

»We explore a possible alternative to PCA-based PSF estimation for direct imaging of exoplanets which works by learning causal pixel-wise noise models.«

Half-Sibling Regression meets Direct Imaging: A Causal Approach for Uncovering Exoplanets

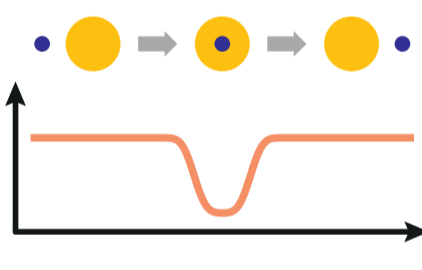
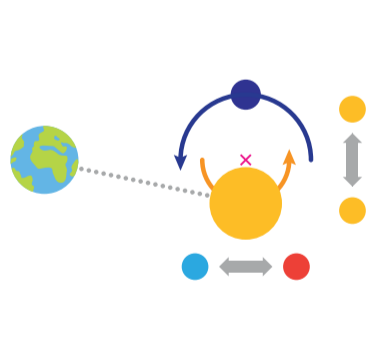
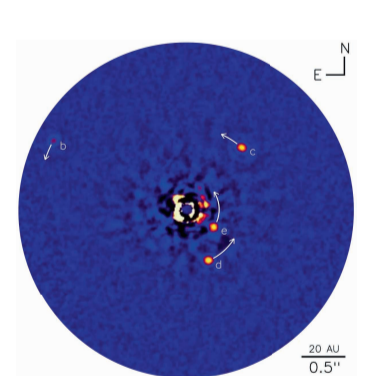
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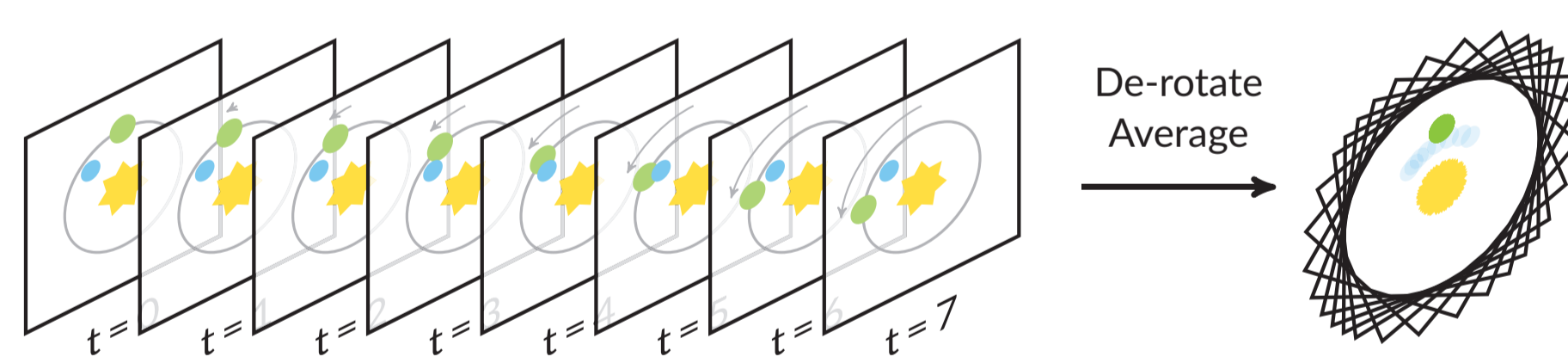
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Detecting Exoplanets

- 1 Transit Photometry**
Identify periodic “dips” in the light curves of stars (→ e.g., Kepler) 
- 2 Astrometry / Radial Velocity**
Observe periodic changes in star position / redshift (→ e.g., Gaia) 
- 3 Direct Imaging**
Capture photons from planet, mostly in the infrared (→ e.g., NACO/VLT) 

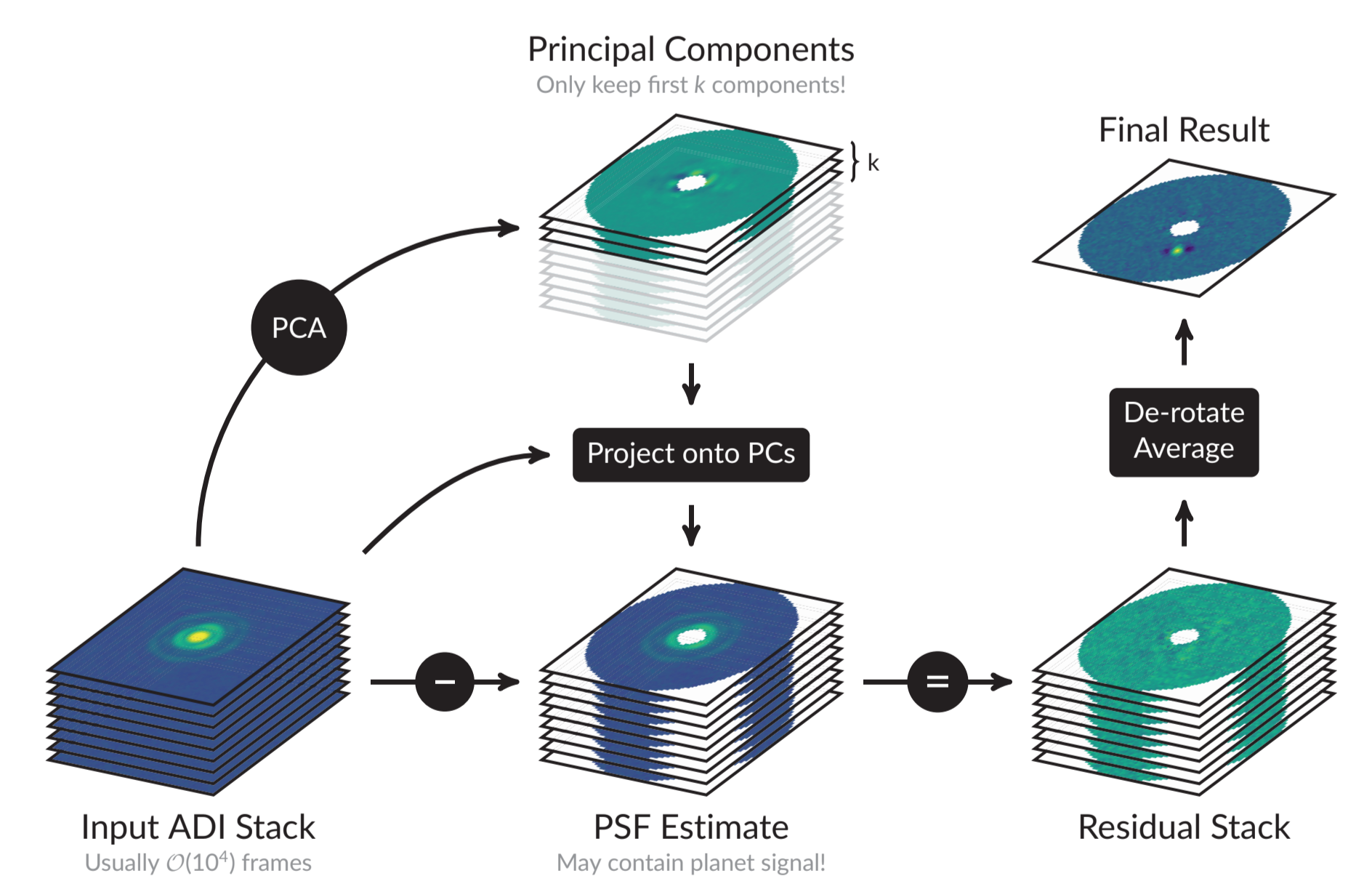
Angular Differential Imaging

Angular Differential Imaging (ADI) is an **observation technique** first introduced by Marois *et al.* (2005):



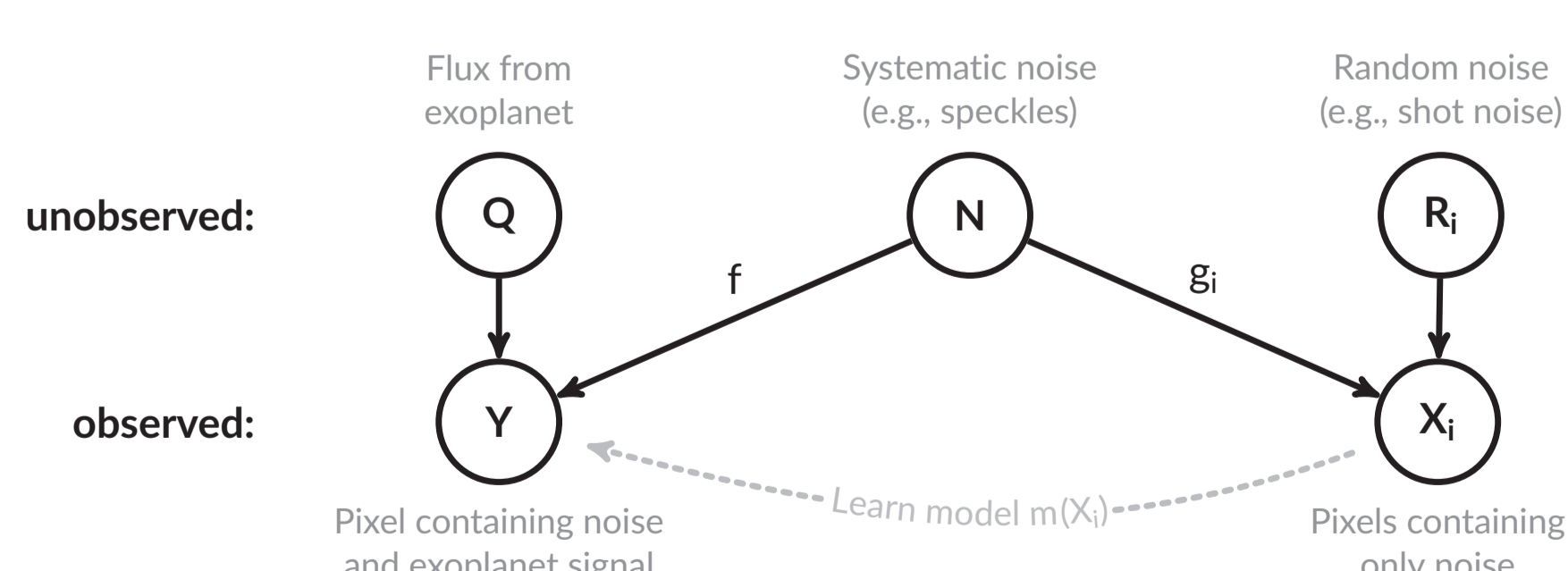
Key idea: **turn off the field derotator** of the telescope. **De-rotating** every frame by its parallactic angle and **averaging** will then **reduce quasi-static speckle noise**.

PCA-based PSF subtraction



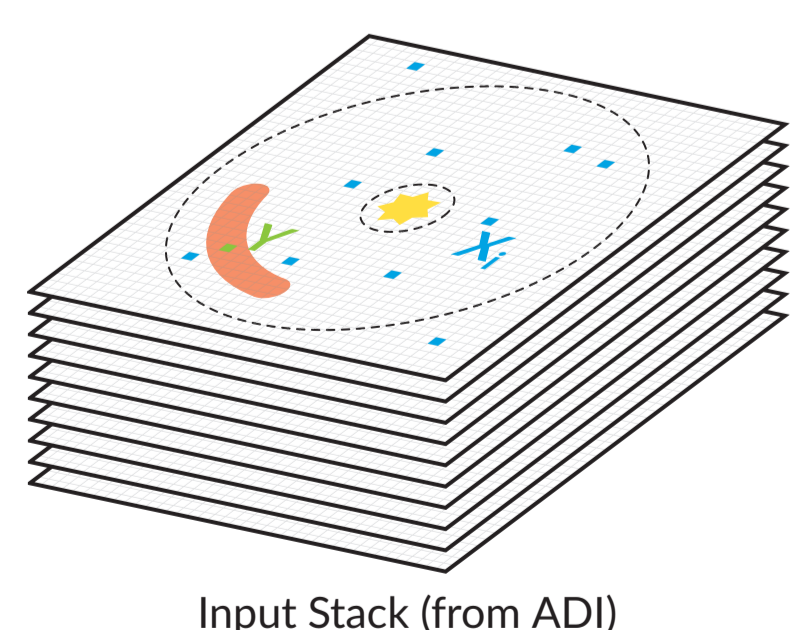
Half-Sibling Regression

Assume we are interested in the latent quantity $Q \perp\!\!\!\perp X_i$:



Schölkopf *et al.* (2016) have shown that by regressing Y onto X_i , one can learn an estimate for $f(N)$, allowing us to reconstruct Q (up to a constant) as $\hat{Q} = Y - f(N)$.

Application to Direct Imaging



For every pixel Y in the ROI, we learn a model m :
$$Y = m(X_i) + \epsilon$$

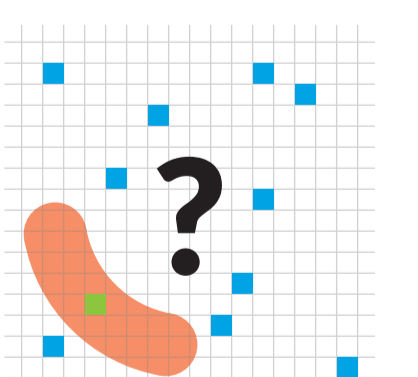
Idea: Use $m(X_i)$ as an estimate for the systematics.
→ ϵ is the estimated planet signal in pixel Y !

Advantage: Lots of freedom in choosing m (and X_i)!
→ Reduce self-subtraction, include time domain, ... ?

Current Questions

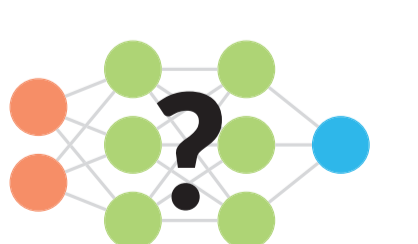
Predictor Selection

- Which pixels should we use?
- How to define the exclusion region?



Machine Learning

- Which model class should we use?
- How to split for training / test?



Evaluation

- Physical metrics vs. ML metrics?
- How to quantify self-subtraction?

