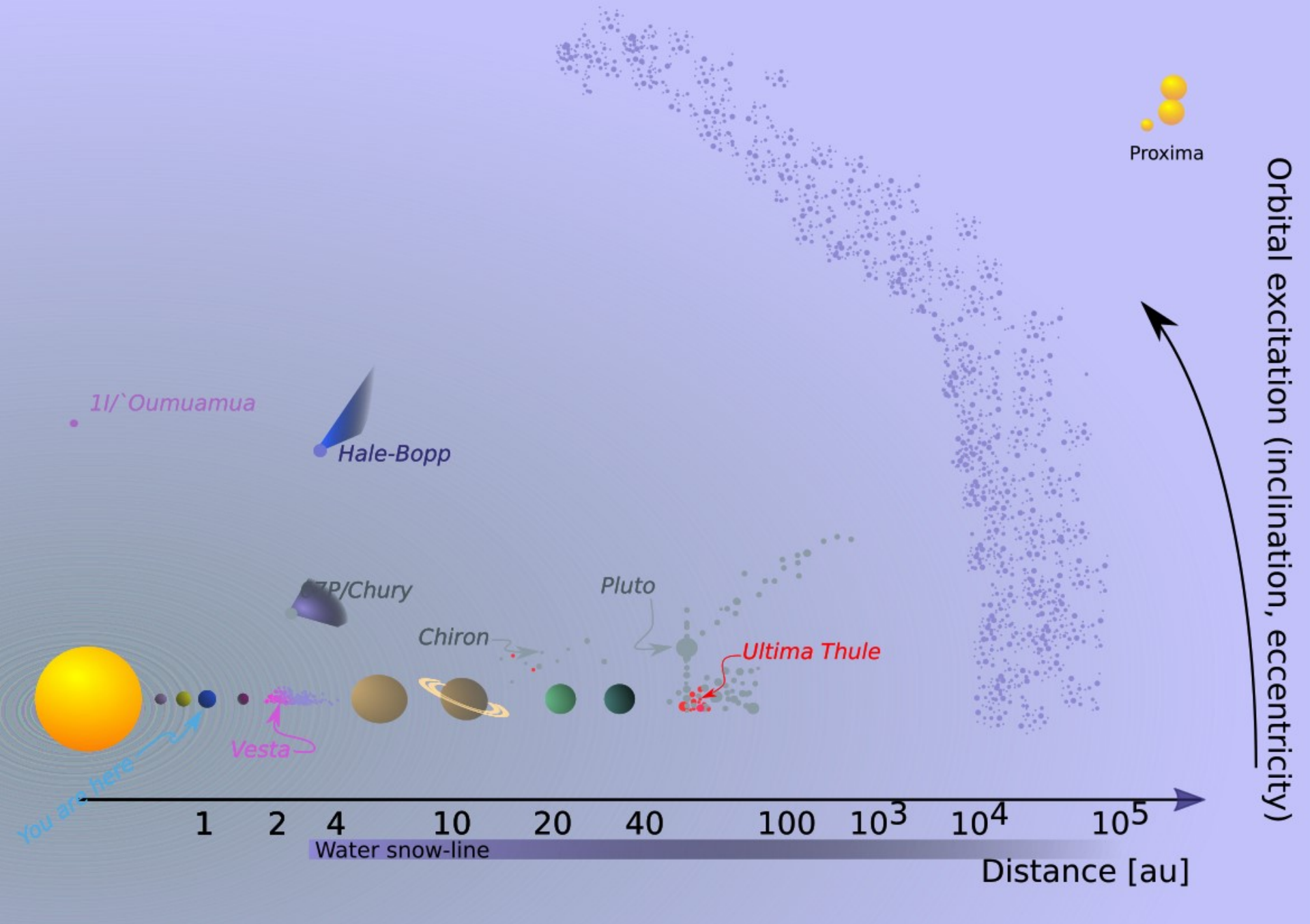
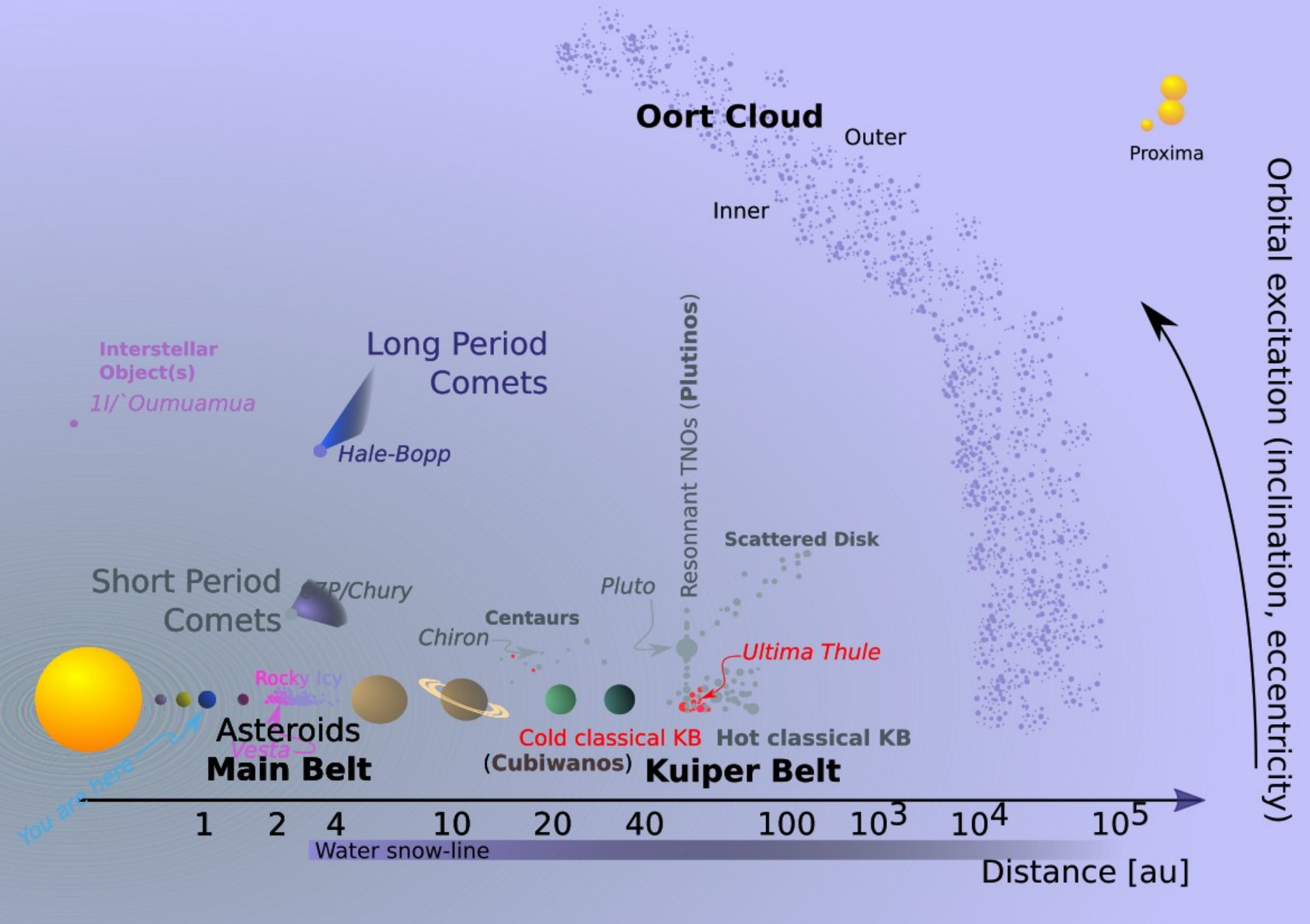
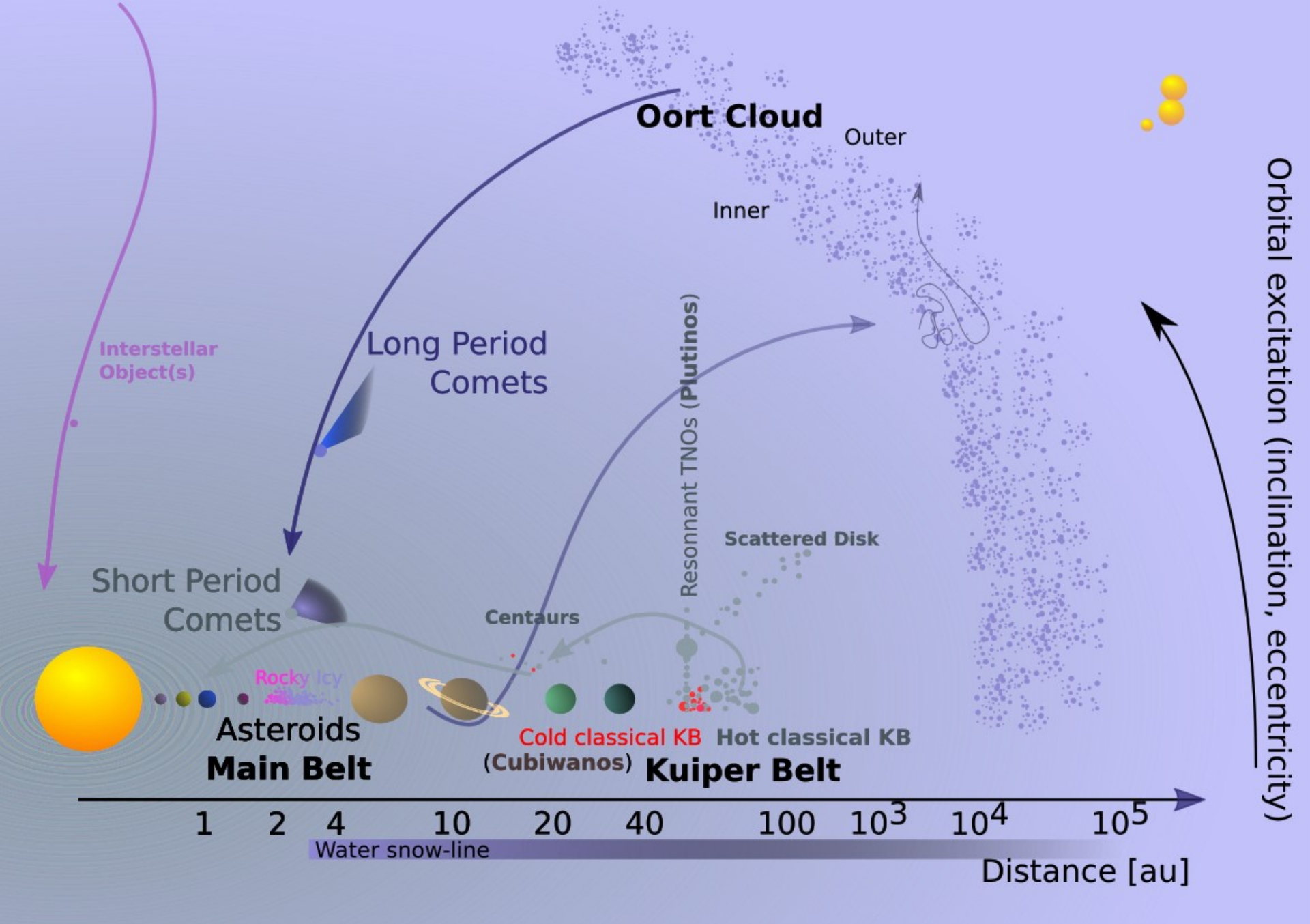


Minor Bodies in our Solar System, as seen by FORS

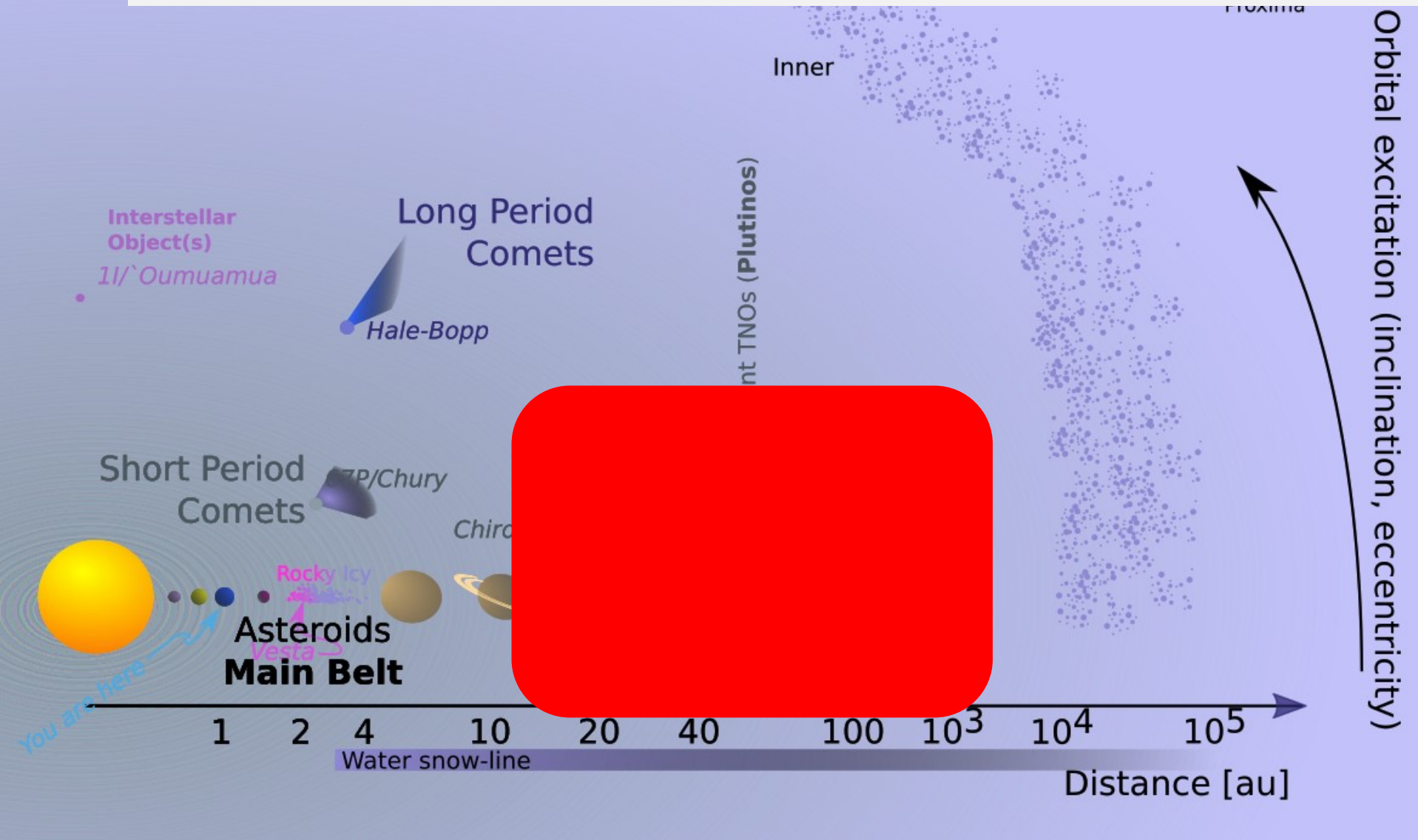
Olivier Hainaut | ESO | Public





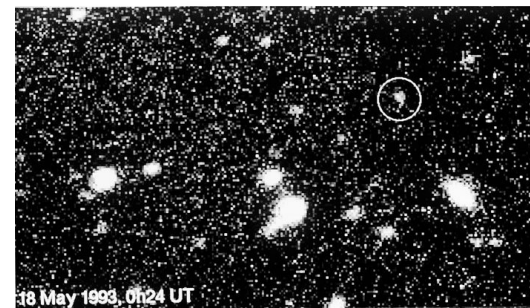
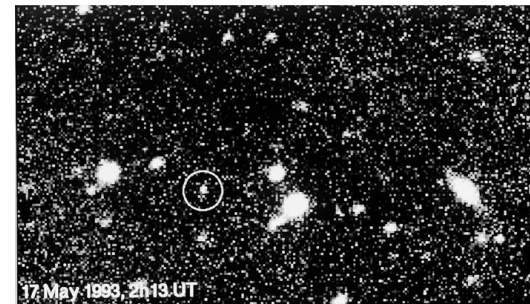
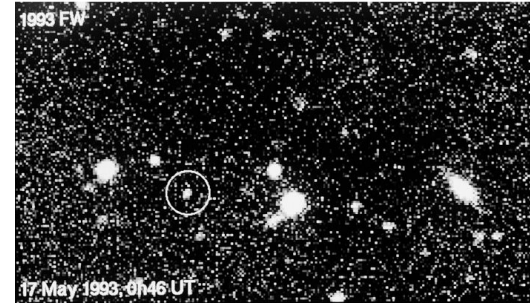


TNO population



TNO population

- First TNO (after Pluto): 1992 QB1, Hawai`i
 - Second one: 1993 FW
 - By 2000: 100 = a population
 - Today: 2000 objects
- ESO large program for the characterization of the TNOs
- 2d ESO large program for the characterization of the TNOs



1993 FW | eso9305

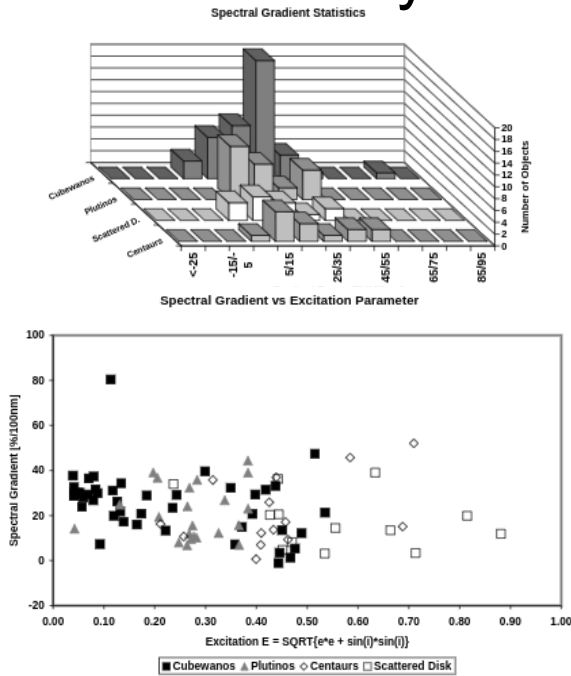
TNO LP1 and LP2

- 2001A&A...380..347D - **BVRI Photometry of 27 Kuiper Belt Objects with ESO/Very Large Telescope** - Delsanti, A. C.; Boehnhardt, H.; Barrera, L. et al.
- 2002A&A...395..297B - **ESO large program on physical studies of Transneptunian Objects and Centaurs: Visible photometry - First results** -Boehnhardt, H.; Delsanti, A.; Barucci, A. et al.
- 2004A&A...421..353F - **ESO Large Program on physical studies of Trans-Neptunian objects and Centaurs: Final results of the visible spectrophotometric observations** - Fornasier, S.; Doressoundiram, A.; Tozzi, G. P. et al.
- 2004A&A...417.1145D - **Simultaneous visible-near IR photometric study of Kuiper Belt Object surfaces with the ESO/Very Large Telescopes** - Delsanti, A.; Hainaut, O.; Jourdeuil, E. et al.
- 2010A&A...510A..53P - **Colors and taxonomy of Centaurs and trans-Neptunian objects** - Perna, D.; Barucci, M. A.; Fornasier, S. et al.
- 2008A&A...487..741A - **Visible spectroscopy of the new ESO large program on trans-Neptunian objects and Centaurs. Part 1** - Alvarez-Candal, A.; Fornasier, S.; Barucci, M. A. et al
- 2009A&A...493..283D - **Visible and near-infrared colors of Transneptunian objects and Centaurs from the second ESO large program** - DeMeo, F. E.; Fornasier, S.; Barucci, M. A. et al
- 2009A&A...508..457F - **Visible spectroscopy of the new ESO large programme on trans-Neptunian objects and Centaurs: final results** - Fornasier, S.; Barucci, M. A.; de Bergh, C. et al
- Etc....

TNO population

■ Beginnings

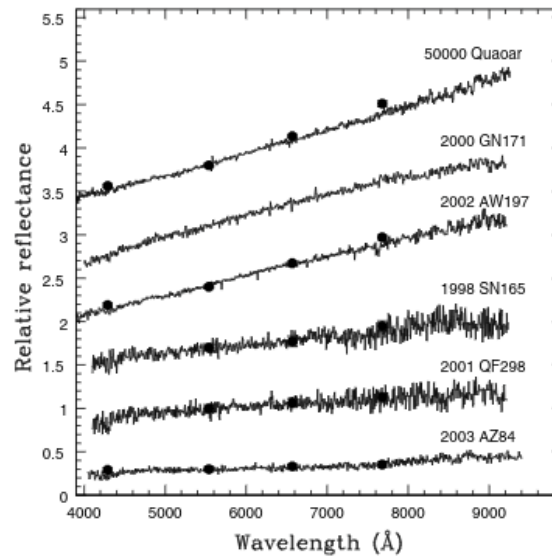
➤ Photometry



Boehnhardt+ 2002A&A...395..297B

Delsanti+ 2004A&A...417.1145D
Perna+ 2001A&A...380..347D

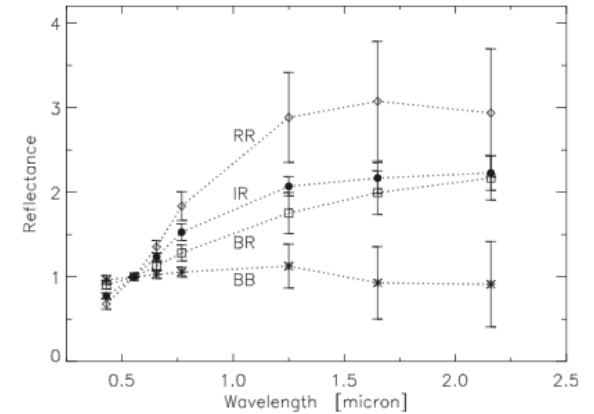
➤ Spectroscopy



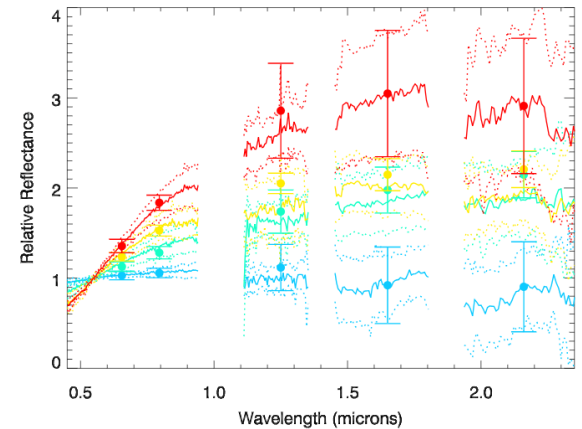
Fornasier+ 2004A&A...421..353F

Alvarez-Candal+ 2008A&A...487..741A
Fornasier+ 2009A&A...508..457F

➤ Taxonomy



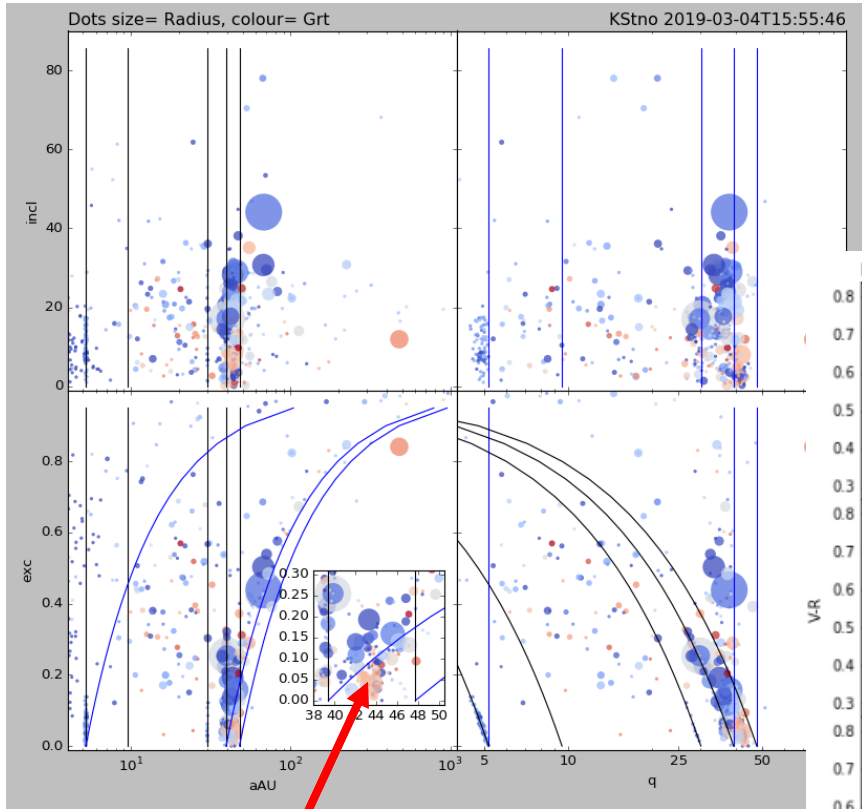
Perna+ 2010A&A...510A..53P



Merlin+ 2017A&A...604A..86M

TNO population

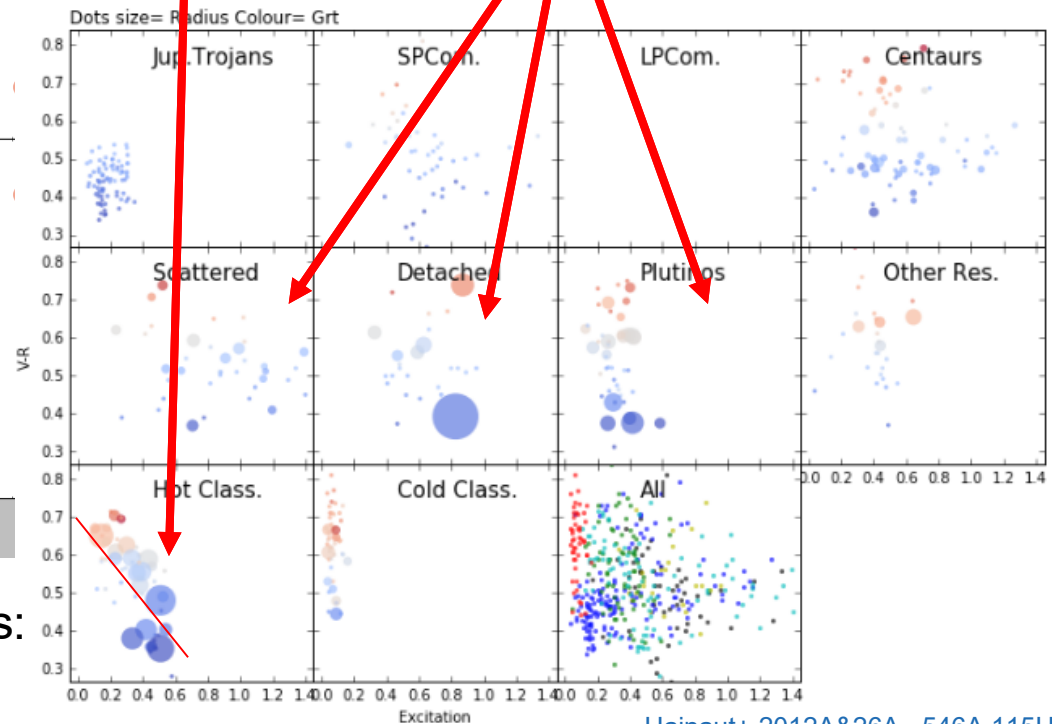
Physical understanding of the population as a whole



A primordial population of red bodies:
Cold Classical TNOs

A trend with orbital excitation

3 populations with the same origin



Physical understanding sub-populations Ultra-wide TNO binaries

"We find that gravitational collapse binary formation models produce an orbital distribution similar to that currently observed, which along with a confluence of other factors supports **formation of the cold Classical Kuiper Belt in situ through relatively rapid gravitational collapse** rather than slow hierarchical accretion."

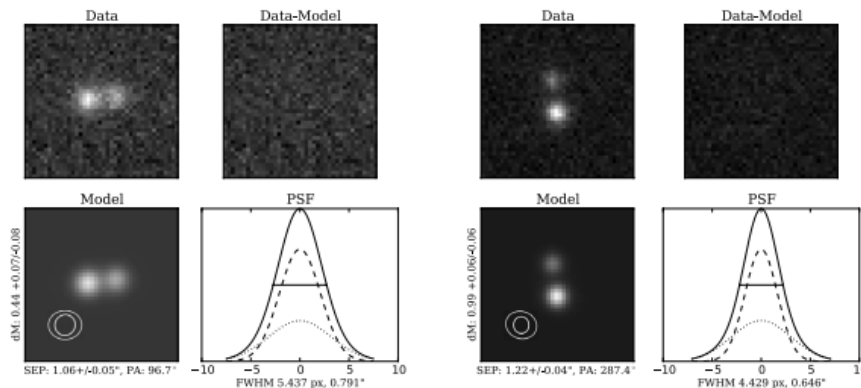
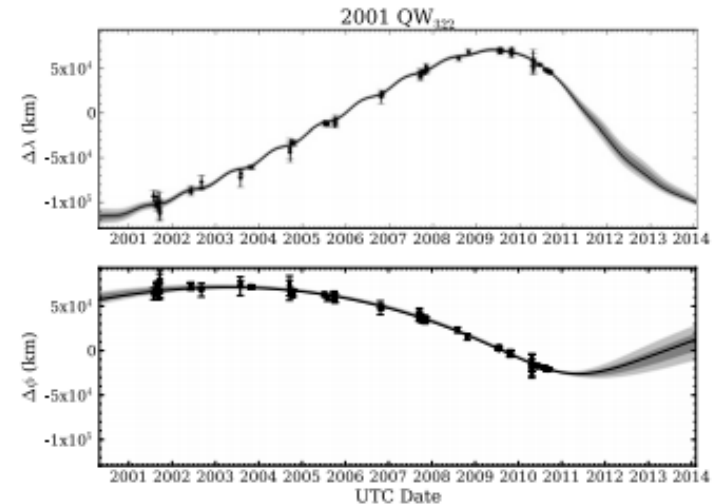


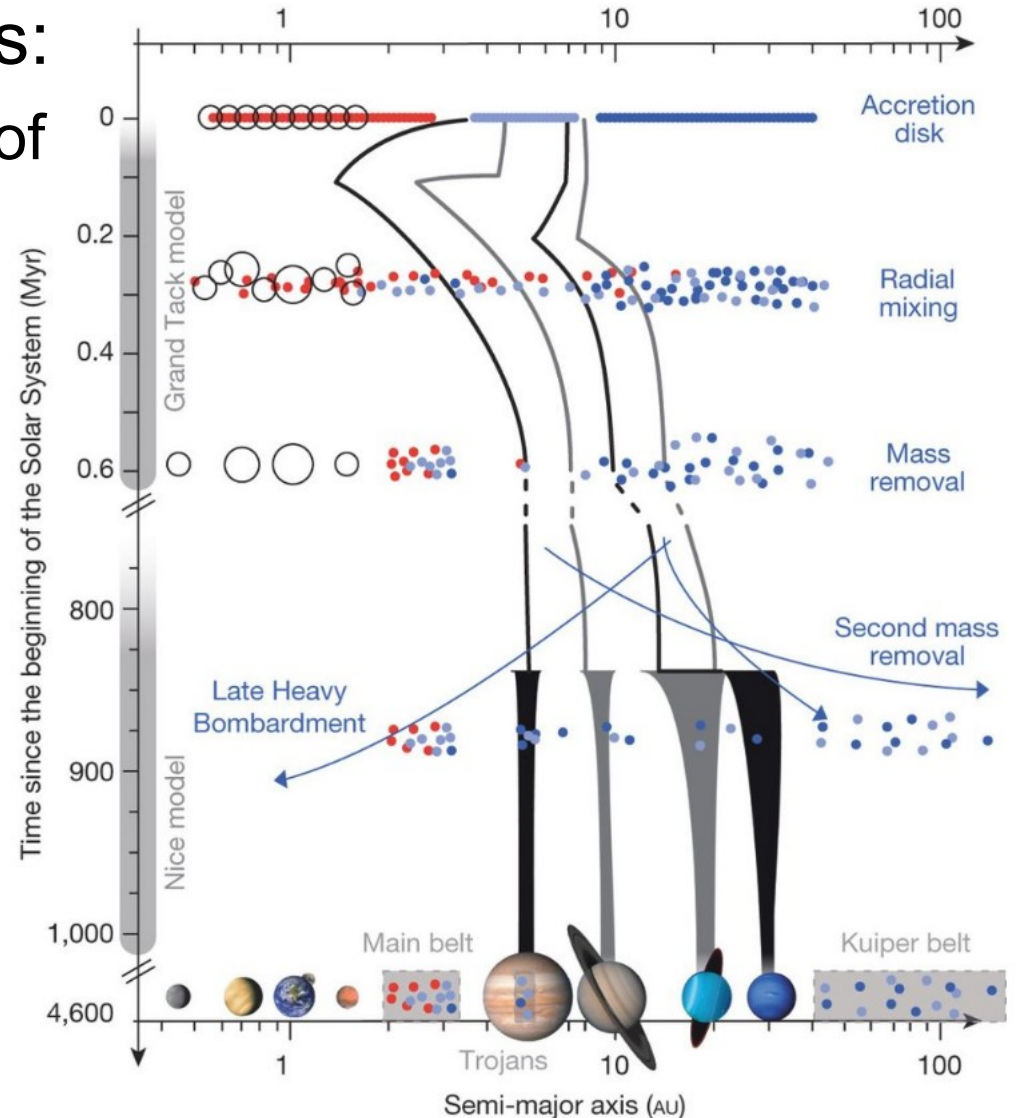
Figure 2. Same as Figure 1, but for CFEPS binaries 2006 BR₂₈₄ (left) and 2006 JZ₈₁ (right).



TNO population

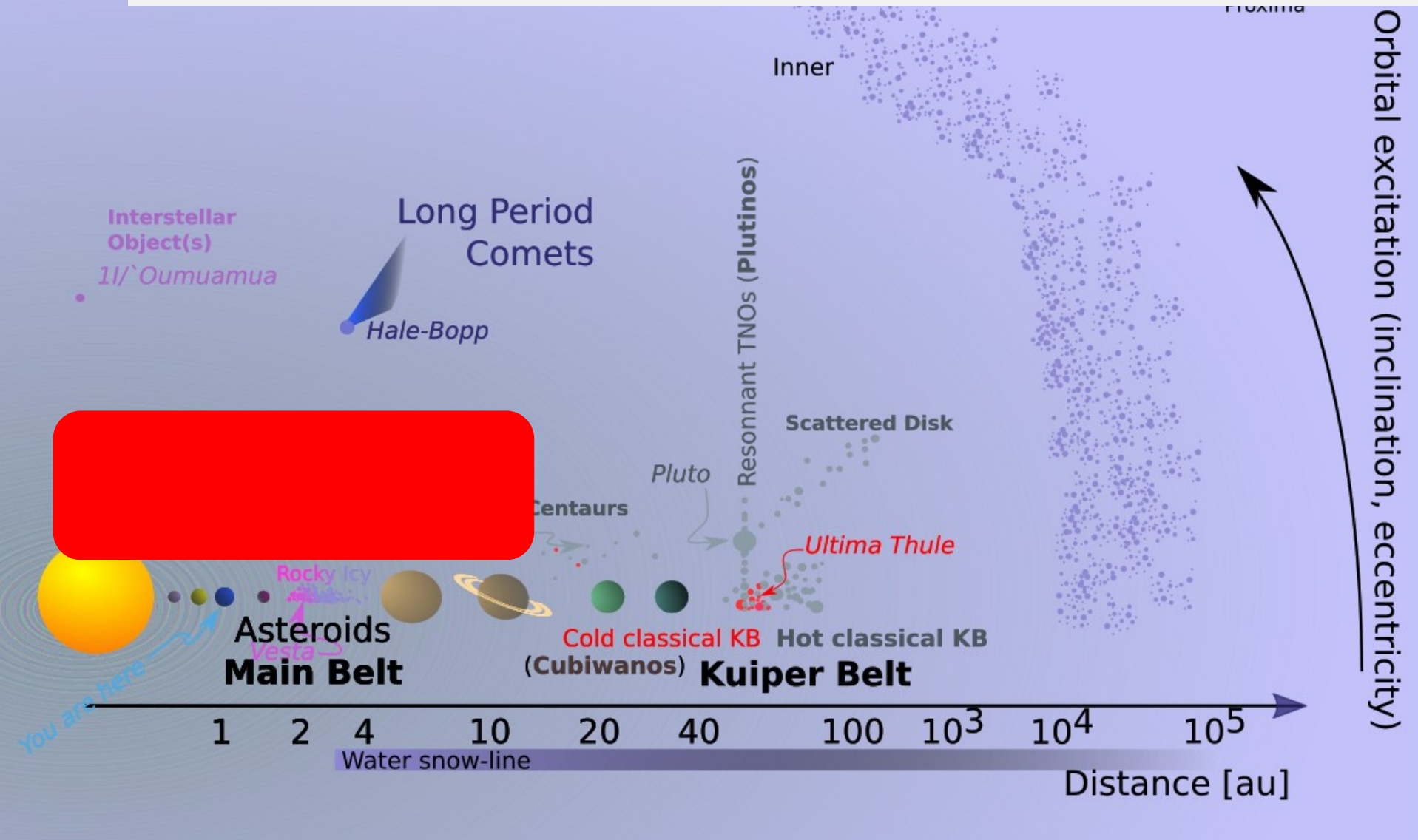
■ Constrain other fields:

- dynamical evolution of early Solar System,
- planet migration,
- architecture of the Asteroid Belt,
- exoplanet system architecture...



DeMeo & Cary 2014 Nature 505, 629

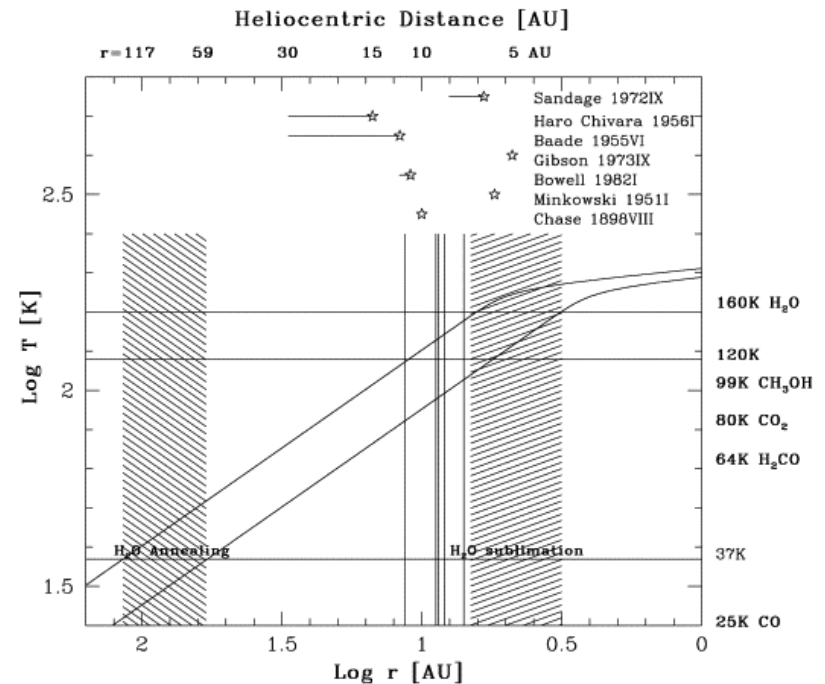
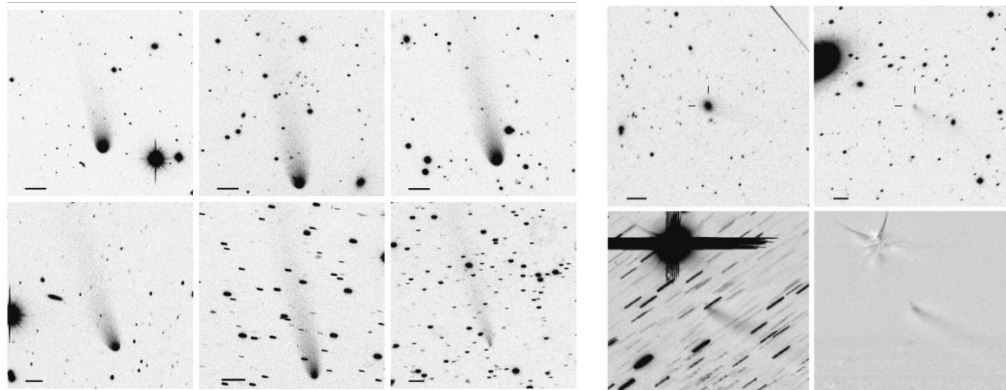
Comets



Comets Population

- Cometary Activity at large heliocentric distances: Start/end of activity (FORS!)
- Thermal models
- ==> activity processes and composition

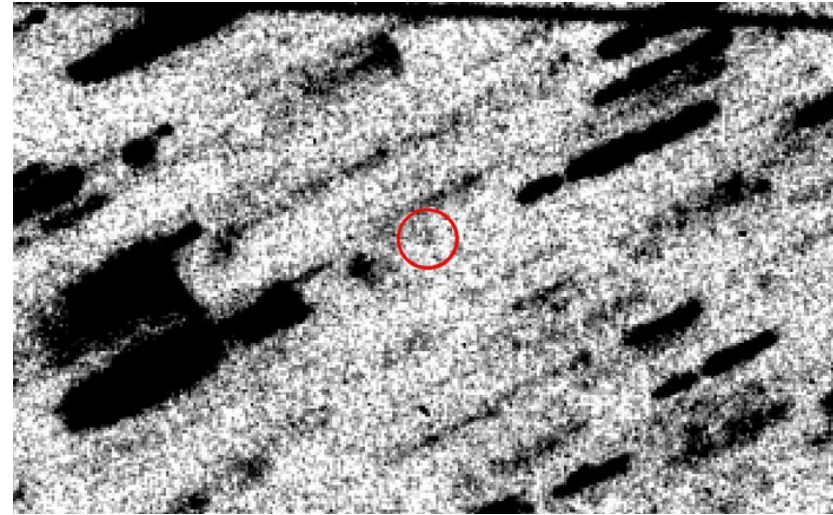
Meech+ 2009 Icar. 201..719M



Comet Highlight

1P/Halley

- FORS1 + FORS2 + VIMOS, 32ks
- $V = 28.2$, $r = 28.1$ au

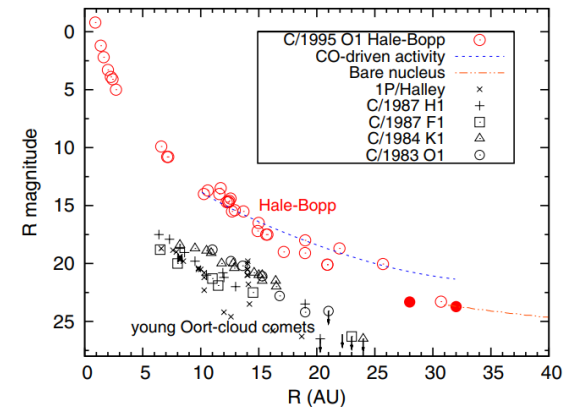
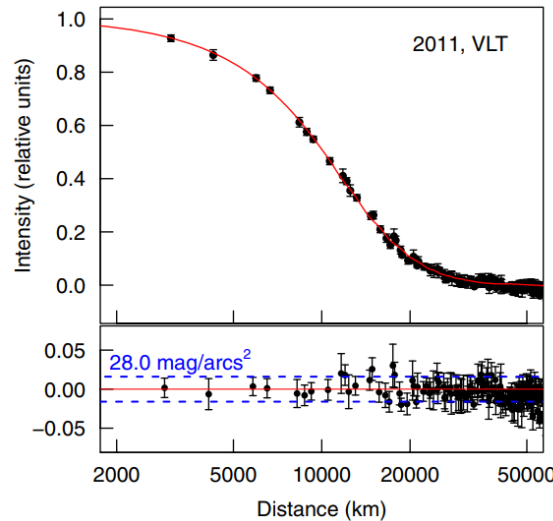


eso0328

C/Hale-Bopp's nucleus

- $r = 37$ km, 8% albedo increased \leftarrow re-depositoin of icy grains

2011, VLT



Szabo+ 2012ApJ...761...8S

Comets / Space Mission

- EPOXI and 103P/Hartley 2
 - Snodgrass, Meech, Hainaut 2010A&A...516L...9S The nucleus of 103P/Hartley 2, target of the EPOXI mission

- Deep Impact / Stardust-NExT and 9P/Tempel 1
 - Meech+ 2011Icar..213..323M Deep Impact, Stardust-NExT and the behavior of Comet 9P/Tempel 1 from 1997 to 2010
 - Weiler+ 2007Icar..190..423W The gas production of Comet 9P/Tempel 1 around the Deep Impact date

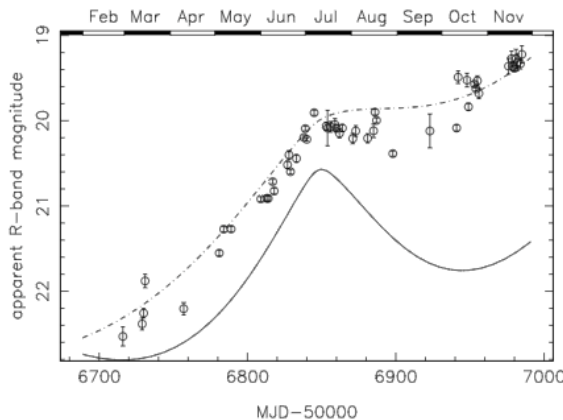
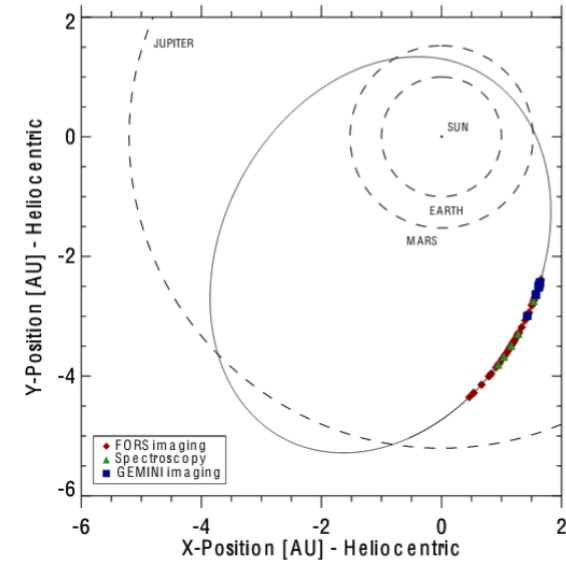
- Rosetta and 67P/Chury
 - Tubiana+ 2008A&A...490..377T - Comet 67P/Churyumov-Gerasimenko at a large heliocentric distance
 - Agarwal+ 2010Icar..207..992A - The dust trail of Comet 67P/Churyumov-Gerasimenko between 2004 and 2006
 - Tubiana+ 2011A&A...527A - 67P/Churyumov-Gerasimenko at large heliocentric distance
 - Tozzi+ 2011A&A...531A..54T - Evolution of the dust coma in comet 67P/Churyumov-Gerasimenko before the 2009 perihelion
 - Vincent+ 2013A&A...549A.121V - Spin and activity of comet 67P/Churyumov-Gerasimenko
 - Snodgrass+ 2016A&A...588A..80S - Distant activity of 67P/Churyumov-Gerasimenko in 2014: Ground-based results during the Rosetta pre-landing phase
 - Moreno+ 2016A&A...587A - The dust environment of comet 67P/Churyumov-Gerasimenko from Rosetta OSIRIS and VLT observations in the 4.5 to 2.9 AU heliocentric distance range inbound
 - Stinson, Bagnulo, Tozzi 2016A&A...594A.110S - Polarimetry of comets 67P/Churyumov-Gerasimenko, 74P/Smirnova-Chernykh, and 152P/Helin-Lawrence
 - Moreno+ 2017MNRAS.469S.186M - The dust environment of comet 67P/Churyumov-Gerasimenko: results from Monte Carlo dust tail modelling applied to a large ground-based observation data set
 - Opitom+ 2017MNRAS.469S.222O - Ground-based monitoring of comet 67P/Churyumov-Gerasimenko gas activity throughout the Rosetta mission

Comets / Space Missions

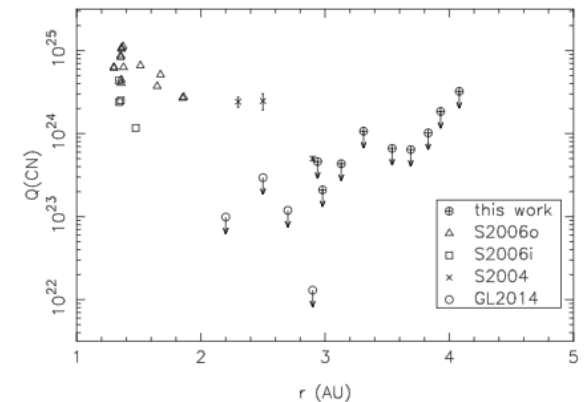
67P/Chury: pre-Rosetta:

Characterization of

- The nucleus,
 - The activity profile, dust properties
 - Gas production rates
- prepared Rosetta's strategy



Photometry: activity profile
Snodgrass+ 2016A&A...588A..80S

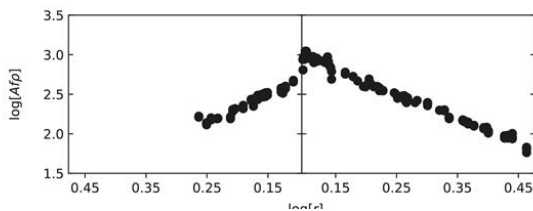
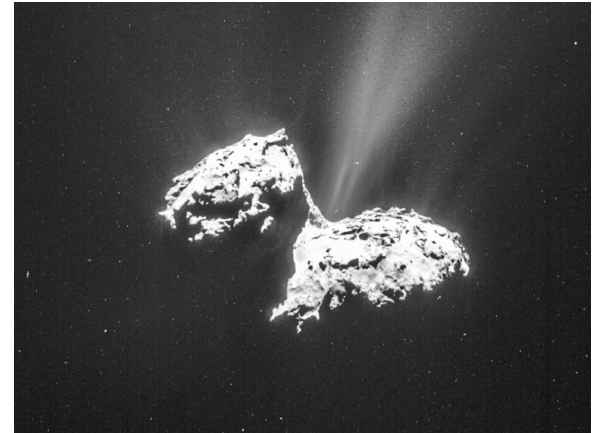


Spectro: gas production rates
Snodgrass+ 2016A&A...588A..80S

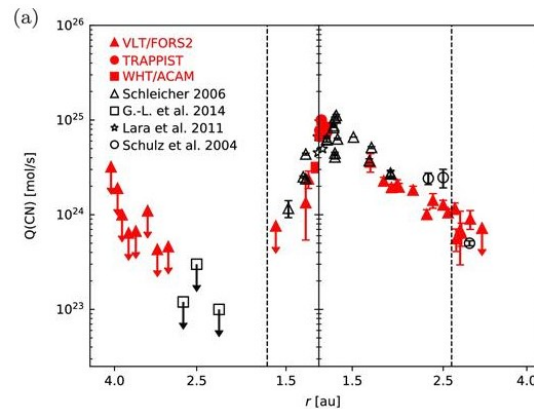
Comets / Space Missions

67P/Chury during Rosetta:

- Gives context
 - Spatial (different scales)
 - Temporal (seasonal effects)
- Link with other comets
 - (observed only from the Ground)

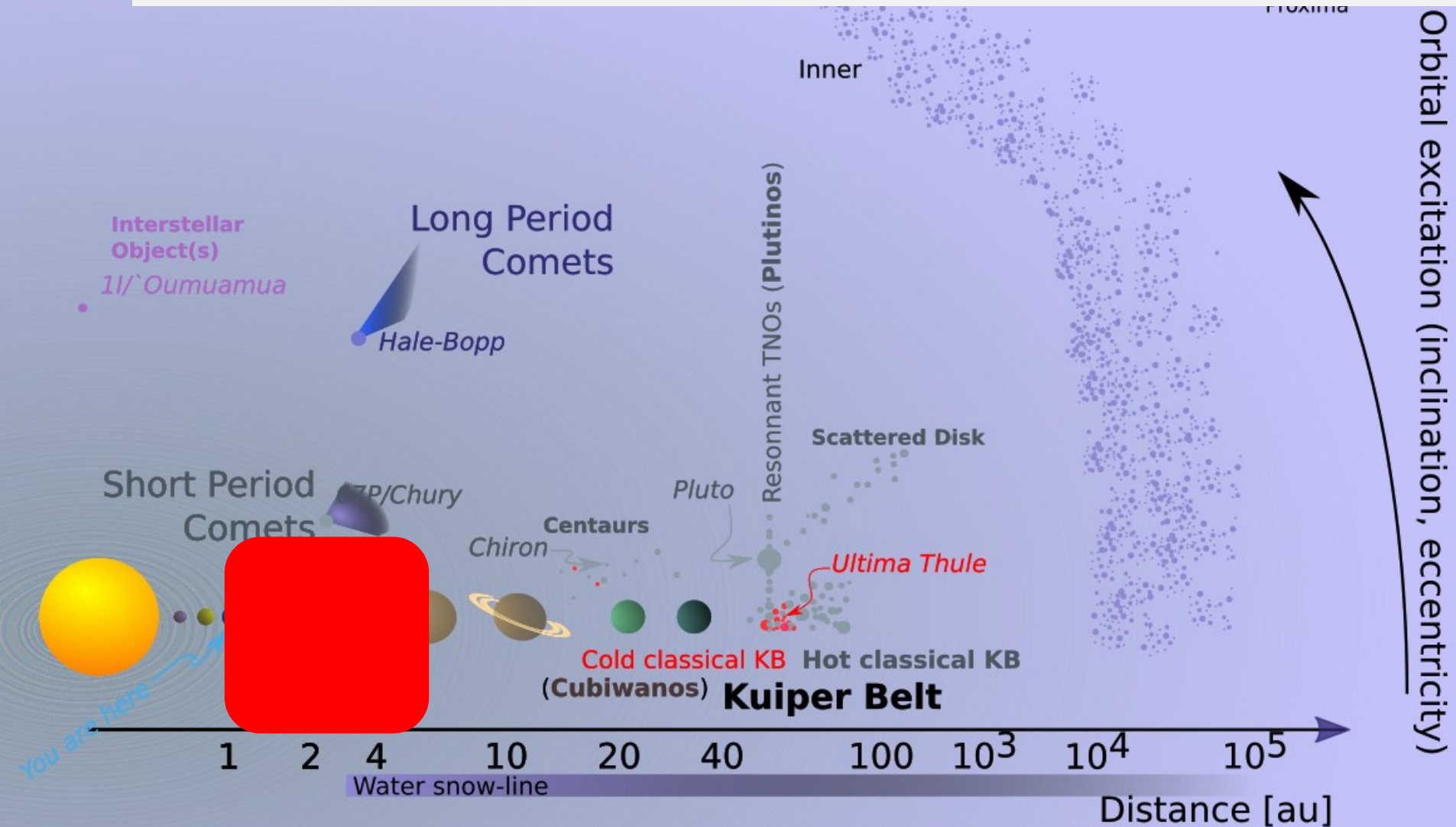


Dust production during Rosetta
Opitom+ 2017MNRAS.469S.2220



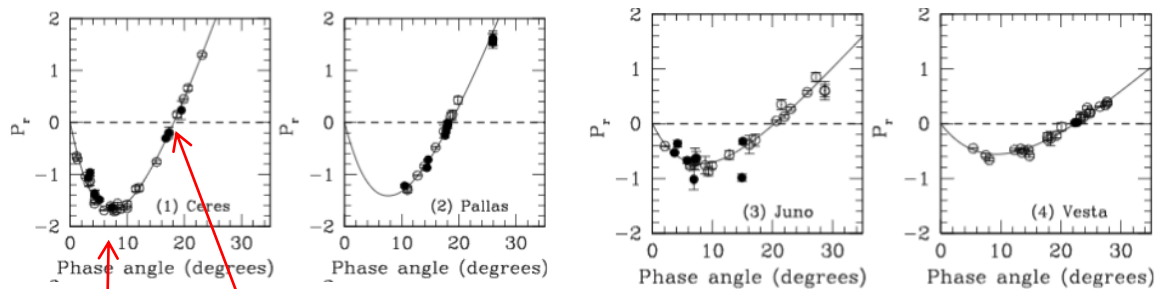
Gas production during Rosetta
Opitom+ 2017MNRAS.469S.2220

Asteroids & Polarimetry



FORS & Polarization

- Boehnhardt+ 2008A&A...489.1337B
Photometry and polarimetry of the nucleus of comet 2P/Encke
- Cellino+ 2014MNRAS.439L..75C
A successful search for hidden Barbarians in the Watsonia asteroid family.
- Bagnulo+ 2015MNRAS.446L..11B
Linear spectropolarimetry: a new diagnostic tool for the classification and characterization of asteroids.
- Bagnulo+ 2016A&A...585A.122B
Broadband linear polarization of Jupiter Trojans
- Bagnulo+ 2017EPJP..132..405B
Polarimetry of small bodies and satellites of our Solar System



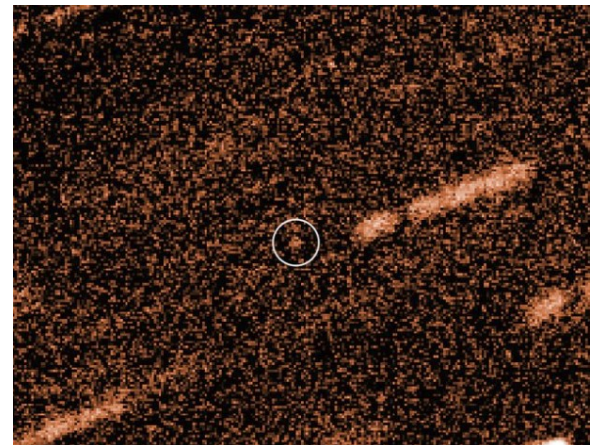
Bagnulo+ 2017EPJP..132..405B

NEO prevention

ESO

in collaboration with ESA Planetary Defence Office,
in the context of the International Asteroid Warning Network,
in UN-Committee for the Peaceful Use of Outer Space:

- Recover / refine the orbit of threatening NEOs that are too faint for NEO facilities (up to 12h/semester)
- Characterize an upcoming impactor to prepare mitigation.

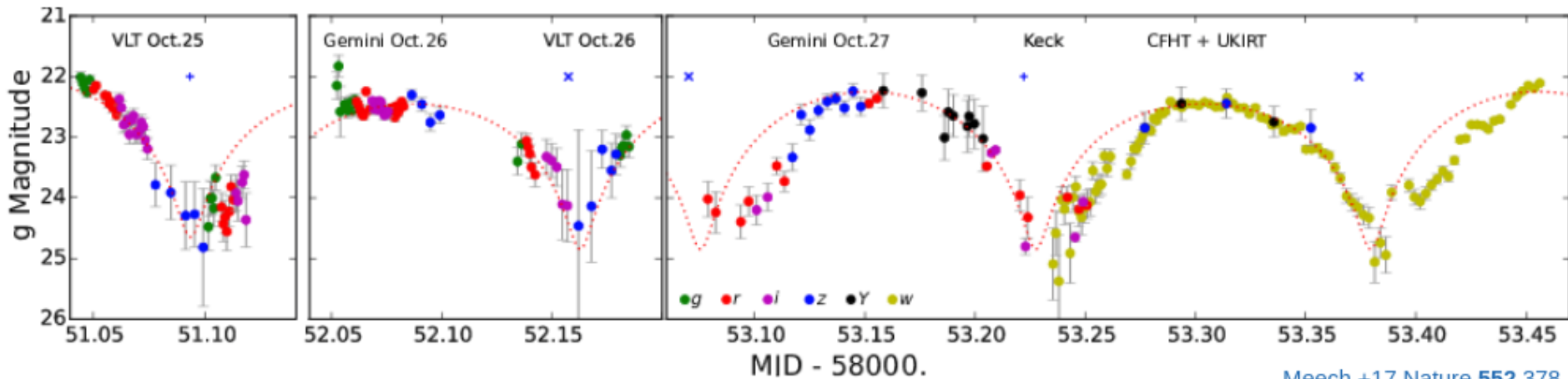
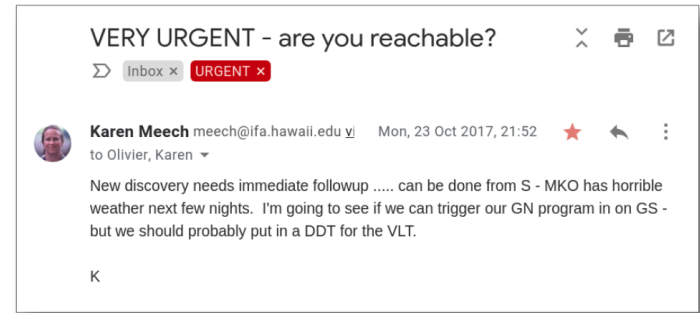


2009 FD, mag=26.5, threat decreased 10x
ESO ann14004

1I/ Oumuamua

■ ESO DDT

- 2017-oct-19 Discovery @ PS1
- 2017-oct-22 Hyperbolic orbit
- 2017-oct-23 Alert
- 2017-oct-24 DDT submitted
- 2017-oct-24 evening: FORS is observing



Meech +17 Nature 552 378

■ First interstellar object

- Huge elongation ~10:1
- Colours are ~comet-like
- Non-gravitational forces are comet-like
- Many questions; 5 papers (2 Nature) out of this DDT



- **Significant contribution to population studies**
 - TNOs
 - Comets
 - Main-Belt Comets and Active Asteroids
 - **Strong support to Space Missions**
 - Rosetta @ 67P
 - Deep Impact @ 9P/T1
 - **A series of unique/ record/ special object characterized**
 - **Special mention:**
 - Versatility well adapted to Solar System objects
 - Power of polarimetry
 - Super DDT tool
- ➔ **When can we get FORS+ ?**