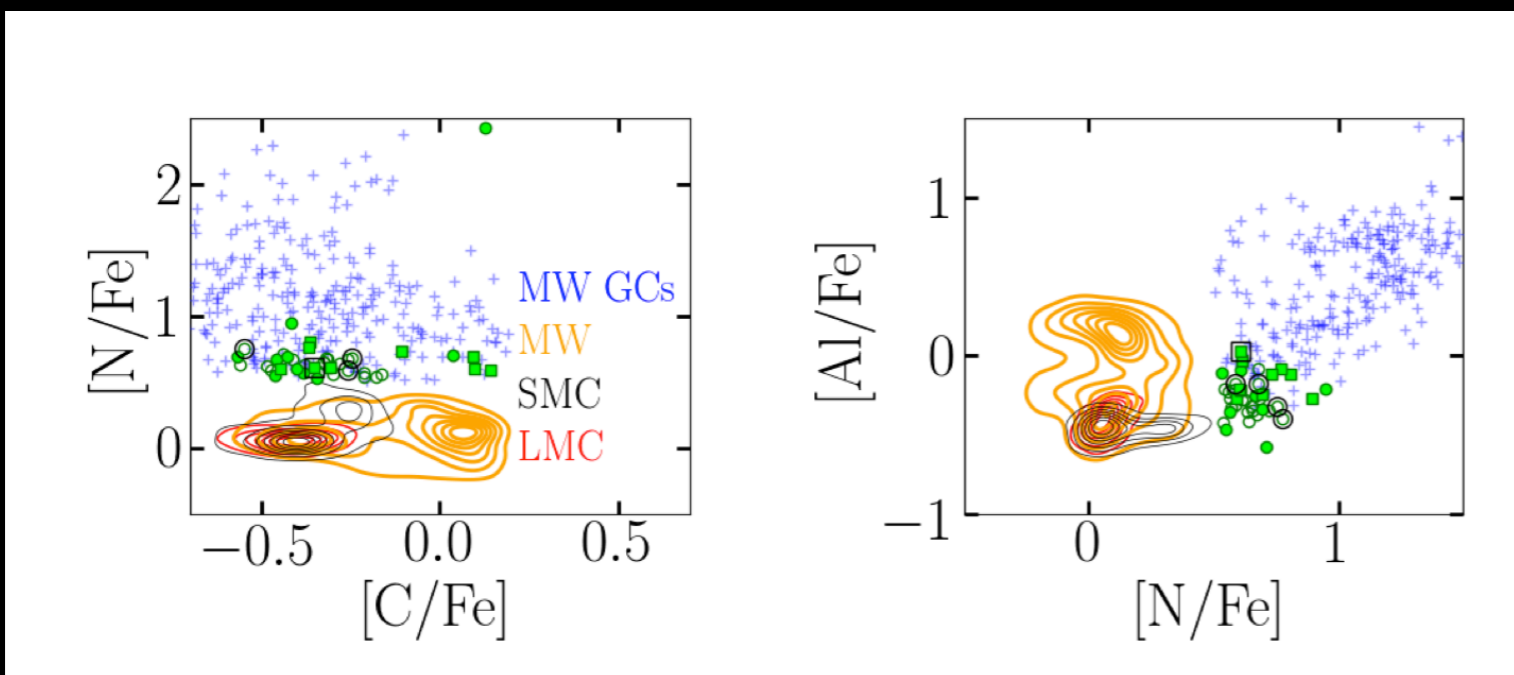
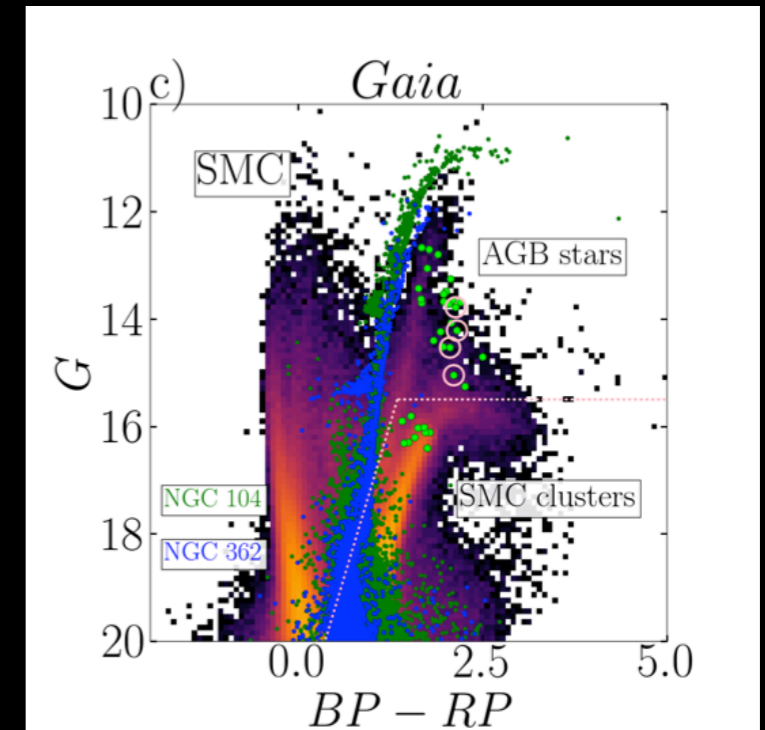
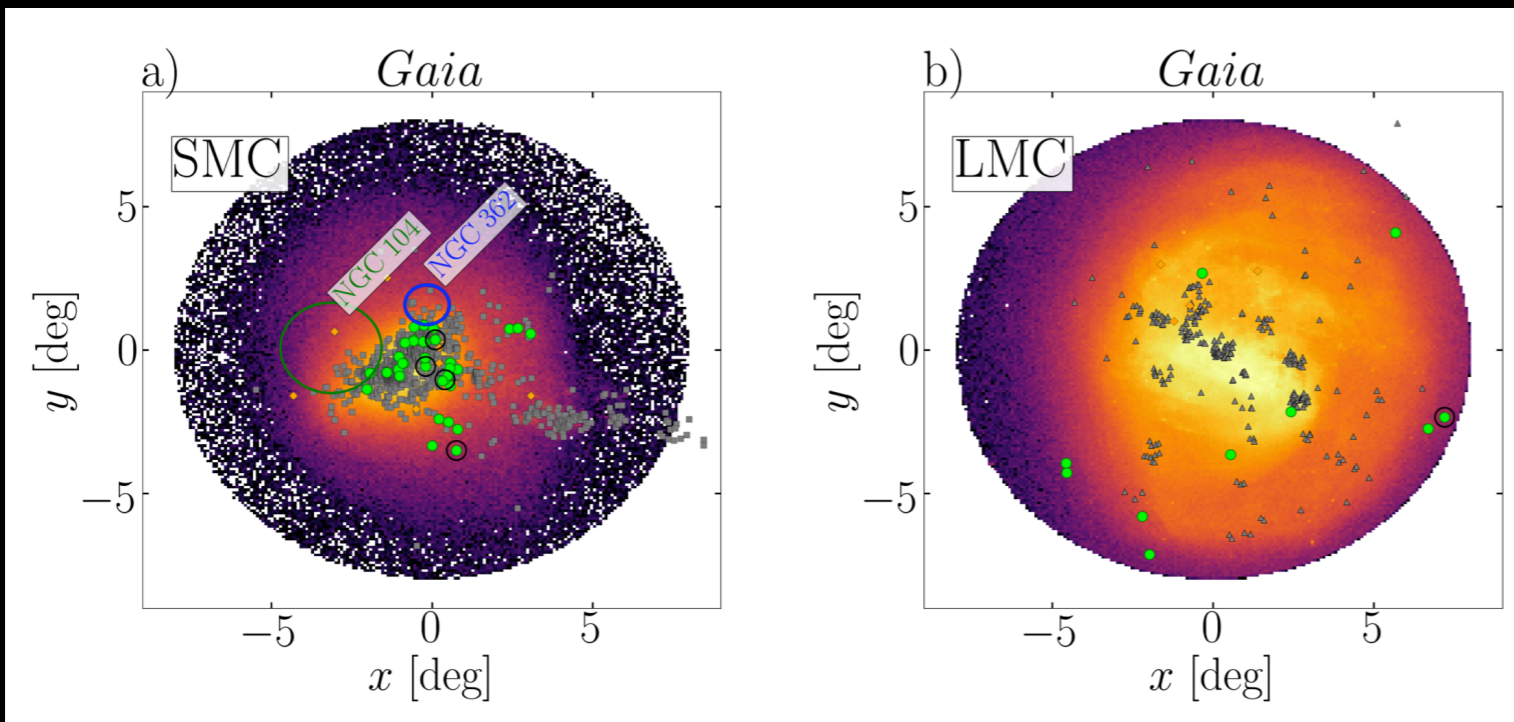
$E(B-V) \sim 0.05$ mag

Extra-tidal
N-rich Star

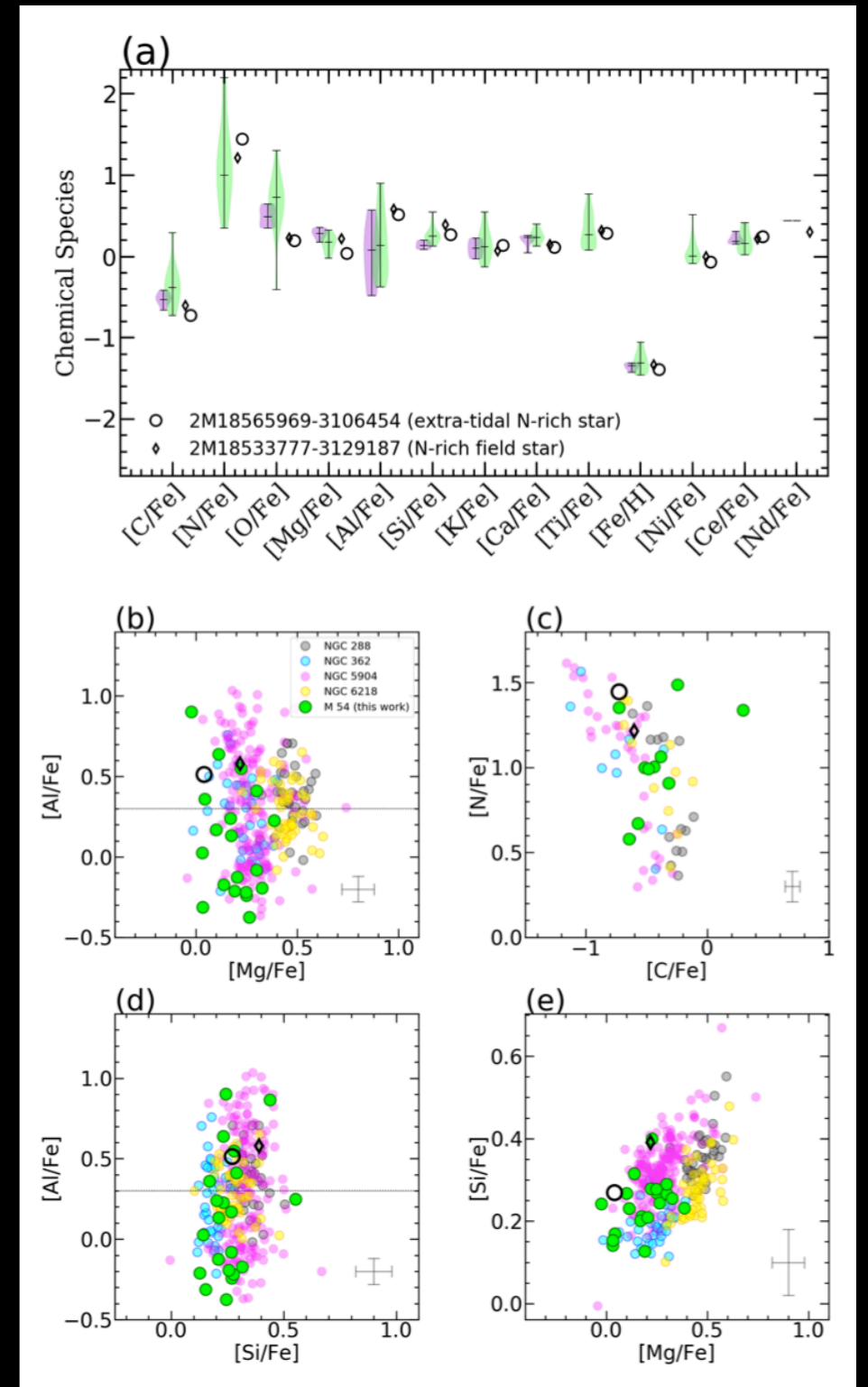
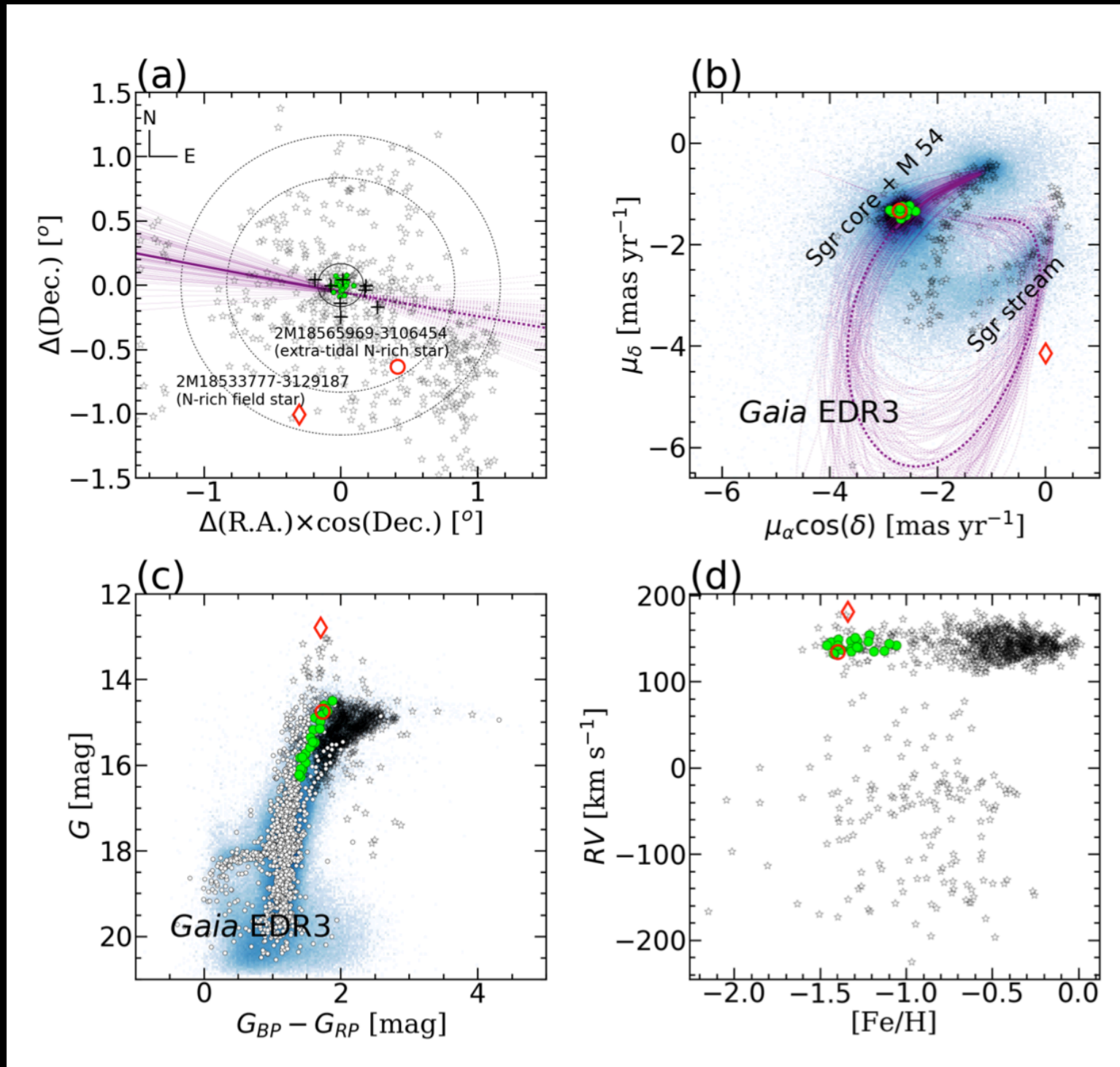


What about Dwarf Galaxies?

Discovery of Nitrogen-rich stars in the Large and Small Magellanic Clouds



Discovery of Nitrogen-rich stars in the Sagittarius Dwarf Galaxy



About ~15% of the Milky Way halo is filled by a Zoo of Chemically unusual giants

Anomalous Metal-Poor Field Stars,
(-1.8, -0.7] dex

N-rich, $> +0.5$

N-normal, $< +0.5$

Al-rich, $> +0.5$

Al-normal, $< +0.5$

Si-rich, $> +0.6$

Si-normal, $< +0.6$

- ◆ *GCs-like abundance patterns*
- ◆ *Extragalactic-like patterns*

- ◆ *GCs/dSph-like abundance patterns*
- ◆ *Evidence for Binary mass transfer*

GC-like abundance patterns

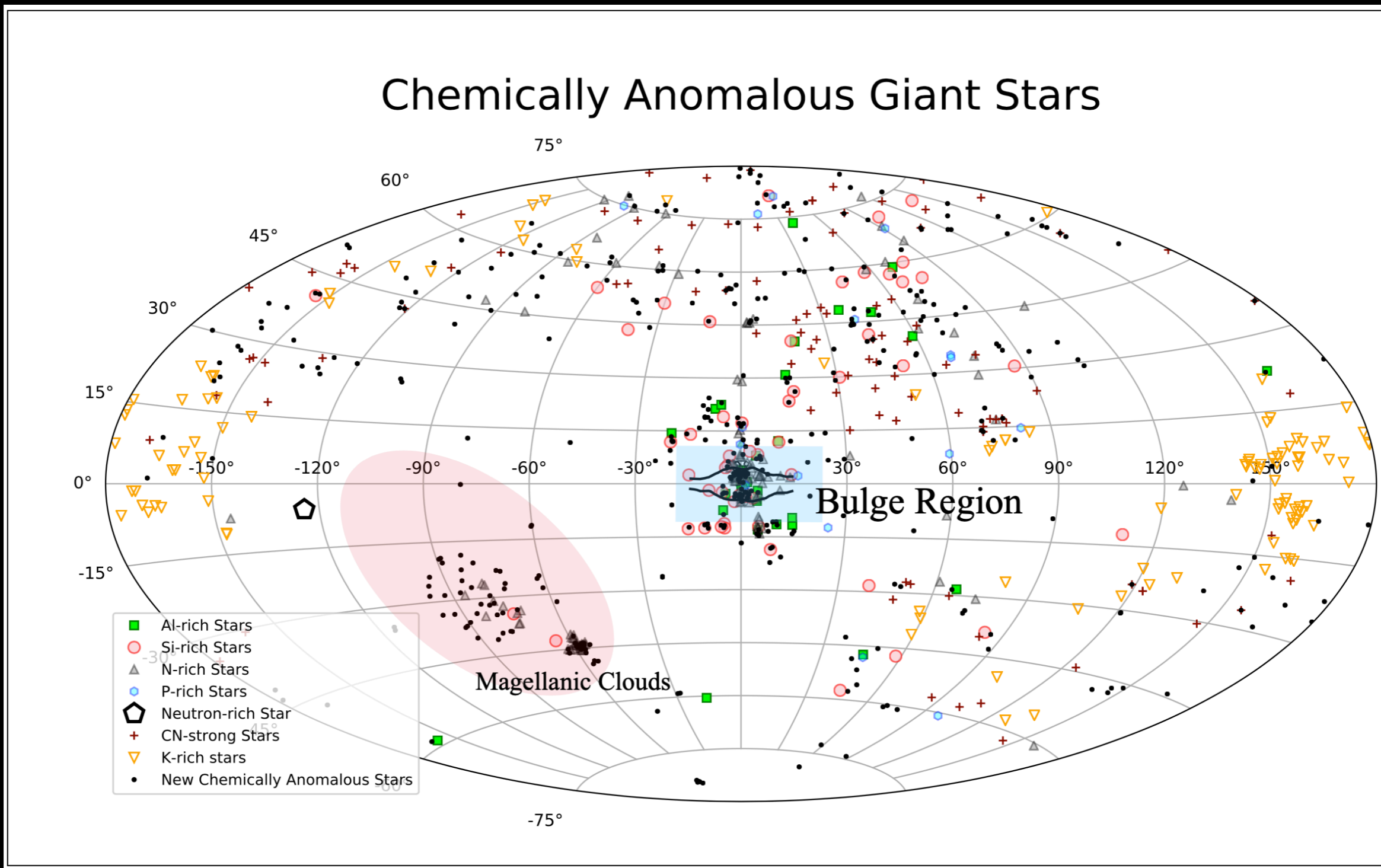
Light-Elements Galactic levels

- ✓ *Such stars may have resulted from the destruction of GC populations formed *in situ* and *ex situ*.*
- ✓ *Such stars were not perforce associated with Galactic GCs, but formed in similar environments, while never being gravitationally bound to the Gas themselves.*

What is Next?

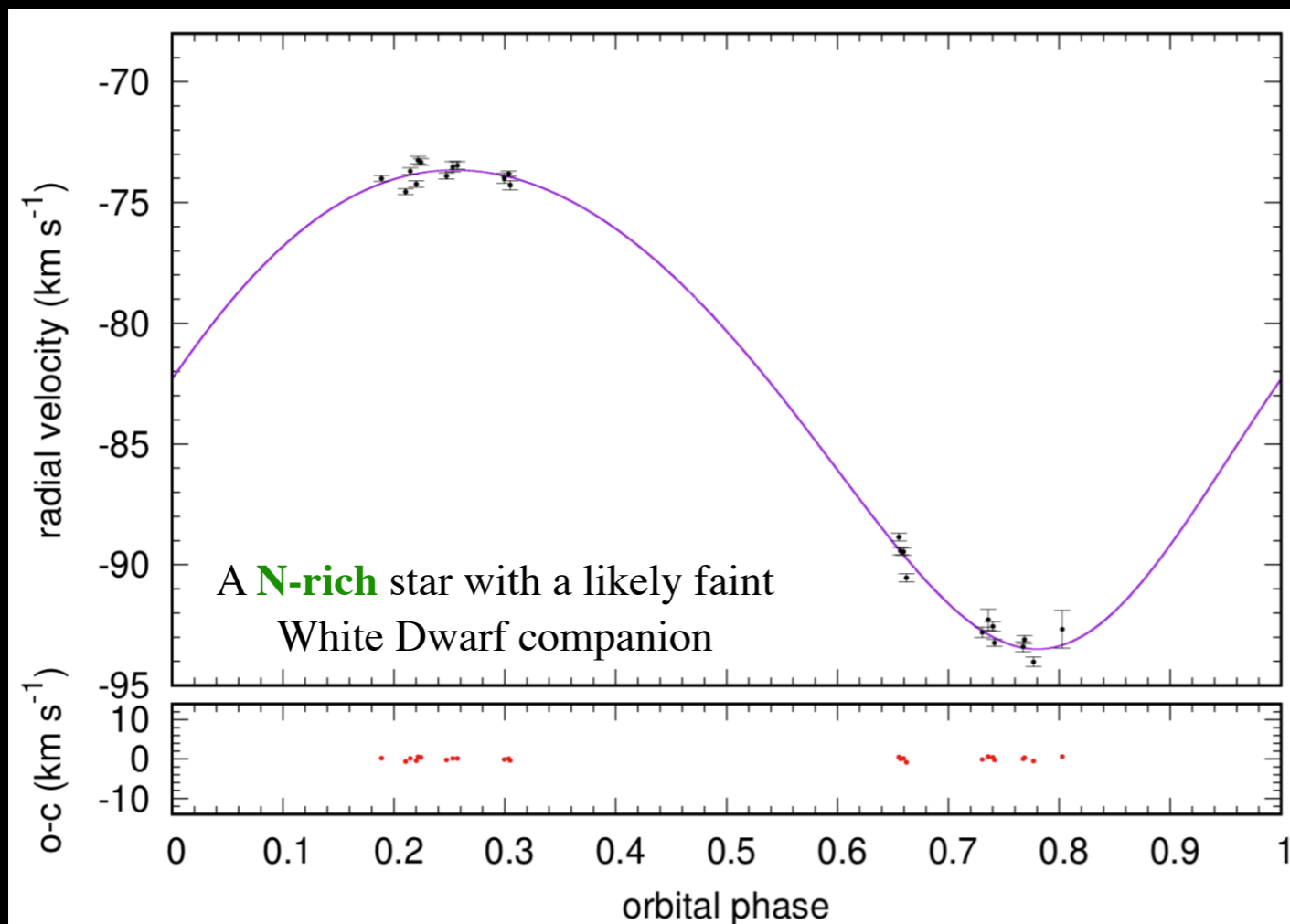
Radial velocities for chemically anomalous stars

PI: José G. Fernández-Trincado

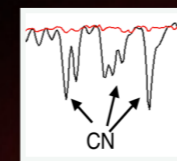
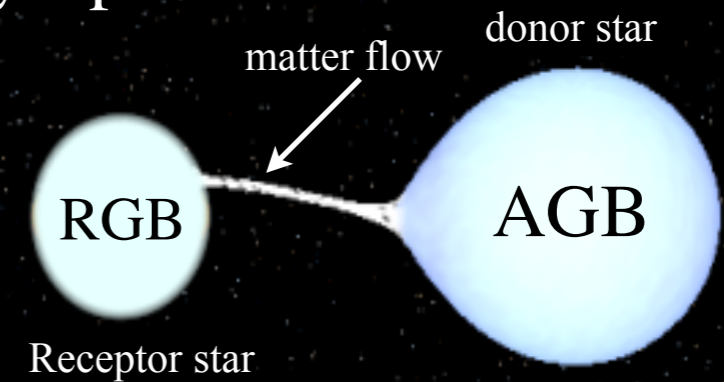


A Binary **N-rich field** companion of evolved star

Large variability in Radial Velocity



Early Epoch



Observed Binary system today!

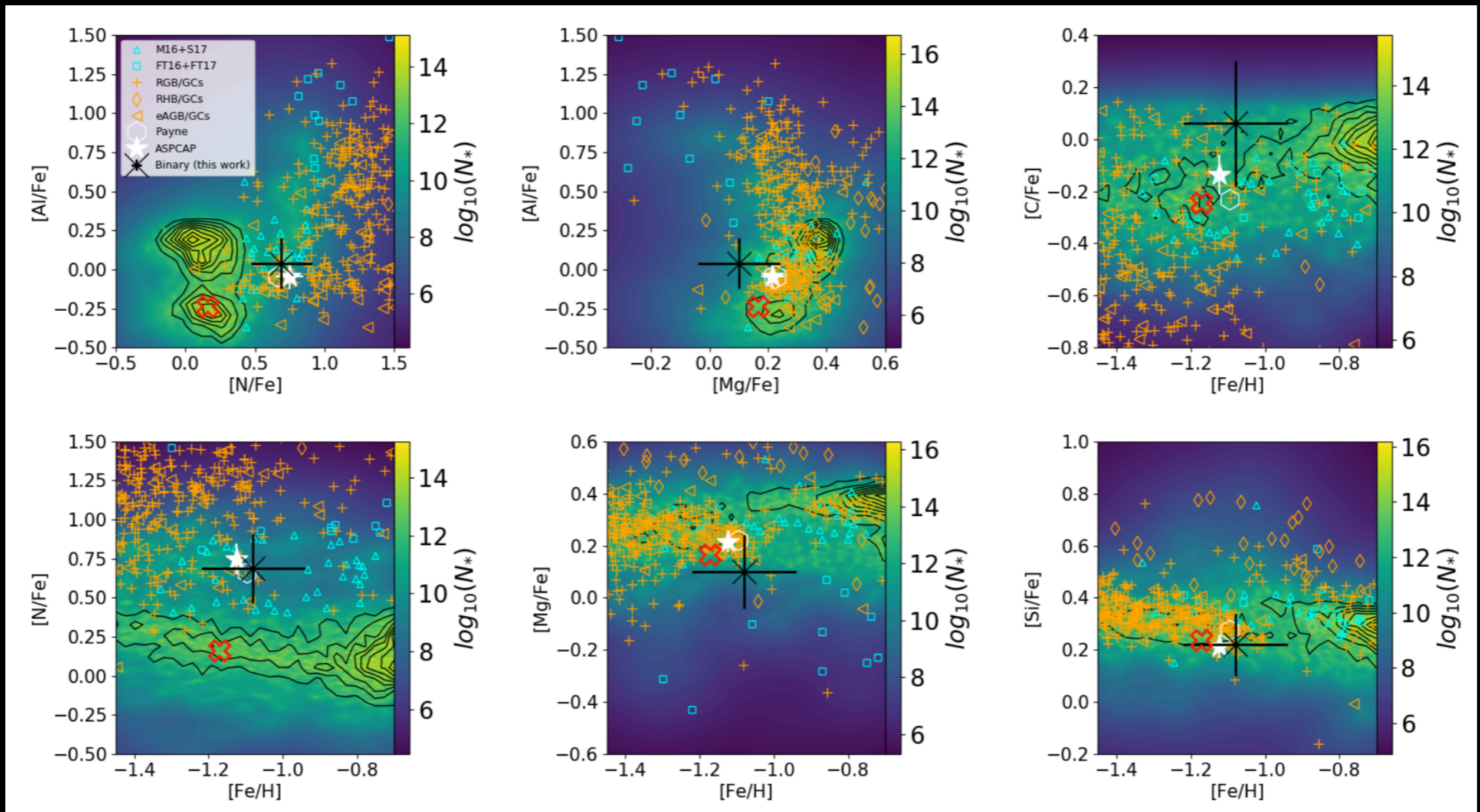


RGB



White dwarf

A Binary **N-rich field** companion of evolved star from the **APOGEE** survey





SDSS-V extra-tidal globular cluster stars

PI: José G. Fernández-Trincado

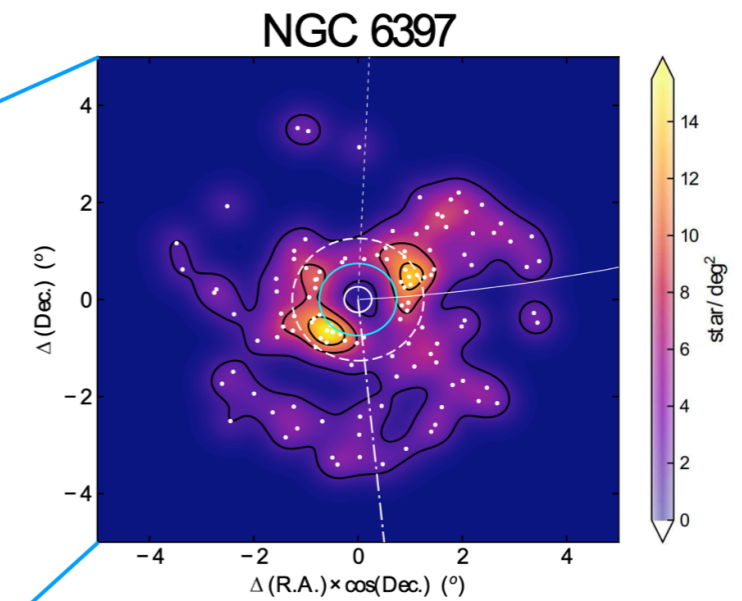
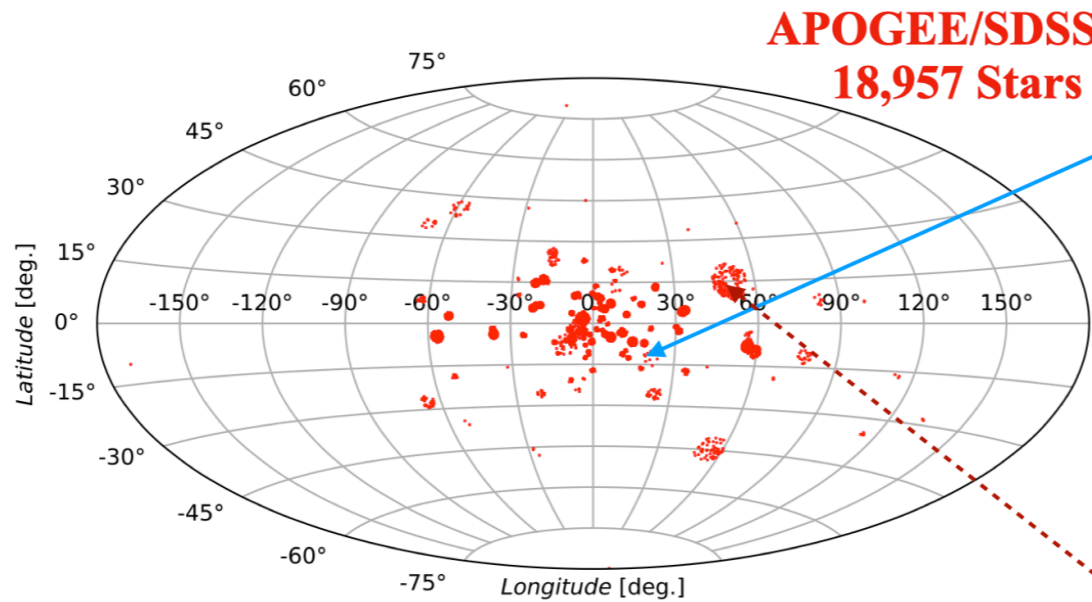


Figure taken from
Kundu et al. 2021, A&A, 645, A116

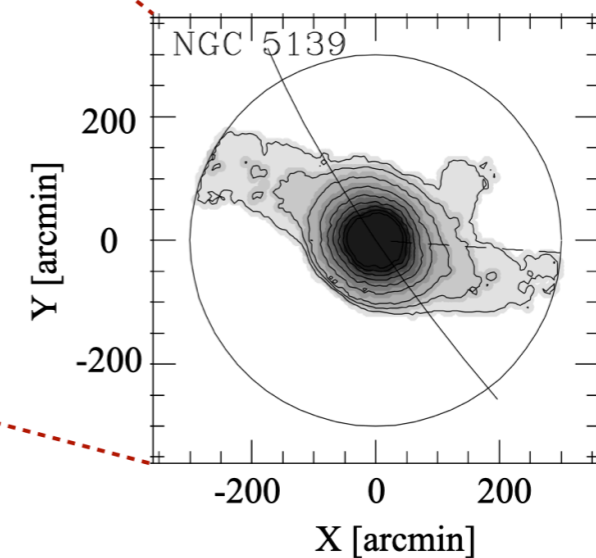
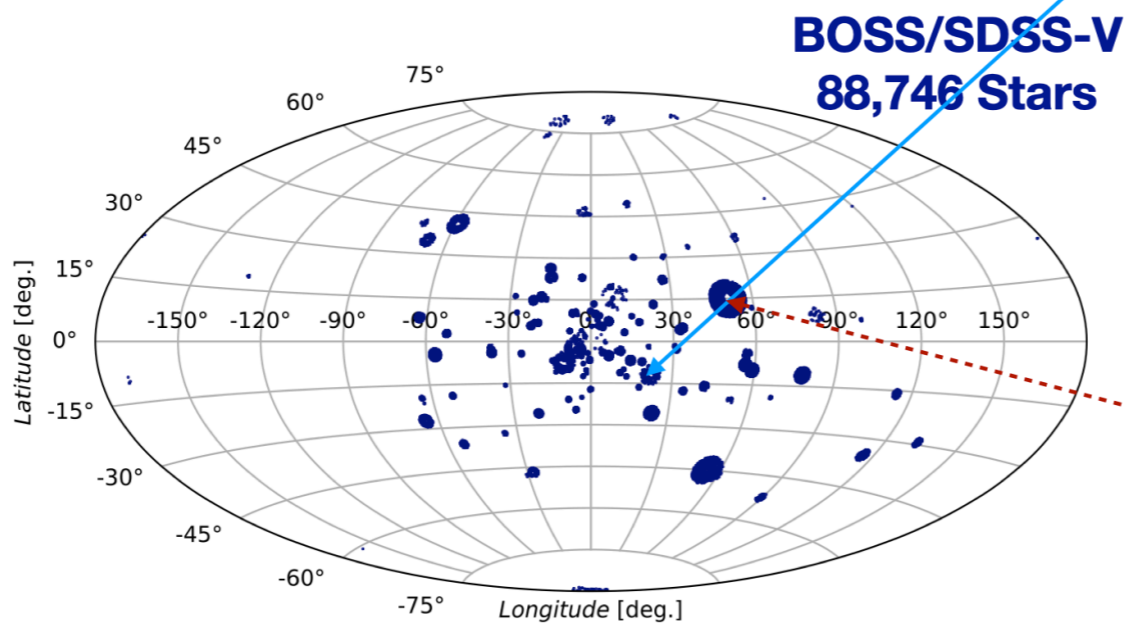


Figure taken from
Sollima et al. 2020, MNRAS, 495, 2222-2233