



**ALMA observations of
 $z > 6.5$ quasar hosts:**

**Massive galaxy formation in
the epoch of reionisation**

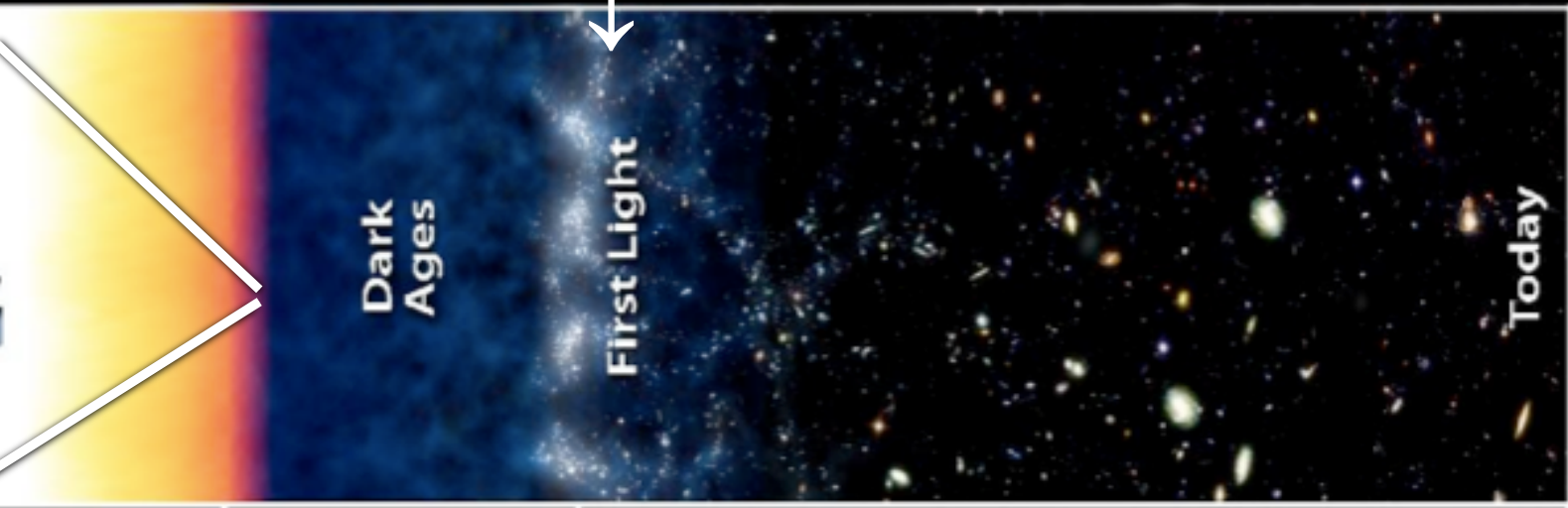
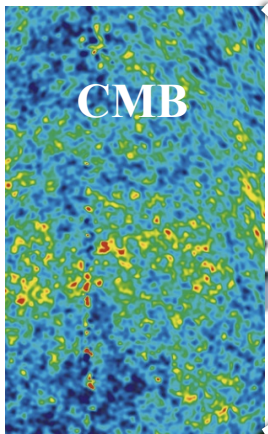
Bram Venemans (MPIA)

The Epoch of Reionisation

Recombination

Epoch of Reionisation:
first luminous sources

Now



0 yrs

400,000 yrs

400 million yrs

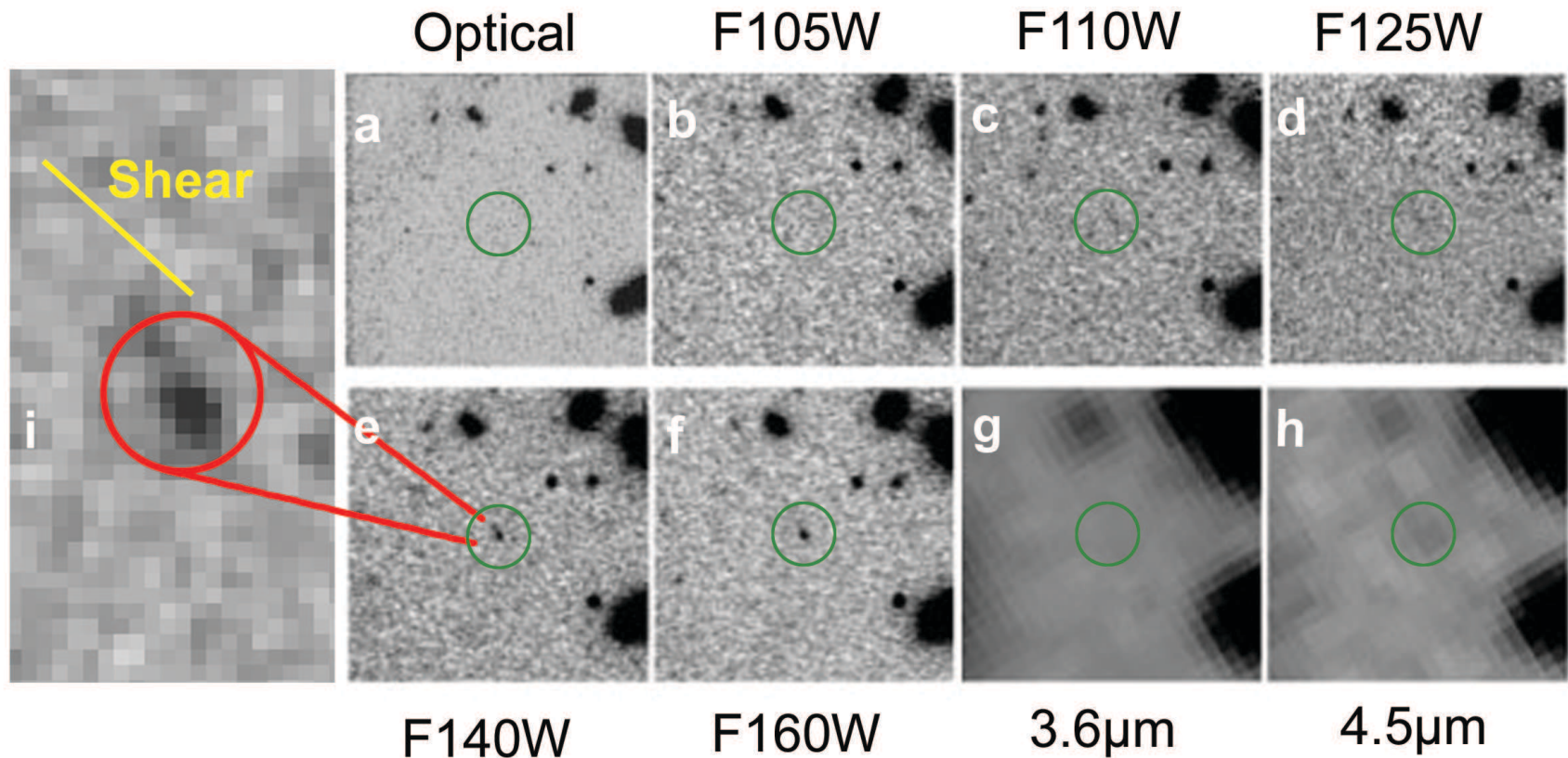
13.7 billion yrs

Time after Big Bang →

Background

- When did the first galaxies and black holes form and what were their characteristics?
- Tremendous progress in discovering galaxies out to $z > 10$ (e.g. Bouwens+ 11,12,13; Mclure+ 10,11,13; Oesch+ 12,13; Ouchi+ 09,10; Zheng+ 12)
- Hard to study these galaxies in detail with current facilities

Example: MACS1149-JD at $z \sim 9.6$ (Zheng+ 12)

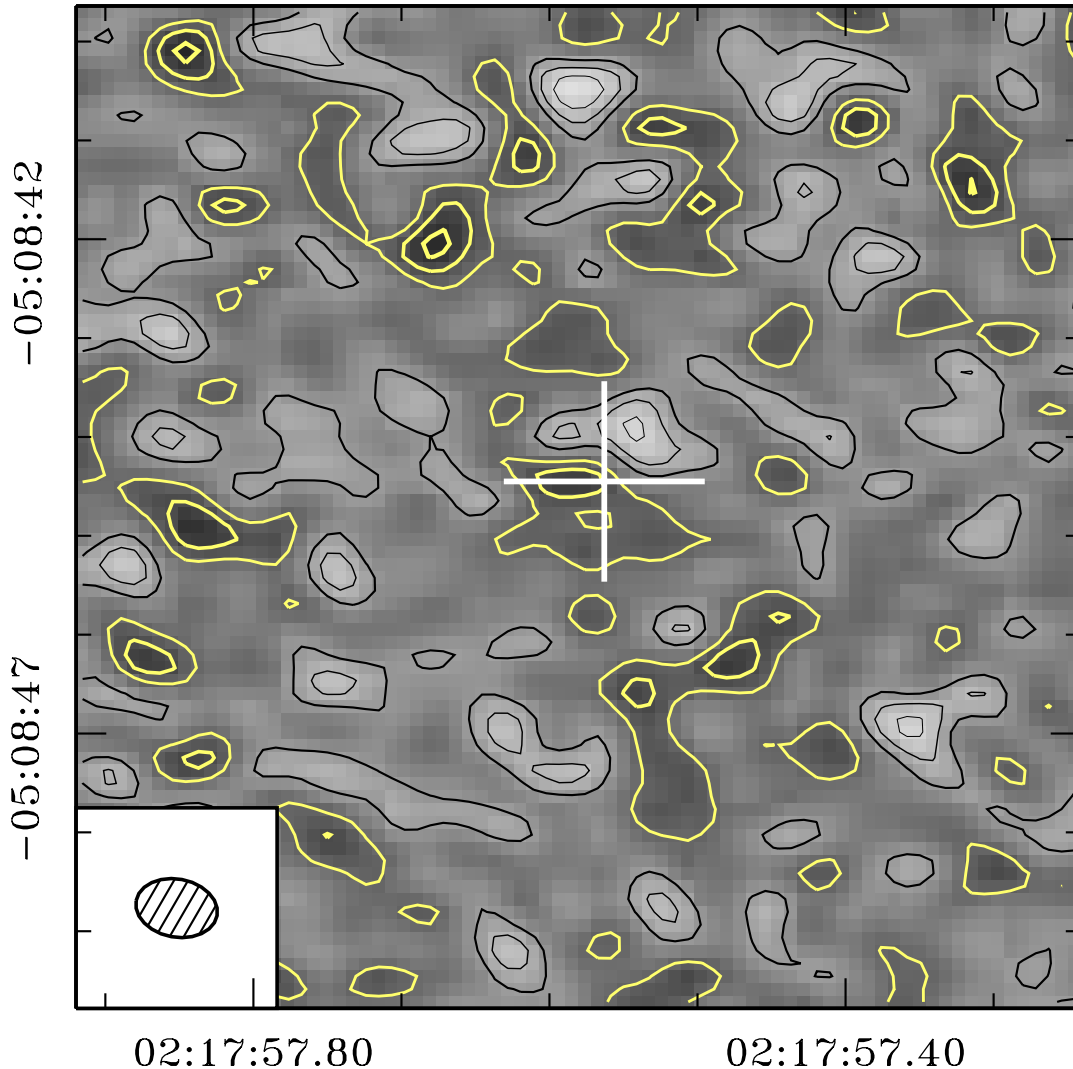


Follow-up faint sources difficult

Even getting redshifts is challenging...

“Himiko” at $z=6.595$: $\text{SFR}_{\text{UV}} \sim 100 M_{\text{sun}}/\text{yr}$

Undetected in the
FIR continuum in
3.2 hrs with ALMA
Cycle 0 (16 ant.)



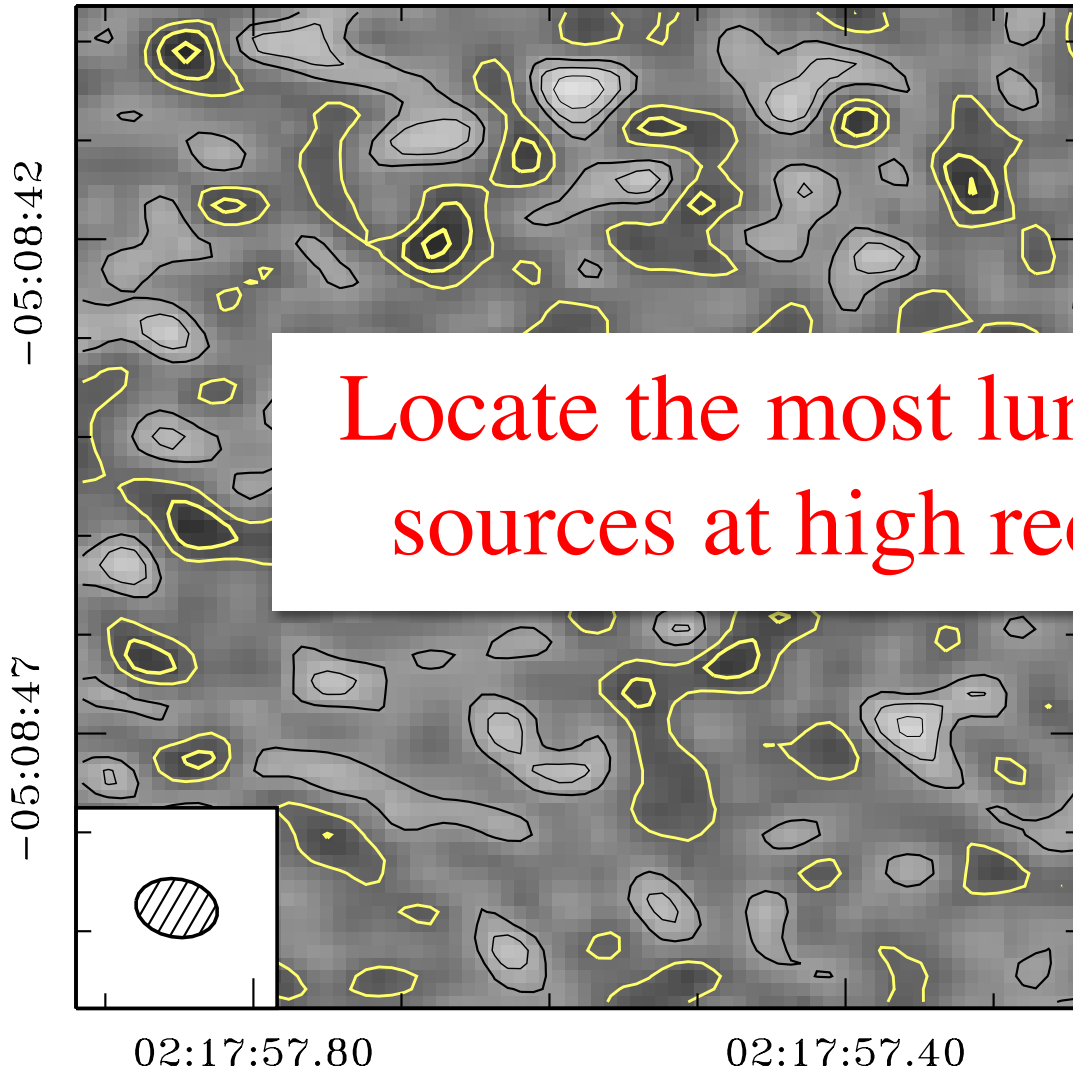
Ouchi+ 2013

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Locate the most luminous
sources at high redshift

Ouchi+ 2013

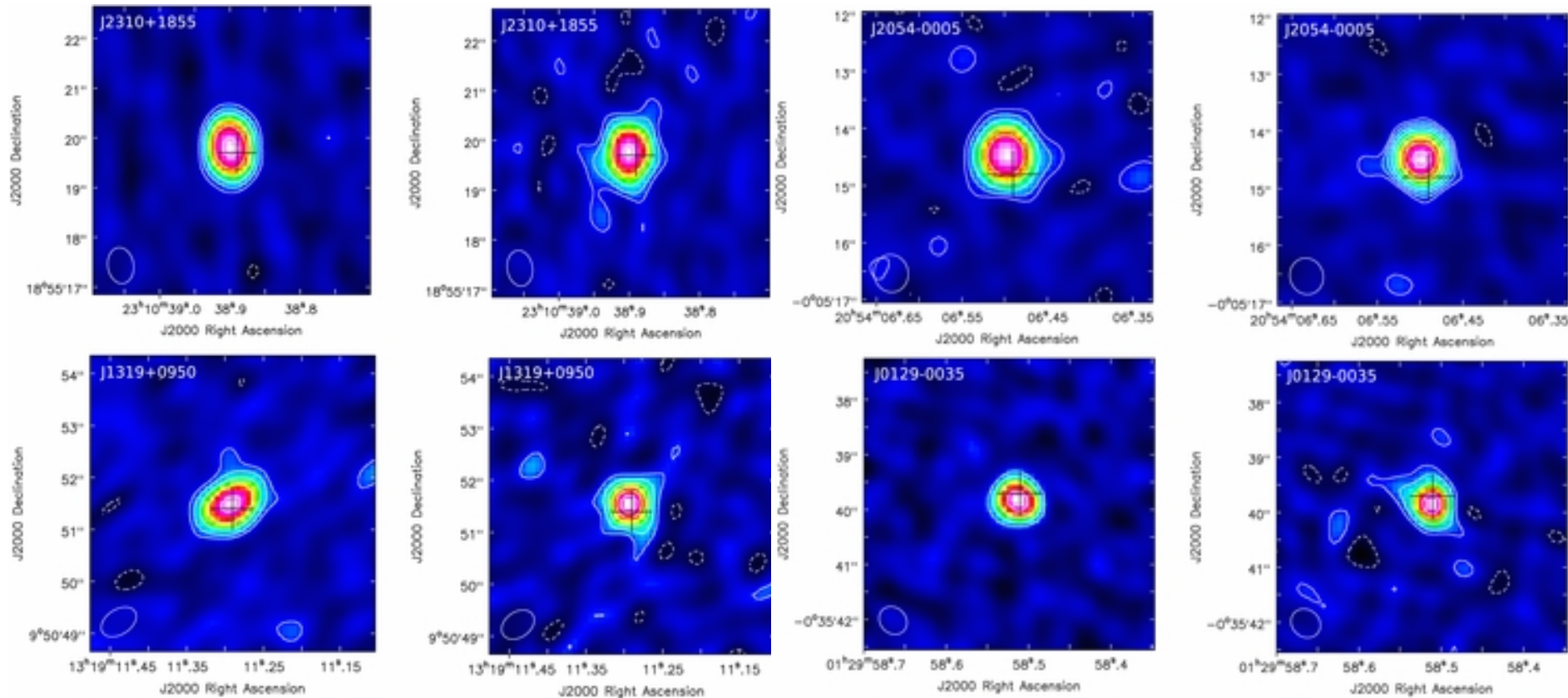


Massive black holes at high redshift

- Find massive black holes: look for quasars
 - Problem: quasars are very rare
 - Need multicolour surveys over large area
- SDSS very successful in discovering many luminous quasars up to $z=6.4$ (J1148 at $z=6.4$)

Dust and [CII] emission in SDSS quasars

ALMA Cycle 0 data of mm bright quasars at $z \sim 6$

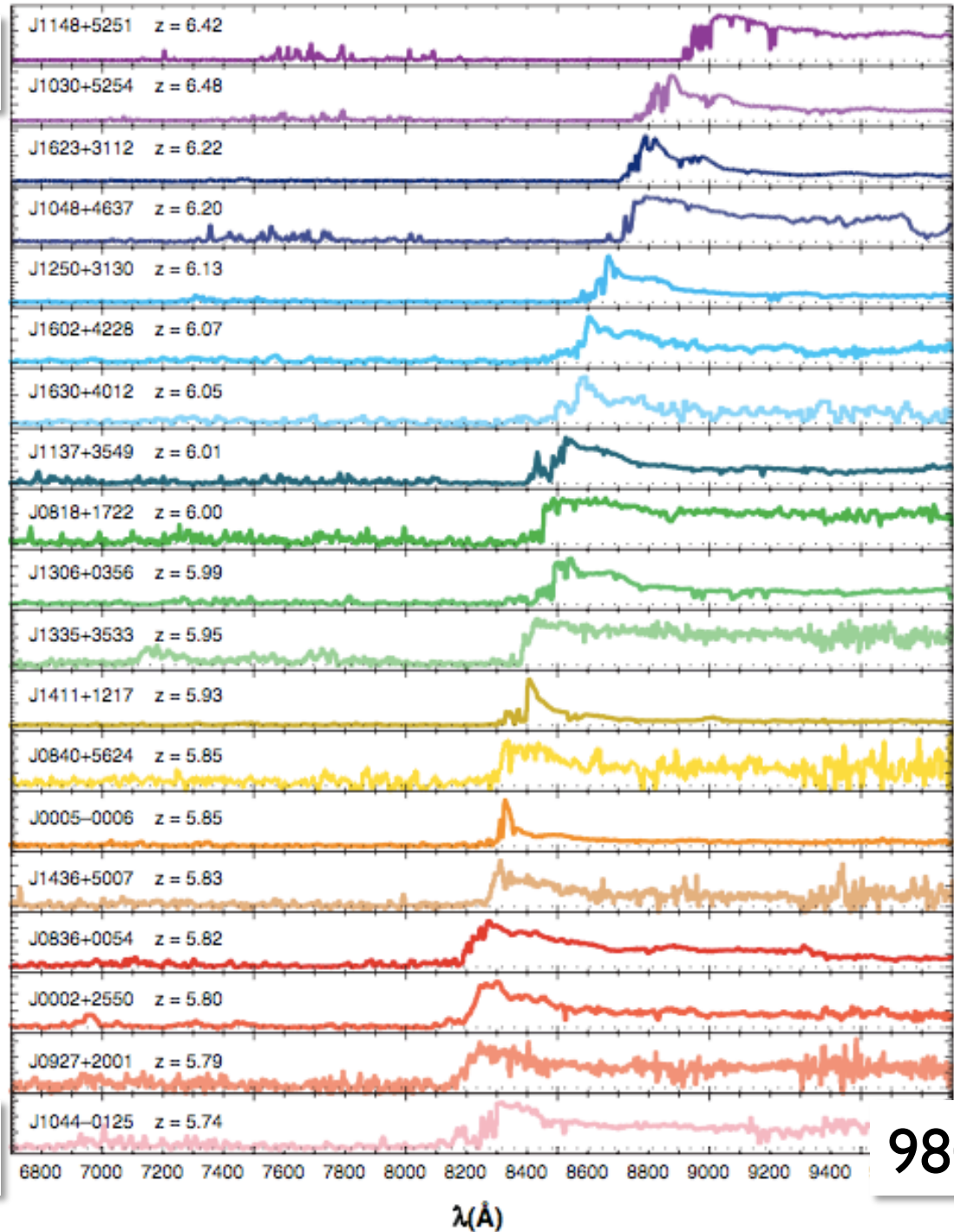


Wang+ 2013; see also Willott+ 2013

$z=6.4$

19 quasars at
 $5.7 < z < 6.4$ from
the SDSS survey

$f\lambda$



9800 Å

Fan, Carilli &
Keating 2006

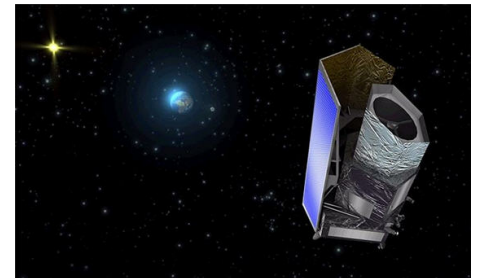
$z=5.7$

Massive black holes at high redshift

- Find massive black holes: look for quasars
 - Problem: quasars are very rare
 - Need multicolour surveys over large area
- SDSS very successful in discovering many luminous quasars up to $z=6.4$ (J1148 at $z=6.4$)
- To find quasars at higher redshifts, wide field NIR surveys are needed

Near infrared wide field surveys

- UK Infrared Deep Sky Survey (UKIDSS)
- ESO VISTA surveys (VIKING, VHS)
- Pan-STARRS: includes y-band (~ 1 micron)
- Upcoming: Dark Energy Survey, Euclid, LSST, ...



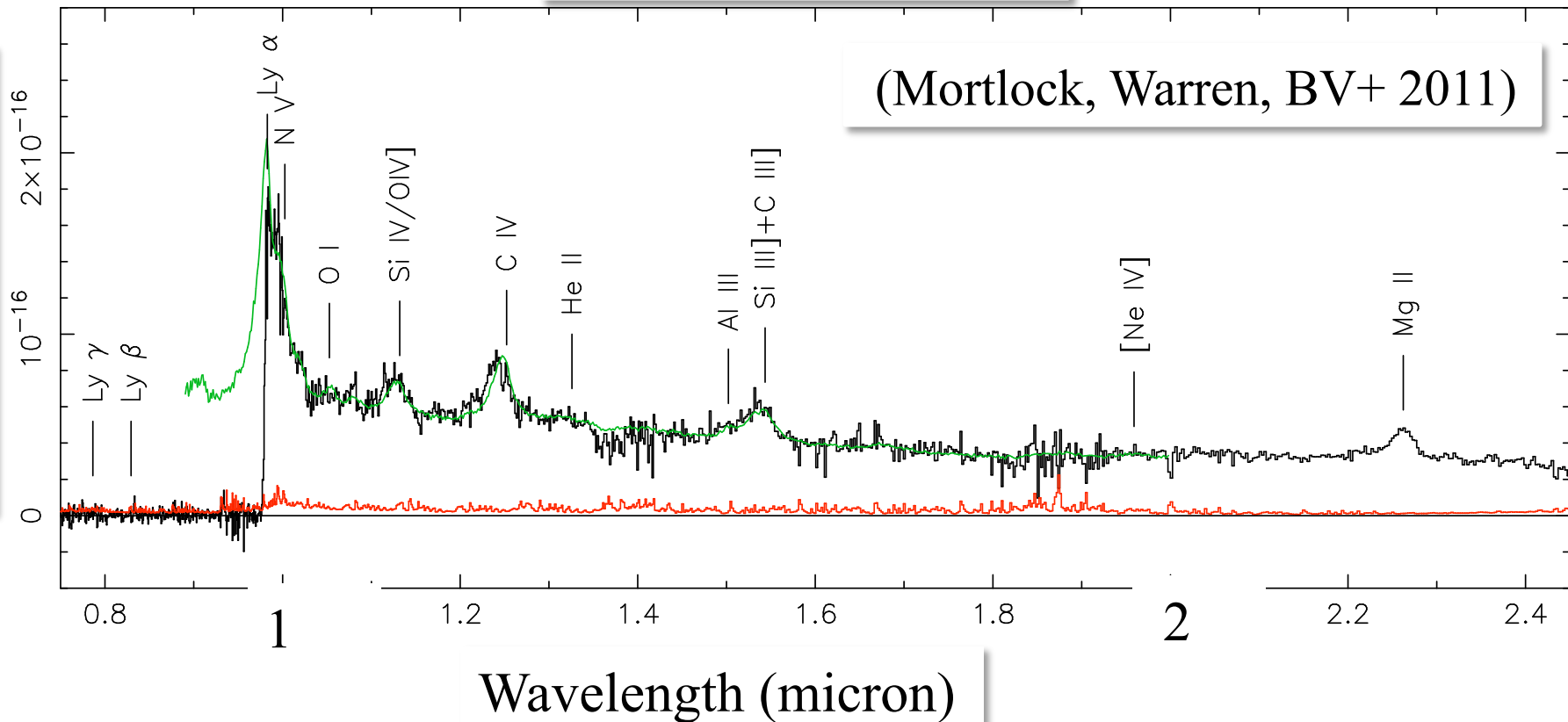
A luminous quasar in UKIDSS at $z=7.1$

Bright quasar: $M_{1450} = -26.6$, $M_{\text{BH}} \approx 3 \times 10^9 M_{\odot}$

ULAS J1120+0641

(Mortlock, Warren, BV+ 2011)

Flux density



Four $z \gtrsim 6.5$ quasars from VISTA/VIKING

BV+ 2013

J2318, $z=6.45$

J0305, $z=6.61$

J0109, $z=6.75$

J2348, $z=6.89$

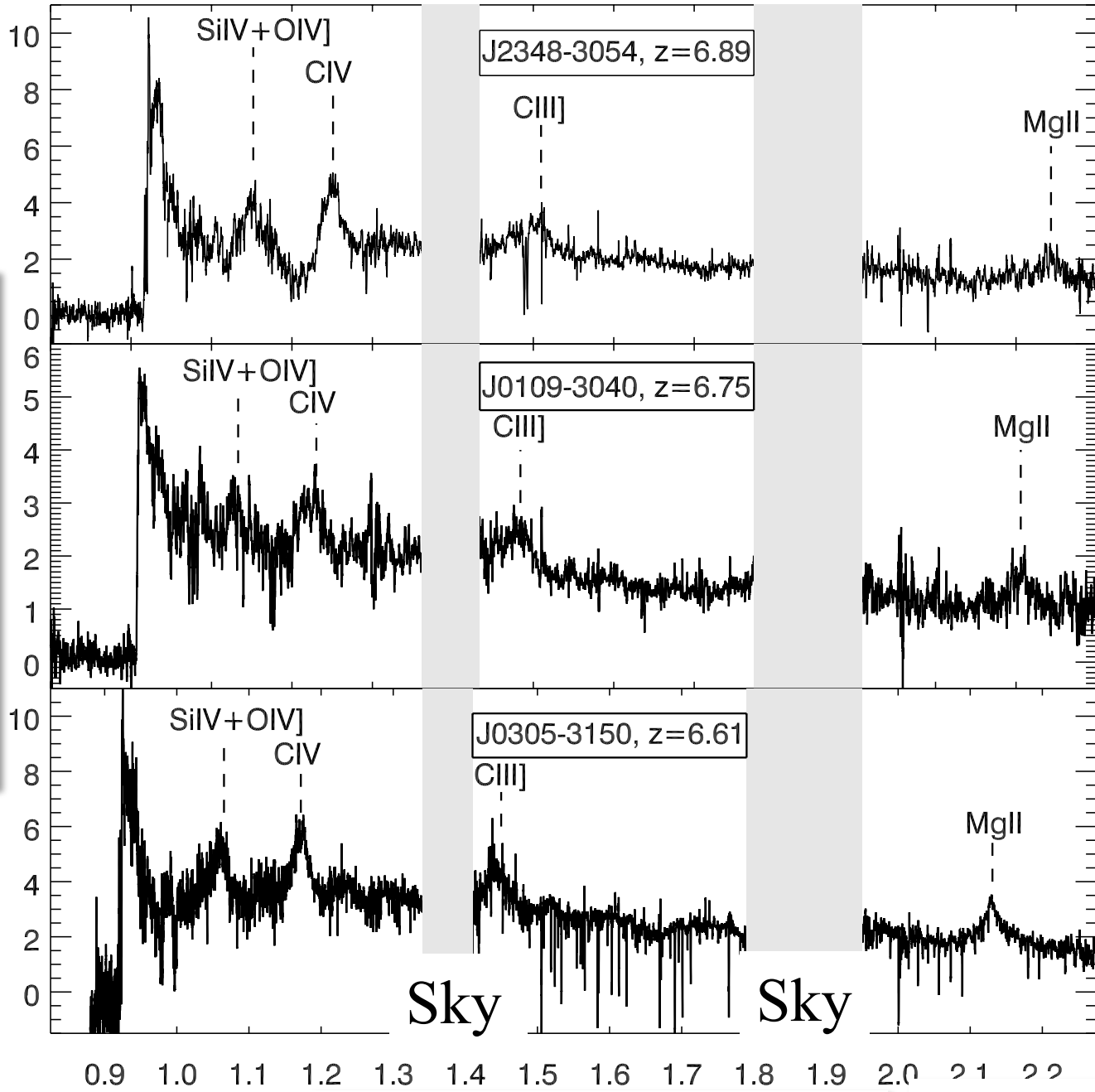
J1120, $z=7.08$

8500 Å

10300 Å

- 4 $z > 6.5$ quasars in VISTA/VIKING survey
 $Y_{AB} \sim 20.8 - 21.3 \rightarrow M_{UV}: -25.5 \text{ to } -26.1$

Flux density



Black hole
mass:
 $2.1 \times 10^9 M_{\odot}$

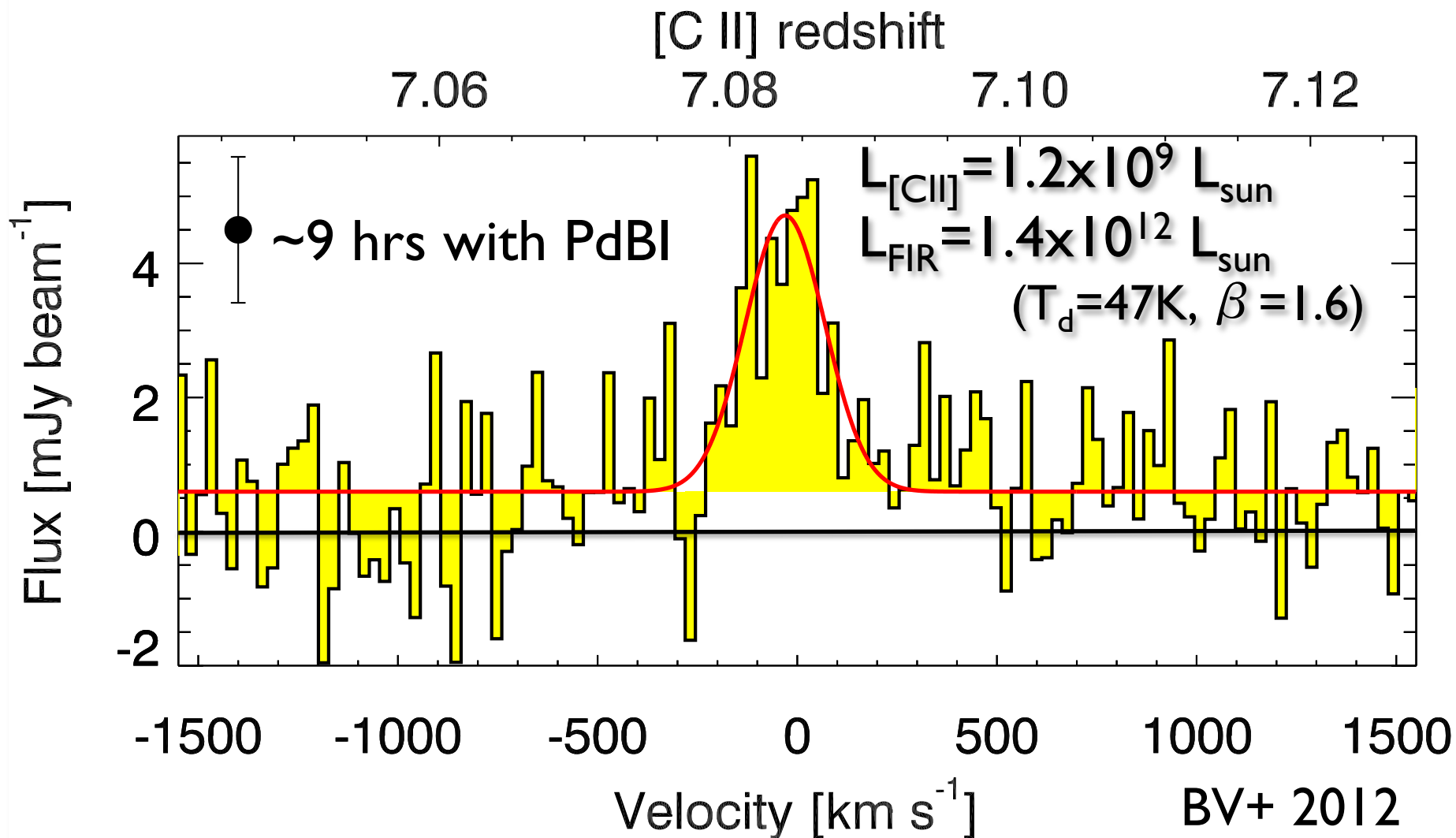
$1.5 \times 10^9 M_{\odot}$

$1.0 \times 10^9 M_{\odot}$

Observed wavelength

De Rosa, BV+ 2014

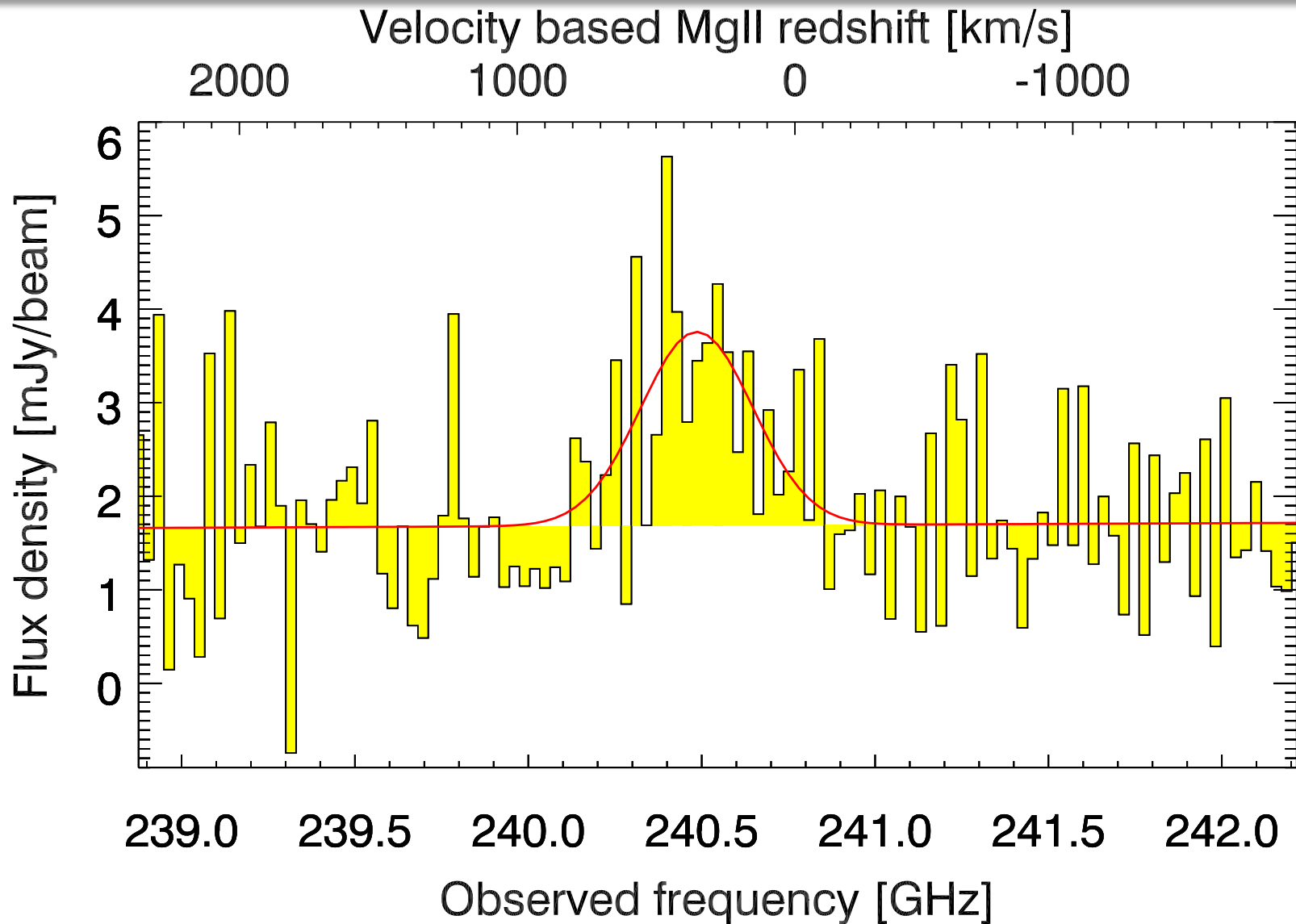
Spectrum of J1120+0641 at $z=7.1$



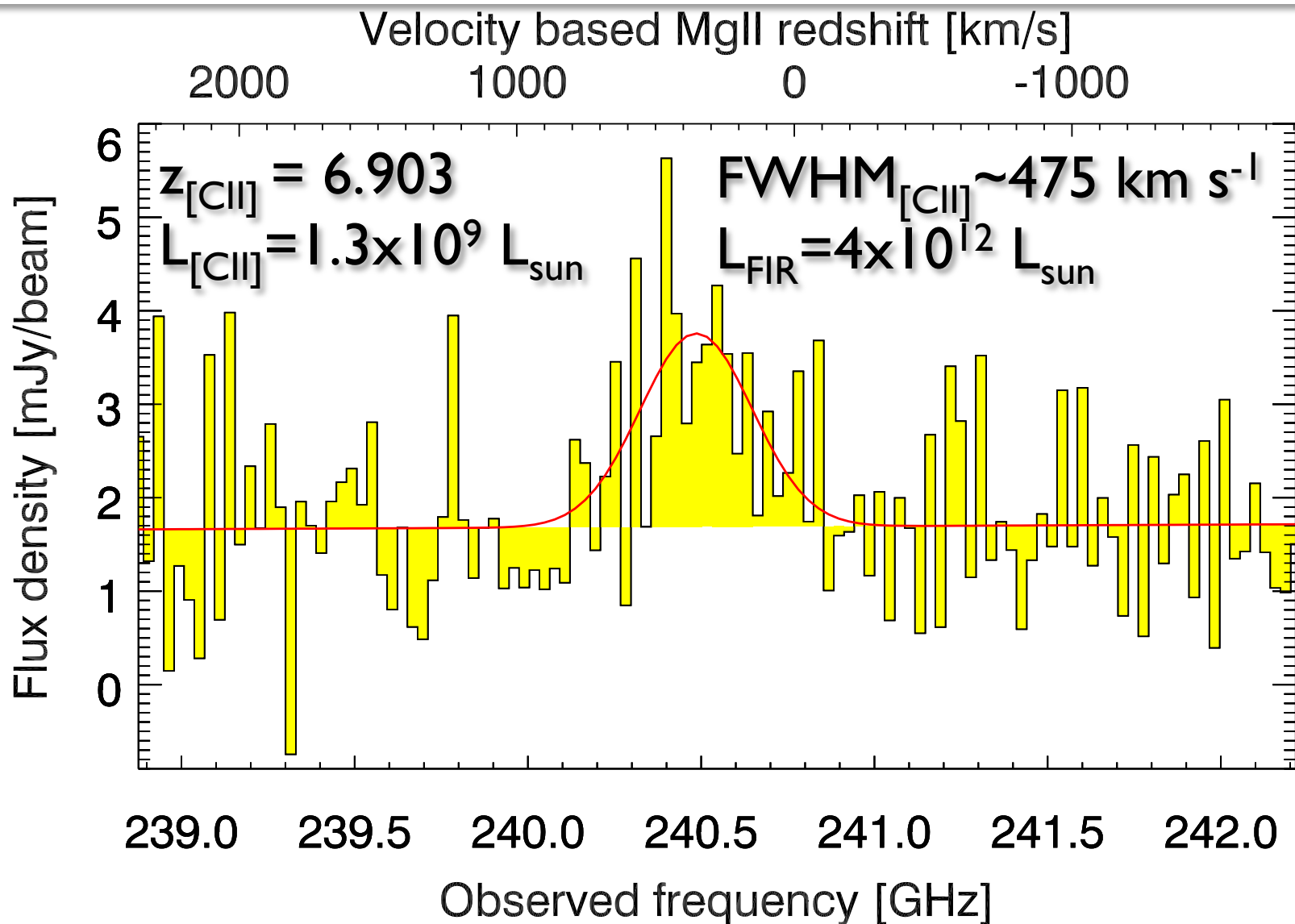
ALMA Cycle I snapshot observations

- Observations in Oct-Nov 2013
- 3 $z > 6.5$ quasar hosts from VIKING
- 21-30 min execution time (per source)
- 23-30 antennas used
- 10-16 min on-source (45-60% efficiency)
- reach same sensitivity as 9 hrs PdBI(!)
- Resolution $\sim 0.5''$ (requested: $1.6''$)

[CII] spectrum of J2348-3054 at $z \sim 6.9$



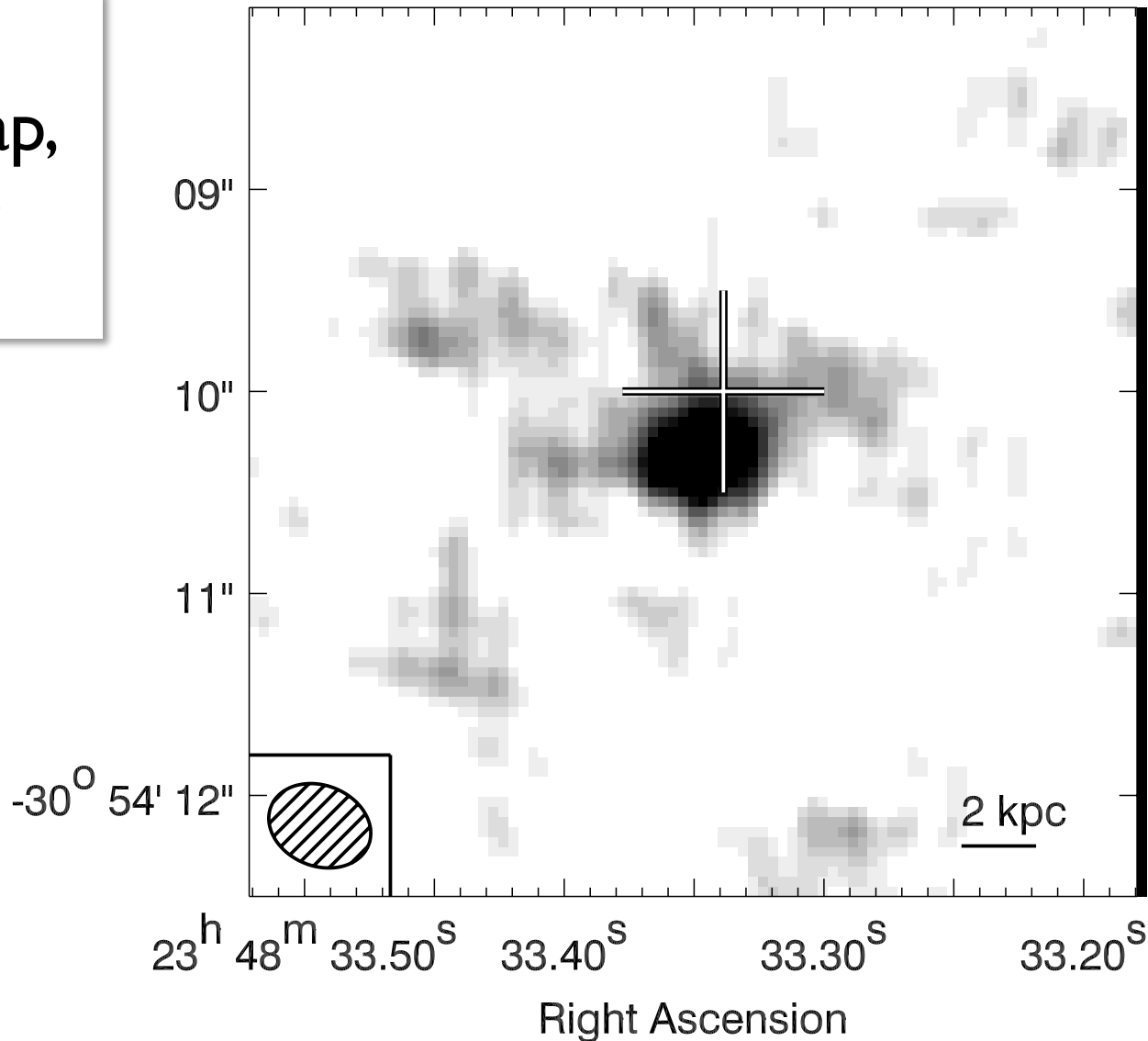
[CII] spectrum of J2348-3054 at $z \sim 6.9$



Map of [CII] emission of J2348-3054

Continuum-subtracted map, averaged over line FWHM

Declination

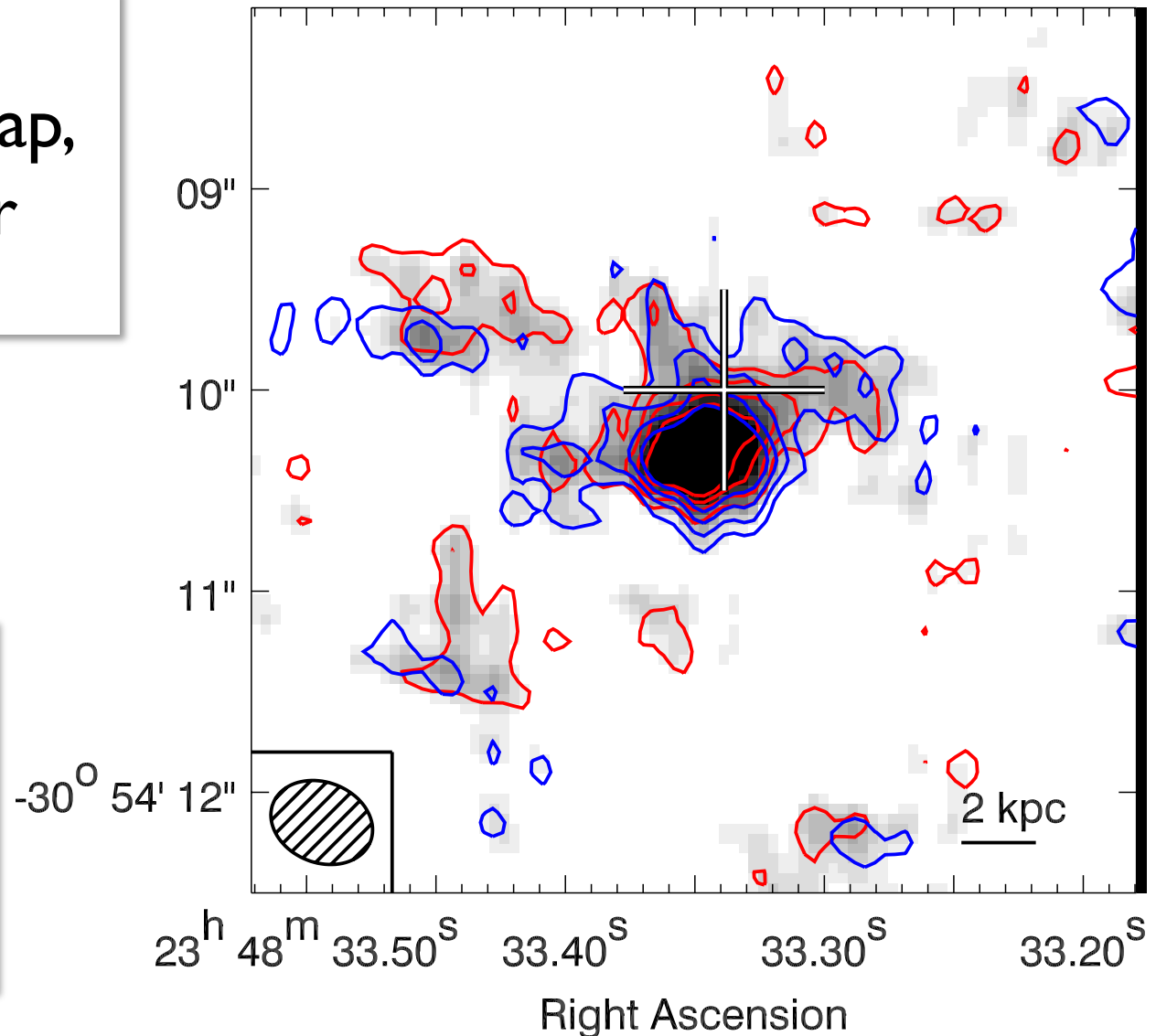


Map of [CII] emission of J2348-3054

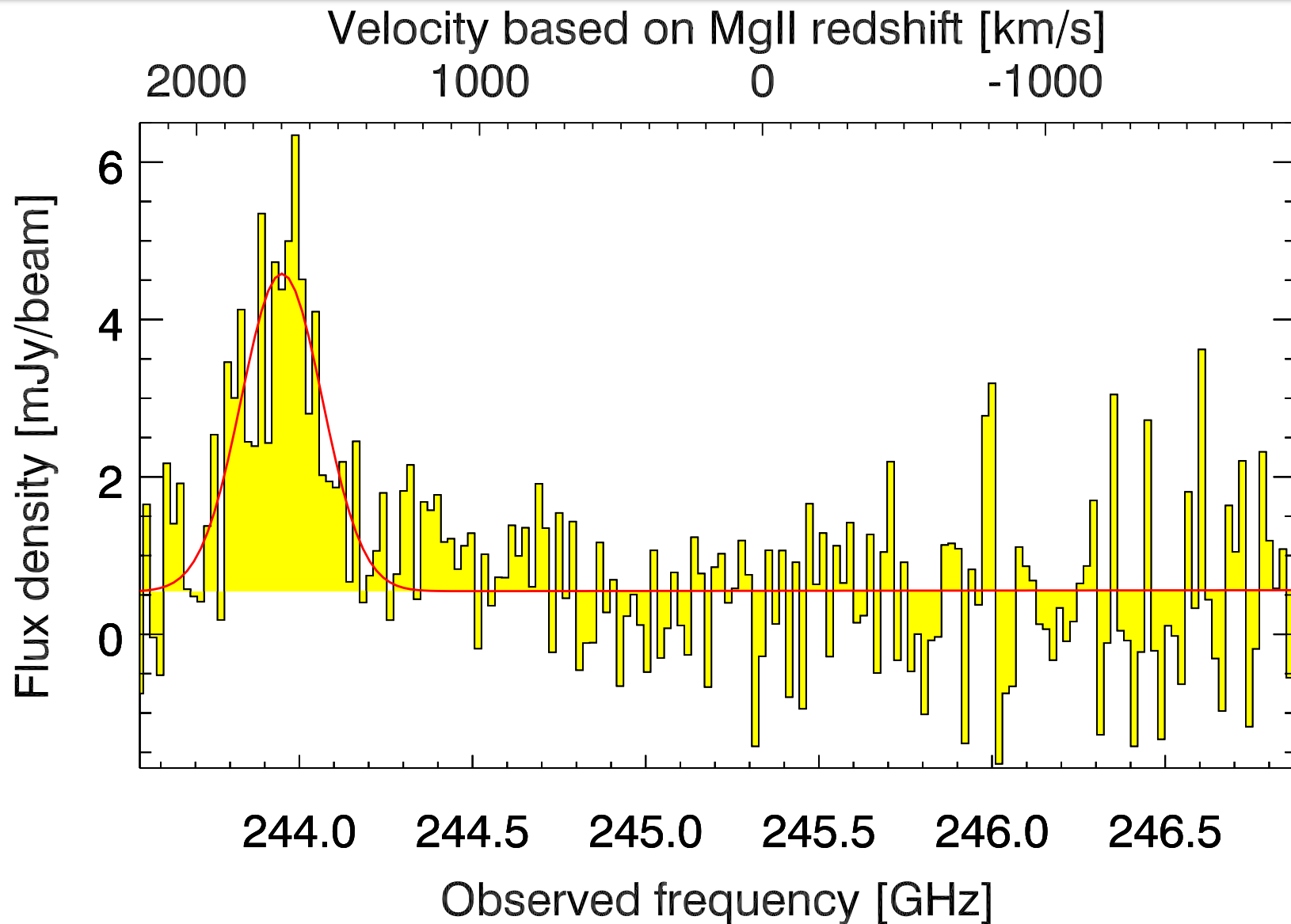
Continuum-subtracted map, averaged over line FWHM

Declination

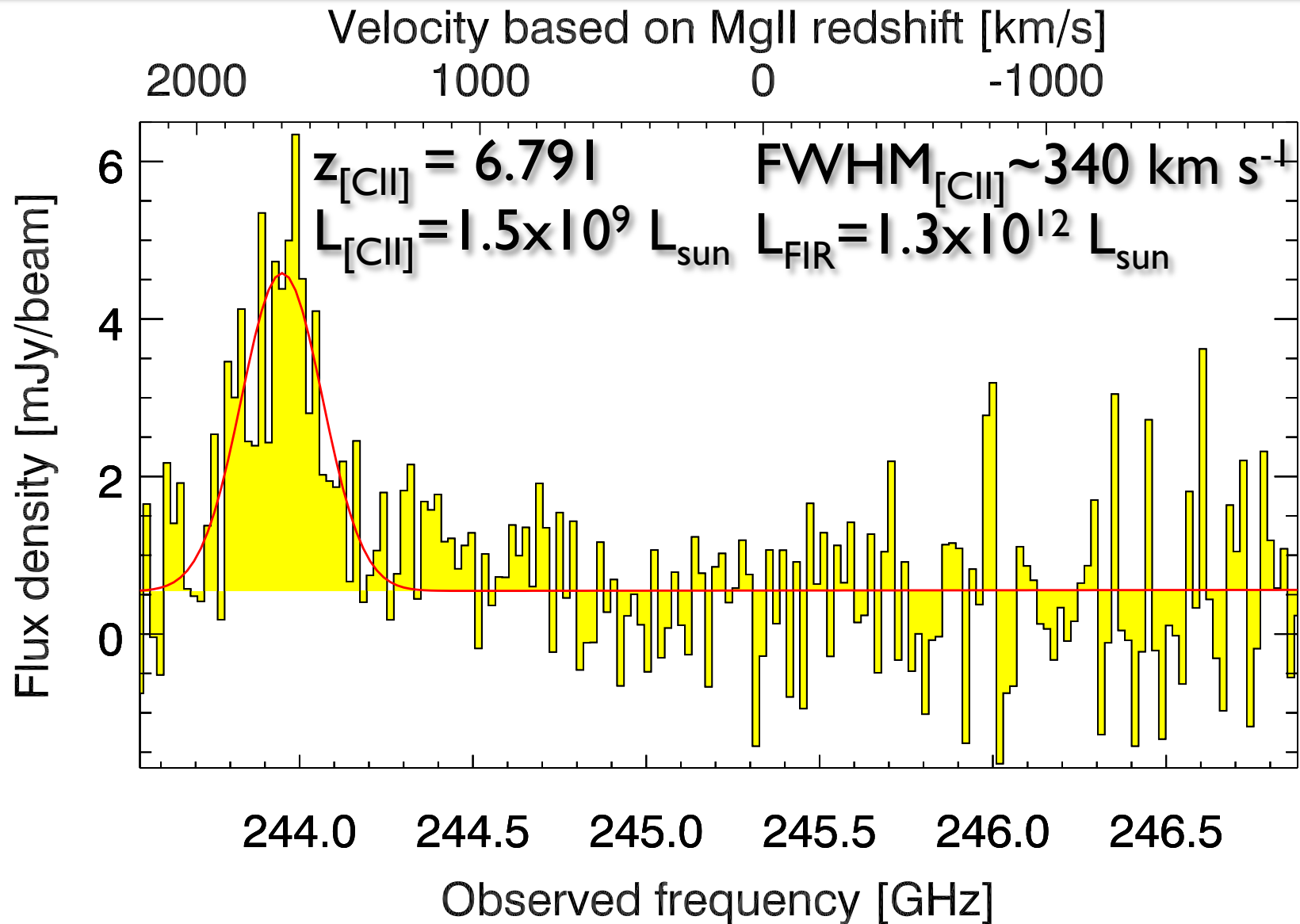
Blue: blue side of line peak
Red: red side of line peak



[CII] spectrum of J0109-3047 at $z \sim 6.75$



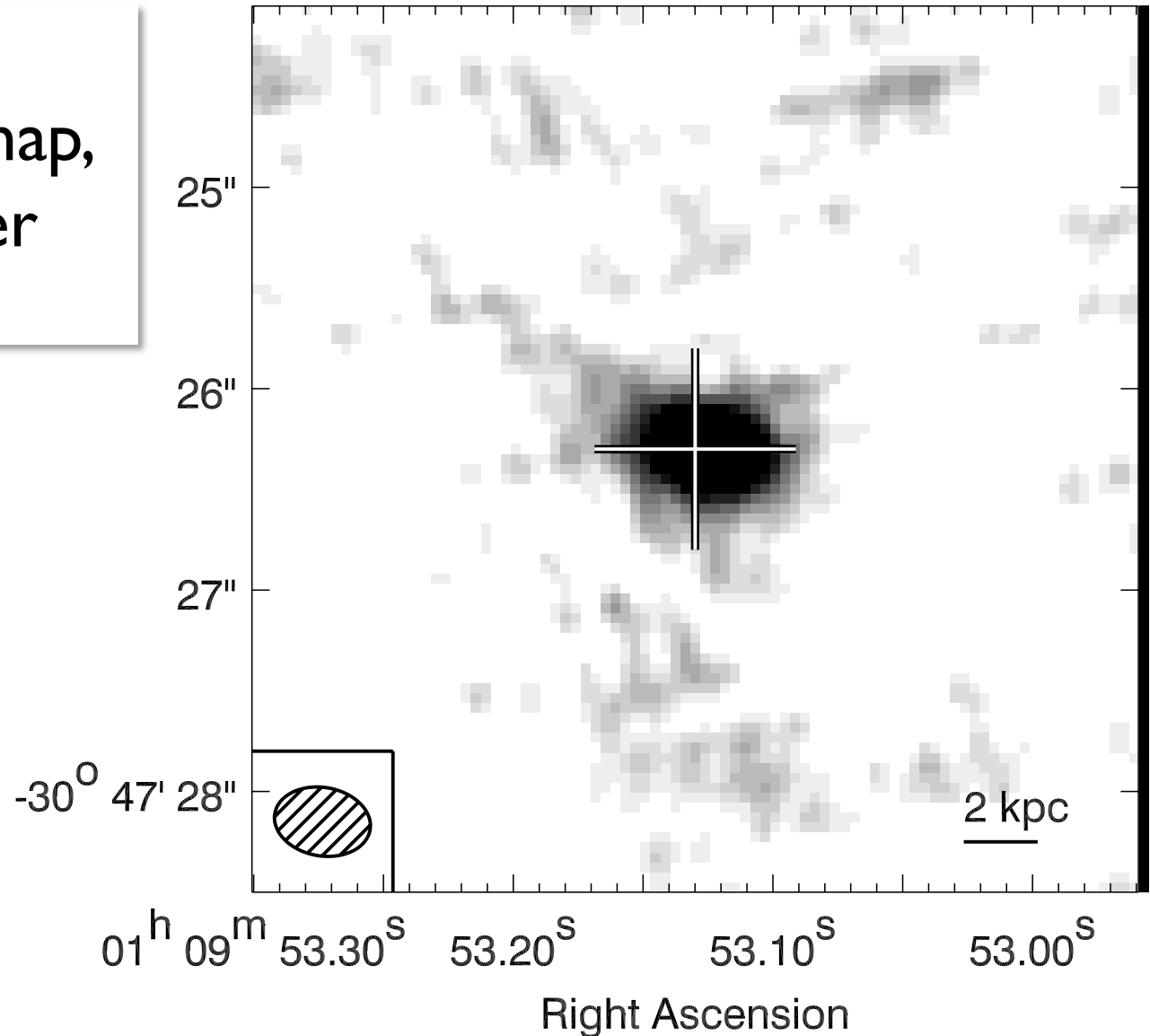
[CII] spectrum of J0109-3047 at $z \sim 6.75$



Map of [CII] emission of J0109-3047

Continuum-subtracted map, averaged over line FWHM

Declination

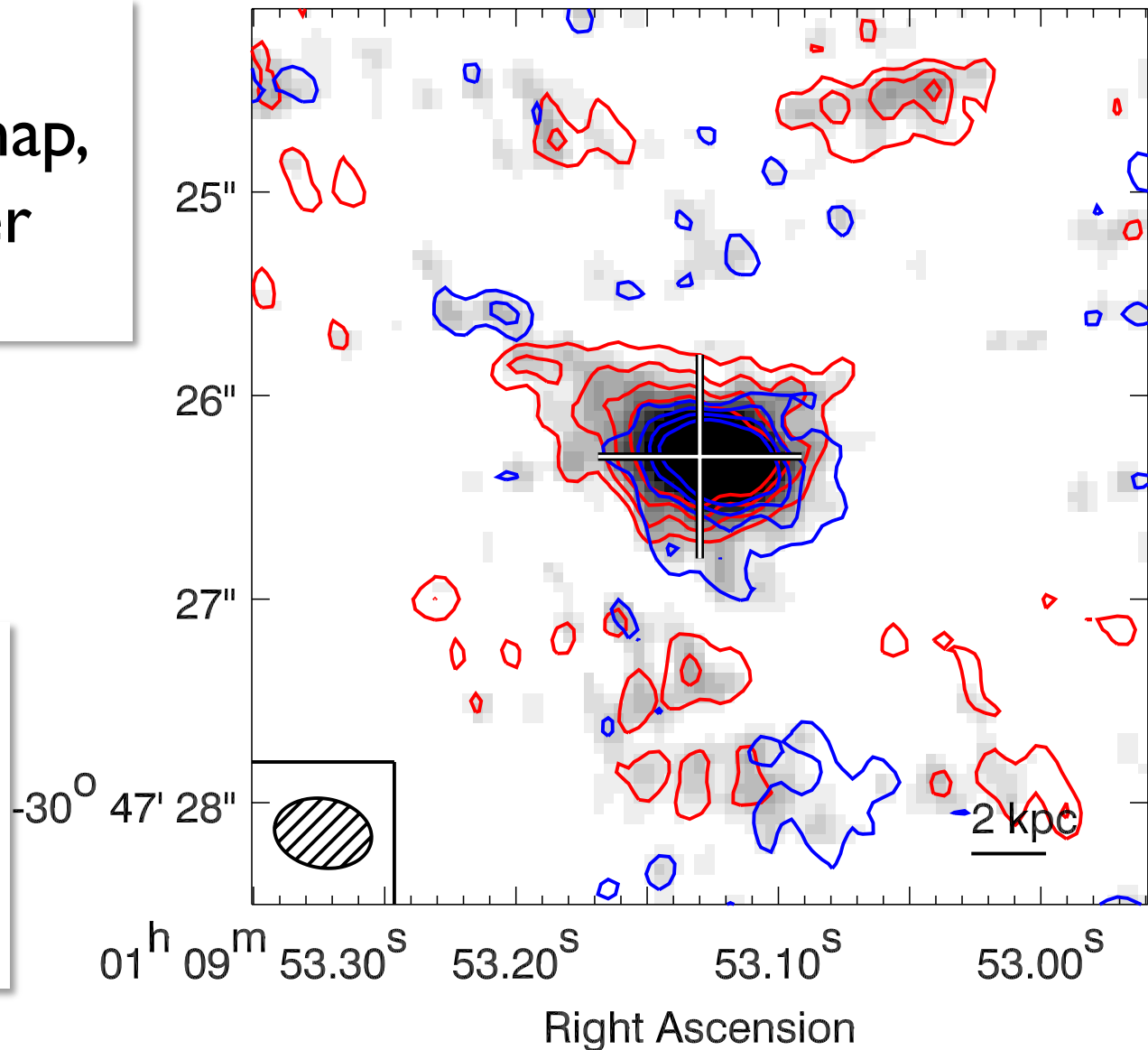


Map of [CII] emission of J0109-3047

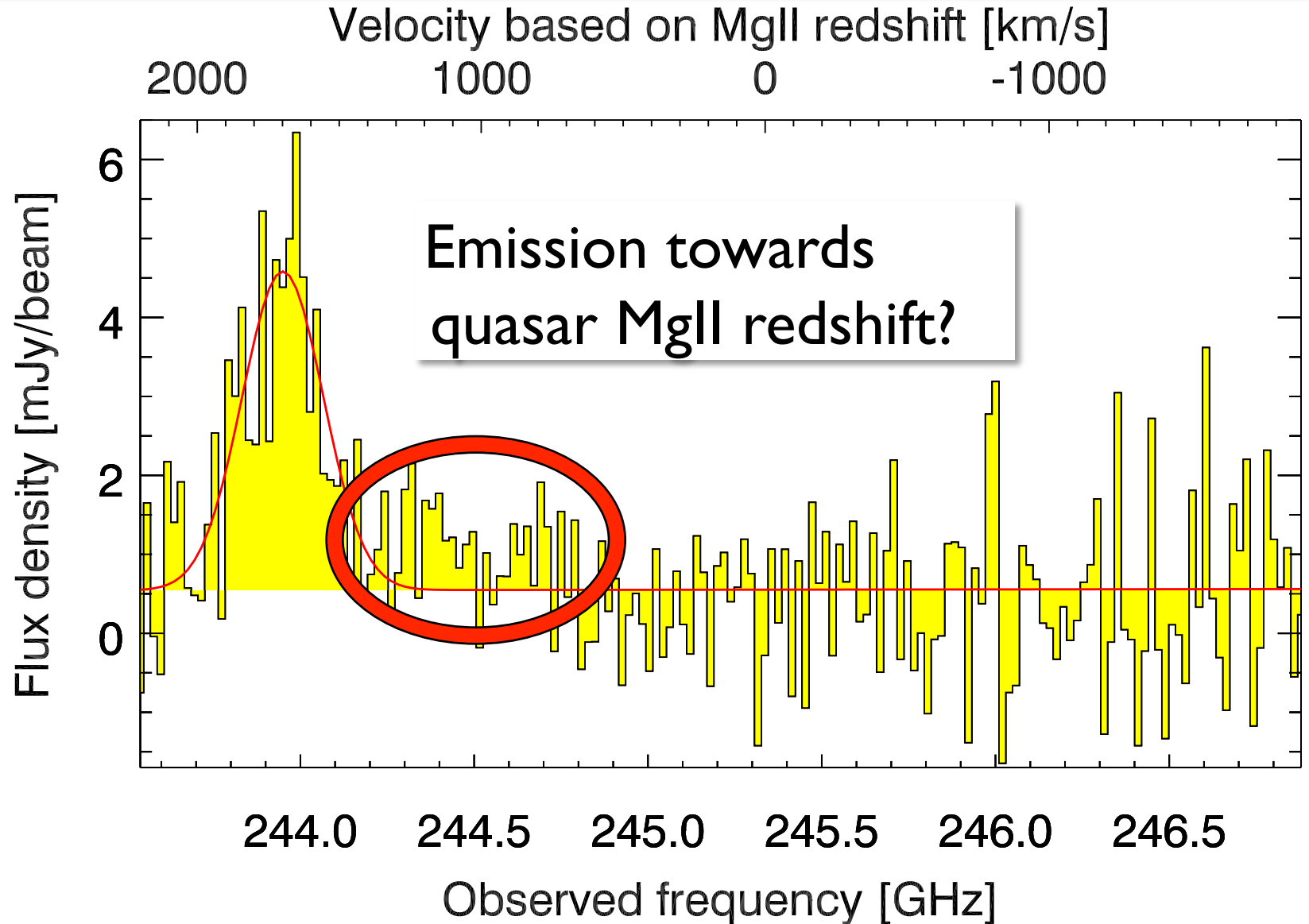
Continuum-subtracted map, averaged over line FWHM

Declination

Blue: blue side of line peak
Red: red side of line peak



[CII] spectrum of J0109-3047 at $z \sim 6.75$

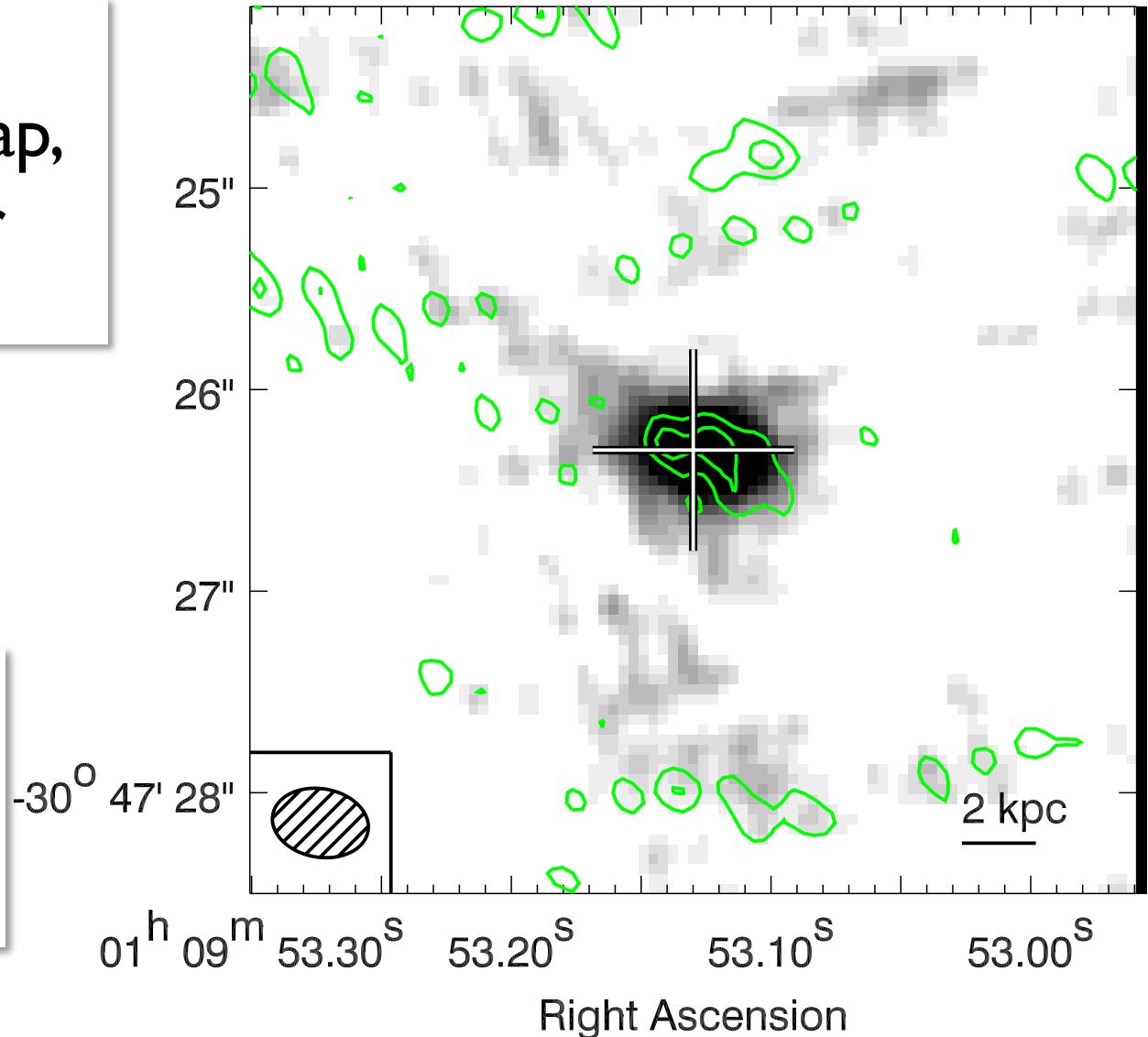


Map of [CII] emission of J0109-3047

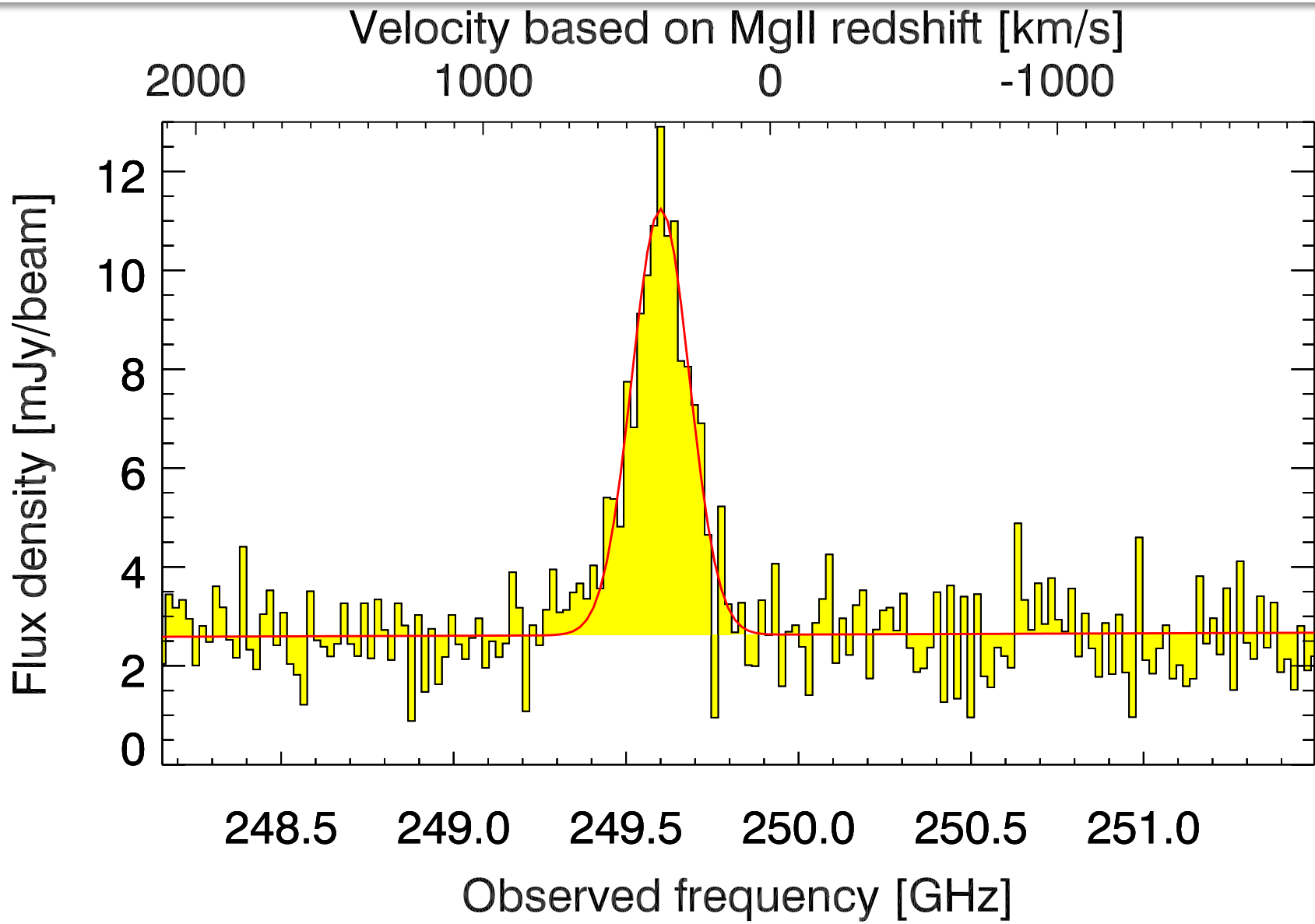
Continuum-subtracted map, averaged over line FWHM

Declination

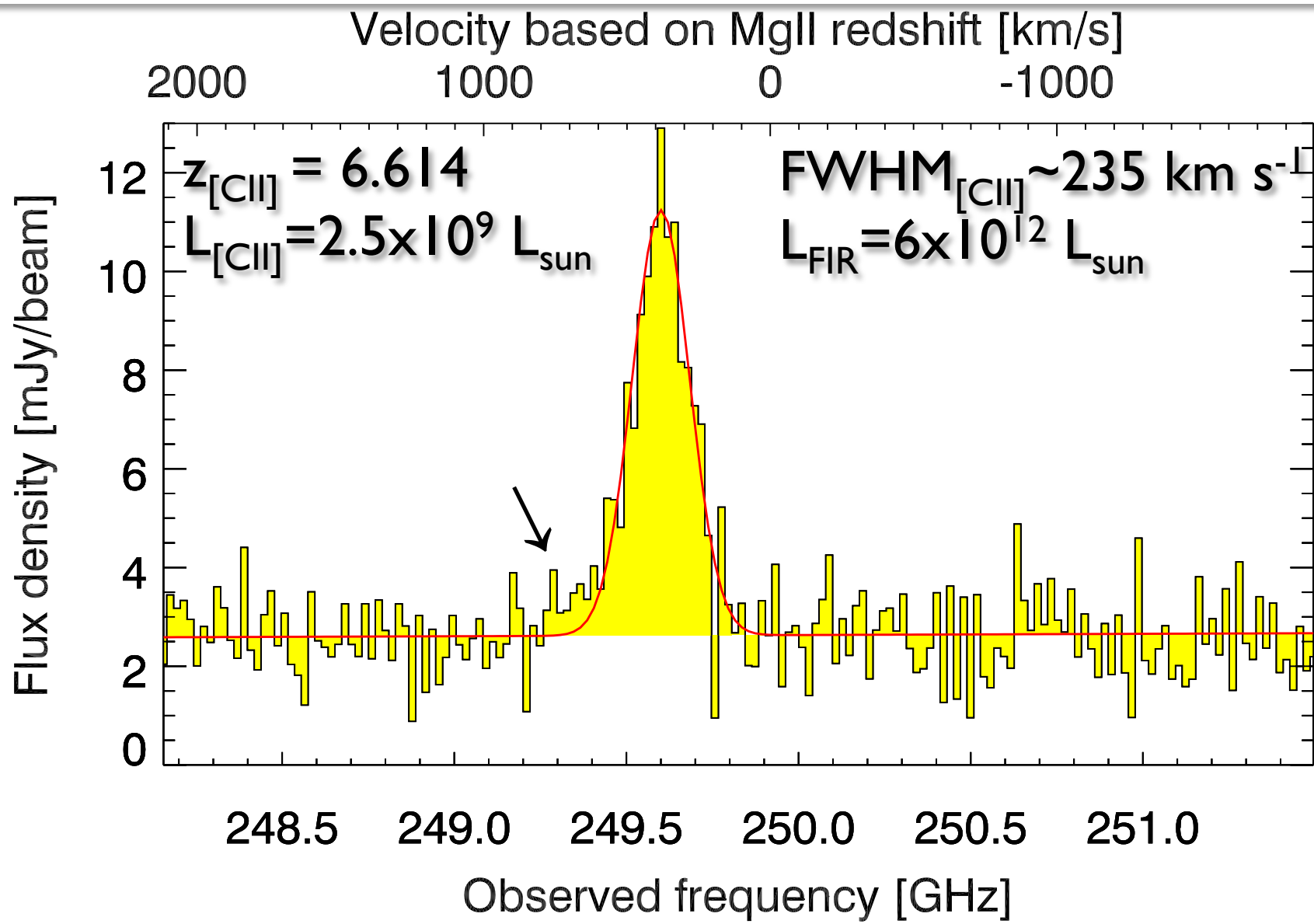
Green: emission towards quasar
MgII redshift



[CII] spectrum of J0305-3150 at $z \sim 6.6$



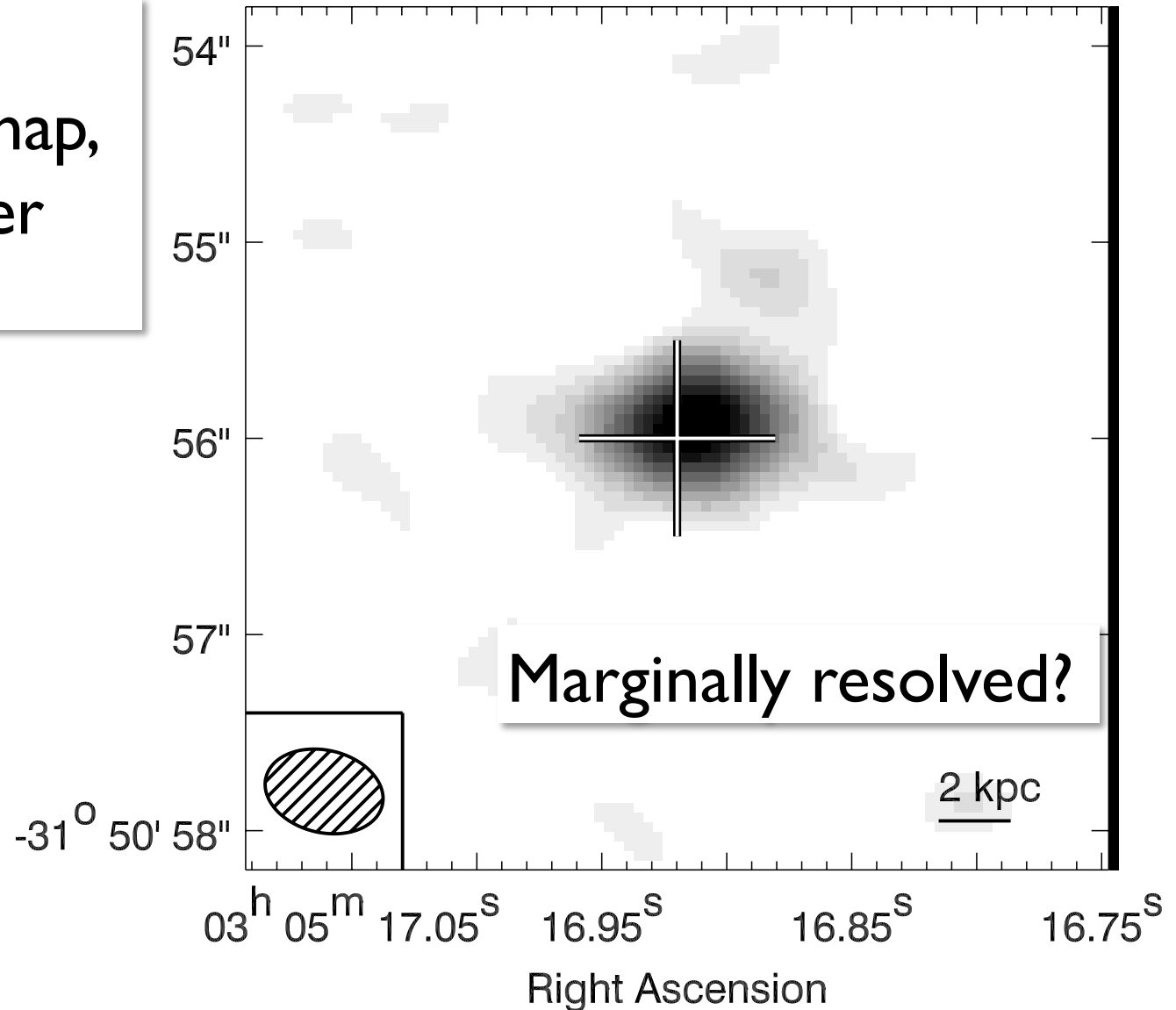
[CII] spectrum of J0305-3150 at $z \sim 6.6$



Map of [CII] emission of J0305-3150

Continuum-subtracted map, averaged over line FWHM

Declination

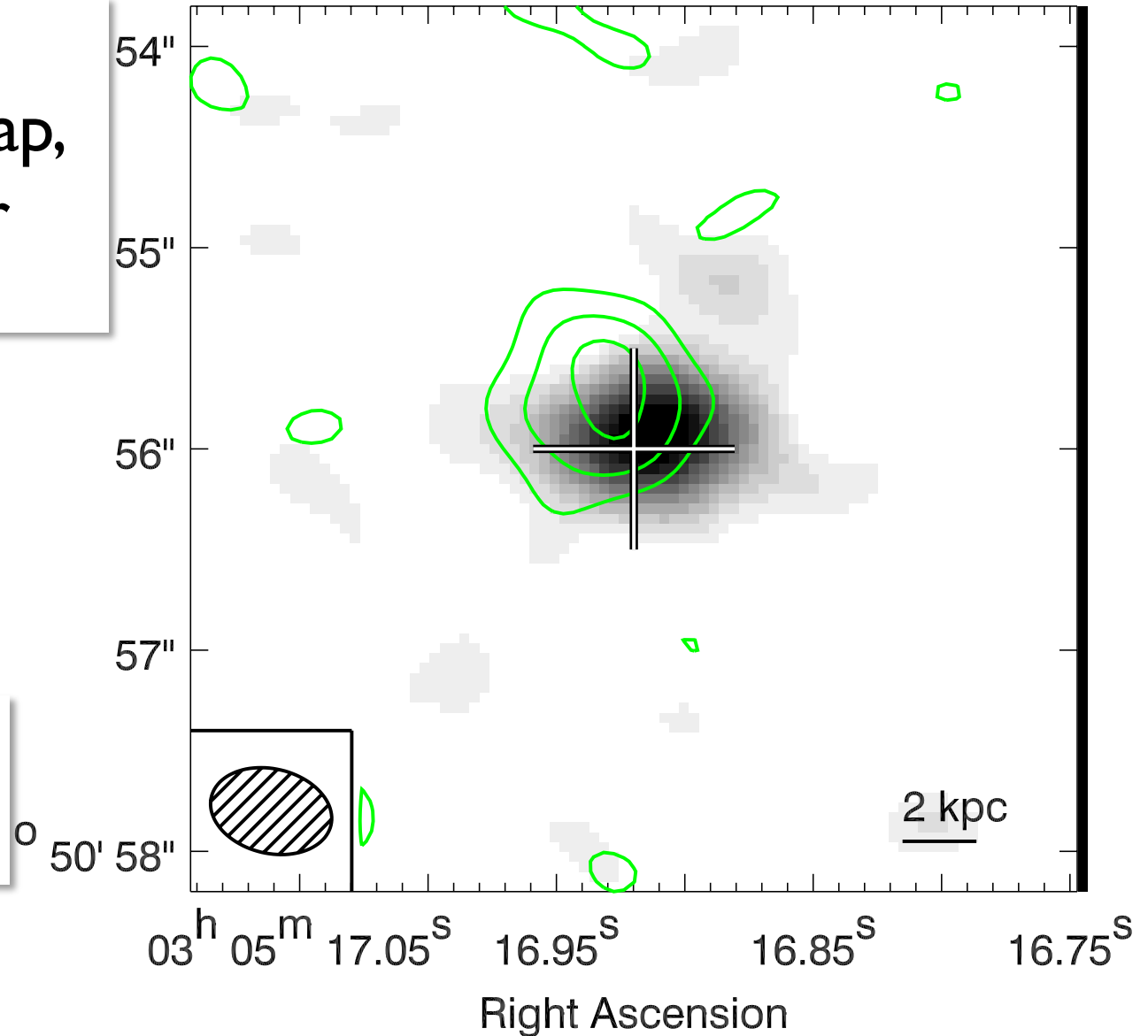


Map of [CII] emission of J0305-3150

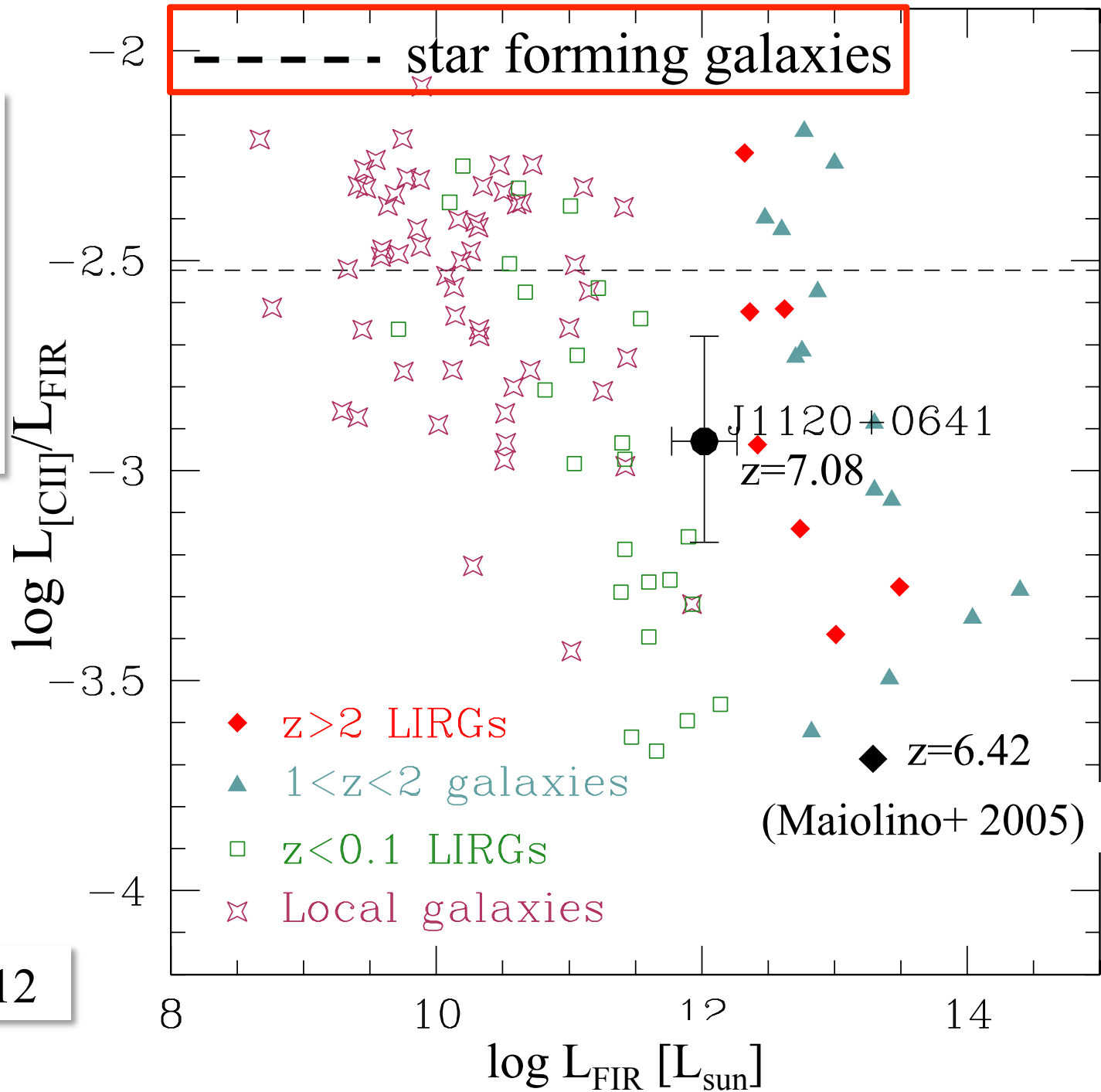
Continuum-subtracted map, averaged over line FWHM

Declination

Green: second component

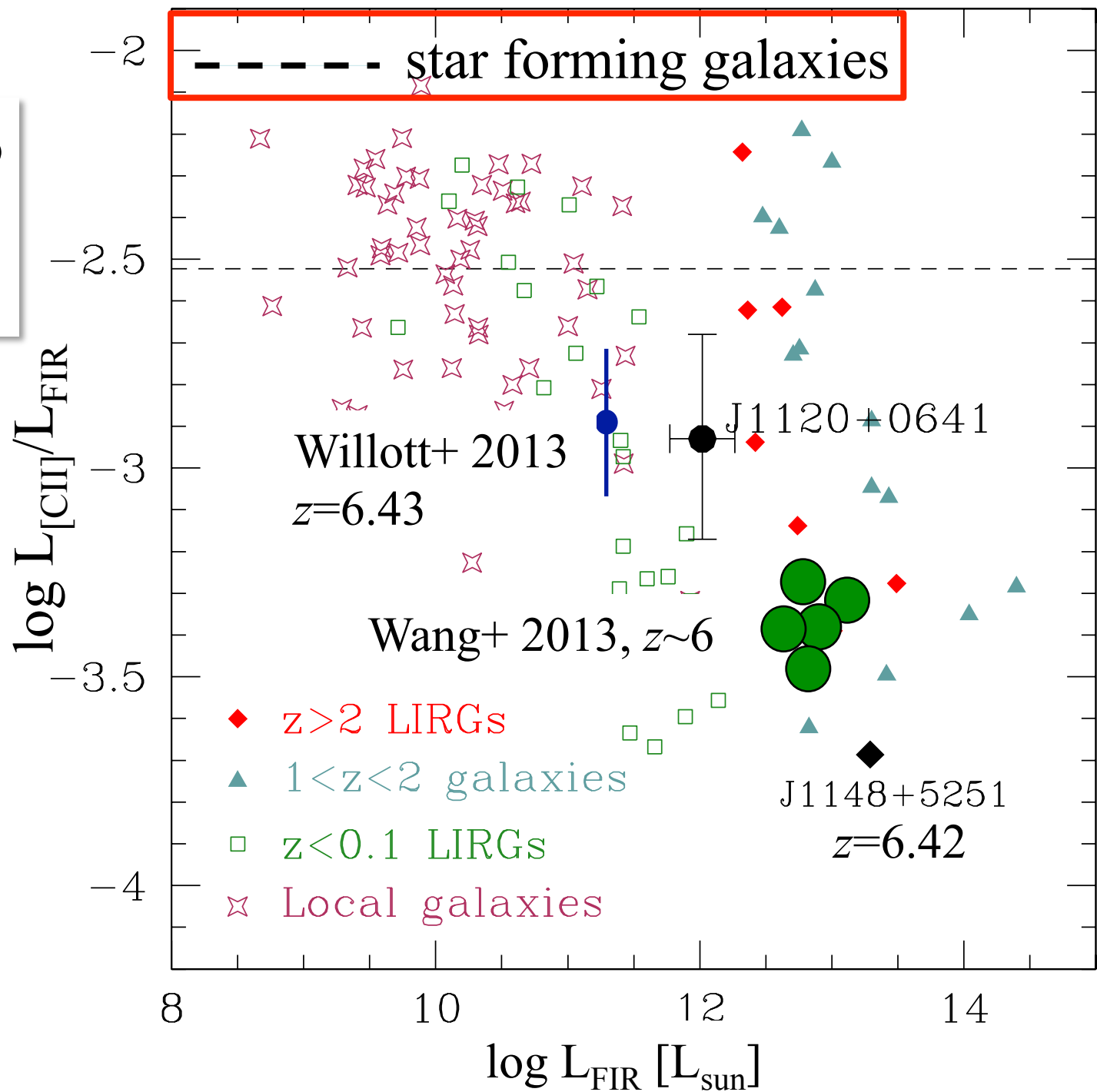


[CII] and FIR
luminosity of
 $z > 6.5$ quasars
compared to
literature

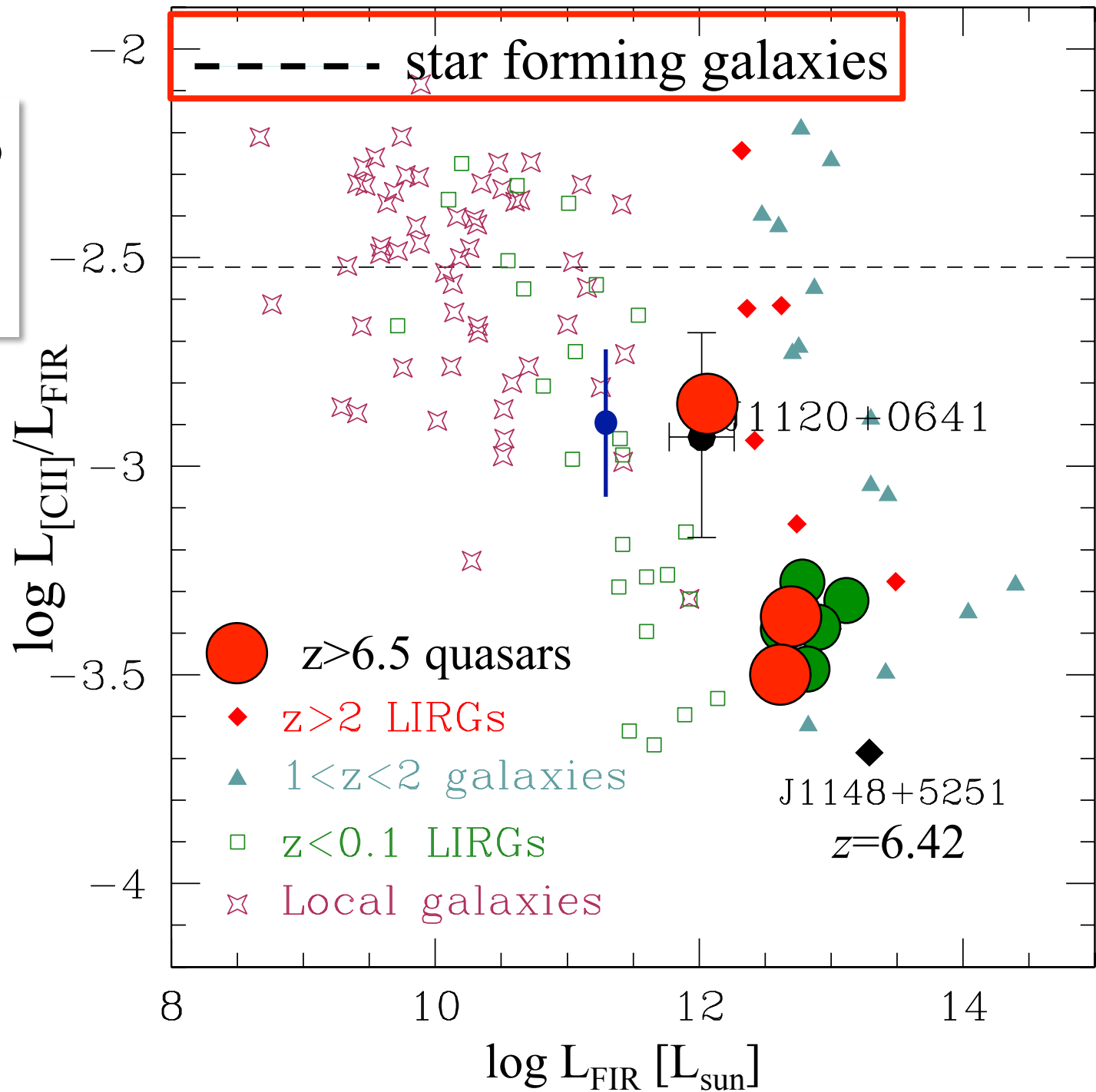


BV+ 2012

Comparison to
lower redshift
objects



Comparison to
lower redshift
objects



Summary

- Quasars ideal for early science observations
- ALMA observations of 3 $z > 6.5$ quasars:
 - [CII] luminosities of $(1.3-2.5) \times 10^9 L_{\text{sun}}$
 - FIR luminosities $\sim (1-6) \times 10^{12} L_{\text{sun}}$
 - Velocity shifts between MgII and [CII]
 - Additional components (spatially / in velocity)
 - Range in $L_{\text{[CII]}}/L_{\text{FIR}}$

Outlook

- More ALMA Cycle 1 observations:
 - high resolution [CII] in $z=7.1$ quasar host
- Proposed ALMA Cycle 2 observations:
 - [CII] in $z>6.5$ quasar hosts in higher resolution
 - CO(7-6) and CI transitions
- Study more $z>6.5$ quasars from VIKING, PS1...