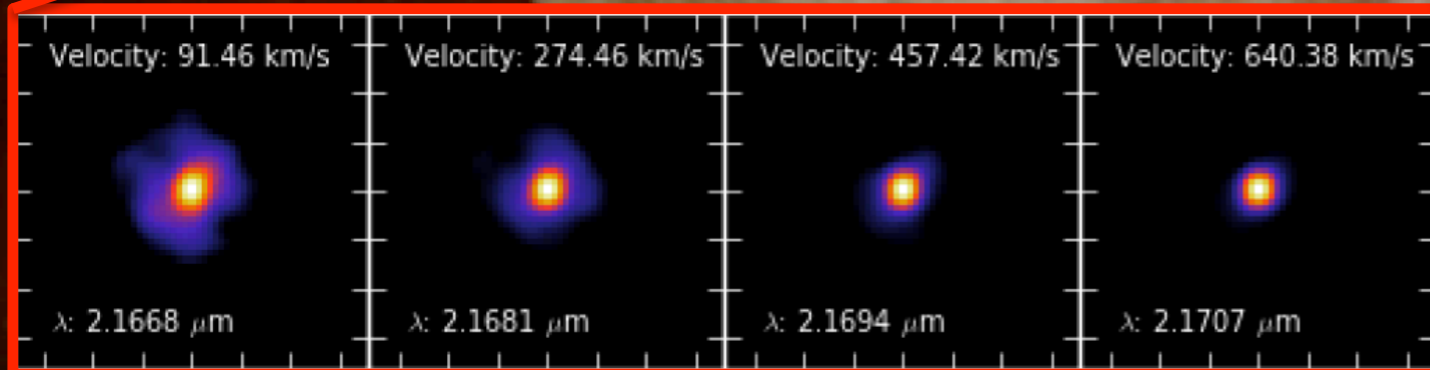
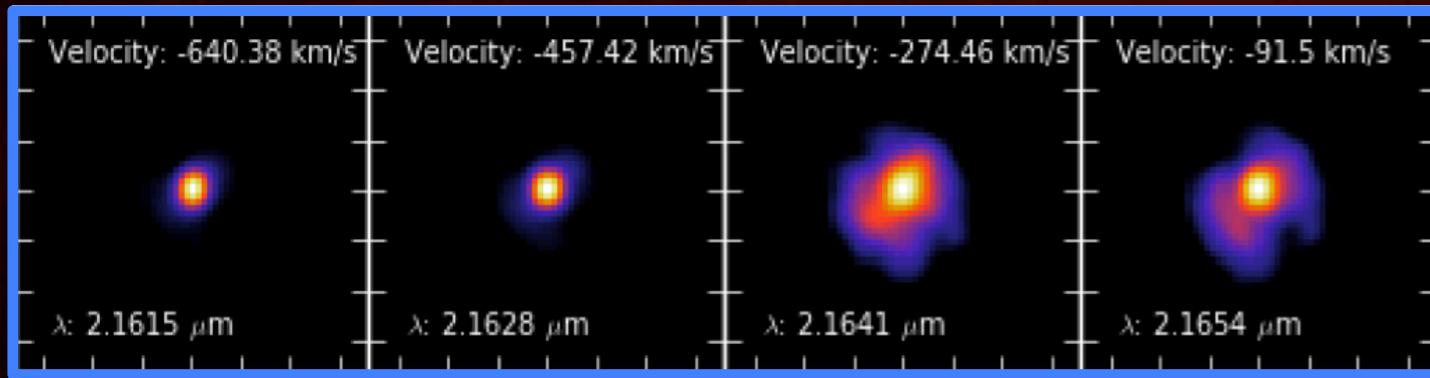
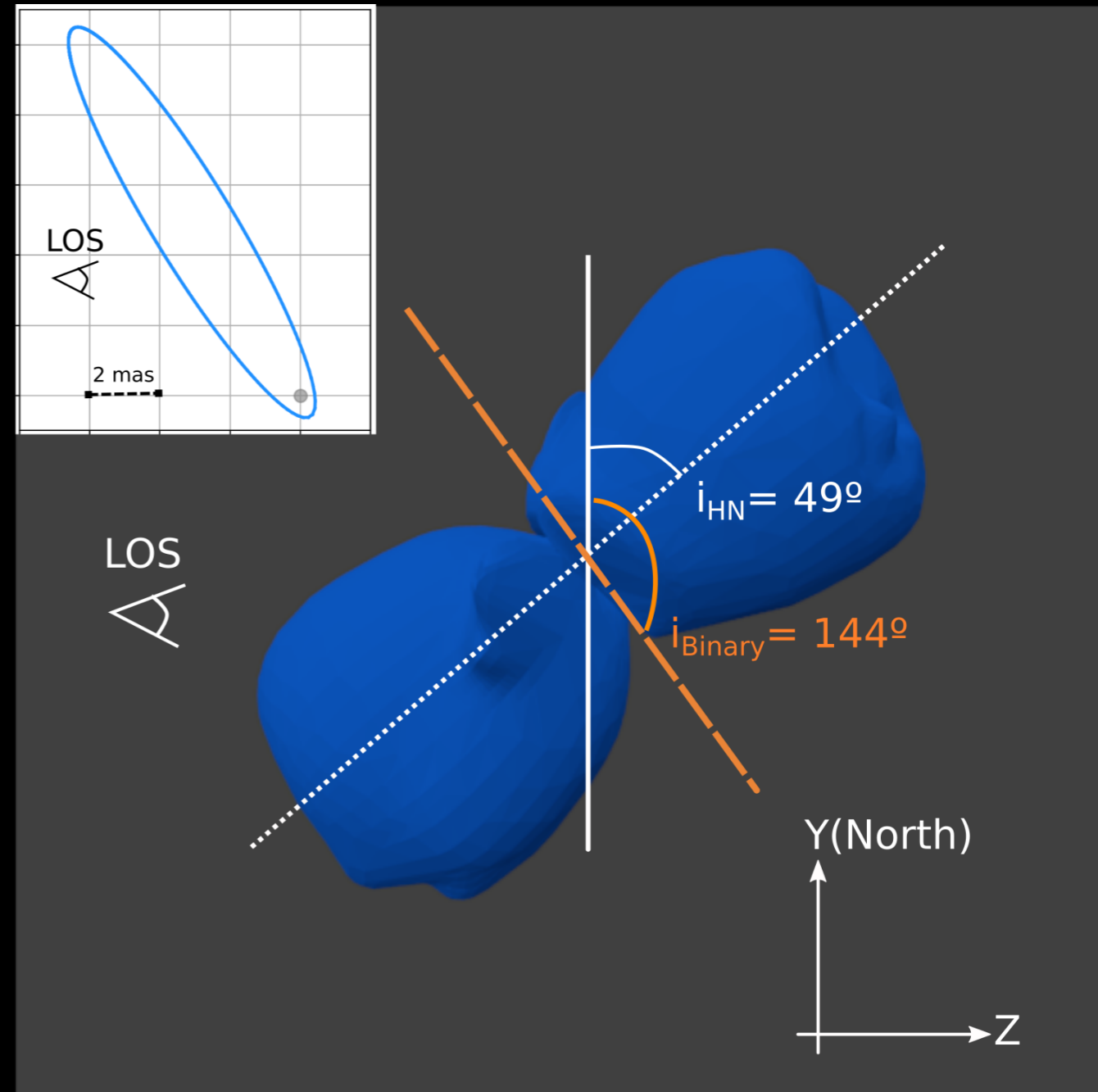
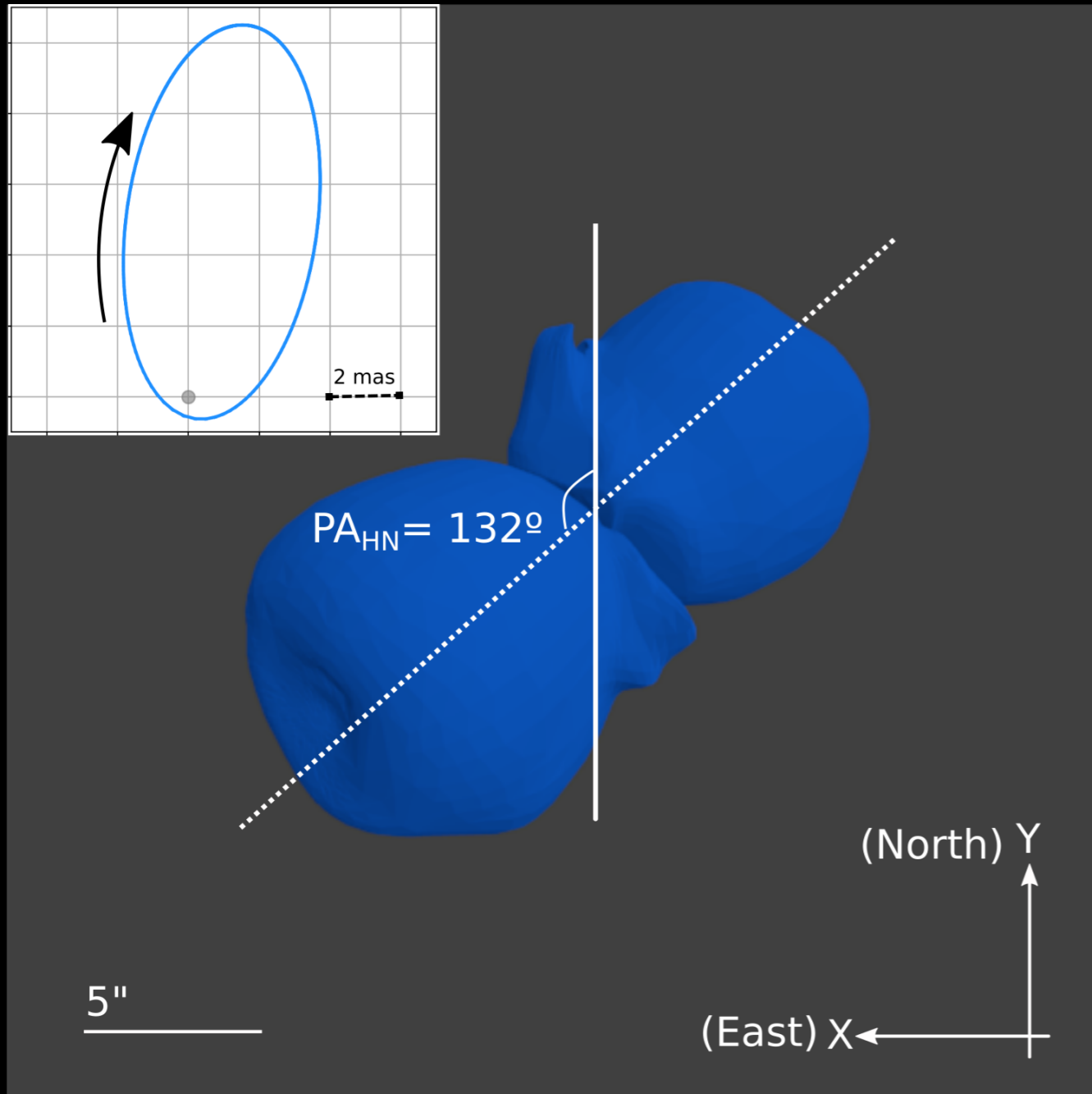


# GRAVITY/VLTI chromatic image reconstruction of the Eta Car wind-wind collision region

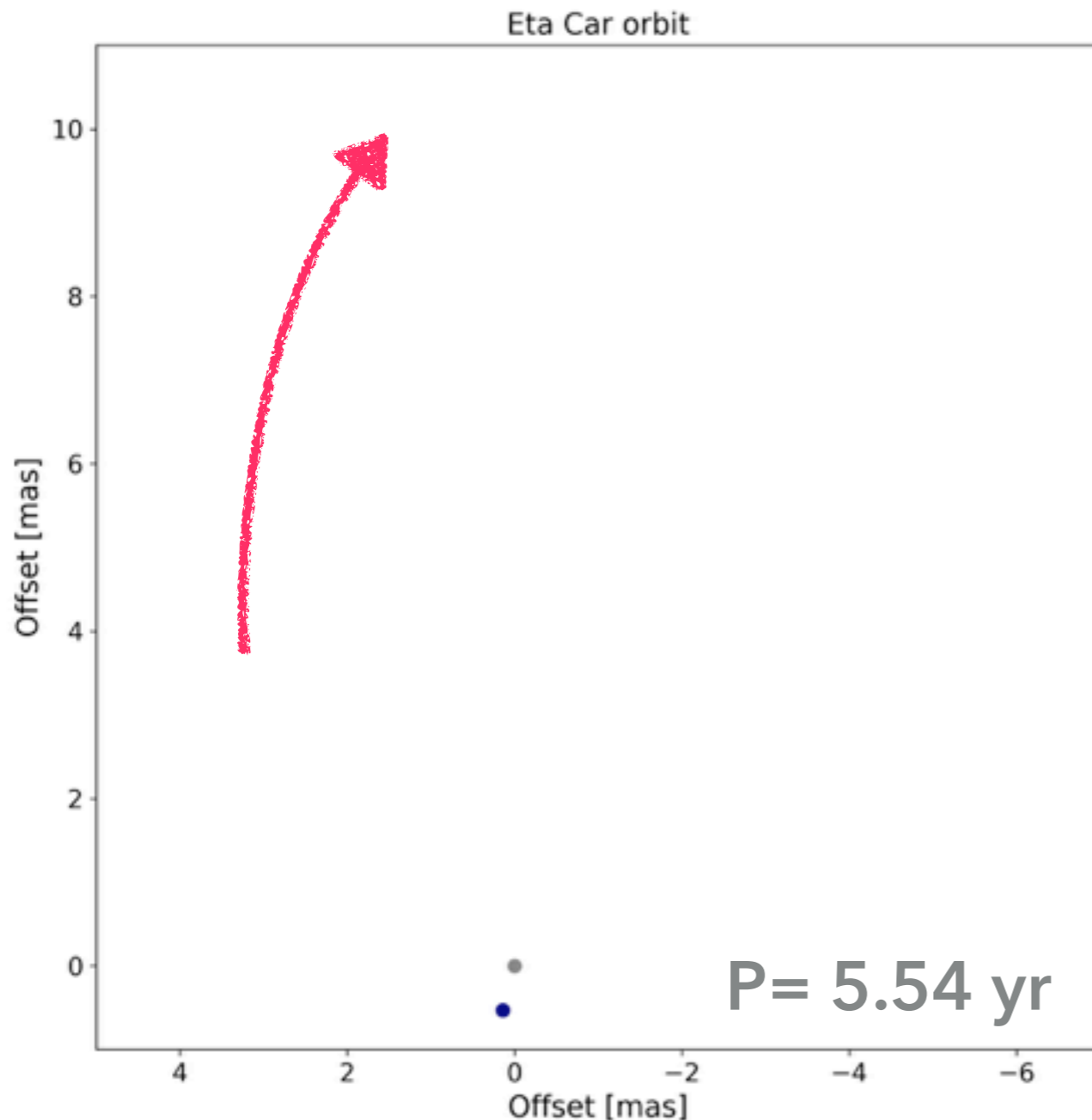


Joel Sanchez-Bermudez (MPIA)  
(On behalf of the GRAVITY-VLTI team)

# ETA CAR: IMAGING THE CORE



# ETA CAR: IMAGING THE CORE



## PROPERTIES $\eta_A$ :

- $M > 100 M_{\odot}$
- $\dot{M} \sim 8.5 \times 10^{-4} M_{\odot}/\text{yr}$
- $V_{\text{wind}} \sim 420 \text{ km/s}$

## PROPERTIES $\eta_B$ :

- $\dot{M} \sim 1 \times 10^{-5} M_{\odot}/\text{yr}$
- $V_{\text{wind}} \sim 3000 \text{ km/s}$

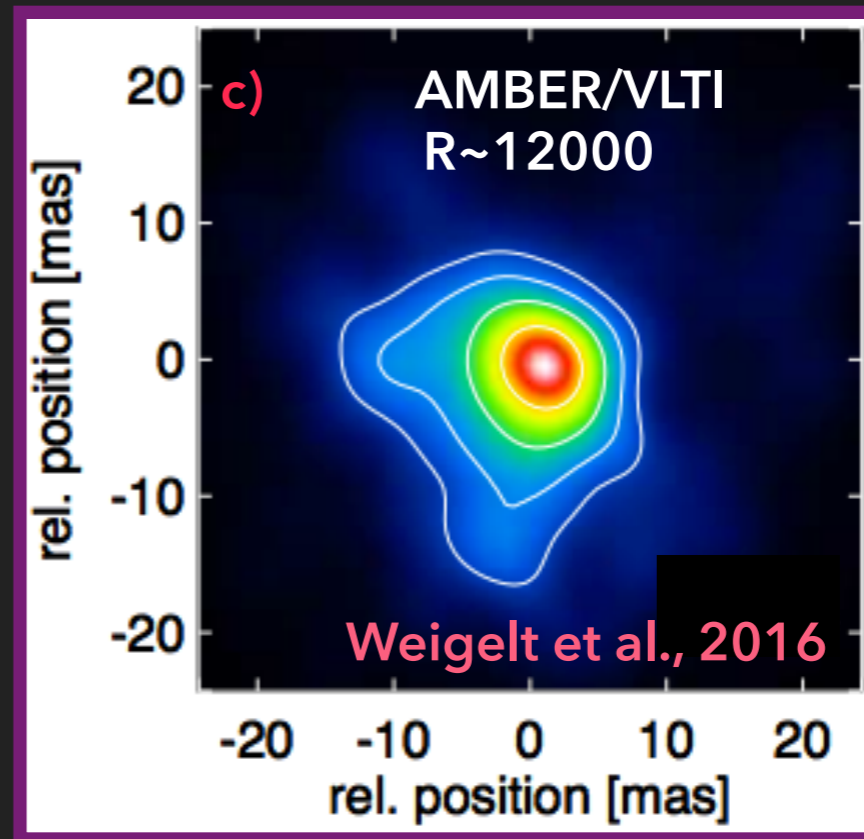
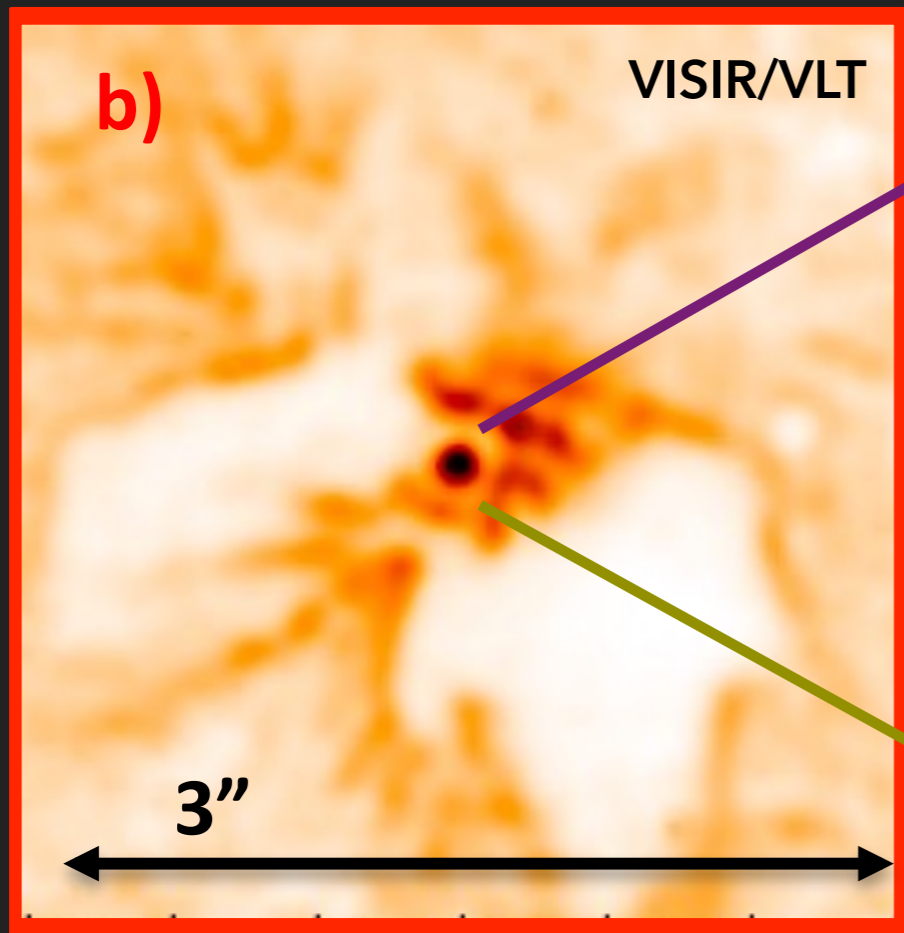
## Orbital solution:

- $e = 0.9$
- $i = 144^{\circ}$
- $\Omega = 47^{\circ}$
- $\omega = 243^{\circ}$

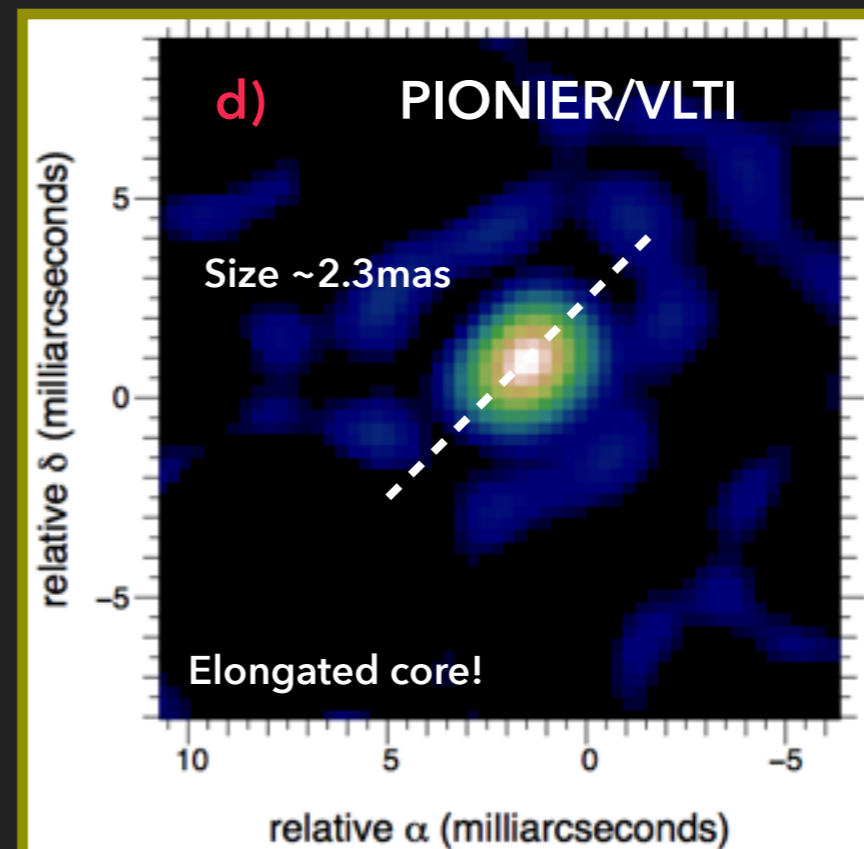
e.g., Daminieli 96; Davidson & Humphreys 97; Davidson et al. 01; Hillier et al. 01, 06; Pittard & Corcoran 2002; Groh et al. 12b

# ETA CAR: IMAGING THE CORE

N-band (8.0  $\mu\text{m}$ )



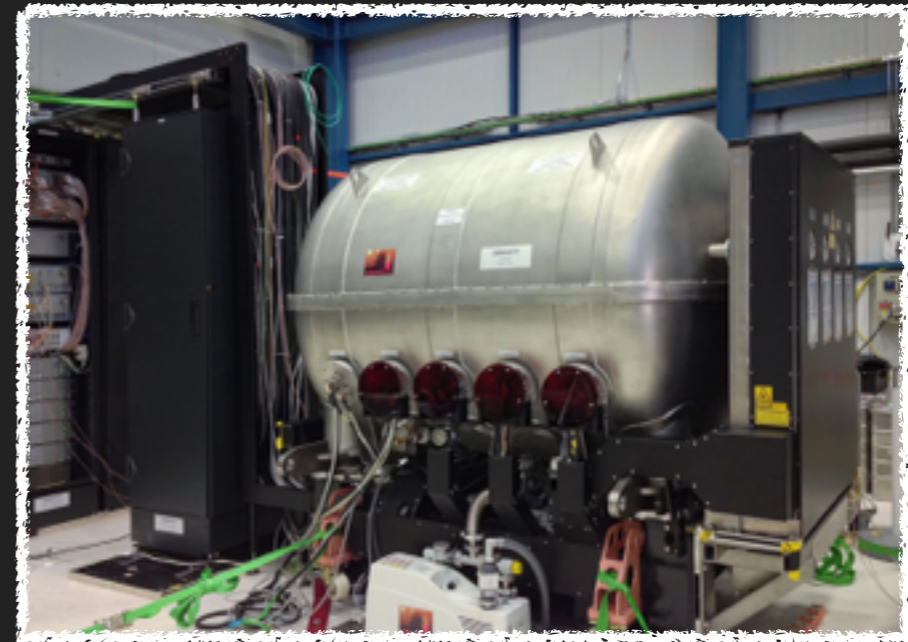
K-band (2.2  $\mu\text{m}$ )



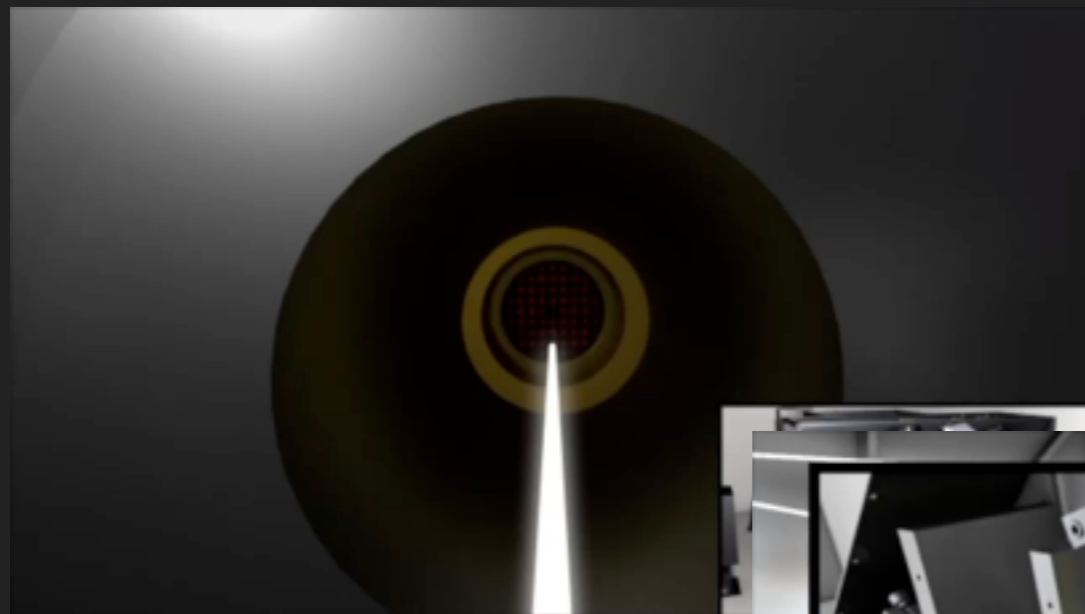
H-band (1.7  $\mu\text{m}$ )

See the poster No. 93!

- K-band ( $\sim 1.9\text{-}2.5\ \mu\text{m}$ )
- Angular Resolution:  $\sim 2\ \text{mas}$
- Phase referenced visibilities
- Narrow-angle astrometry:  $10\ \mu\text{as}$
- Simultaneous interferometric obs. of two objects (**DUAL mode**)
- Spectral resolution: **50, 500, 4000**
- Limiting  $K_{\text{mag}}$ : **16.5** (objective: 18)

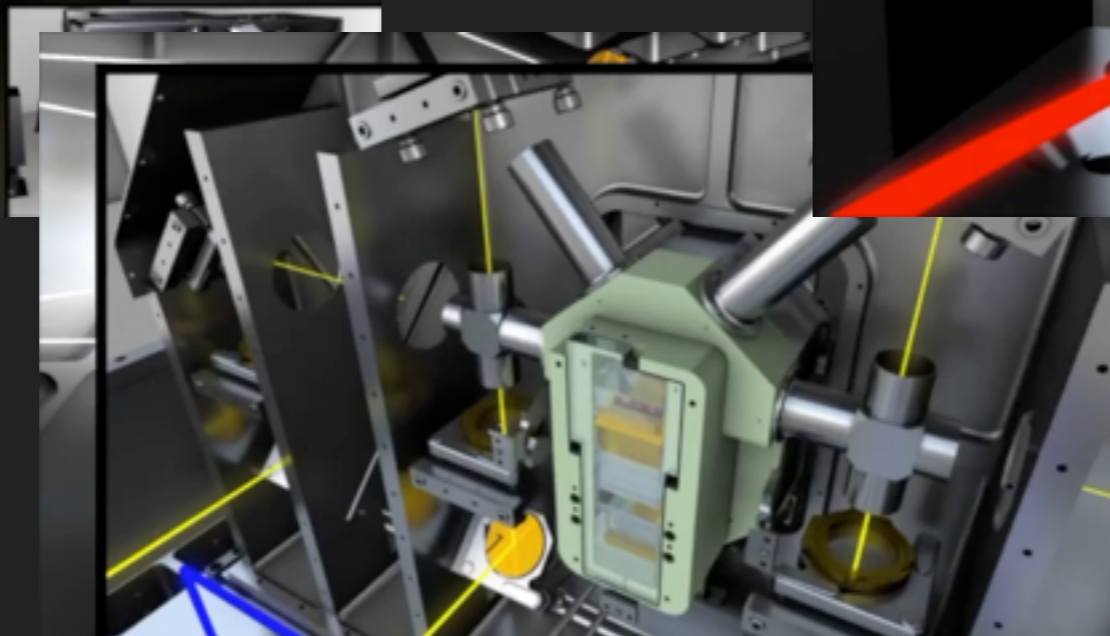


## GRAVITY collaboration et al., 2017

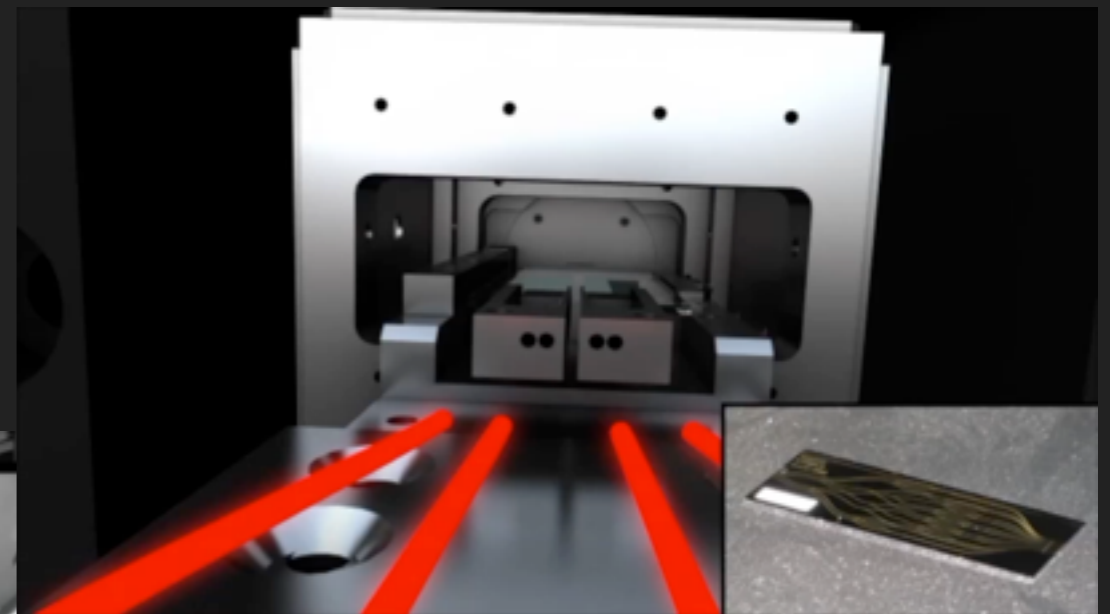


Infrared Wavefront  
Sensor

**CIAO units!**



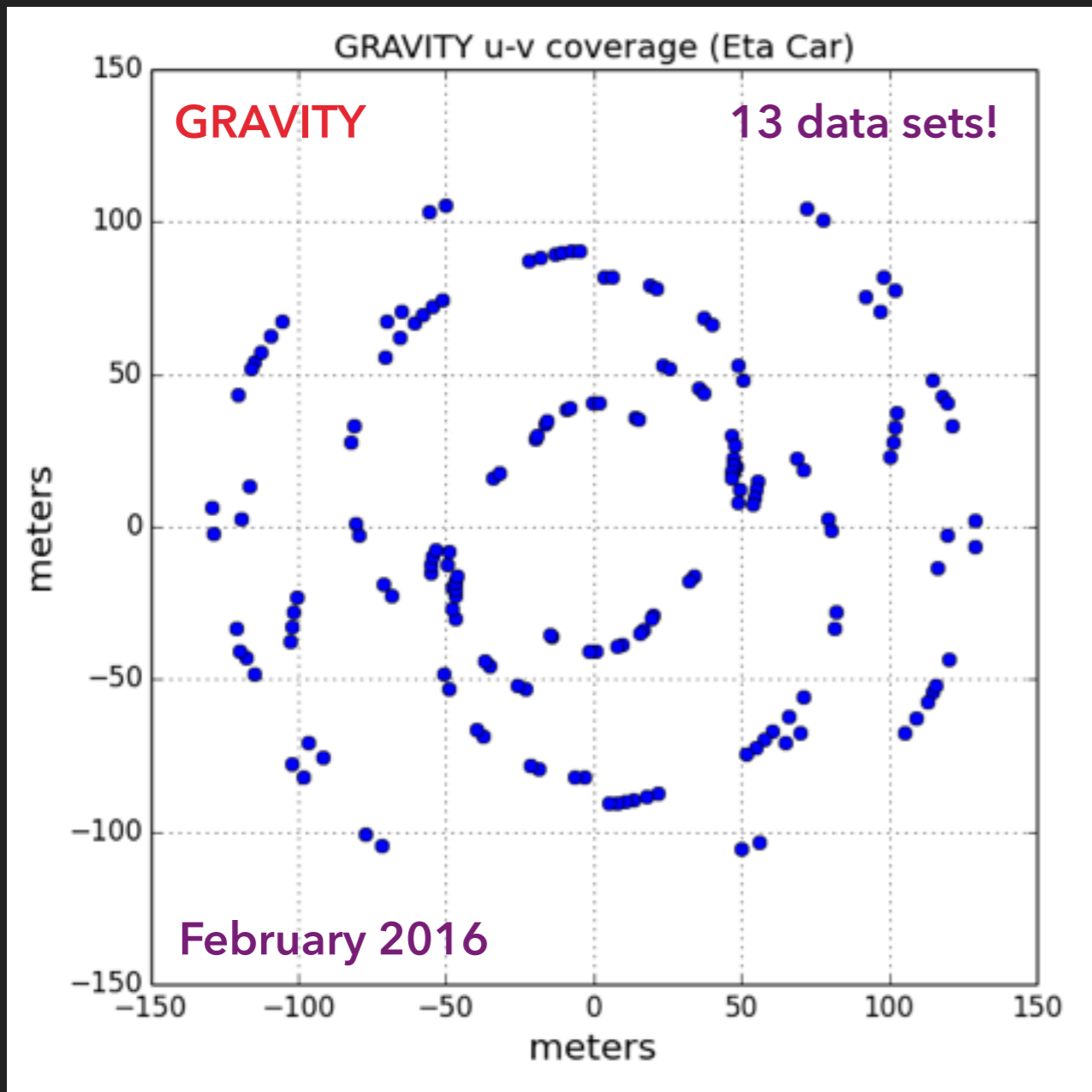
Star Separator



IC Beam Combiner

# ETA CAR: IMAGING THE CORE

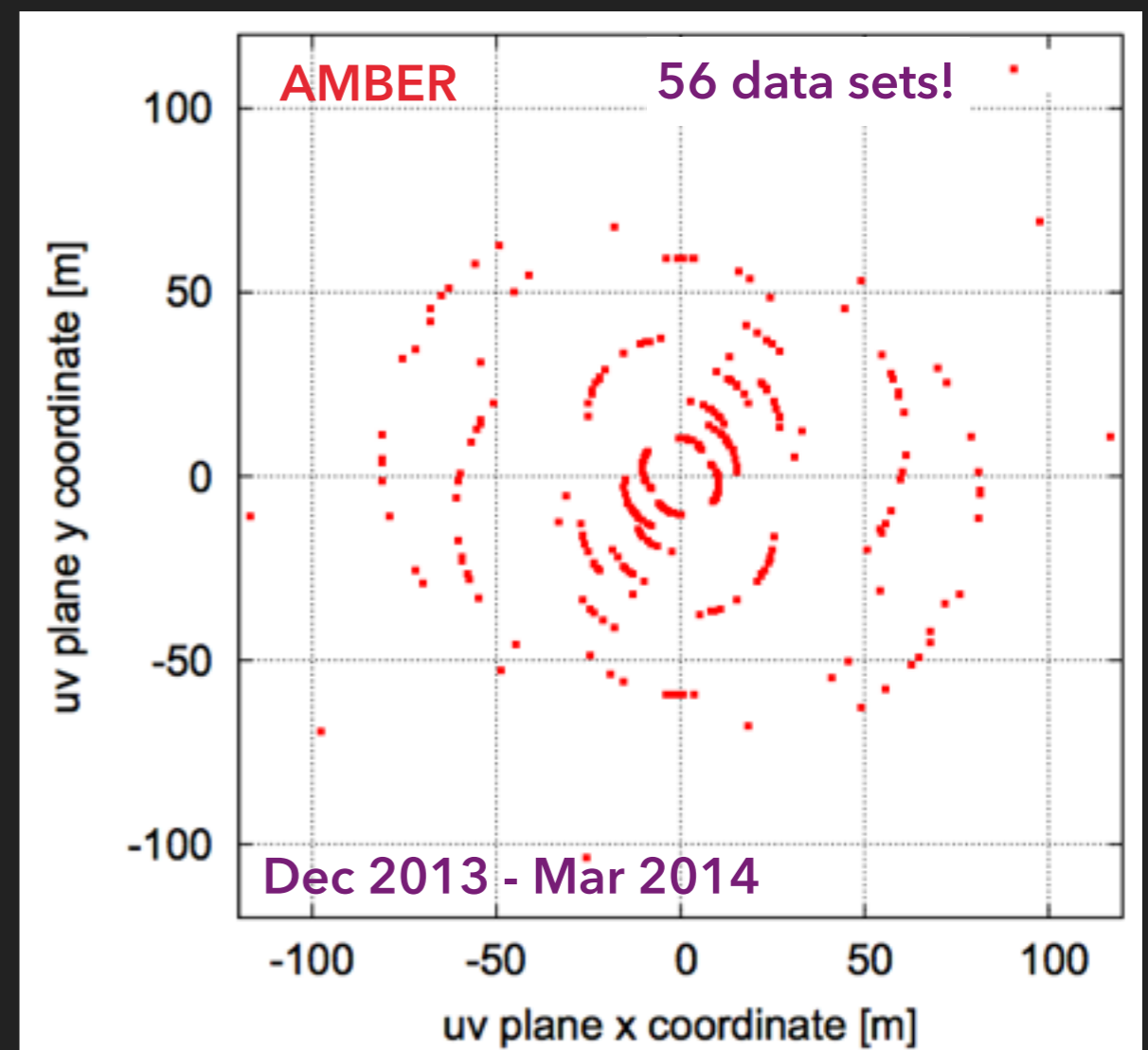
u-v coverage



**GRAVITY**  $\rightarrow \theta \sim 3.49$  mas ( $\lambda/D$ )

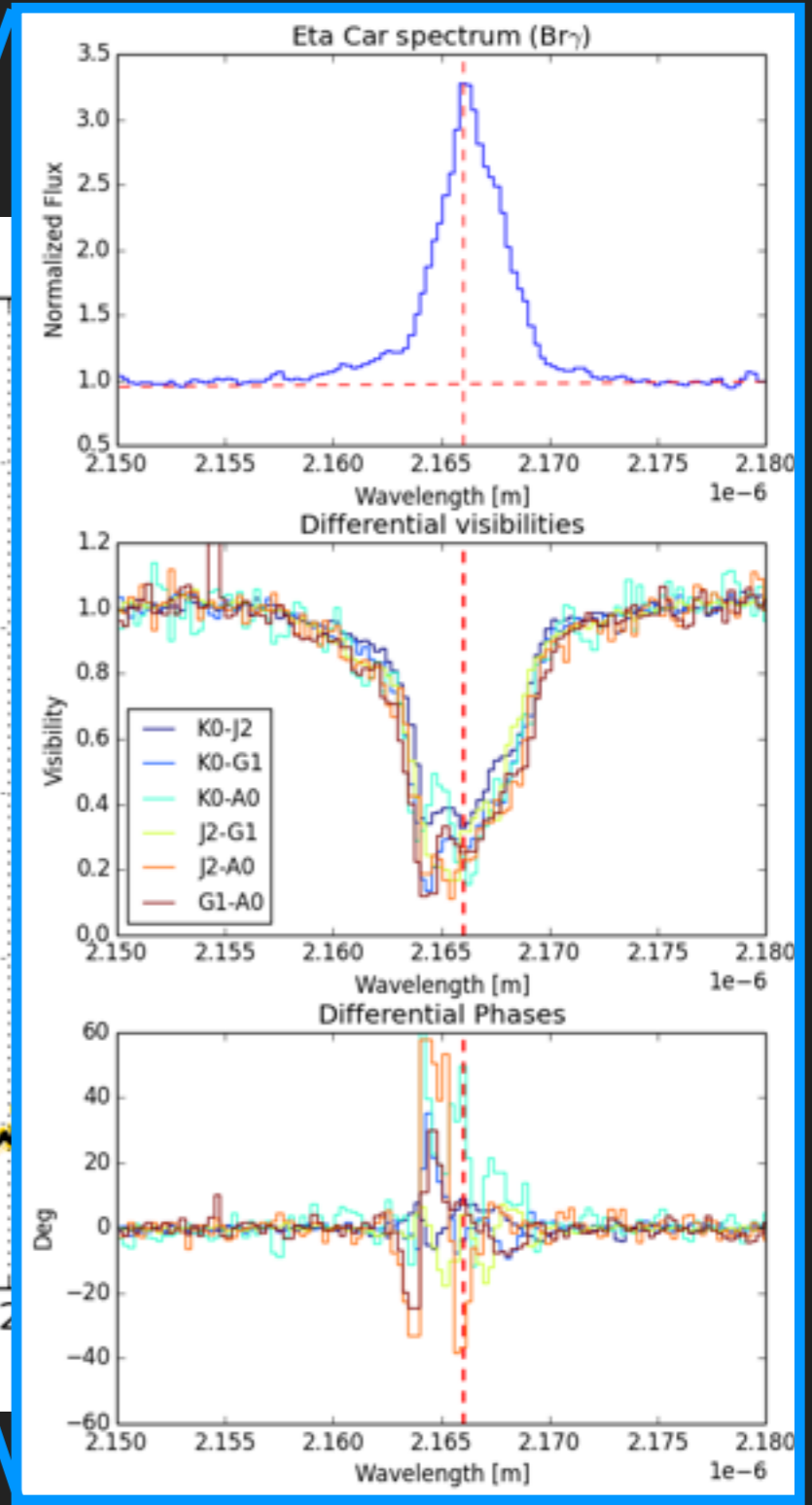
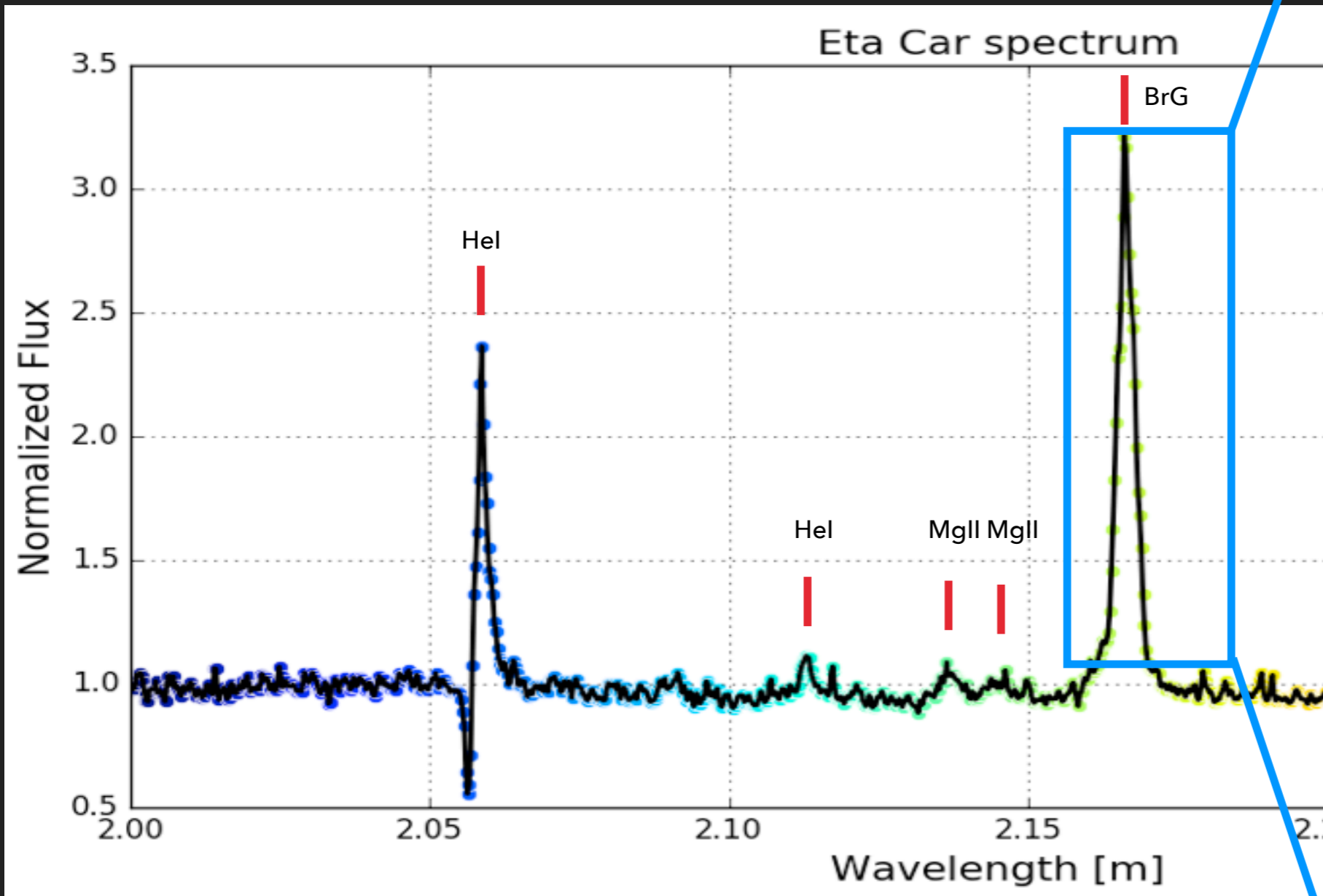
**AMBER**  $\rightarrow \theta \sim 6.00$  mas

- High-Res observations ( $R \sim 4000$ )
- 13 data sets (February 2016)
- $B_{\max} \sim 130$  m,  $B_{\min} \sim 40$  m



# ETA CAR: IMAGING THE CORE

GRAVITY Science Beam Combiner: Calibrated Spectrum



# ETA CAR: IMAGING THE CORE

Image reconstruction in optical interferometry

$$\mathbf{x}_{ML} = \underset{\mathbf{x}}{\operatorname{argmin}} [1/2\chi^2(\mathbf{x}) + \mu R(\mathbf{x})]$$

Likelihood

Hyperparameter

Regularizer

- ◉ **SQUEEZE**  
(Monte Carlo)  
F. Baron

- ◉ **BSMEM, MiRA, PAINTER, IRBis**  
(Gradient Descent)  
J. Young, E. Thibaut, A. Schutz, K.-H. Hofmann

## Regularizers

L0-norm

L2-norm

Entropy

Total Variation

Sanchez-Bermudez et al., *Imaging capabilities of the VLT/MATISSE spectro-interferometric instrument*  
*Proc. SPIE 9907, Optical and Infrared Interferometry and Imaging V, 99070B, 2016*



# ETA CAR: IMAGING THE CORE

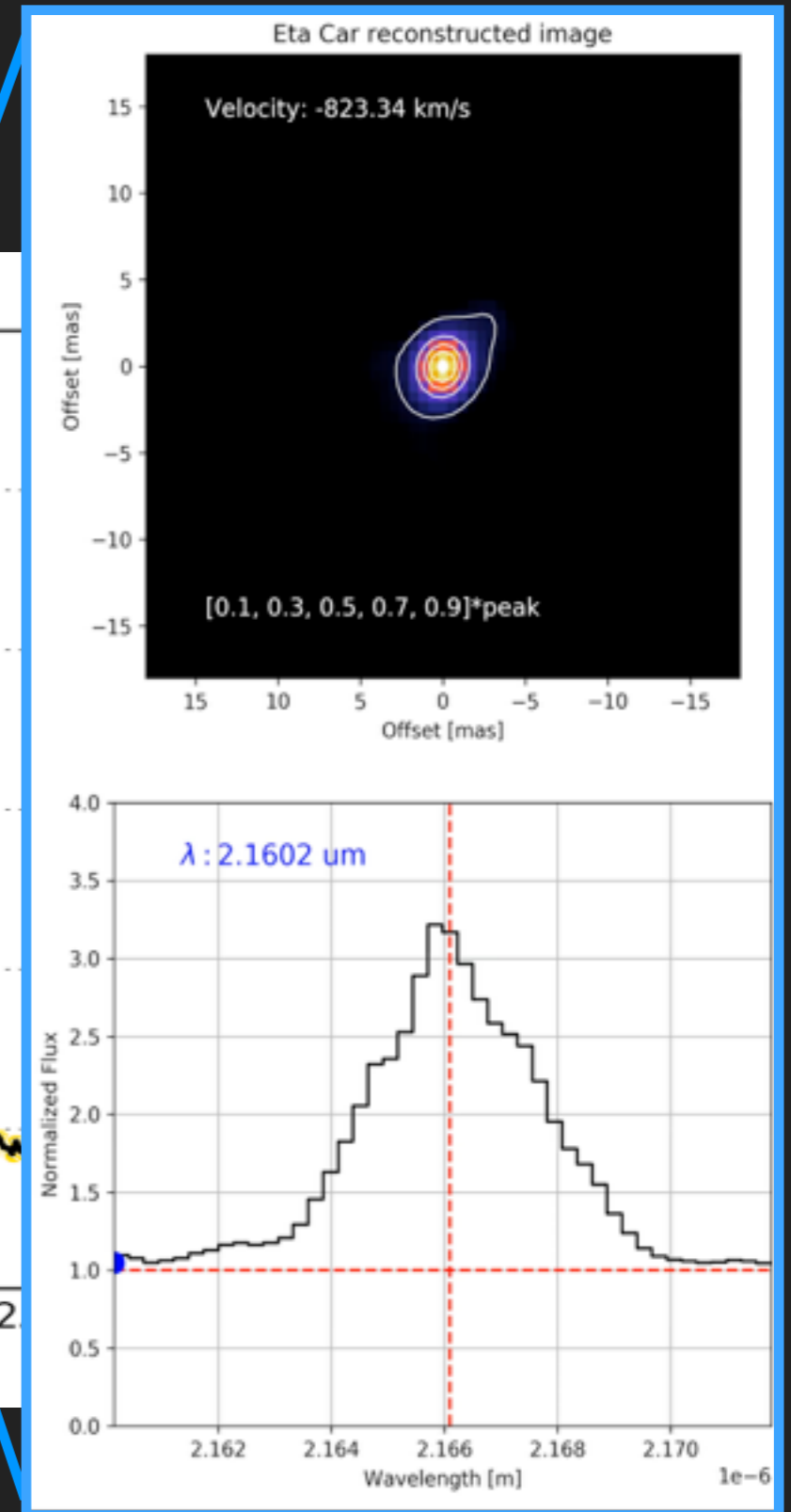
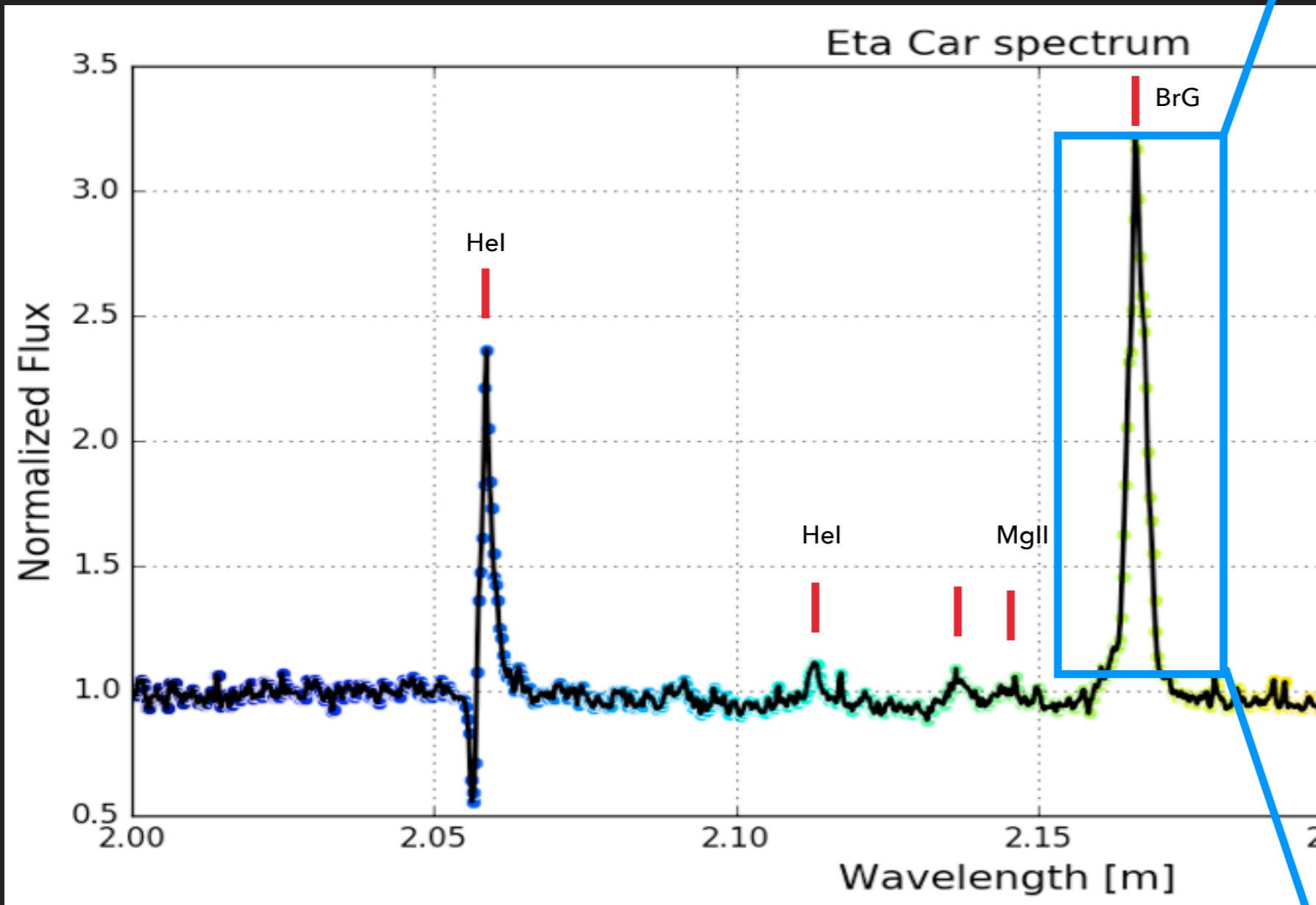
Image reconstruction parameters

**SQUEEZE; Baron+2010b**

Band	B $\gamma$ (2.160 - 2.172 $\mu$ m)
Pixel Grid	167x167 (100 mas)
Pixel Scale	0.6 mas/pixel
Chains	50
Iterations	250
Initial Image	Gaussian with 50% of the total flux
Observables	V <sup>2</sup> + CPs + Differential Phases
Regularizers	L0-norm (avoid point-like sources), Laplacian (favours extended sources), Transpectral reg. (L2-norm across the spec.)

# ETA CAR: IMAGING THE CORE

GRAVITY Science Beam Combiner: Calibrated spectrum



# ETA CAR: IMAGING THE CORE

Image fitting to V2 and closure phases

Image visibilities

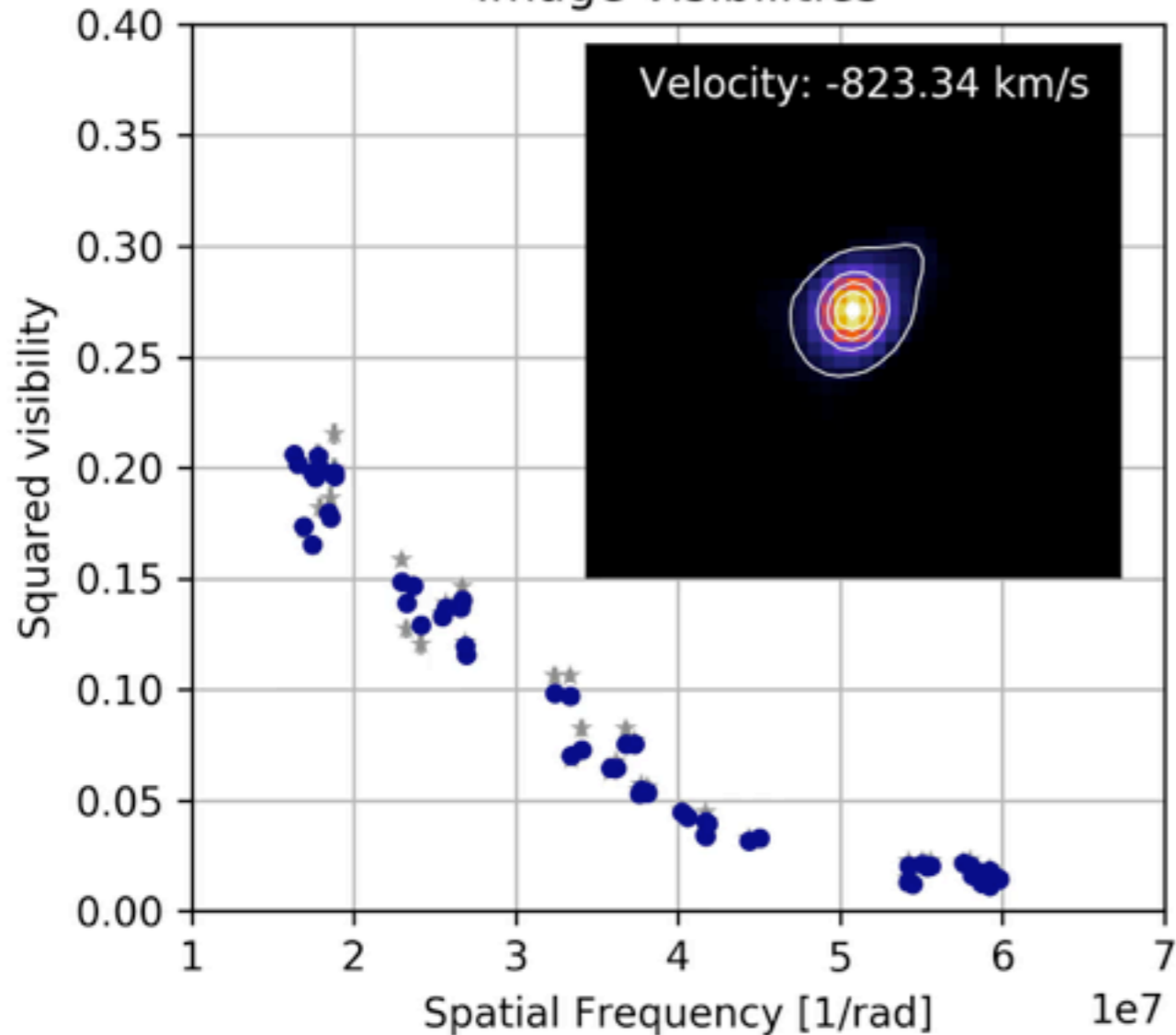
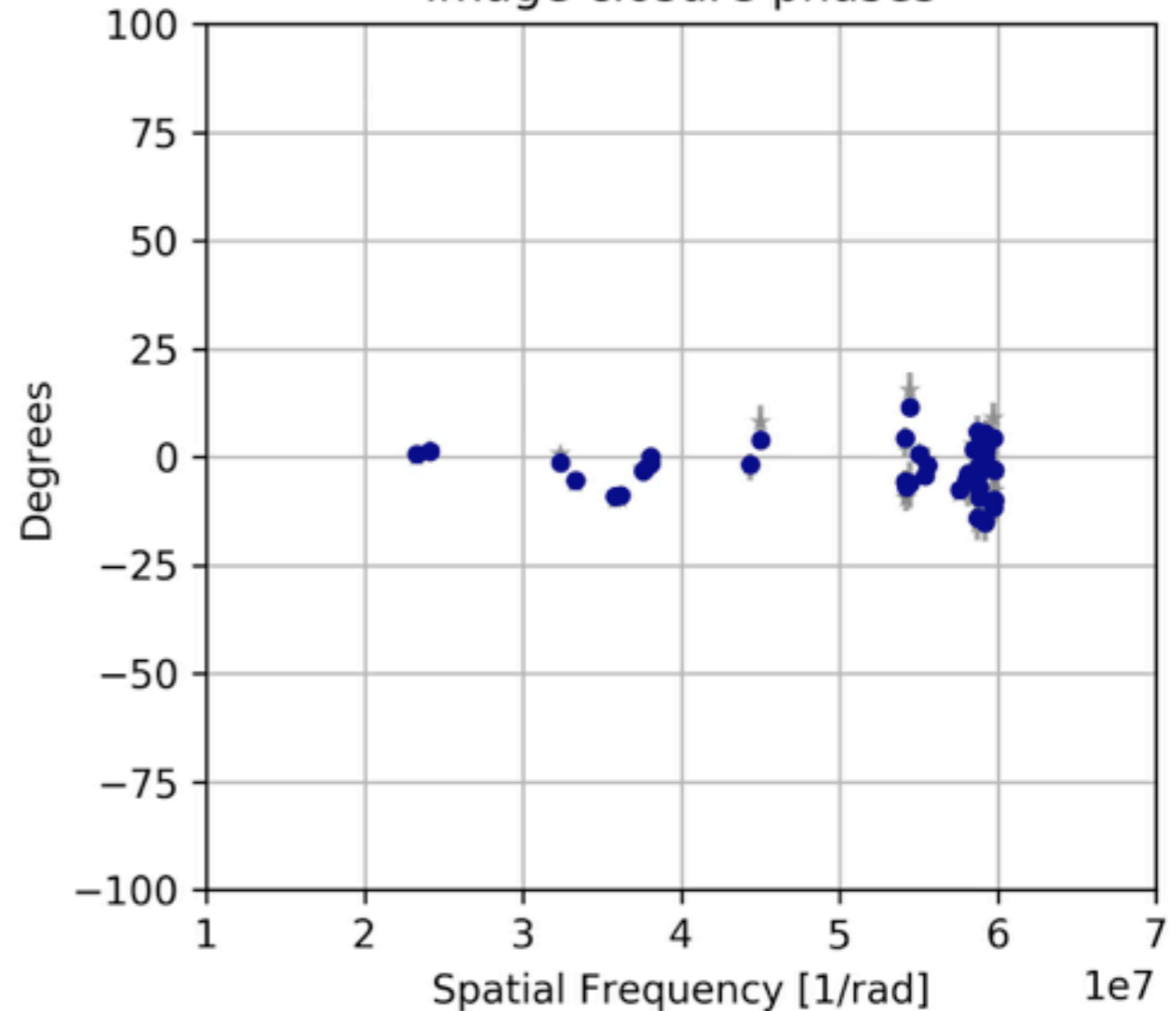
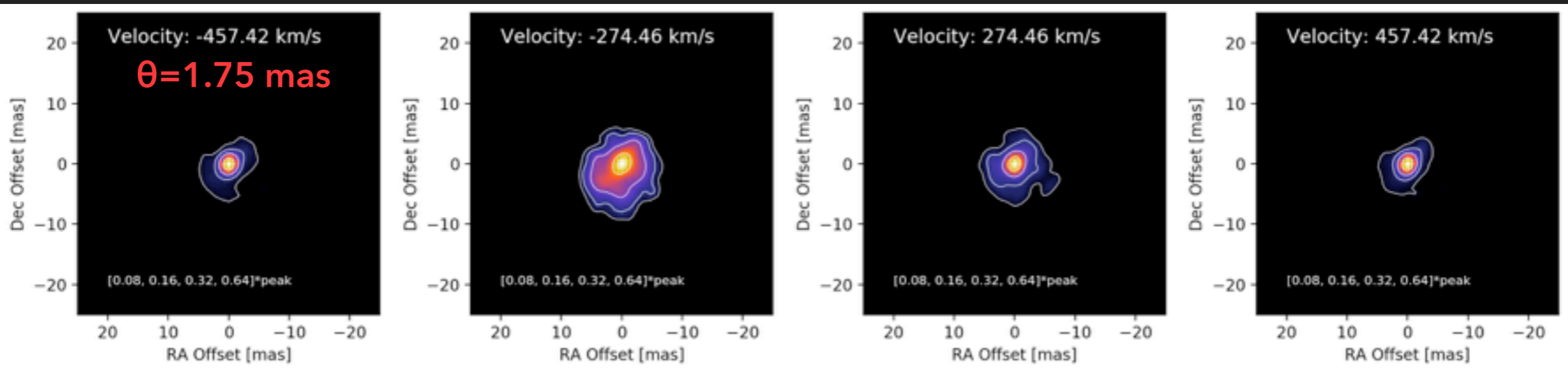
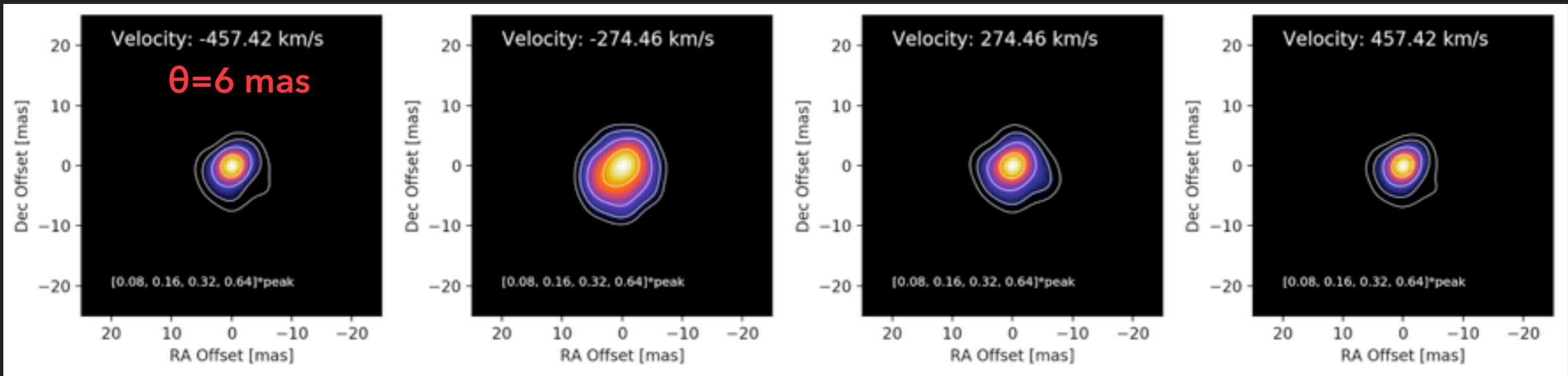
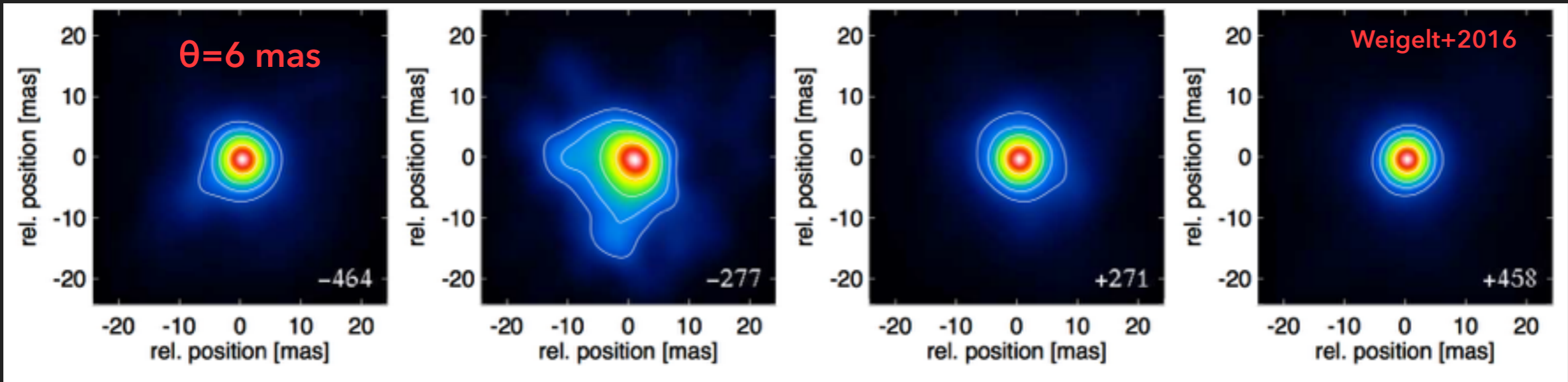
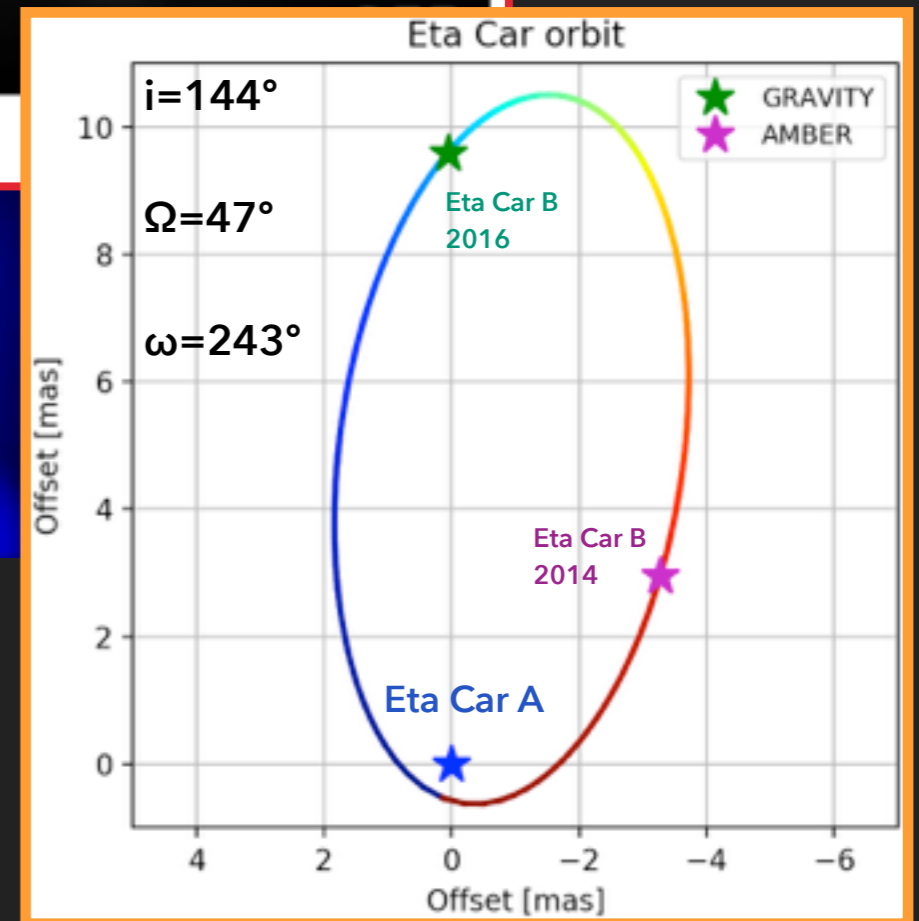
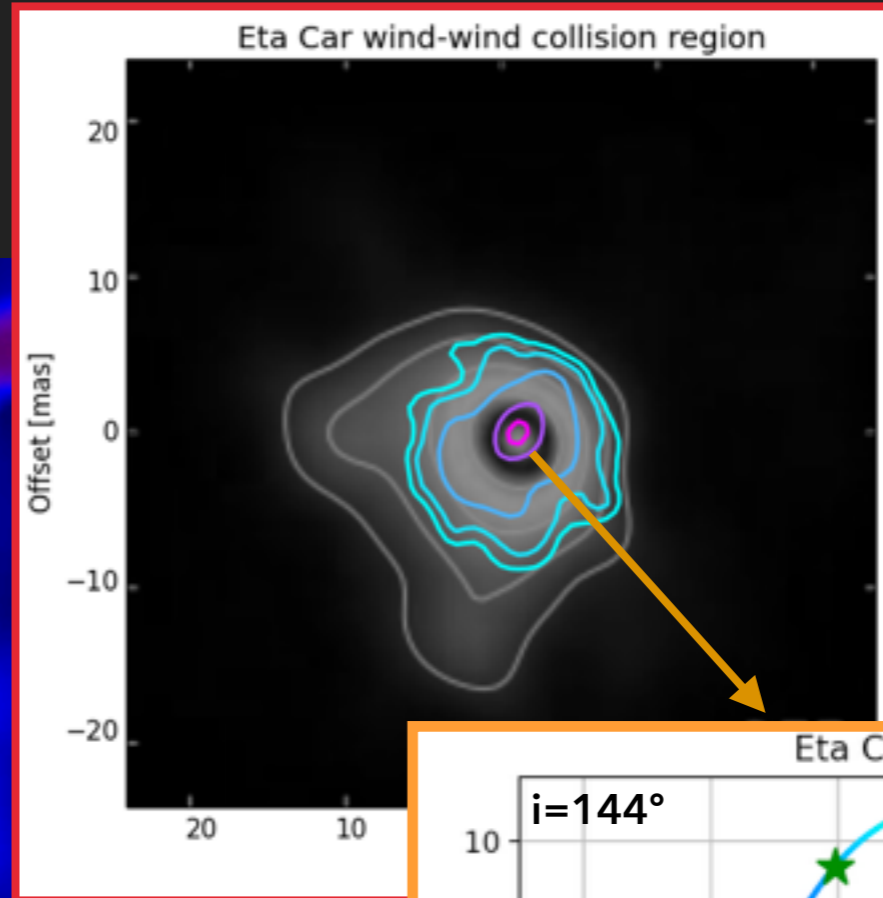
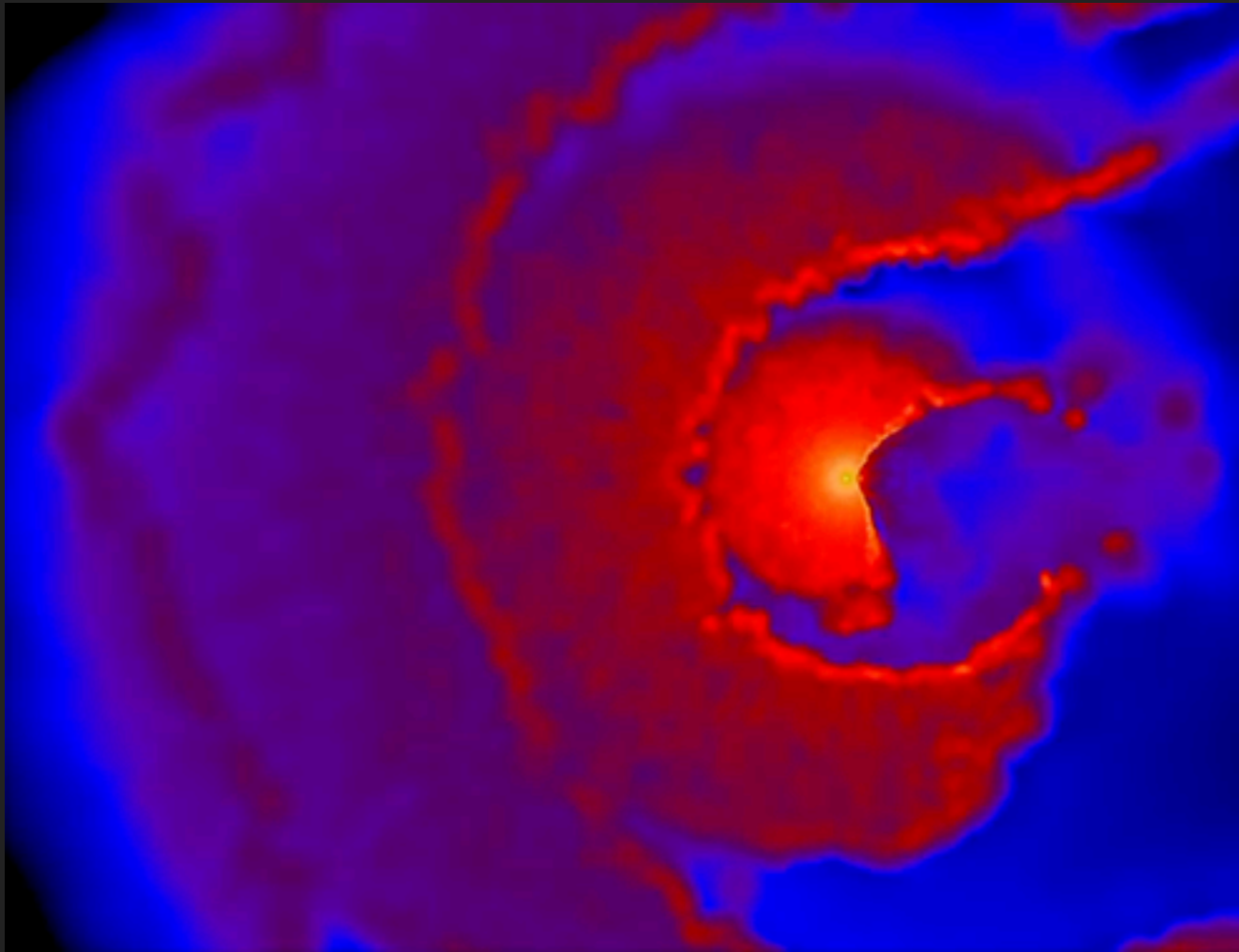


Image closure phases



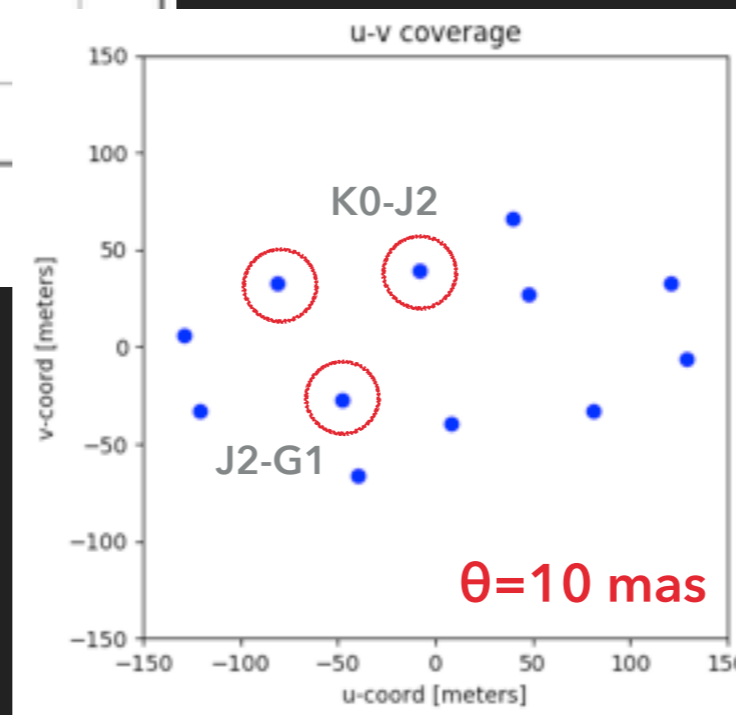
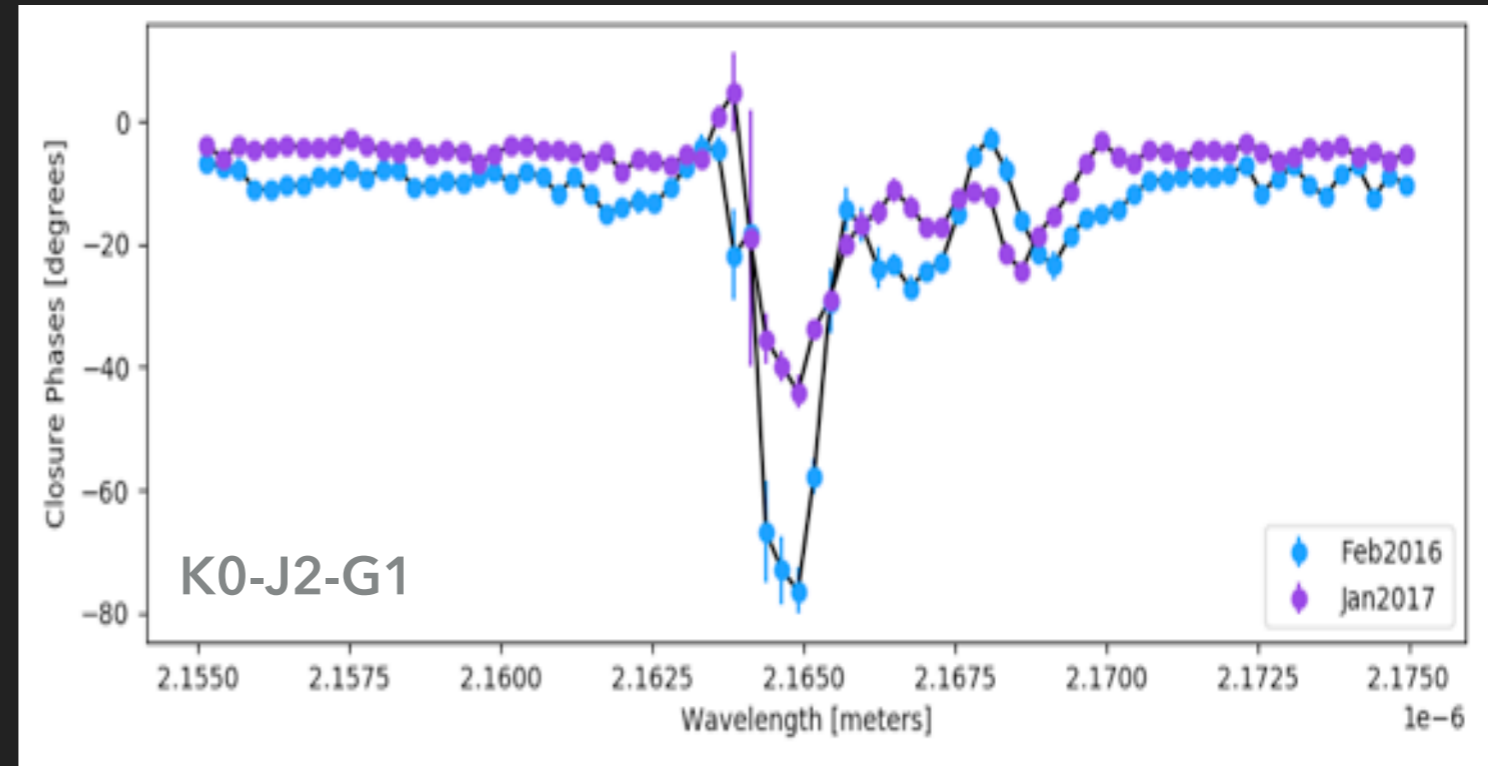
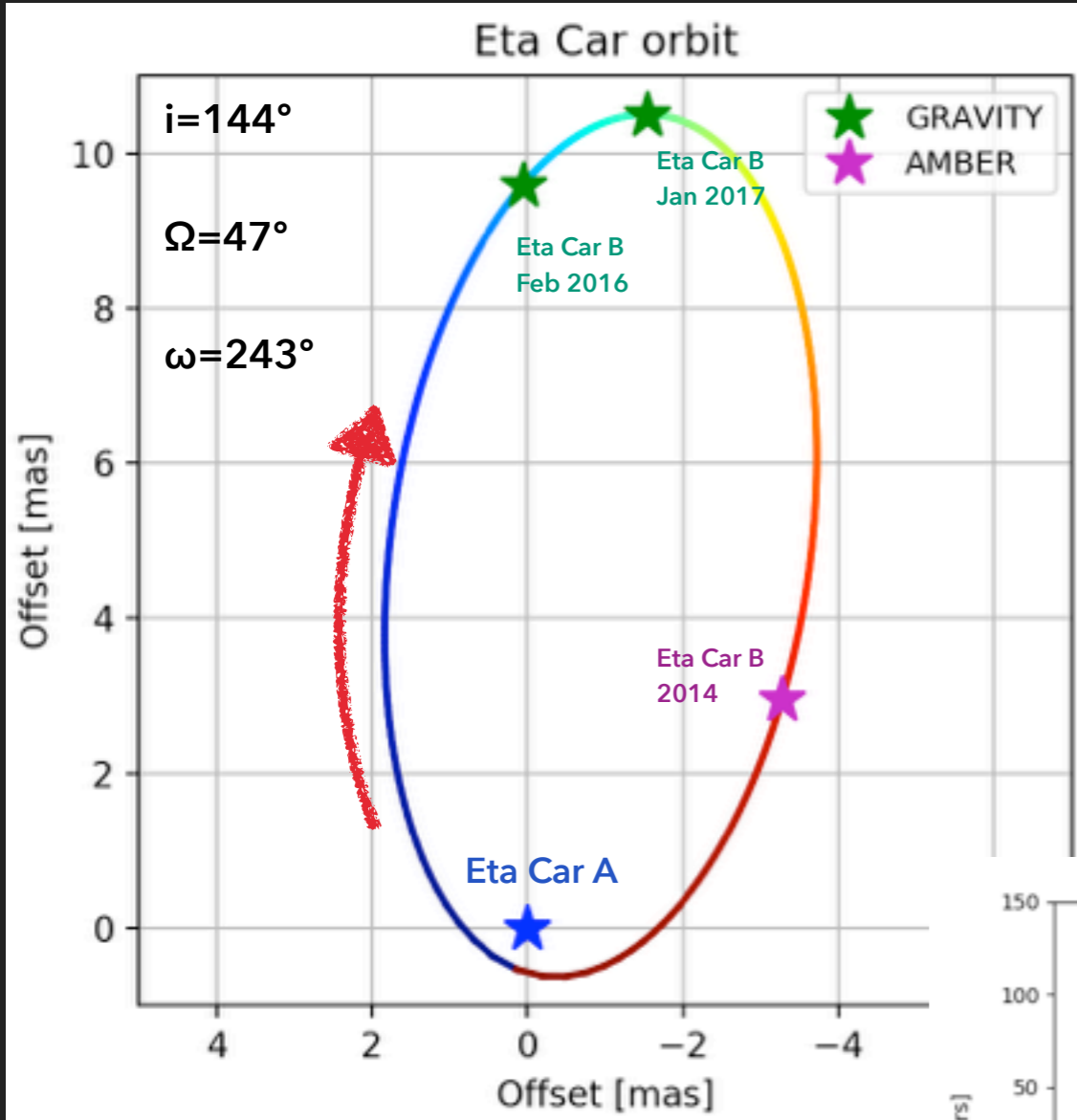


# WIND-WIND COLLISION REGION

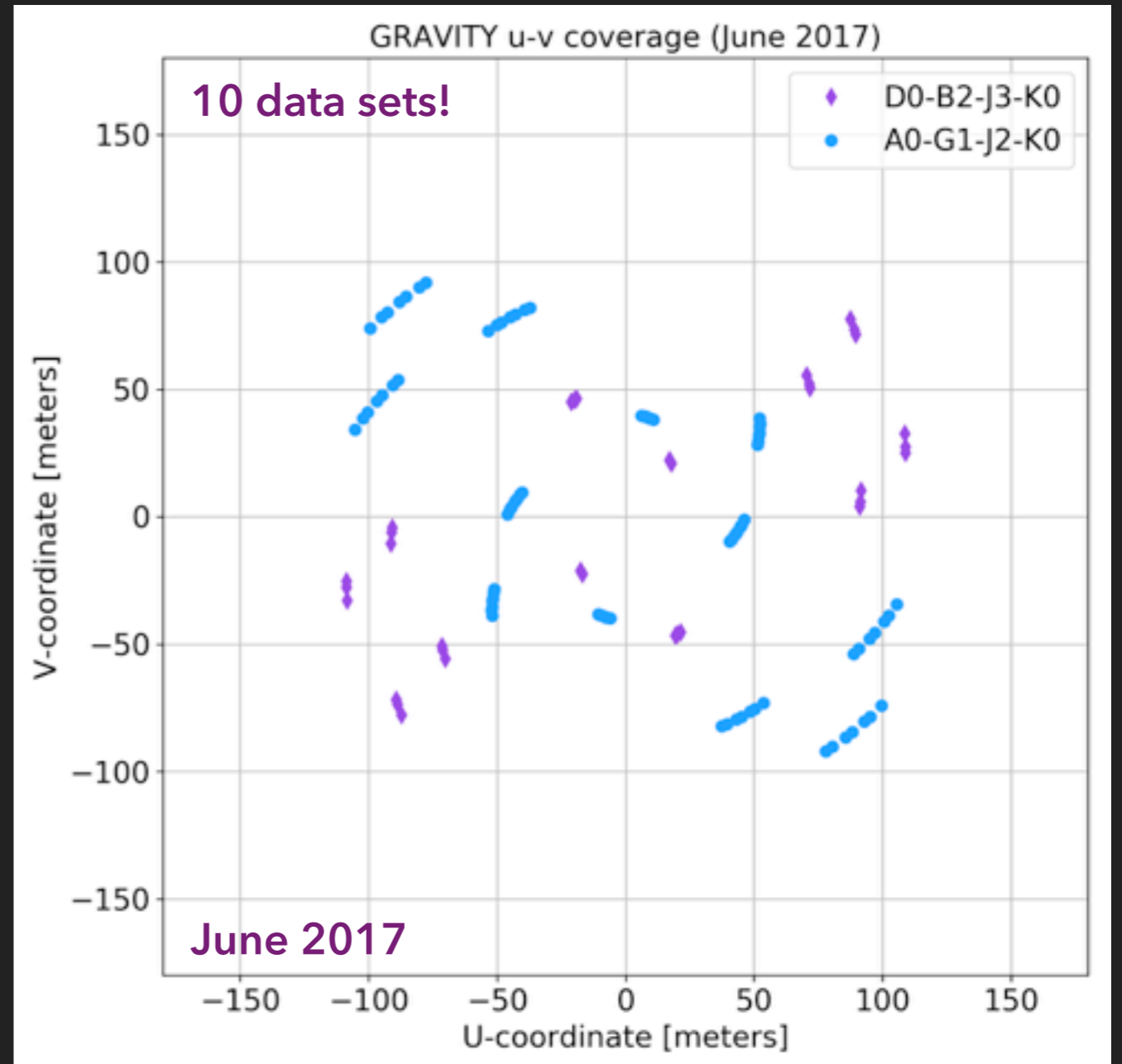
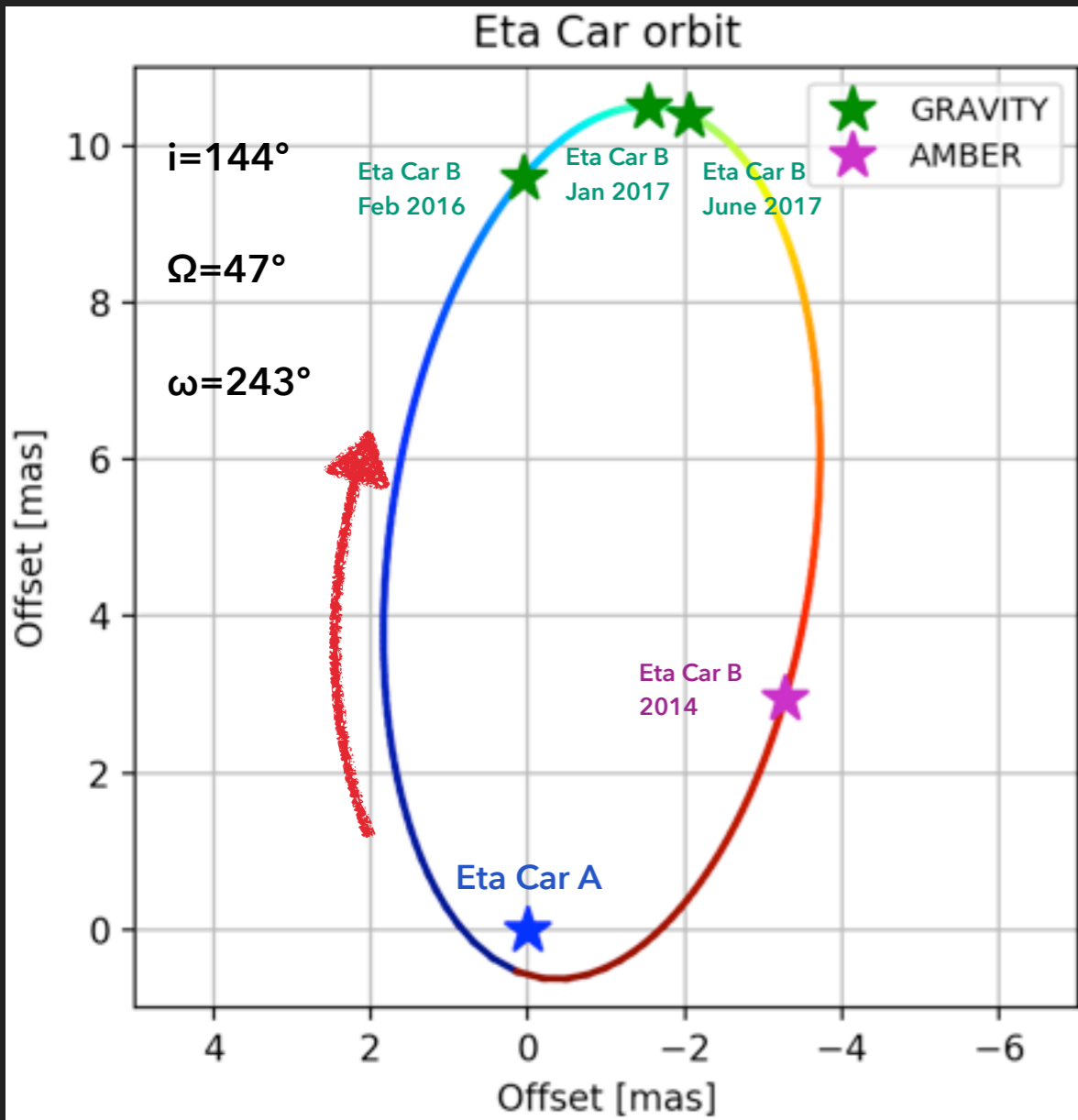


Hydrodynamical Simulations: Madura+2012; Madura +2013; Clementel+2015b

# WIND-WIND COLLISION REGION



# WIND-WIND COLLISION REGION



Second imaging epoch of GRAVITY data!

# WIND-WIND COLLISION REGION

Model fitting to V2 and closure phases

Image visibilities

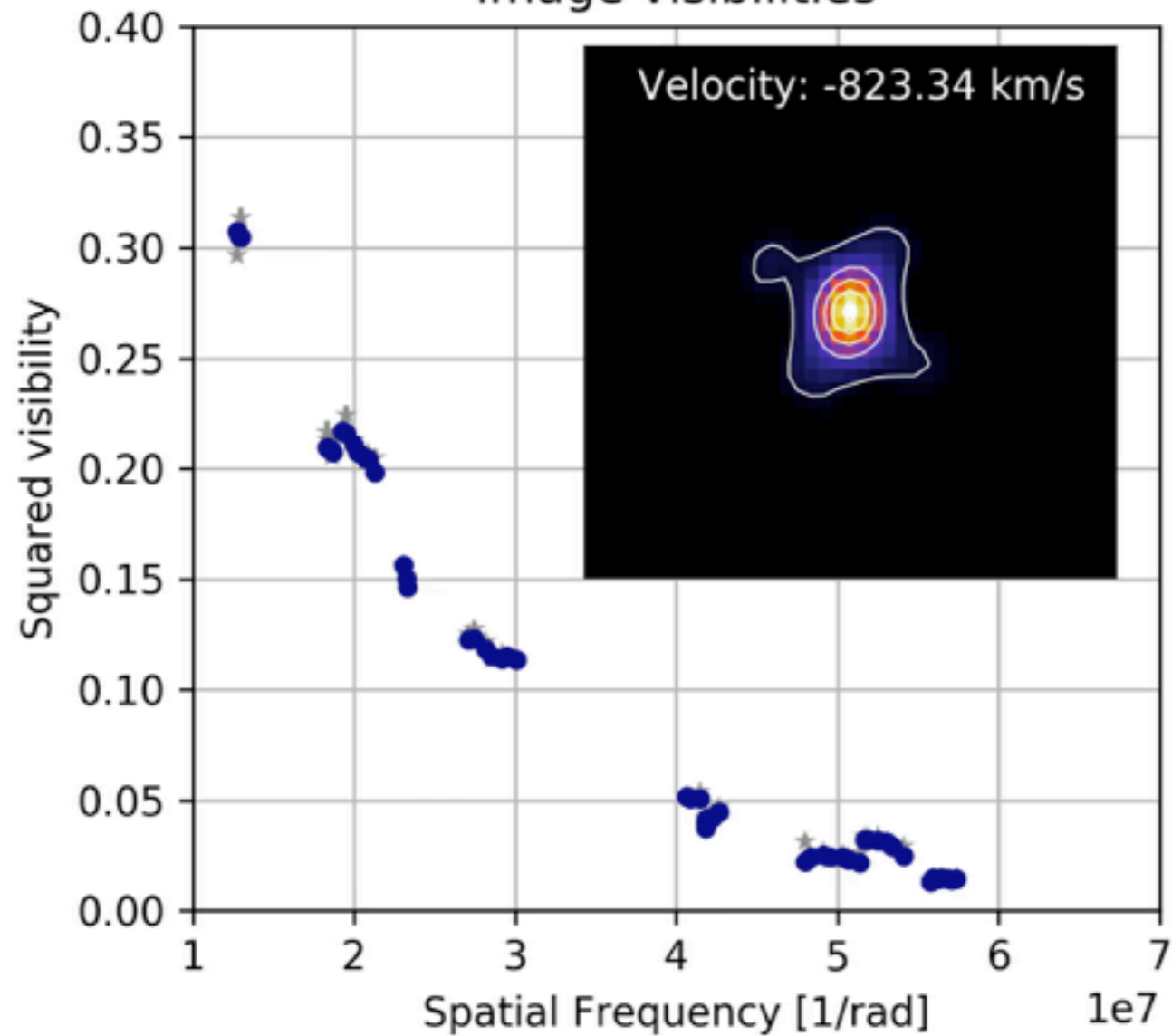
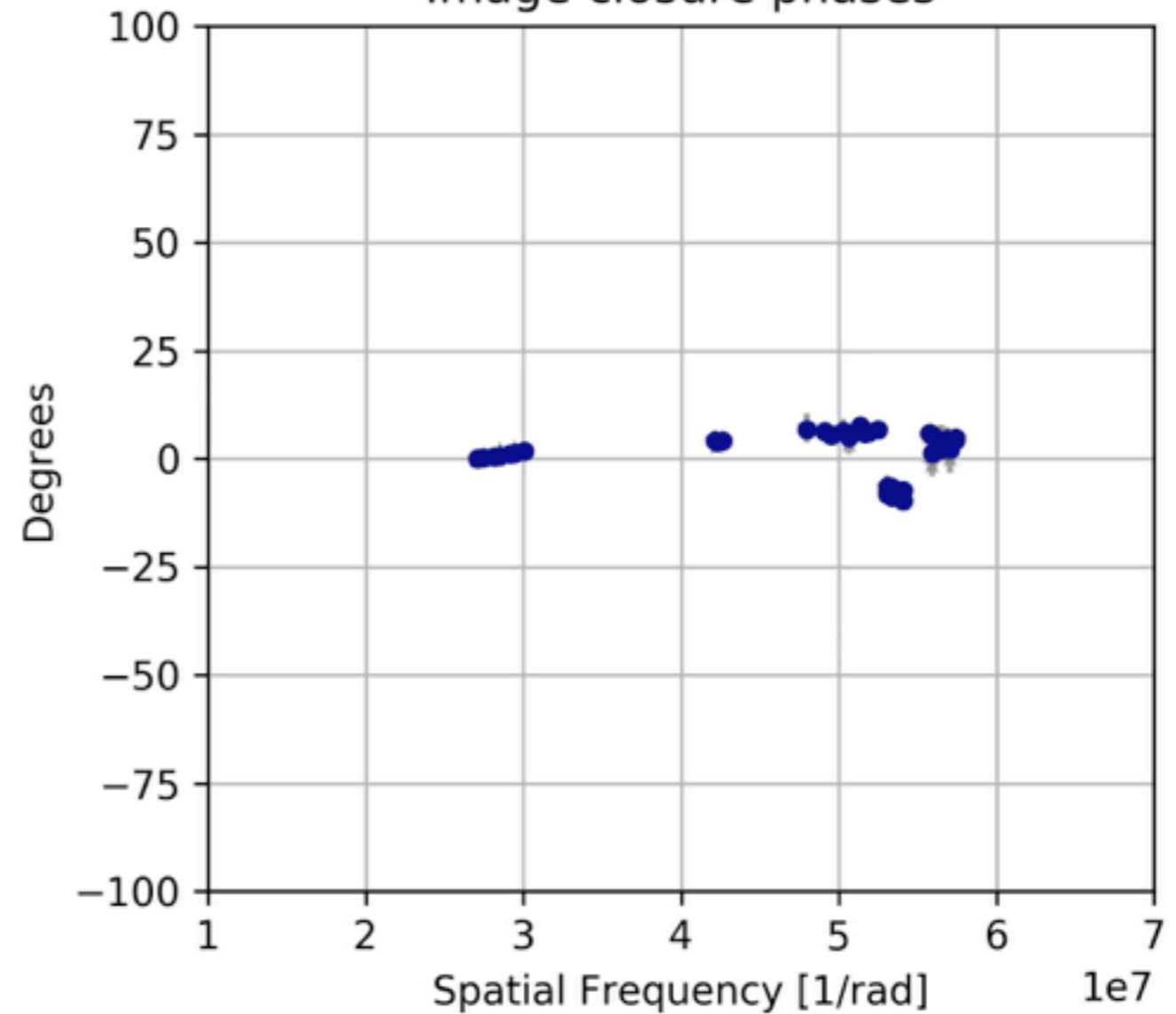


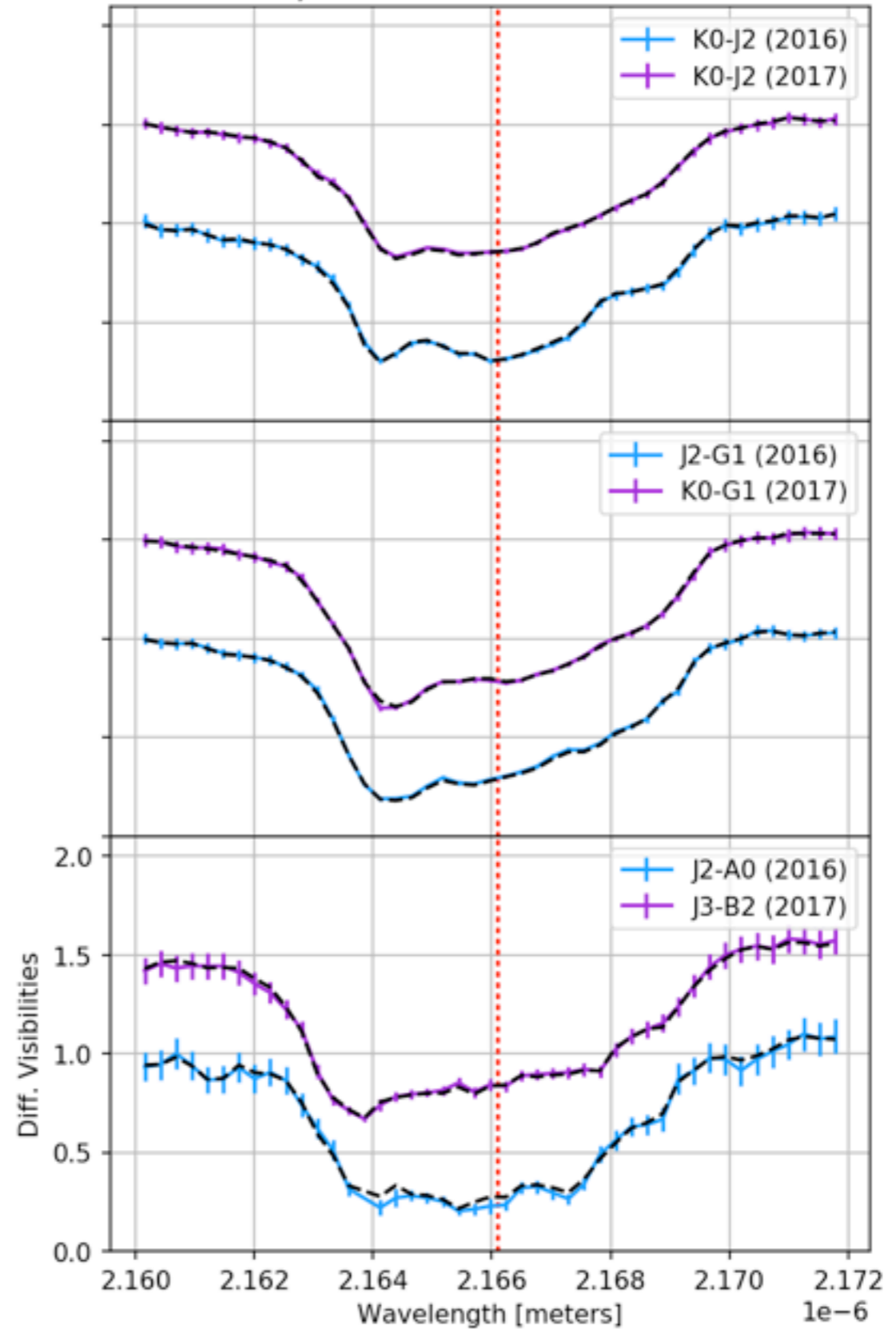
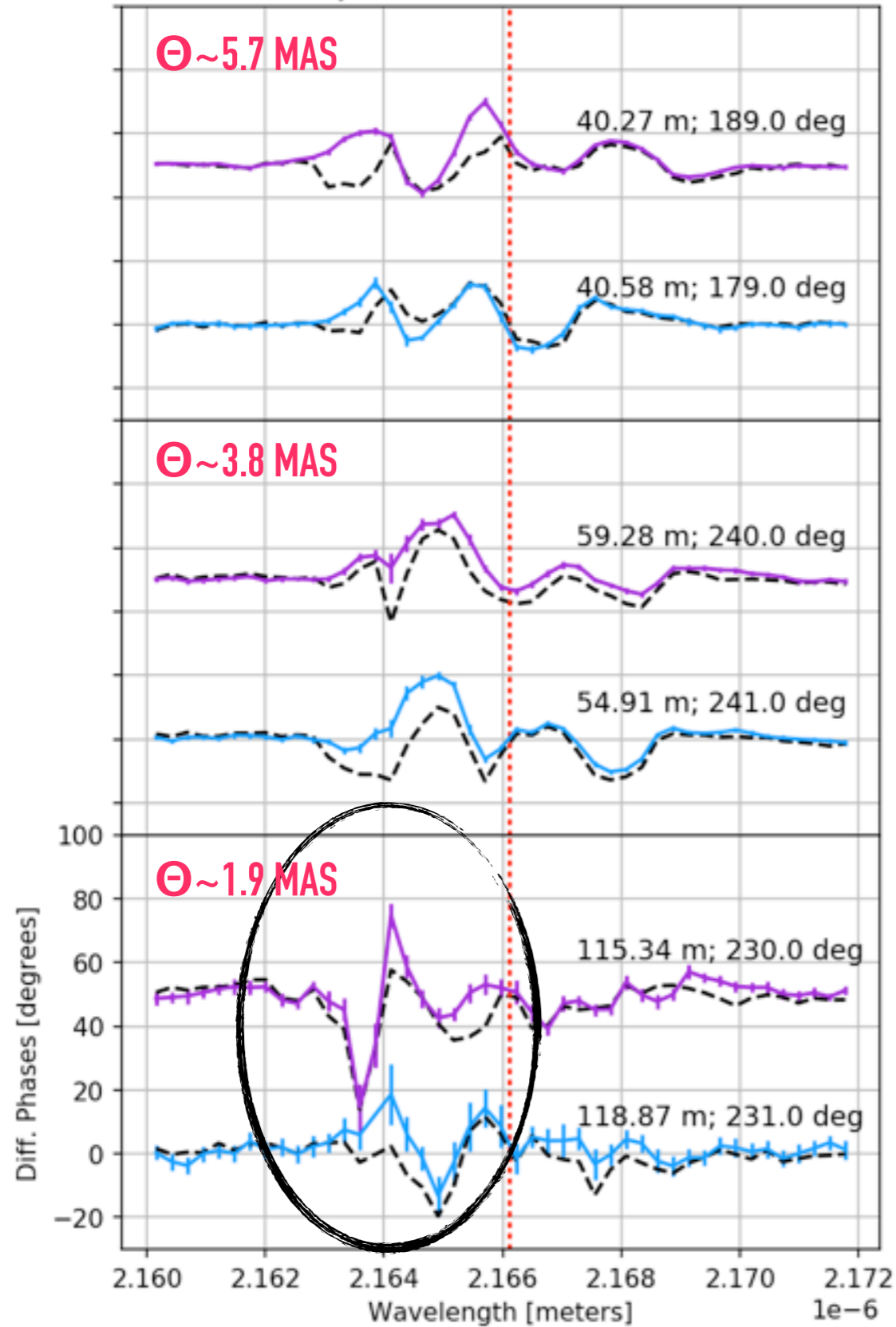
Image closure phases

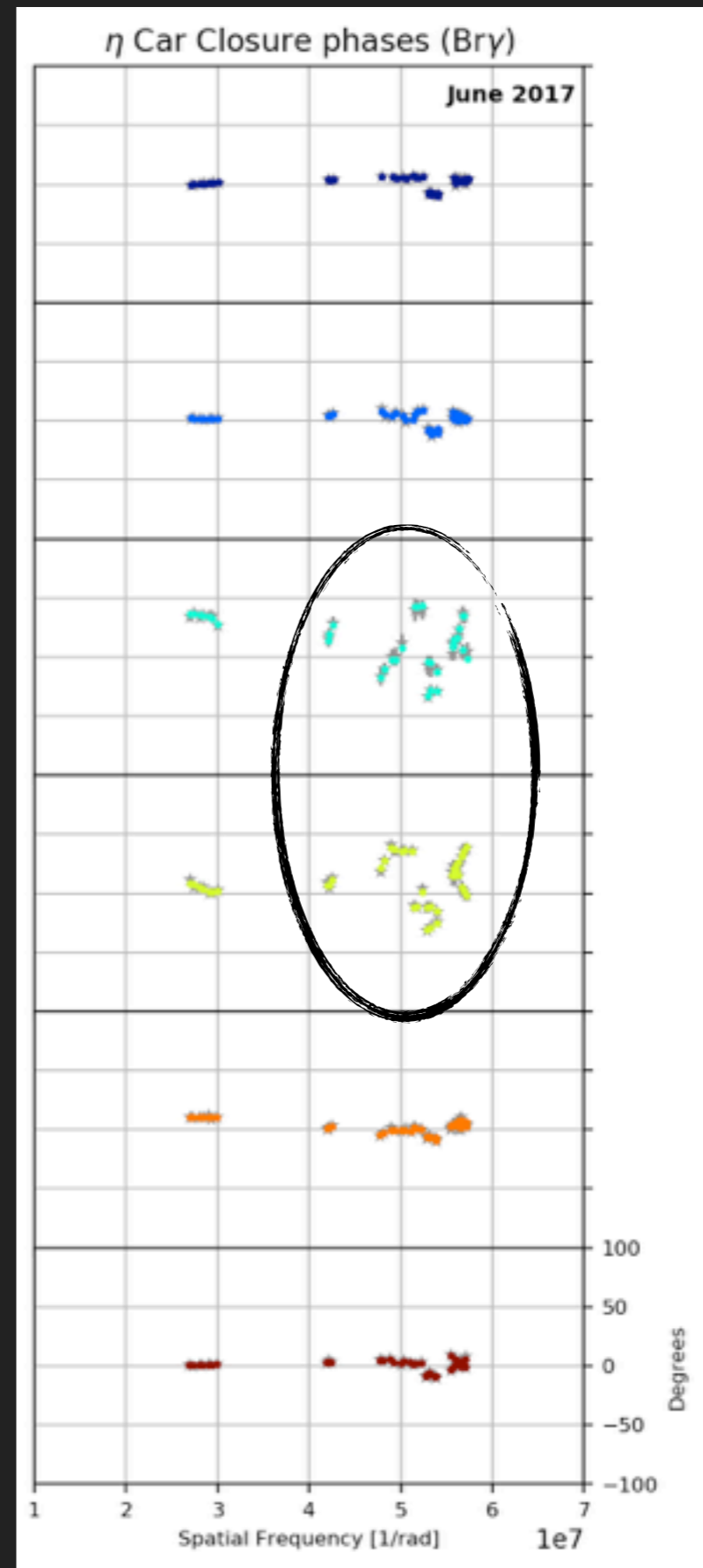
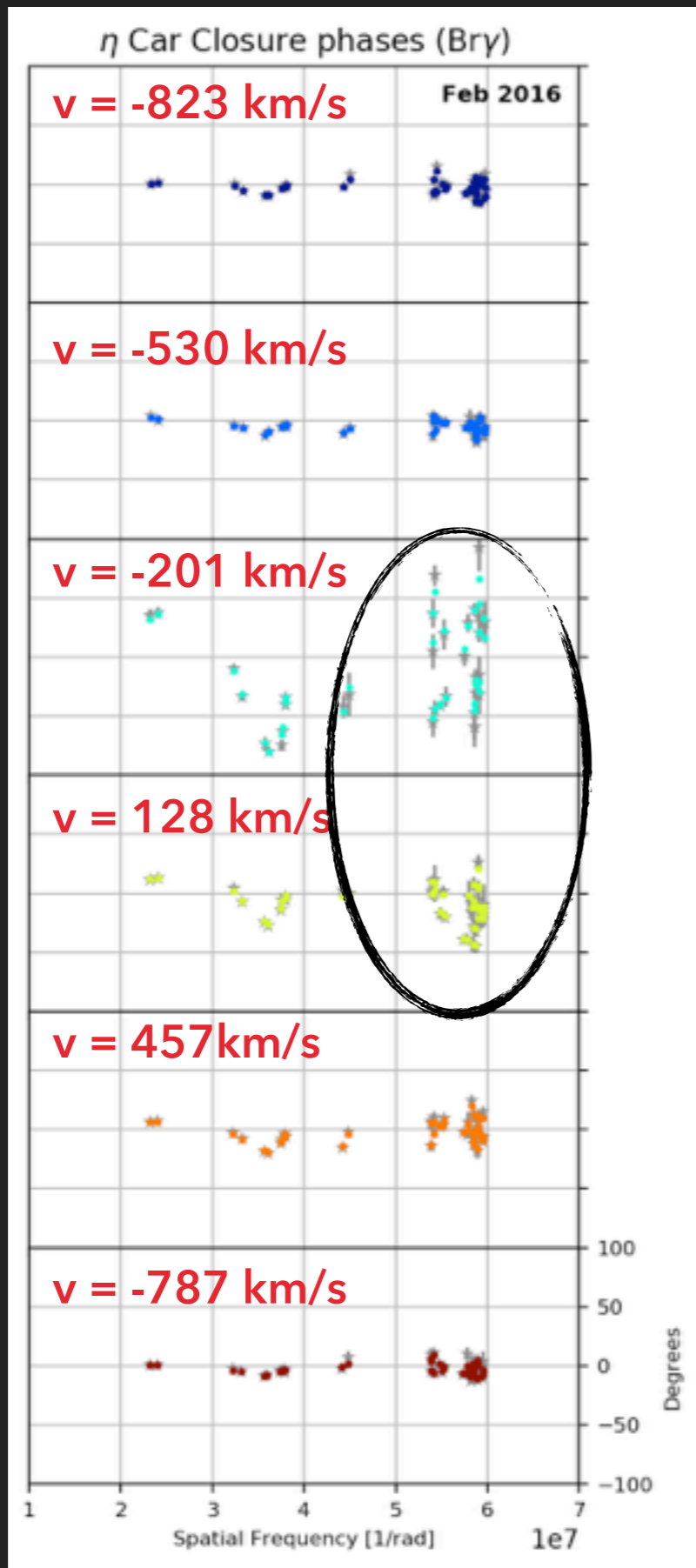




$\eta$  Car Diff. Phases

$\eta$  Car Diff. Visibilities





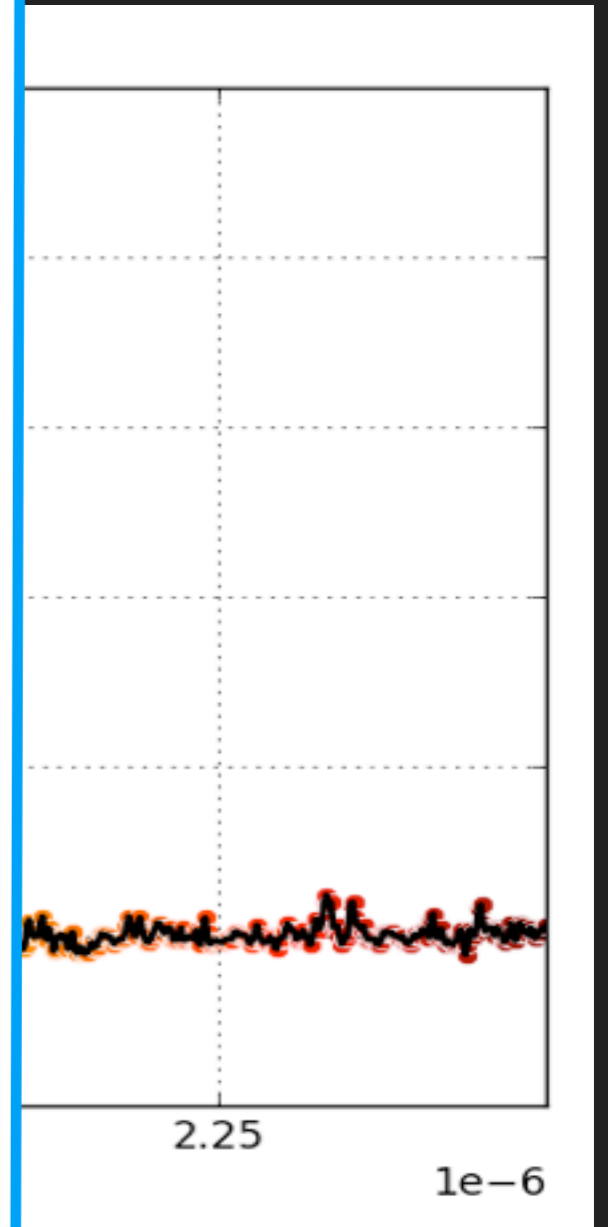
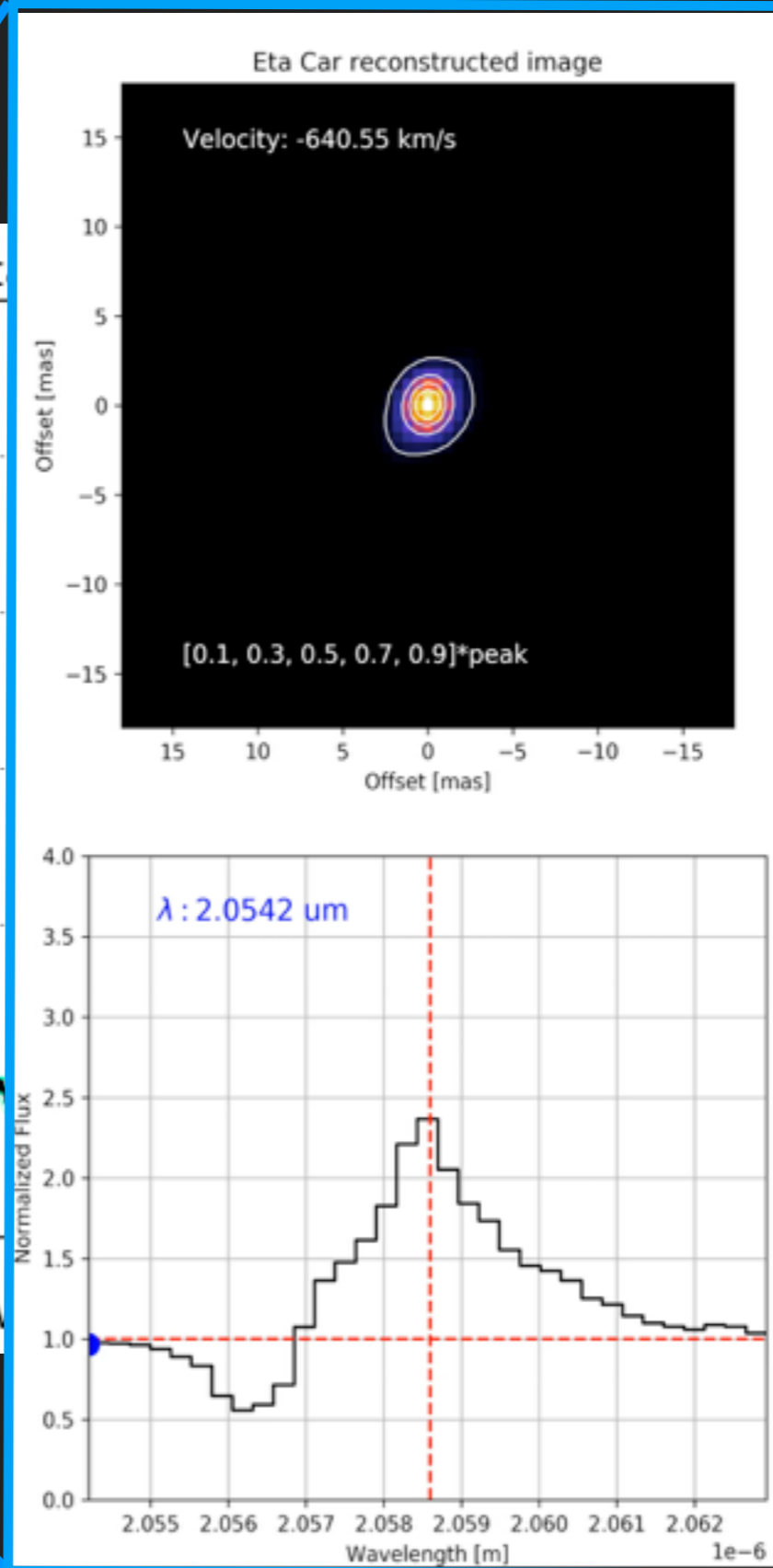
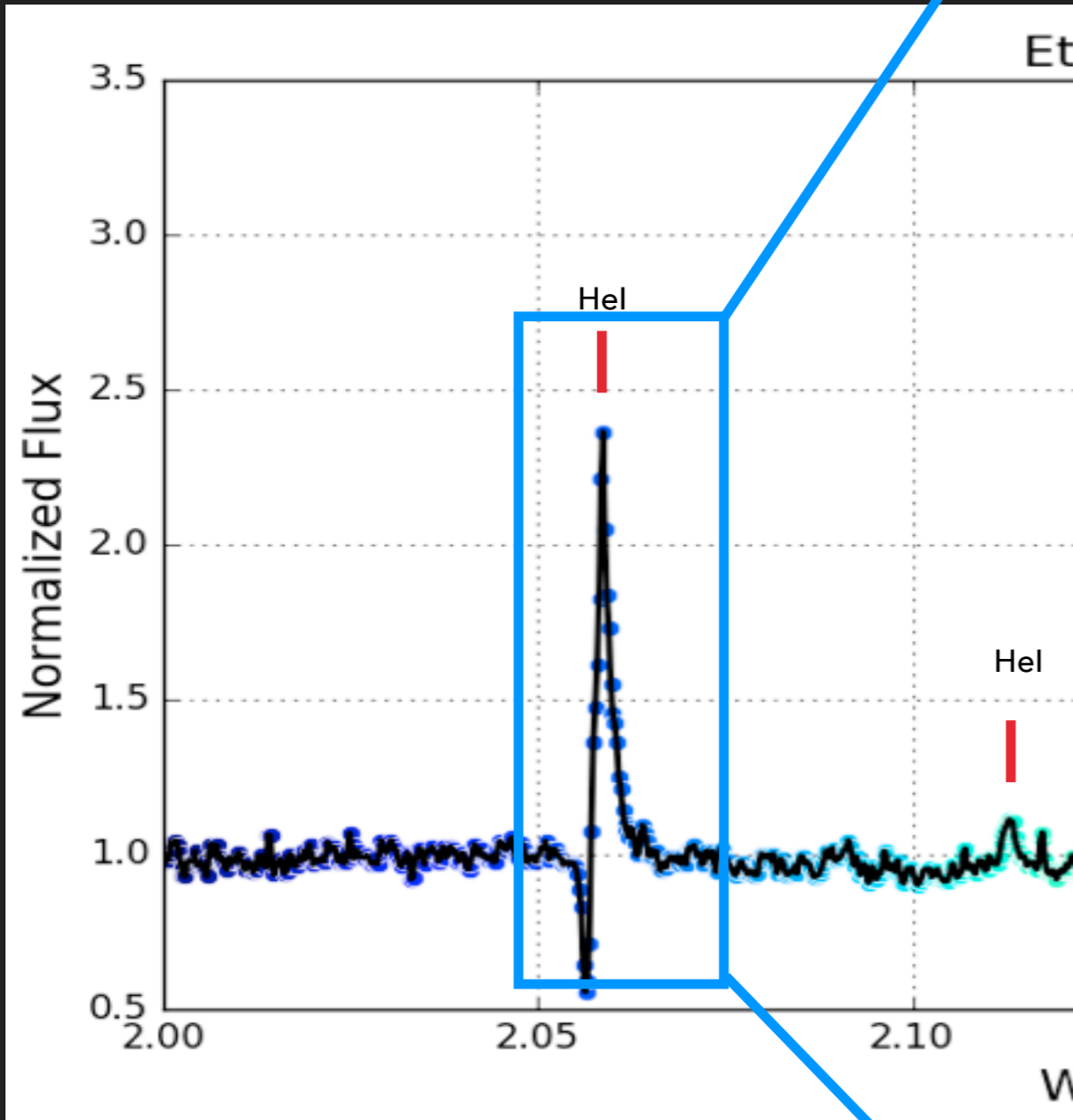
# ETA CAR: IMAGING THE CORE

## GRAVITY Science Beam Combiner: The Hel line

Band	Hel (2.054 - 2.063 $\mu\text{m}$ )
Pixel Grid	167x167 (100 mas)
Pixel Scale	0.6 mas/pixel
Chains	50
Iterations	250
Initial Image	Gaussian with 50% of the total flux
Observables	$V^2$ + CPs + Differential Phases
Regularizers	L0-norm (avoid point-like sources), Laplacian (favours extended sources), Transpectral reg. (L2-norm across the spec.)

# ETA CAR: IMAGING THE CORE

GRAVITY Science Beam Combiner

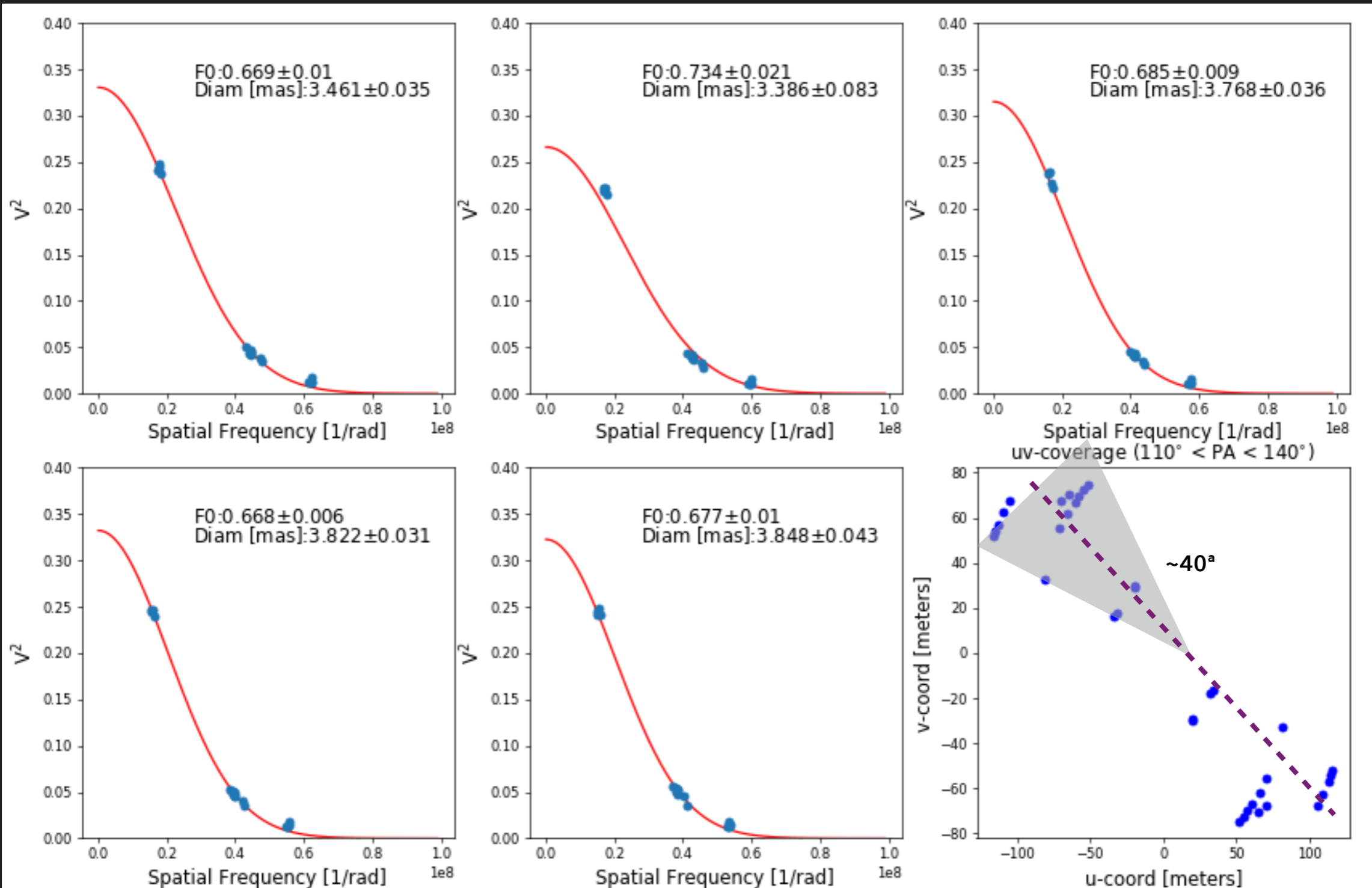


## SUMMARY

- 1** Continuum elongated core consistent with previous NIR Interferometric observations.
- 2** The extended structure in the images is consistent with the wind-wind collision scenario.
- 3** The most extended emission is observed at blue-shifted velocities (WWCZ in the LOS).
- 4** Time-dependent changes in the observables at all spatial scales (particularly for compact structures)
- 5** New reconstruction of images of the HeI (2.054) line
- 6** Future RT and Hydro simulations will help to constrain the physical parameters of the observed structures.

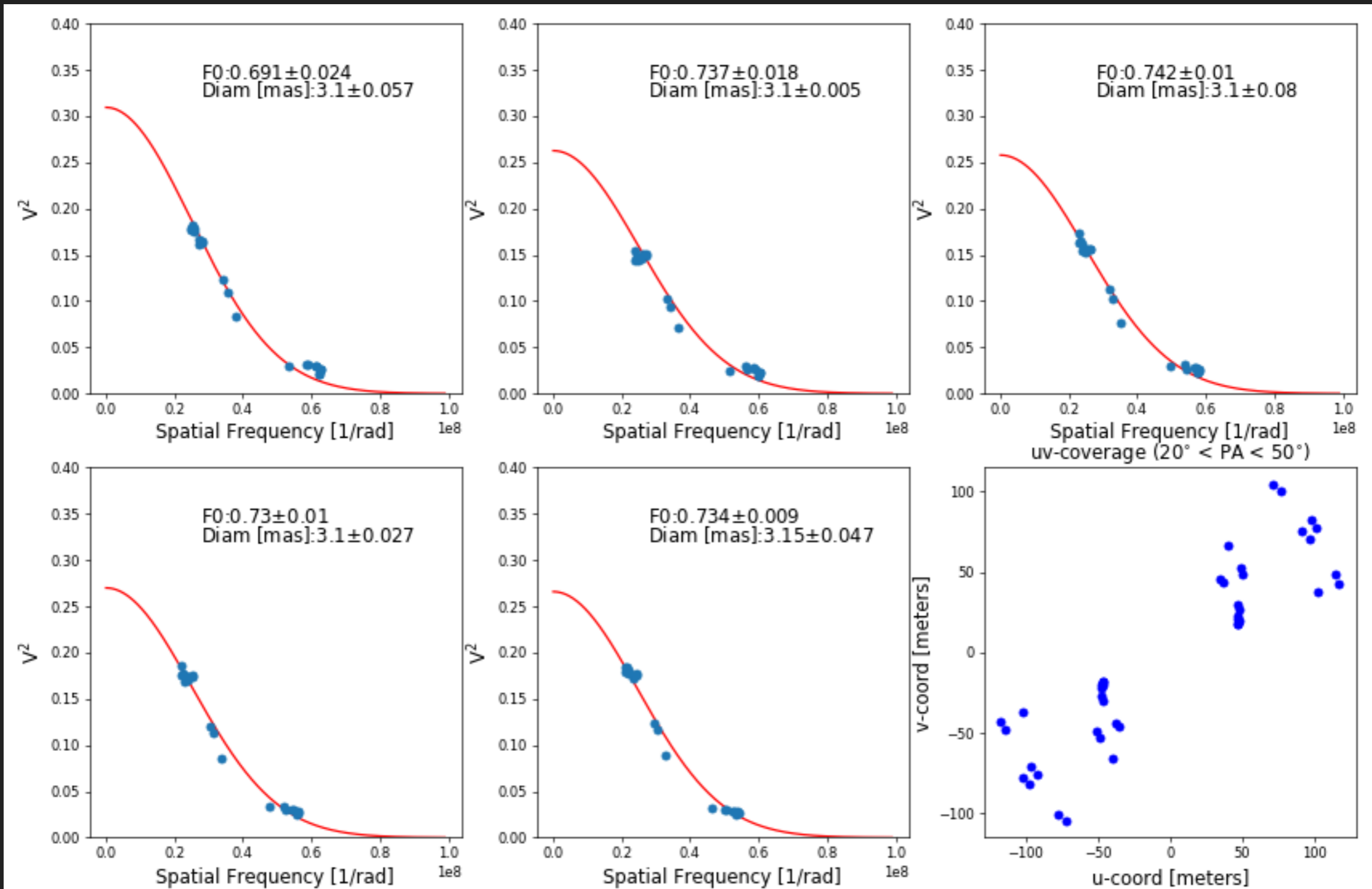
# ETA CAR: IMAGING THE CORE

GRAVITY Fringe tracker: Estimating the size of the core



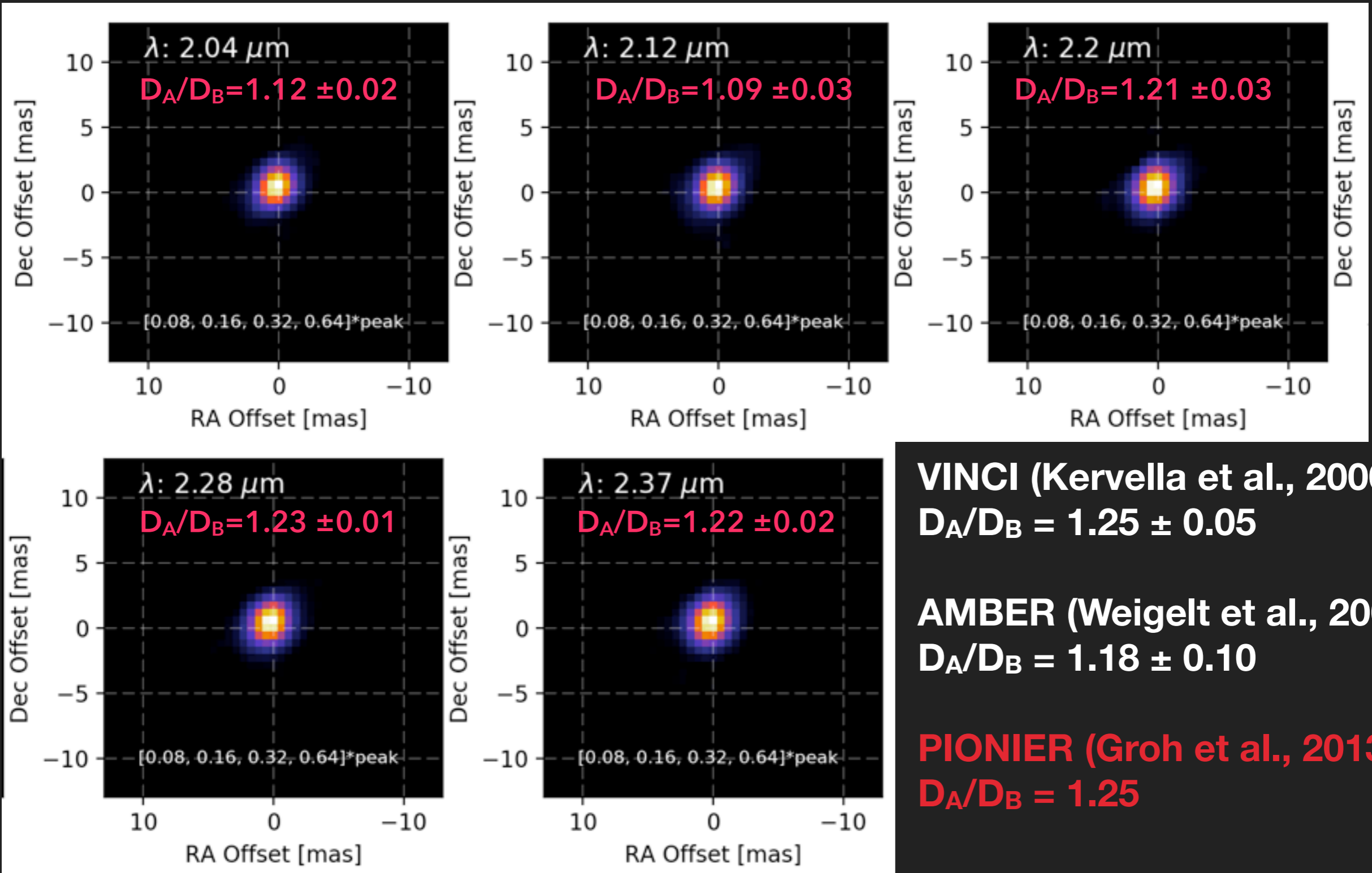
# ETA CAR: IMAGING THE CORE

GRAVITY Fringe tracker: Estimating the size of the core



# ETA CAR: IMAGING THE CORE

GRAVITY Fringe tracker: Estimating the size of the core



**VINCI (Kervella et al., 2000):**  
 $D_A/D_B = 1.25 \pm 0.05$

**AMBER (Weigelt et al., 2007):**  
 $D_A/D_B = 1.18 \pm 0.10$

**PIONIER (Groh et al., 2013):**  
 $D_A/D_B = 1.25$



CMFGEN

Groh+2012: HST spectrum (Visible)

$\dot{M}=8.5e-4 M_{\odot}/yr$

$v_{\infty}=420 \text{ km/s}$

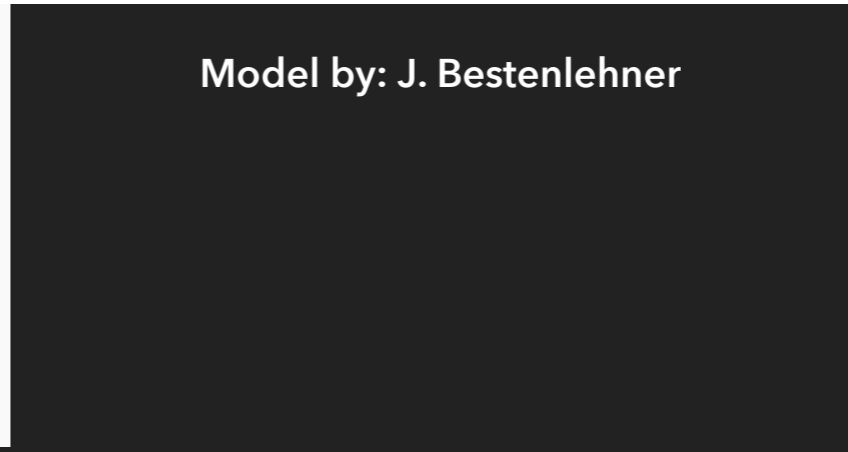
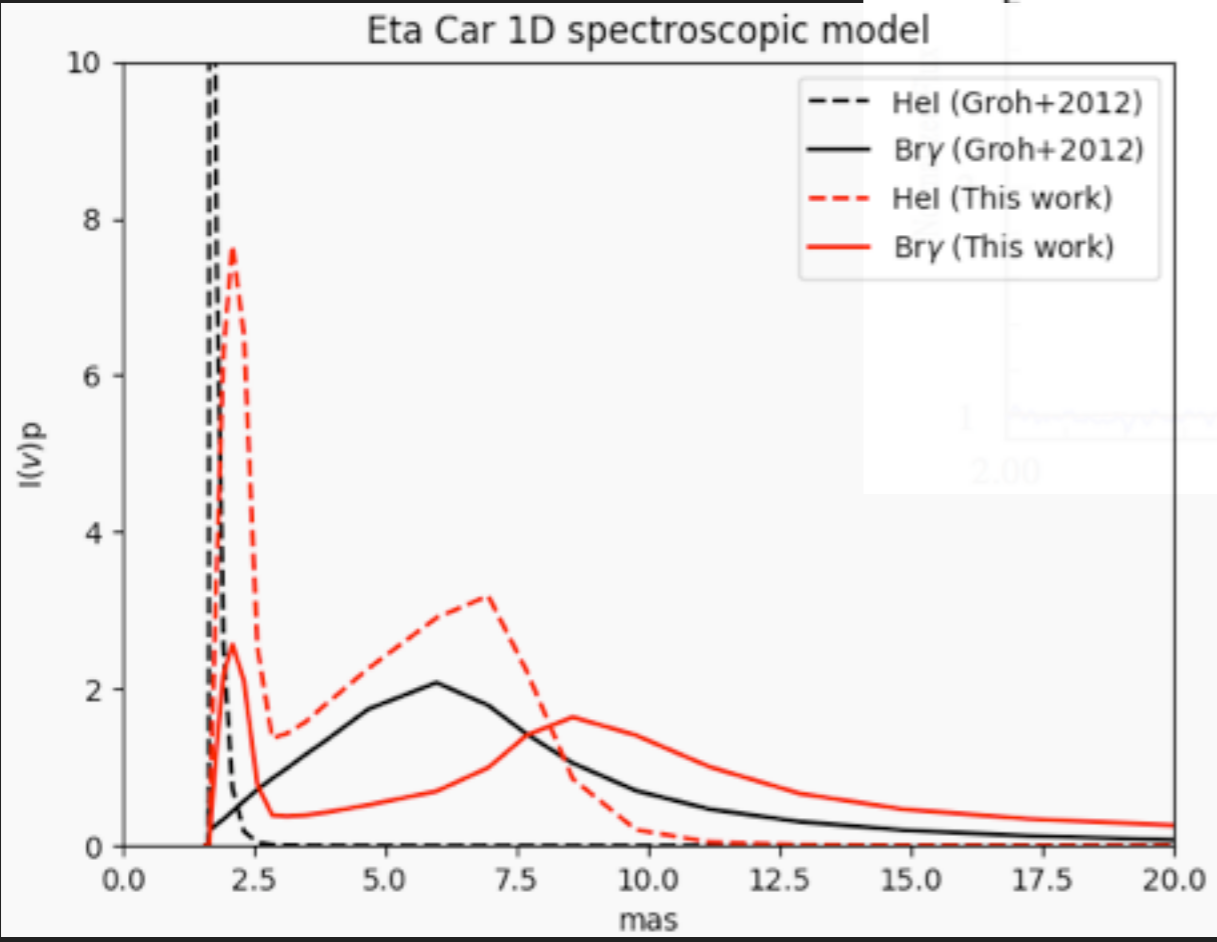
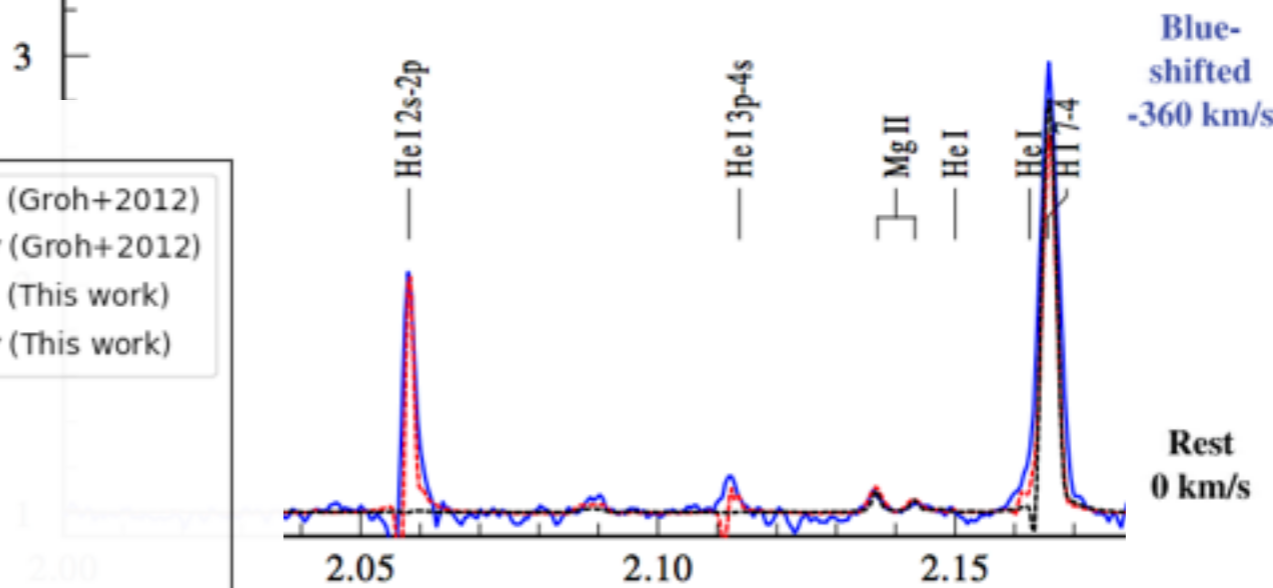
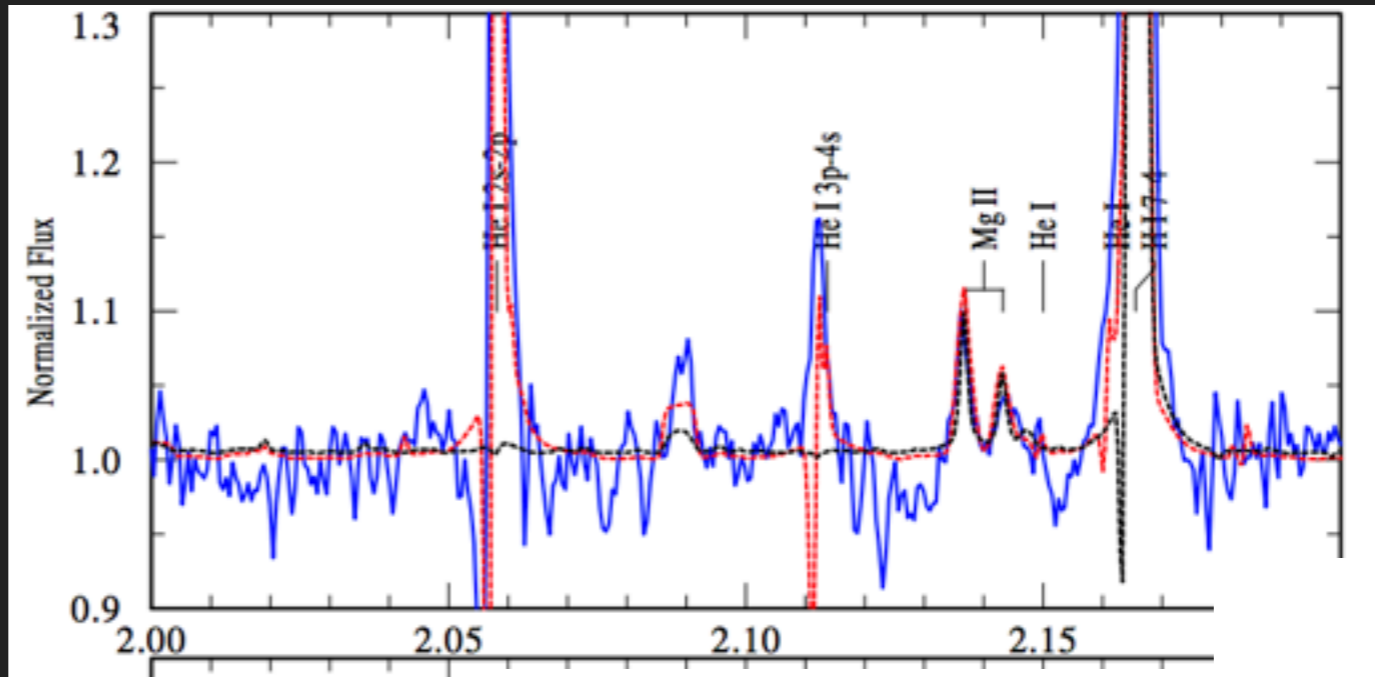
$T_{eff}= 9400 \text{ K}$

This work:

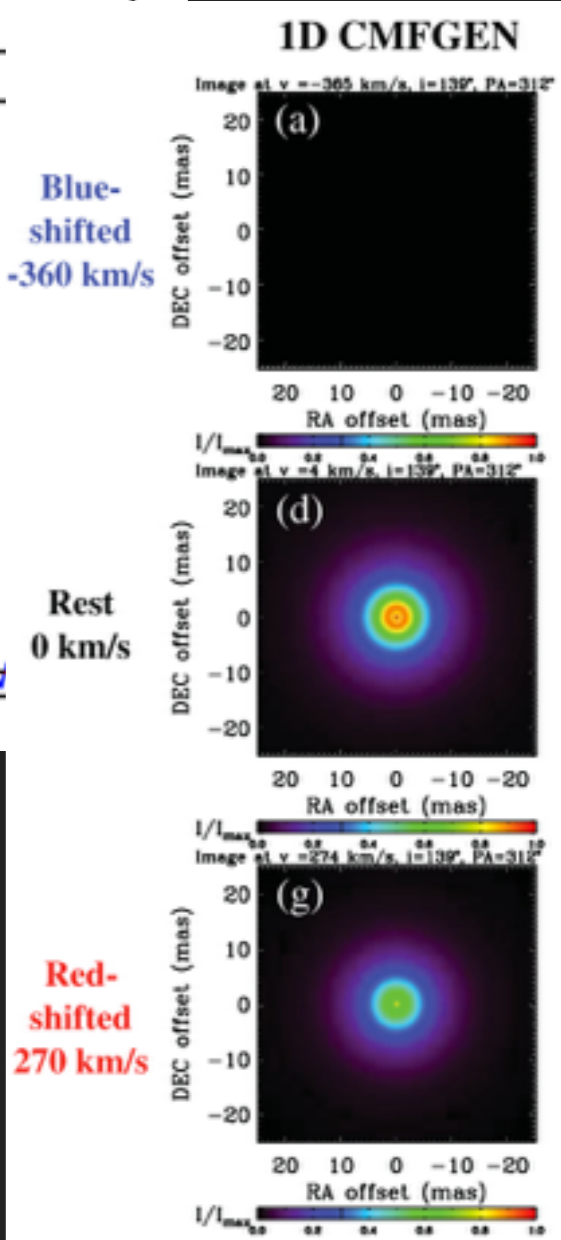
$\dot{M}=1.7e-4 M_{\odot}/yr$

$v_{\infty}=420 \text{ km/s}$

$T_{eff}= 13000 \text{ K}$



Model by: J. Bestenlehner





**THANK YOU!**