

Serendipitous UV source catalogues for 10 years of XMM and 5 years of Swift

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On behalf of the UVOT and OM teams:

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- P. W. A. Roming, S. T. Holland, F. E. Marshall, M. Chester, S. Koch, P. J. Brown,
- A. Talavera, E. Ojero, and others



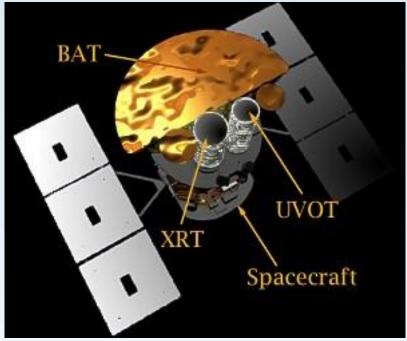
V.N. Yershov
Serendipitous UV source catalogues for 10 years of XMM and 5 years of Swift

Outline

- UV telescopes of XMM and Swift
- Input data
- Source catalogue processing chains
- Catalogue features
- Validation







Mission	XMM Newton (ESA)	Swift (NASA)
Start	10.12.1999	20.11.2004
End	~ 2019 (?)	~ 2021 (?)
UV telescope	ОМ	UVOT

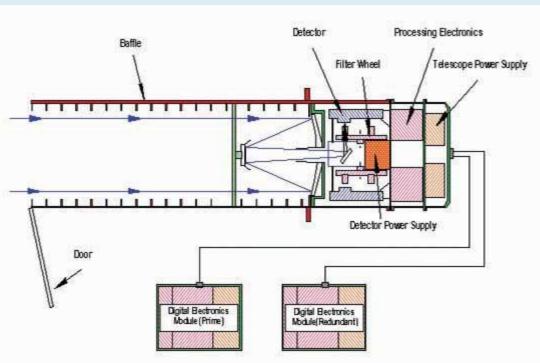
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Both observatories have almost identical UV telescopes







OM / UVOT is a relatively small 0.3-m Ritchey - Chretien UV telescope

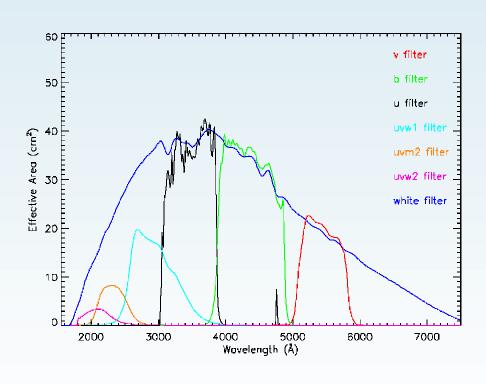
Focal length: 3.8 m Field of view: 17' x 17'

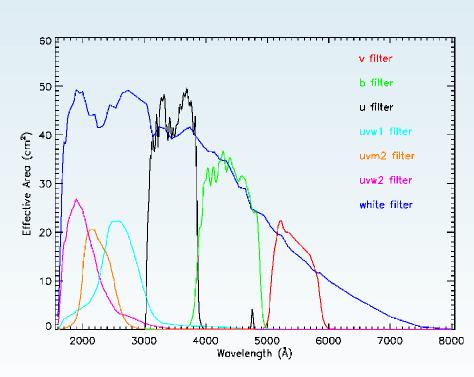
Detector: photocathode with a MCP photomultiplier

Filters: UVW2, UVM2, UVW1, U, B, V



Effective areas of the UV filters:



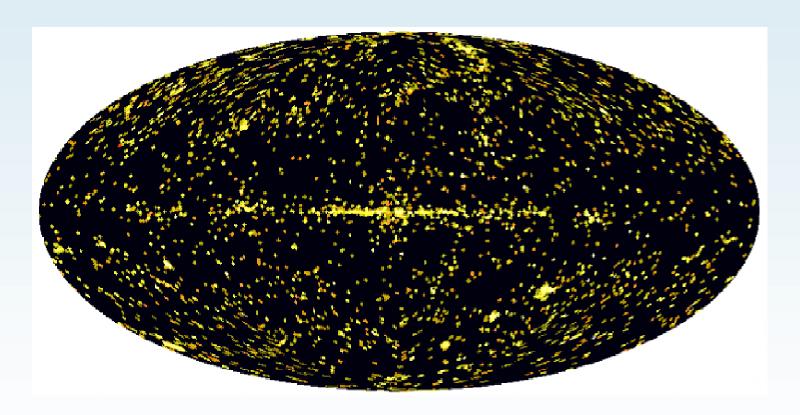


OM filters

UVOT filters

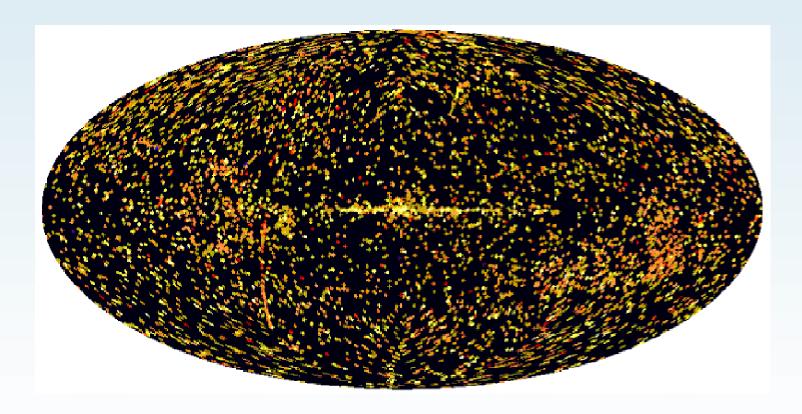


OM observations in the Galactic coordinates





UVOT observations in the Galactic coordinates



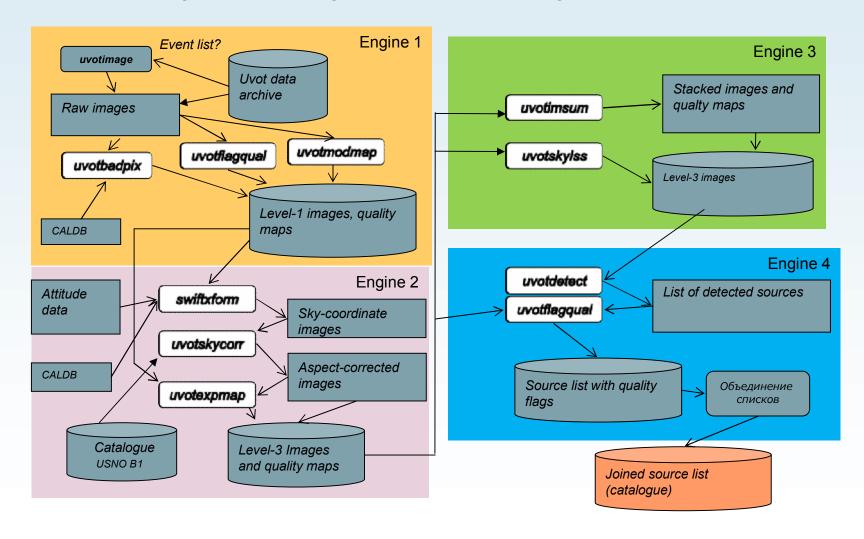


Data used for creating the OM and UVOT catalogues

	Catalogue	Catalogue
	SUSS-2	UVOTSSC
Period of observations	2000 - 2011	2005 - 2010
Total observations	6240	23428

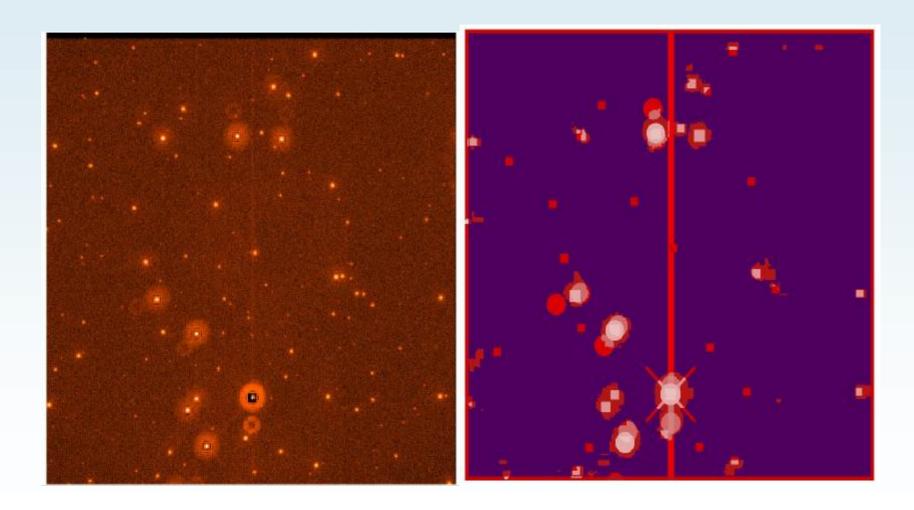


Catalogue processing chart (UVOT catalogue)

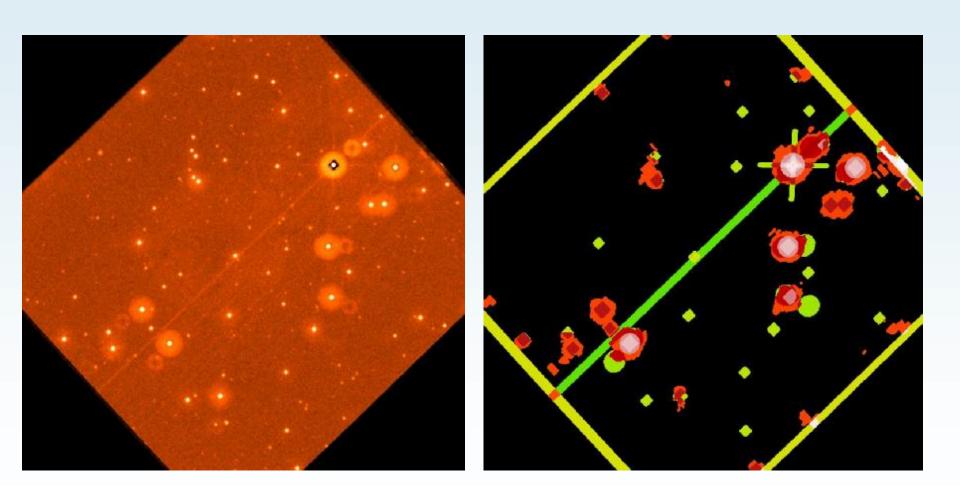




Example of the UVOT image with its corresponding quality map







Products of the Level-3 processing engine – stacked images and quality maps



Main features of the OM and UVOT catalogues

	Catalogue SUSS-2	Catalogue UVOTSSC
Period of observations	2000 - 2011	2005 - 2010
Total observations	6240	23428
Total sources	3 312 623	6 375 495
Repeated obs.	694 715	2 065 036
Total entries	4 918 082	15 015 935



FITS files

File Edit Tools Help

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■ 2 SUMMARY Binary 35 cols X 6246 rows Header Hist Plot All Select	= 2	SUMMA	RY	Binary	35 cols X 6246 rows	Header	Hist Plo	All	Select	

XMMOM J000000.4+623844

XMMOM J0000000.5+622153

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	■ IAUNAME	■ N_SUMMARY	UVW2_VEGA_MAG	UVM2_VEGA_MAG	UVW1_VEGA_MAG
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2	XMM0M J000000.1+622734	1	1.747867E+001	1.724070E+001	1.600135E+001
3	XMMOM J000000.1+623201	1	NULL	NULL	NULL
4	XMMOM J000000.1-250257	2	NULL	NULL	NULL
5	XMM0M J000000.2+622610	1	1.586626E+001	1.571957E+001	1.488039E+001
6	XMM0M J000000.2+624059	1	NULL	NULL	1.894545E+001
7	XMM0M J000000.2+623734	1	NULL	NULL	1.946458E+001
8	XMM0M J000000.2+623330	1	NULL	NULL	NULL
9	XMM0M J000000.2+622336	1	NULL	NULL	NULL
10	XMM0M J000000.2+622932	1	NULL	NULL	NULL
11	XMM0M J000000.3+443805	3	NULL	NULL	1.972667E+001
12	XMM0M J000000.3+443836	3	NULL	NULL	2.029916E+001
13	XMM0M J000000.4+622823	1	NULL	NULL	NULL
14	XMM0M J000000.4-321030	6246	NULL	NULL	2.107738E+001
15	XMM0M J000000.4+622248	1	NULL	NULL	NULL

NULL

1.416193E+001

NULL

1.414053E+001

1.889623E+001

1.327825E+001

Source table



FITS files

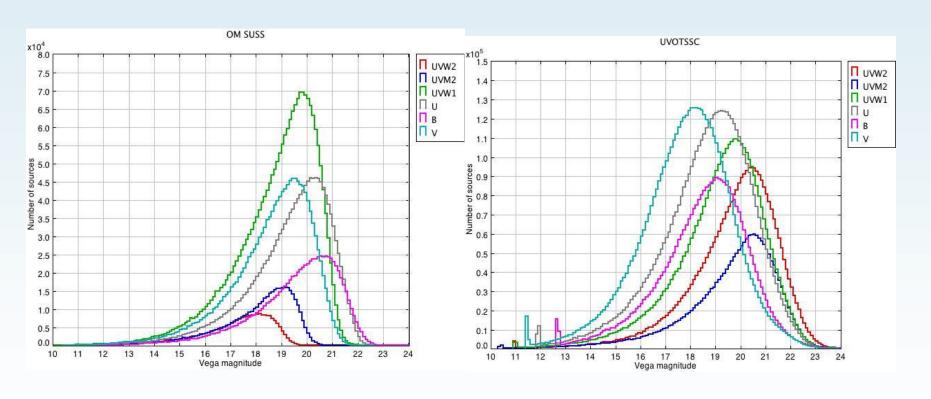
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	N_SUMMARY	FILTERS	RA_PNT	DEC_PNT	
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Invert					
1	1	UVW2, UVM2, UVW1, U, B	3.208360E-002	6.249802E+001	
2	2	U, B, V	1.221564E-001	-2.512193E+00	
3	3	UVW2, UVM2, UVW1, U	1.964872E-001	4.461314E+001	
4	4	UVM2, UVW1, U	4.726913E-001	-3.480373E+00	1
5	5	UAMT .	4.955880E-001	-1.545555E+00	1
6	6	U, V	6.948951E-001	-3.000822E+00	1
7	7	UVM2, UVW1, U	7.762773E-001	-2.004830E+00	1
8	8	UVM2, UVW1, U	7.768553E-001	-1.805427E+00	1
9	9	UVM2, UVW1, U	7.948853E-001	-6.100173E+00	1
10	10	UVW2, UVW1, U	8.112291E-001	-3.592956E+00	1
11	11	UVM2, UVW1	9.576669E-001	2.055688E+000	1
12	12	U	1.021296E+000	7.032944E+001	1
13	13	UVM2	1.176344E+000	1.290394E-001	1
14	14	UVM2, UVW1, B	1.351010E+000	-8.758560E+00	1
15	15	UVW2, UVW1	1.509402E+000	-3.473212E+00	1
4.0		777750		0 4800858.00	1

Summary table



Number of sources as a function of source magnitude

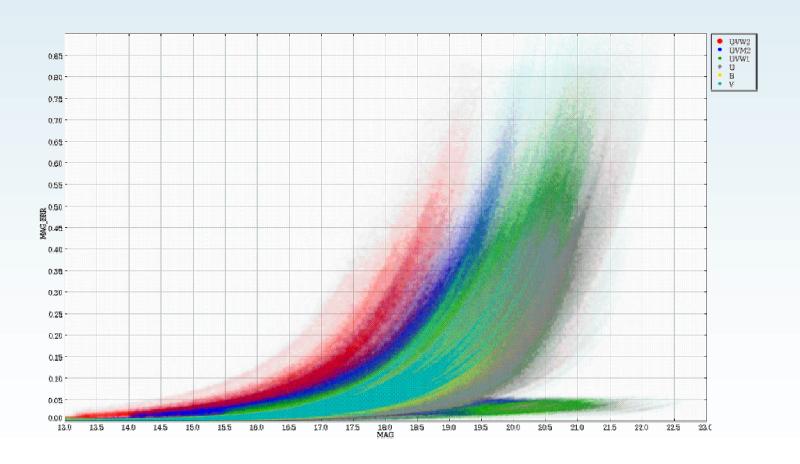


UVOT

OM

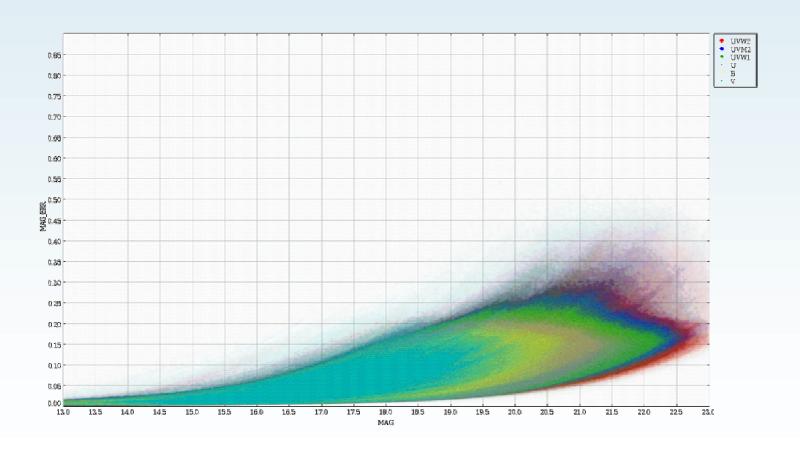


Magnitude errors as a function of source magnitude (OM catalogue)



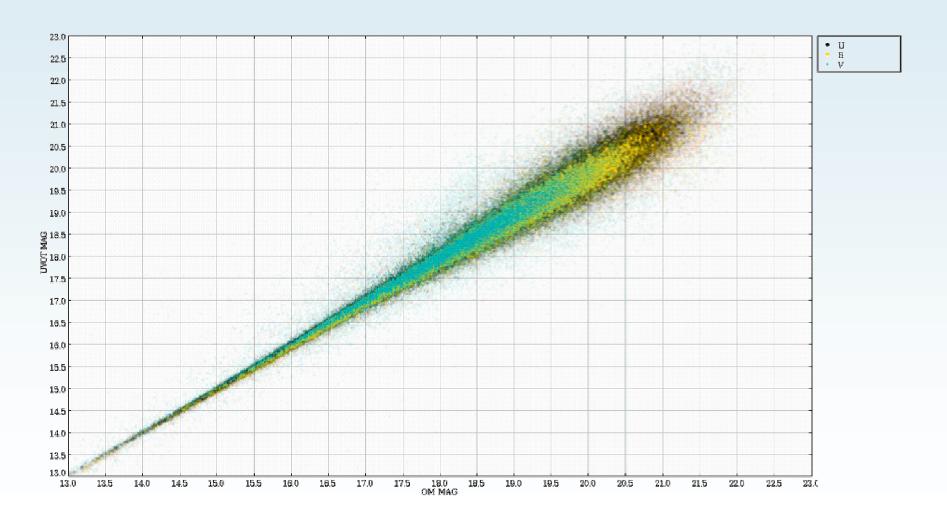


Magnitude errors as a function of source magnitude (UVOT catalogue)



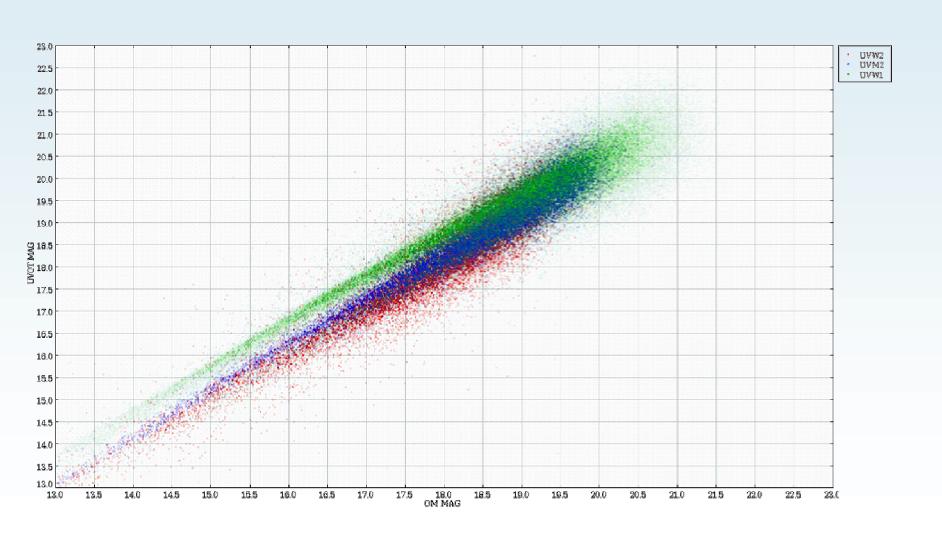


Comparison of the OM and UVOT catalogue for three optical filters

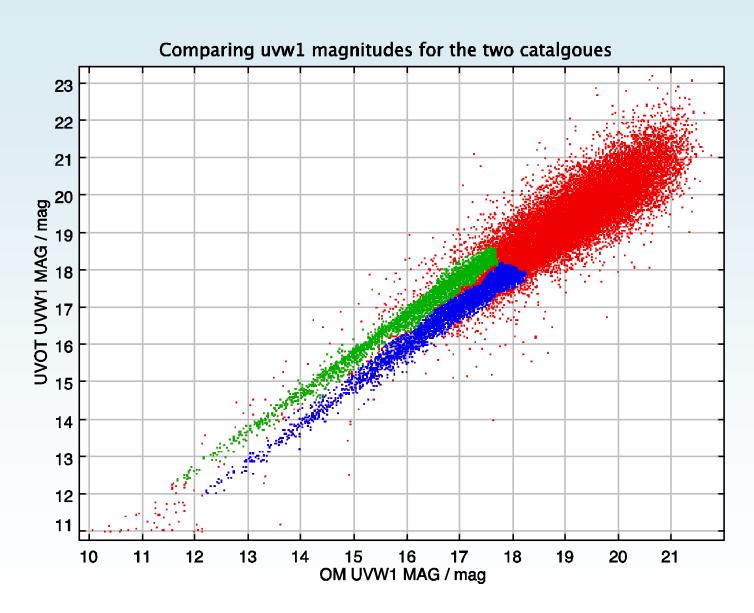




Comparison of the OM and UVOT catalogue for three UV filters

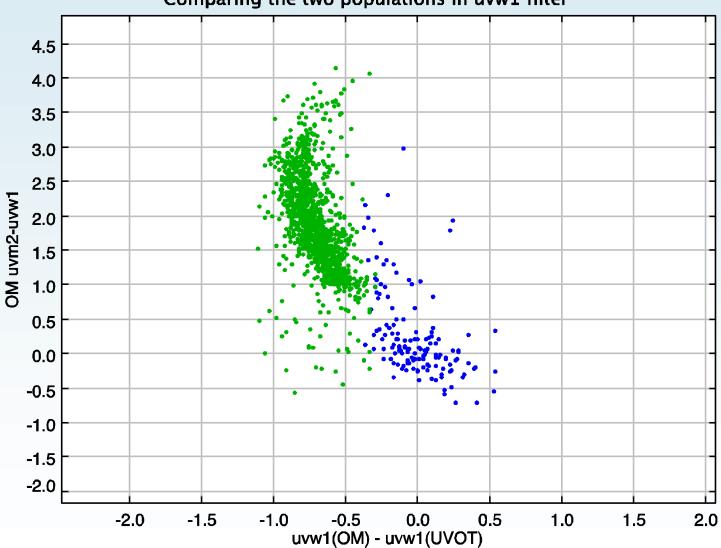






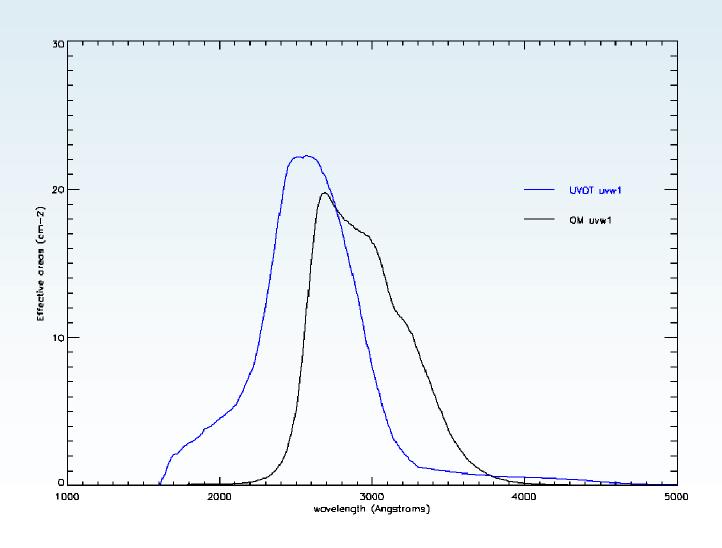






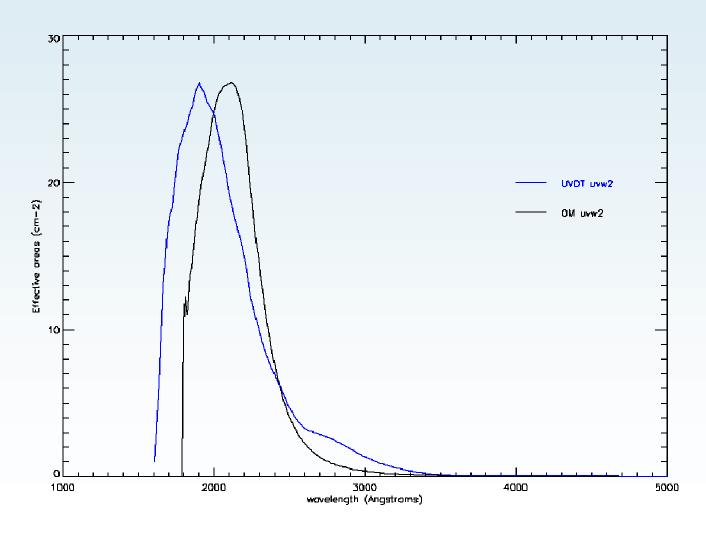


Comparison of the filter UVW1 for UVOT and OM





Comparison of the filter UVW2 for UVOT and OM





Conclusion notes

- ❖ The UV source catalogues for 10+ years of XMM Newton and 5+ years of Swift observations could be powerful tools, in particular, for:
- studying extreme physical processes in compact objects;
- selecting UV sources for additional follow-up;
- finding and studying variable UV sources;
- characterising dust clouds in different regions and regimes of star-formation;
- detecting dust in stellar winds;
- determining the fraction of young generation stars in galaxies with high SFR;
- constraining the IMF
- * etc. etc.





