

Theorem :

$$\text{dSph} = \text{T DG} \implies \text{LCDM}$$

Proof . . .

Pavel Kroupa

Two Zwicky Conjectures of fundamental importance :

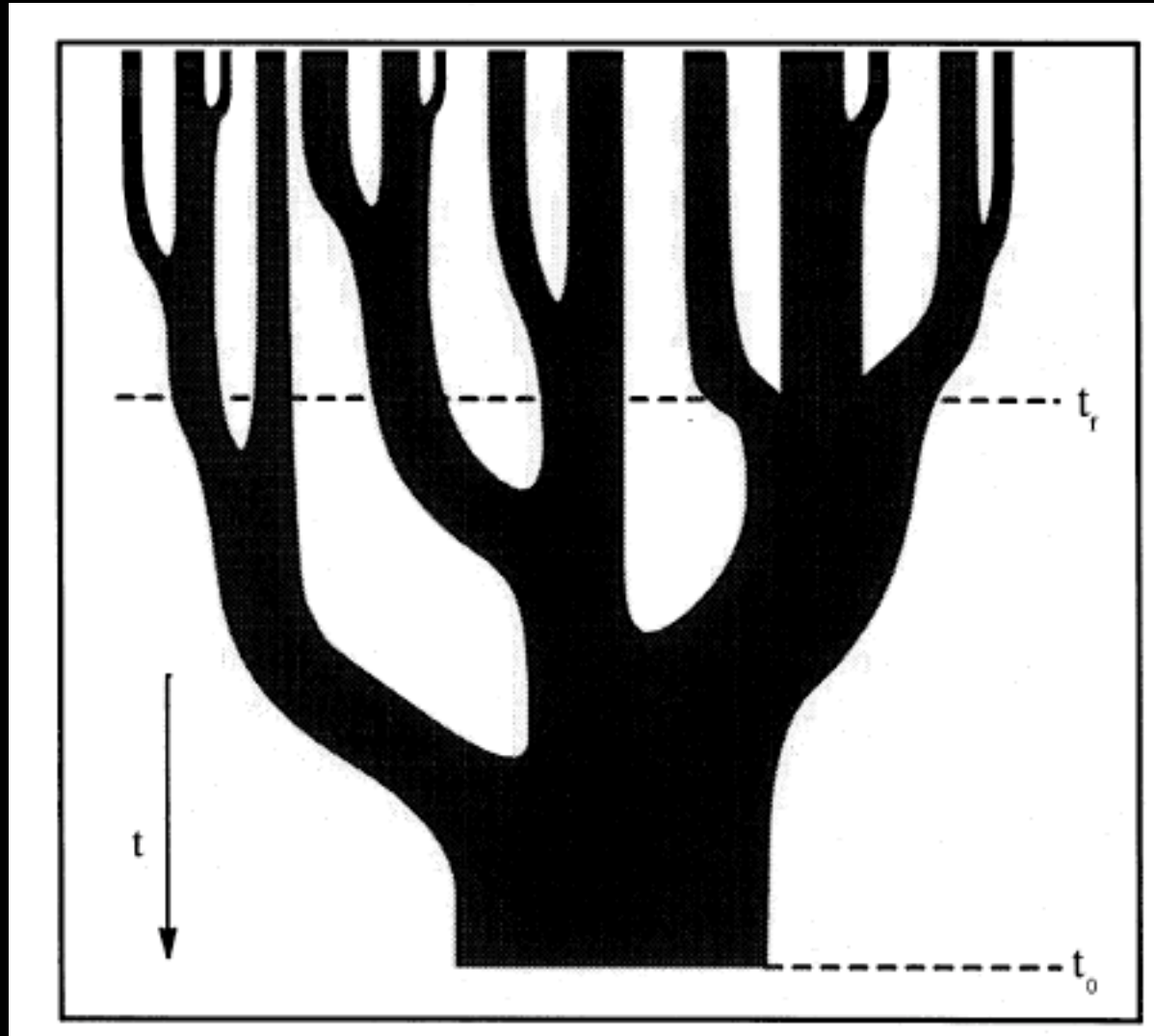
1. Zwicky (1937) galaxies are about **500 times heavier** in the Coma galaxy cluster than judged from their light emission. This is his famous conjecture that there must be **dark matter**.

2. Zwicky (1956) : when galaxies interact (e.g. when they collide), the expelled matter can re-condense in regions and form new smaller (dwarf) galaxies. This is his famous conjecture that **tidal-dwarf galaxies** can form out of the collisional debris of other galaxies.

Both have to be true . . .

→ the LCDM model →
Structures form according to the cosmological merger tree

Lacey & Cole
(1993)



the
beginning

today

Pavel Kroupa: AfA, University of Bonn

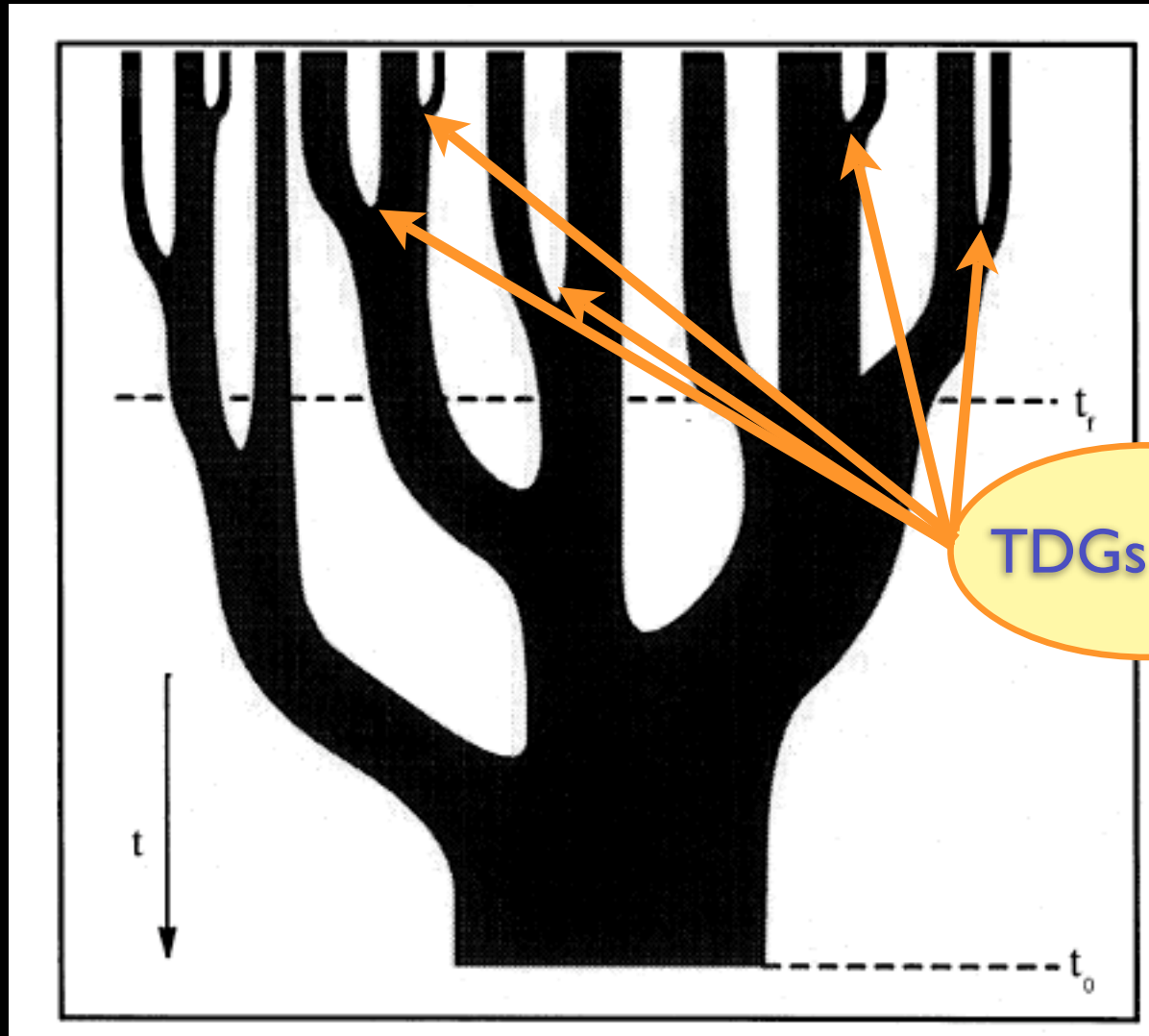


$\approx 250 \text{ kpc}$

by Zwicky's 1st conjecture

Structures form according to the cosmological merger tree

Lacey & Cole
(1993)



the
beginning

by
**Zwicky's
2nd
conjecture**

TDGs form

today

Pavel Kroupa: ALFA, University of Bonn

(Weilbacher et al. 2000)

$$N_{\text{TDG}} \approx 14$$

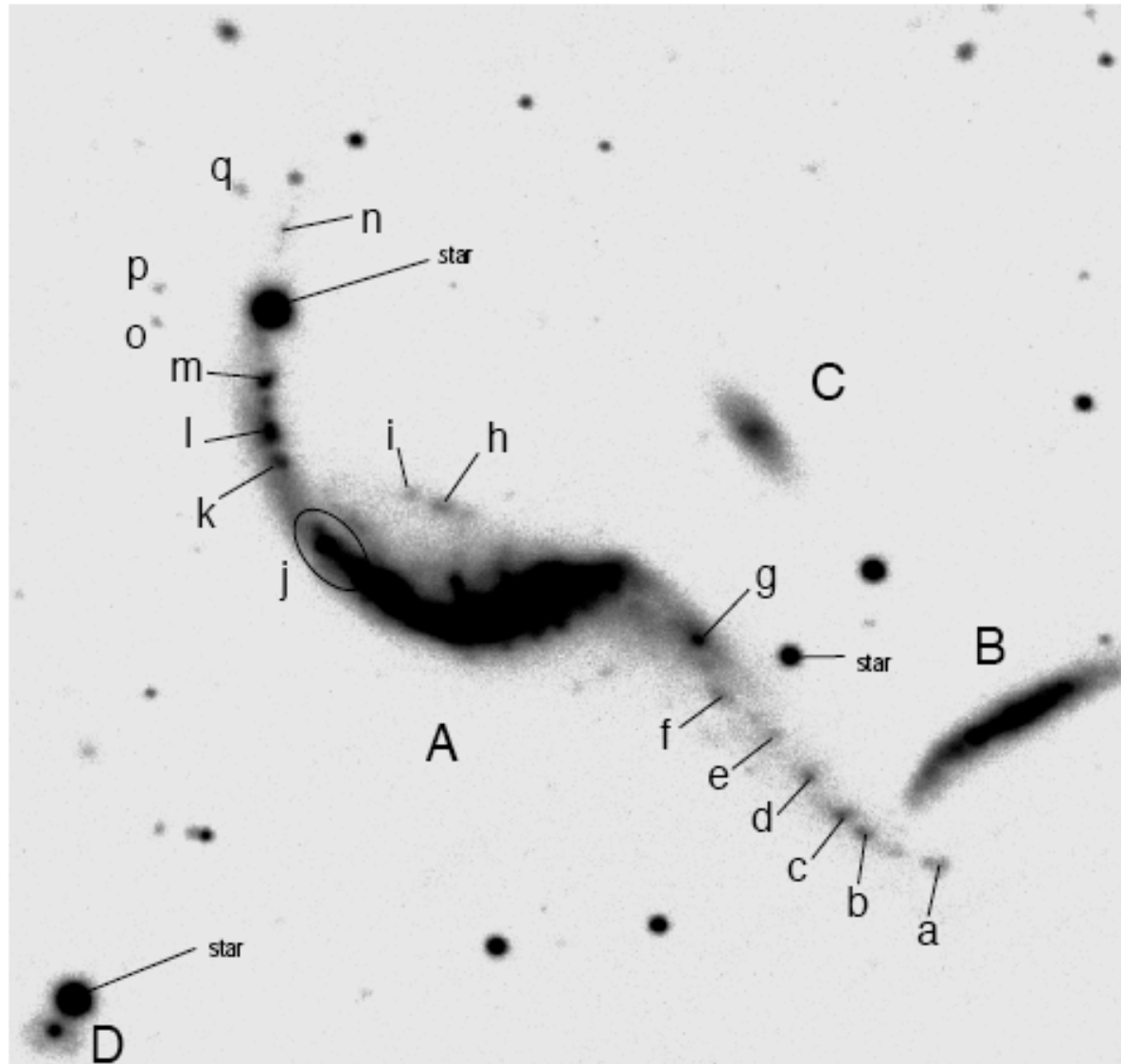


Fig. 21. Identification chart of field 10 around AM 1353-272.

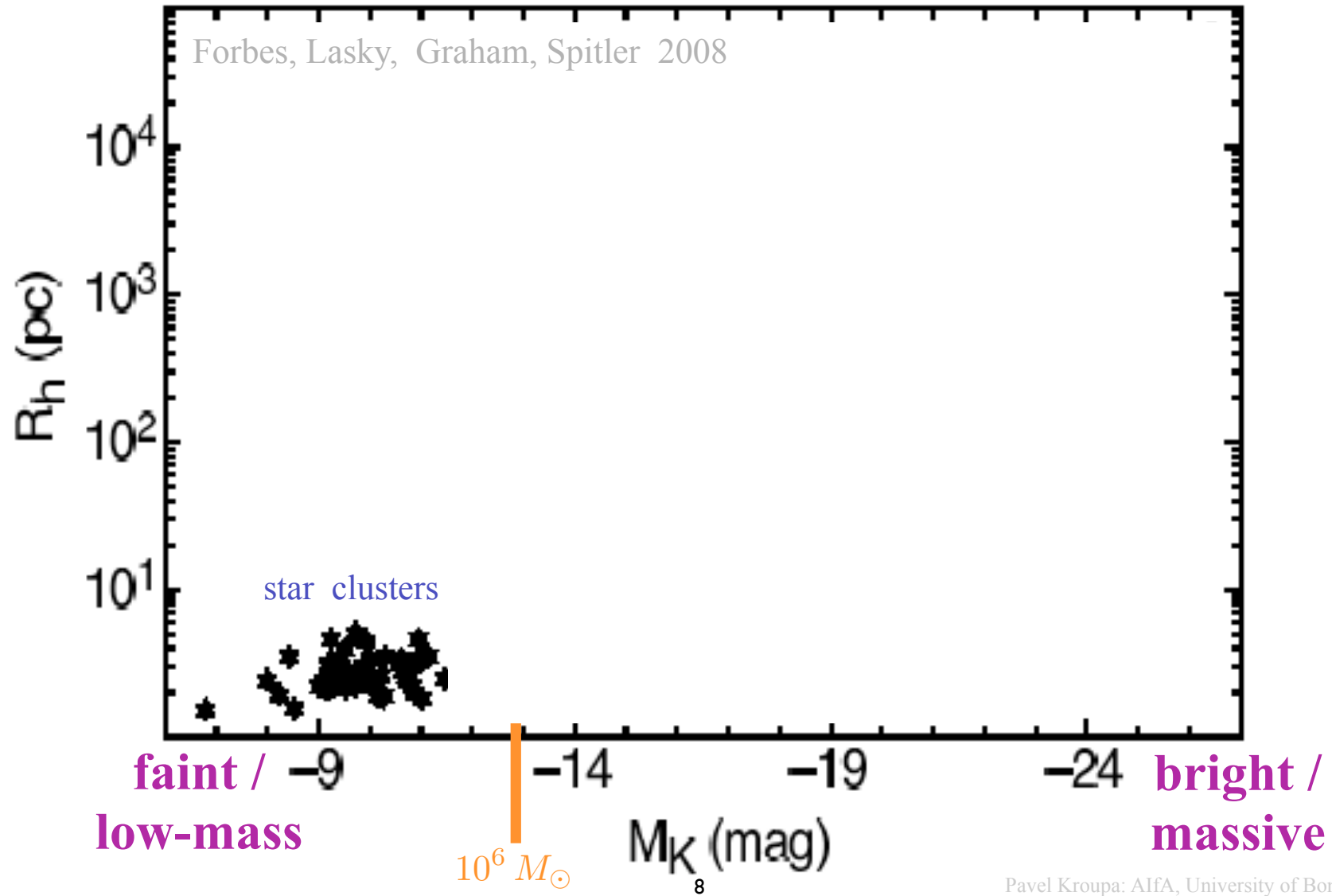
Fritz Zwicky's two conjectures
have
two immediate implications :

1. There exist large numbers of *dark-matter dominated satellite galaxies* (e.g. Moore et al., Klypin et al.).
2. There exist large numbers of *newly formed (tidal-dwarf) satellite galaxies* (they do not contain dark matter) (Okazaki & Taniguchi 2000).

This is OK, but are there
two different types of dwarf galaxy?

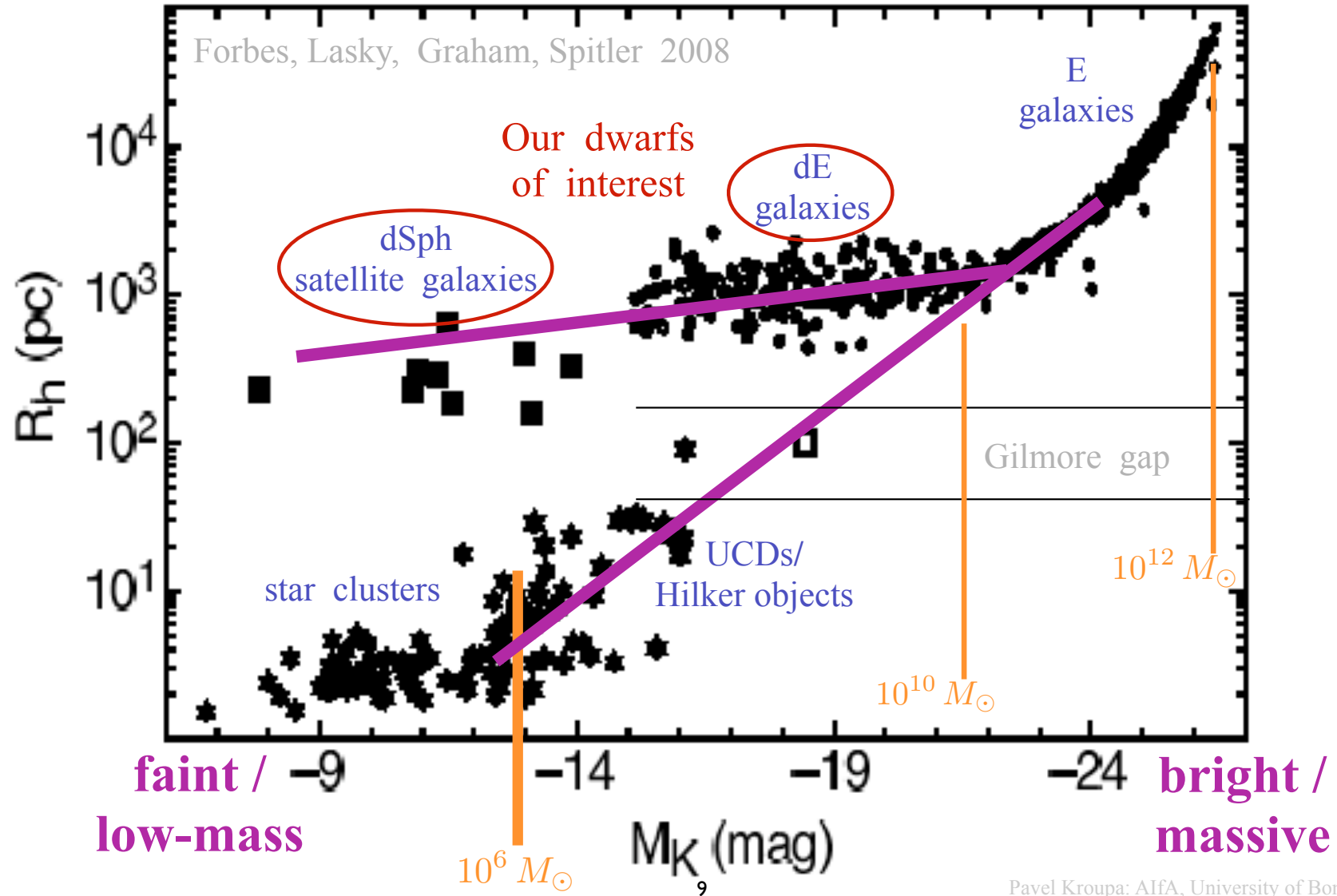
Radius vs luminosity :

Dabringhausen, Hilker, Kroupa 2008;
Misgeld & Hilker 2011



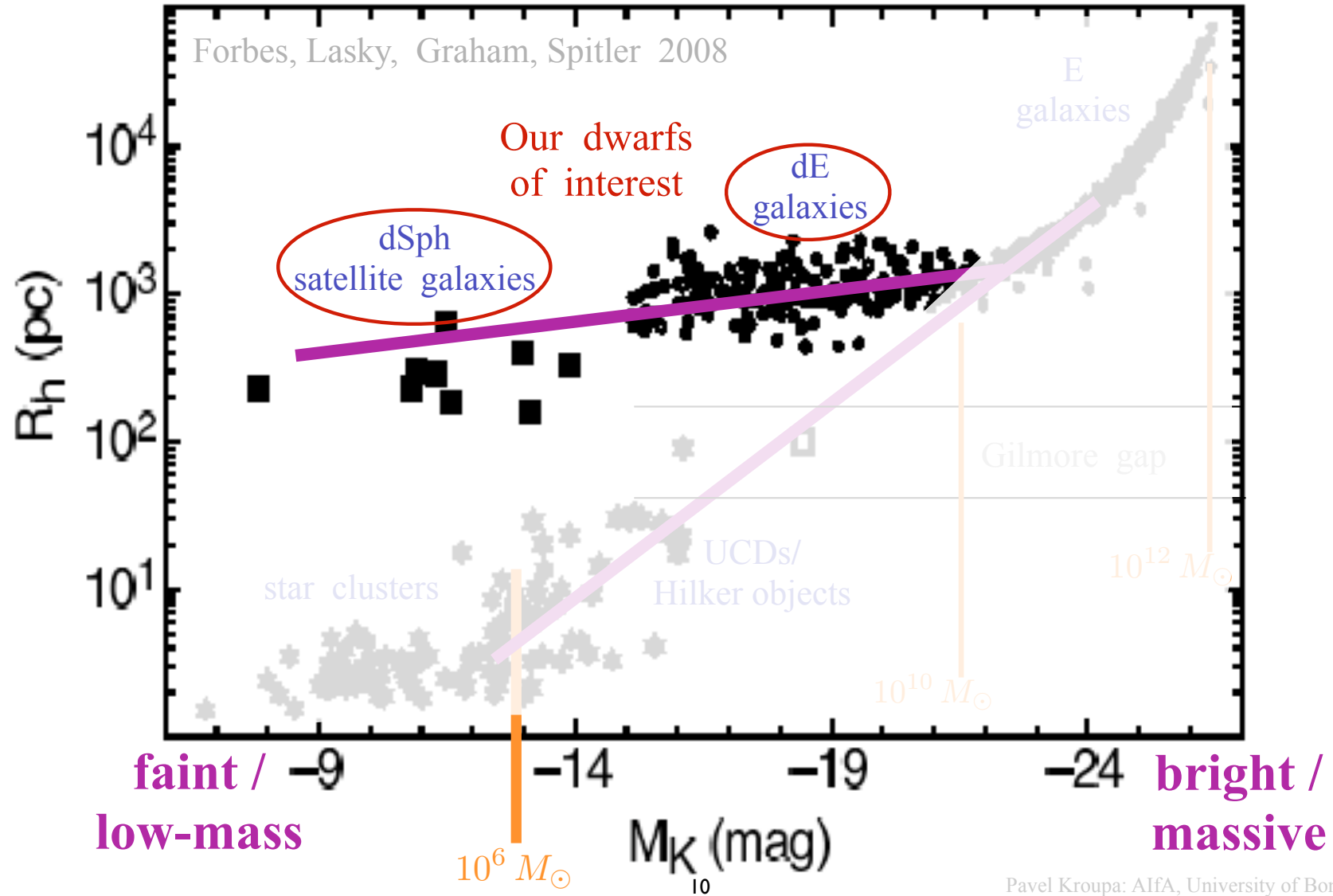
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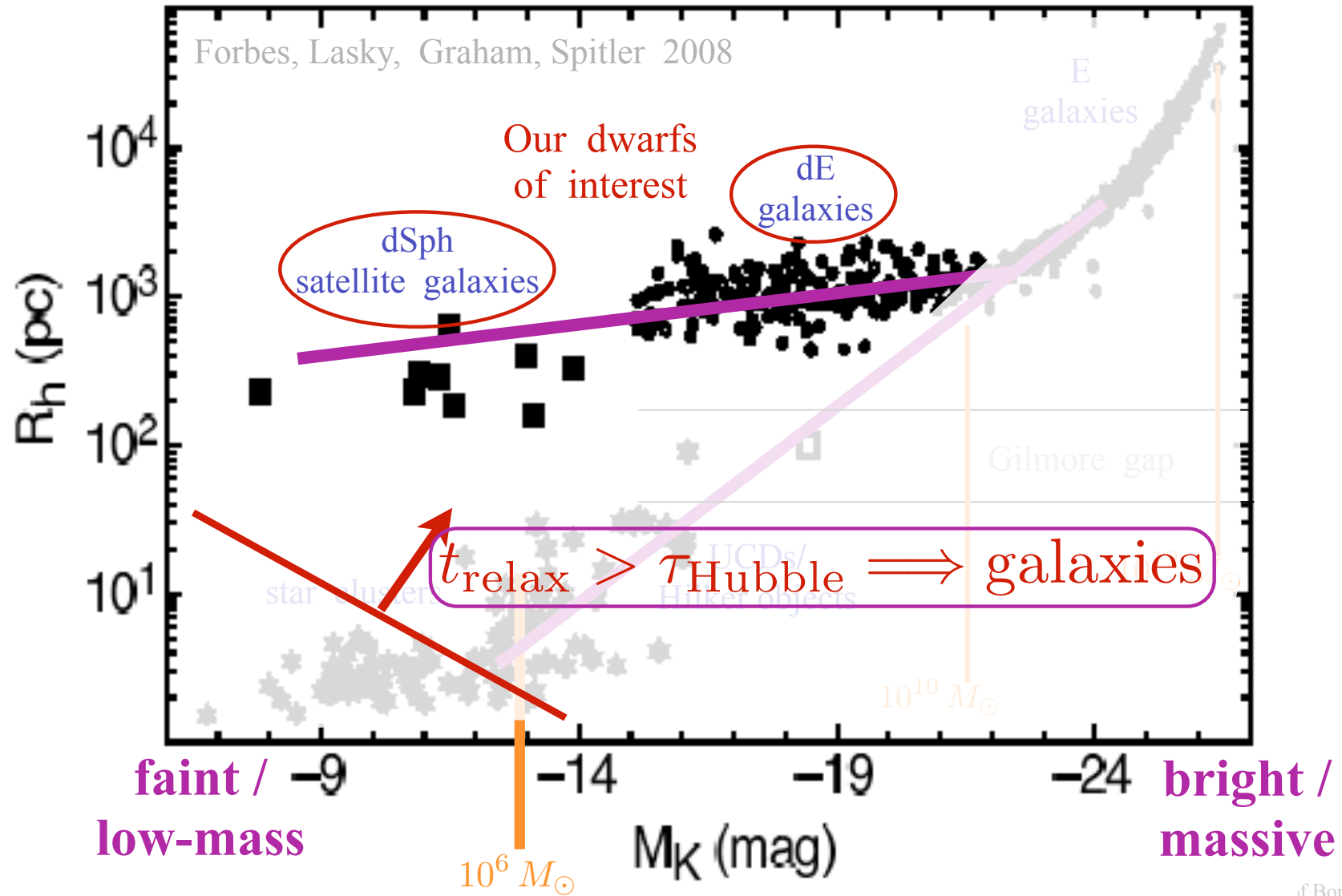
Radius vs luminosity :

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This is OK, but are there
two different types of dwarf galaxy?

No, there is *only one type* of dwarf galaxy !

But, *which one* ? And *why* only one ?

Lets consider first the
dark-matter type satellite
dwarf galaxy.

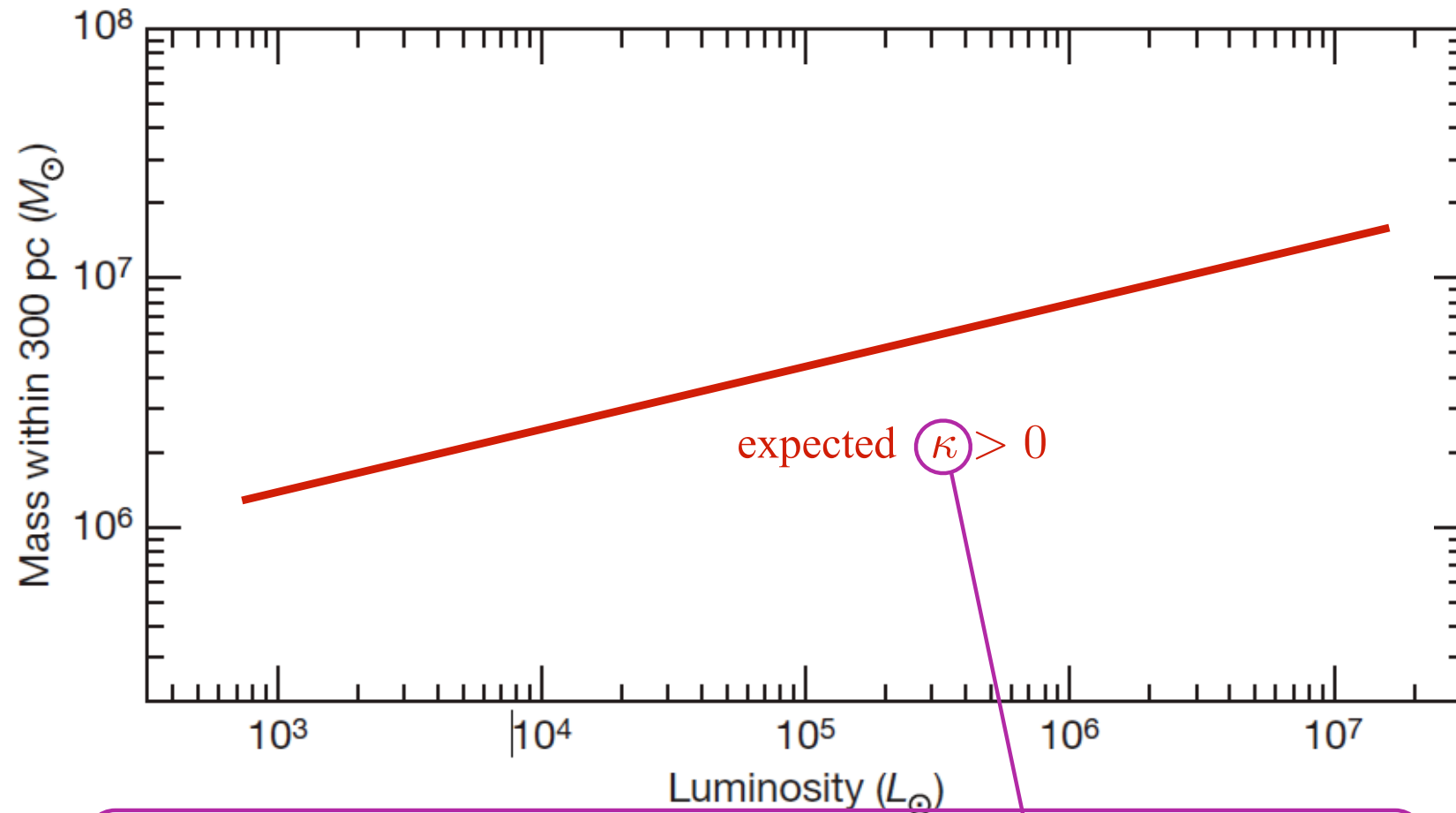
Test this idea with calculations
within the LCDM framework:

The mass-luminosity relation :

energy balance:

more gravitating mass,
more luminous mass ?

The mass - luminosity relation of satellite galaxies

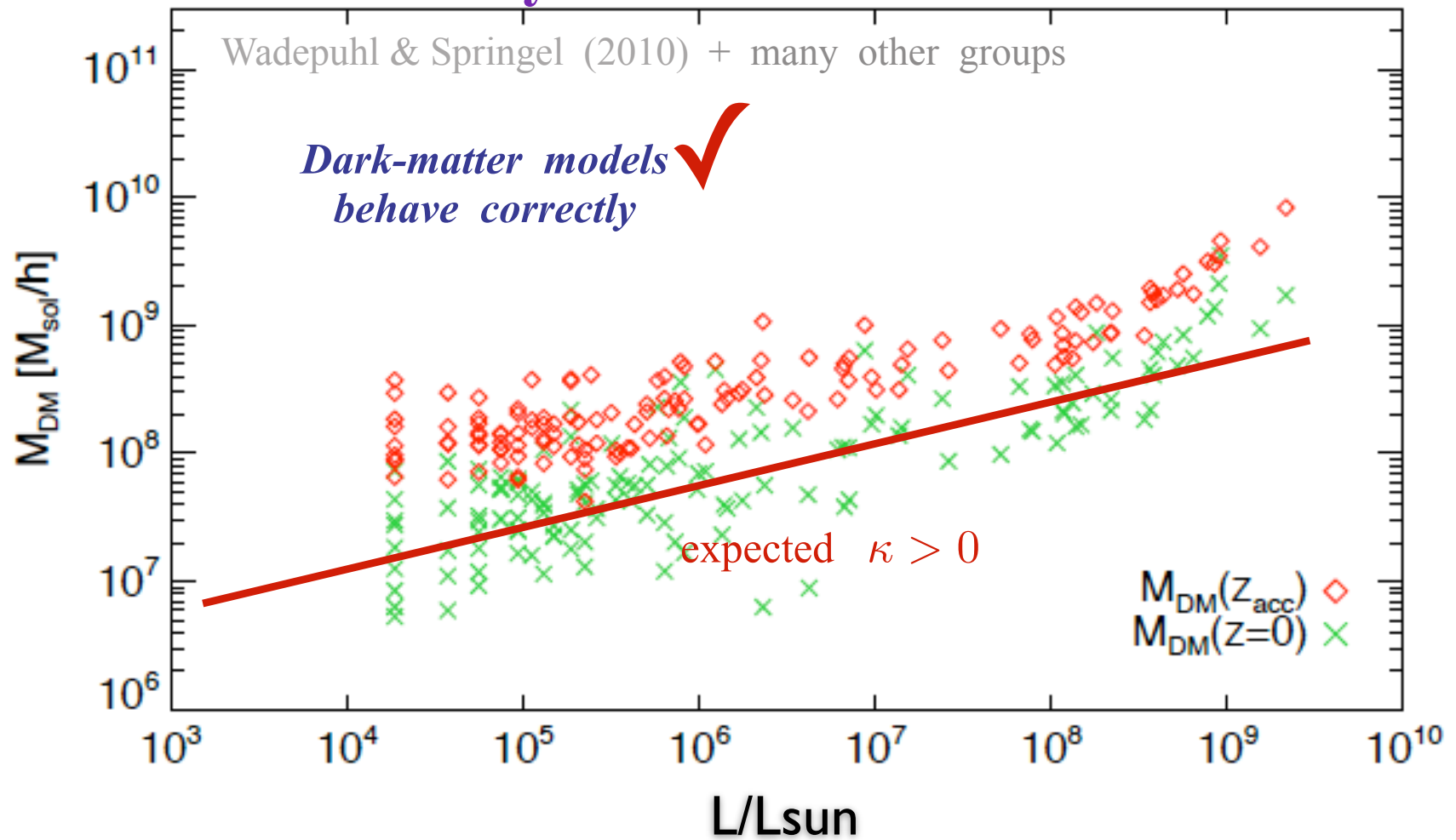


$$\log_{20}(M_{0.3 \text{ kpc}}/M_{\odot}) = \log_{10}(M_{03}/M_{\odot}) + \kappa \log_{10}(L_V/L_{V,\odot})$$

Pavel Kroupa: AIfA, University of Bonn

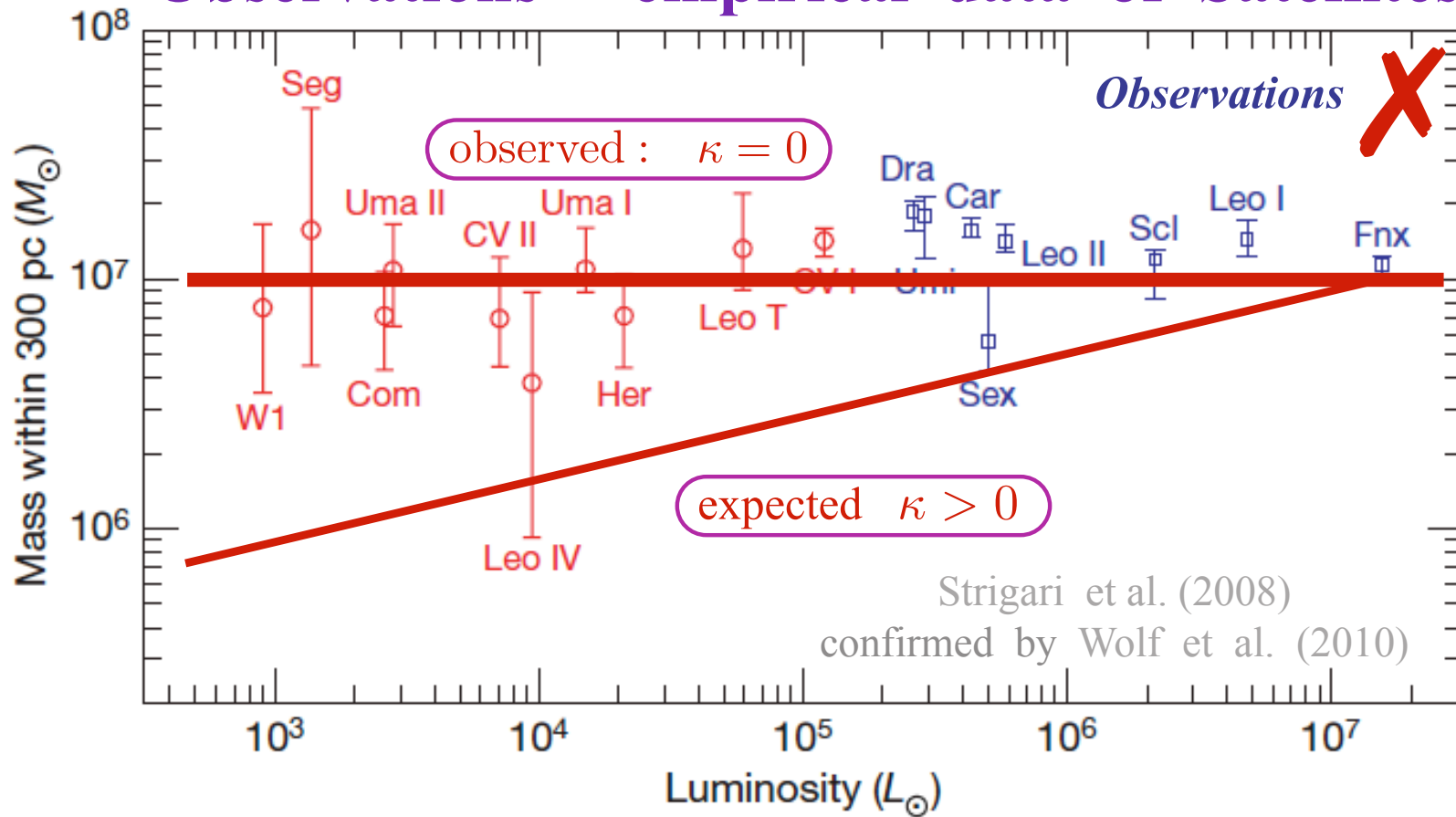
The DM-mass--luminosity relation of satellite galaxies

Theory - Models of Satellites



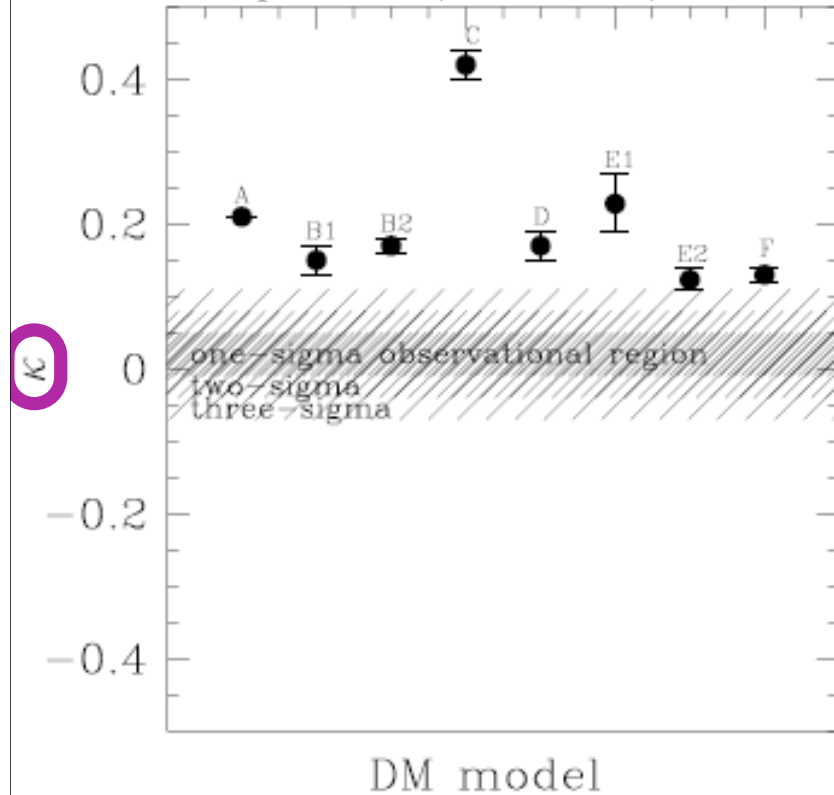
The DM-mass--luminosity relation of satellite galaxies

Observations - empirical data of Satellites



Pavel Kroupa: AIfA, University of Bonn

Kroupa et al. (2010, A&A)



$$\log_{20}(M_{0.3 \text{ kpc}}/M_{\odot}) = \log_{10}(M_{03}/M_{\odot}) + \kappa \log_{10}(L_V/L_{V,\odot})$$

Pavel Kroupa: AlfA, University of Bonn

All model satellite populations have

$$\kappa > 0$$

The *lack* of an *observed* mass-luminosity relation

$$(\kappa \approx 0)$$



nature apparently does not care
about the existence of the
putative dark matter halo.

Thus, the *concept*
of dark-matter halos
appears to be
unphysical
for dSph satellites

~~dark-matter halos~~

Individual dSph
morphology :

DM gravitating potential :
smooth luminous morphology ?

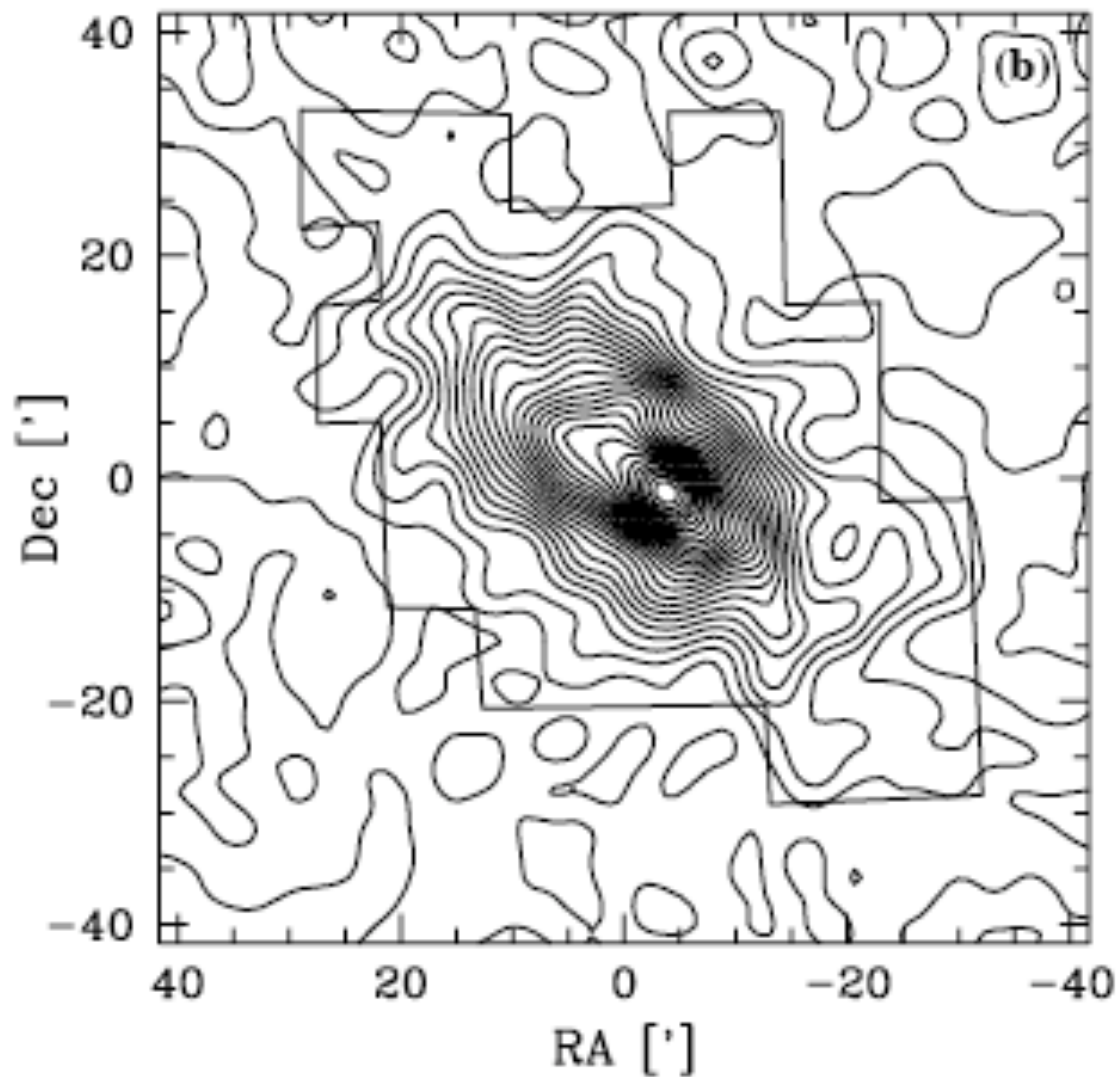
Significant isophote structure is present in many dSph satellites despite a large

$$\sigma \approx 700 \text{ pc}/100 \text{ Myr}$$

Substructure should smear-out if σ is really due to a DM halo, unless it has a harmonic core.



consistent with DM halo ?



UMi

$D=65\text{kpc}$

(Kleyna et al. 1998)

$$\left(\frac{M}{L}\right)_{0,V} = 60$$

$$\left(\frac{M}{L}\right)_{\text{tot},V} = 79$$

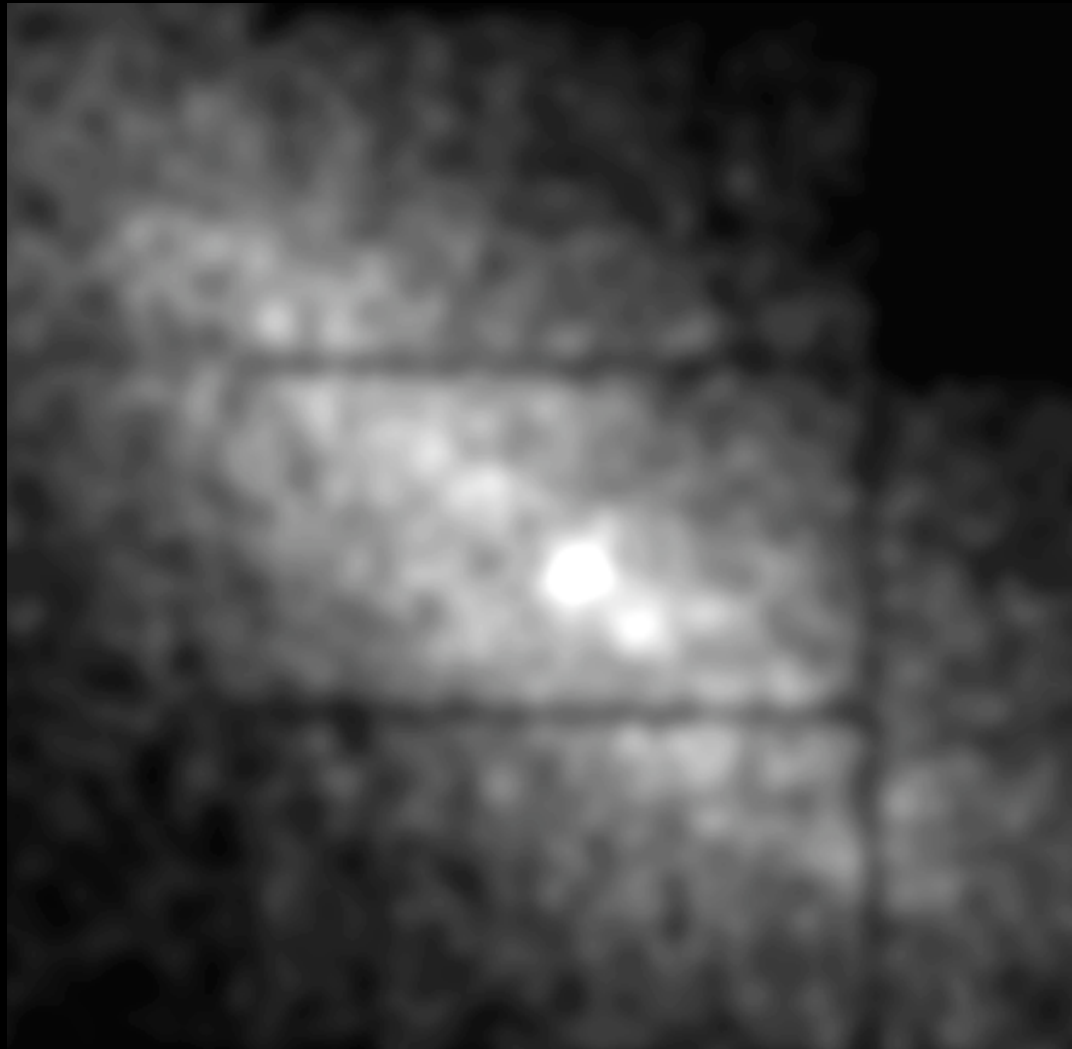
(Mateo 1998)

But :

$$\left(\frac{M}{L}\right)_{\text{particles}} = 12$$

(Gomez-Flechoso & Martinez-Delgado 2003)

Pavel Kroupa: AIfA, University of Bonn



UMi

D=65kpc

(Martinez-Delgado et al.,
in prep)

Substructure
significant :

(Kleyna et al. 2003)



UMi

D=65kpc

(Martinez-Delgado et al.,
in prep)

S shape :
strong evidence for
extra-tidal stars



Massive CDM

halo ?

Pavel Kroupa: AIfA, University of Bonn

Fornax

D=140kpc

(Demers et al. 1994)

$$\left(\frac{M}{L}\right)_{0,V} = 4.8$$

$$\left(\frac{M}{L}\right)_{\text{tot},V} = 4.4$$

(Mateo 1998)

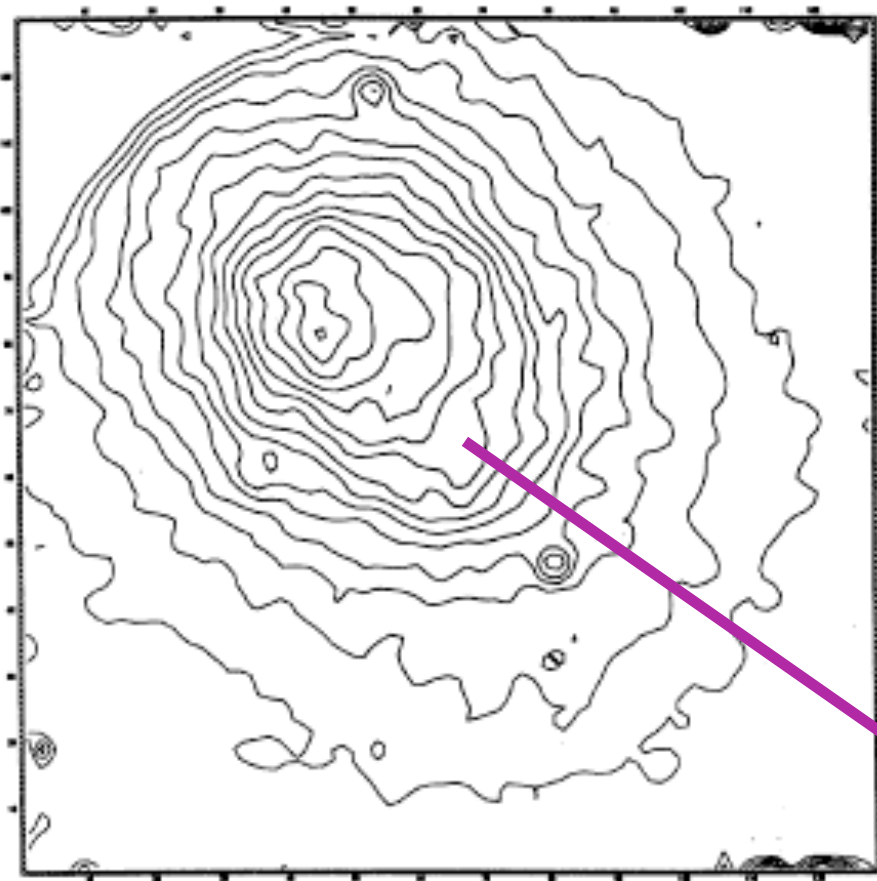


FIG. 4. Isophotes of equal stellar density of field 1 reveal that the structure of Fornax, near the center, is not symmetric.

Not consistent with being embedded / shielded by an extensive dark-matter subhalo !

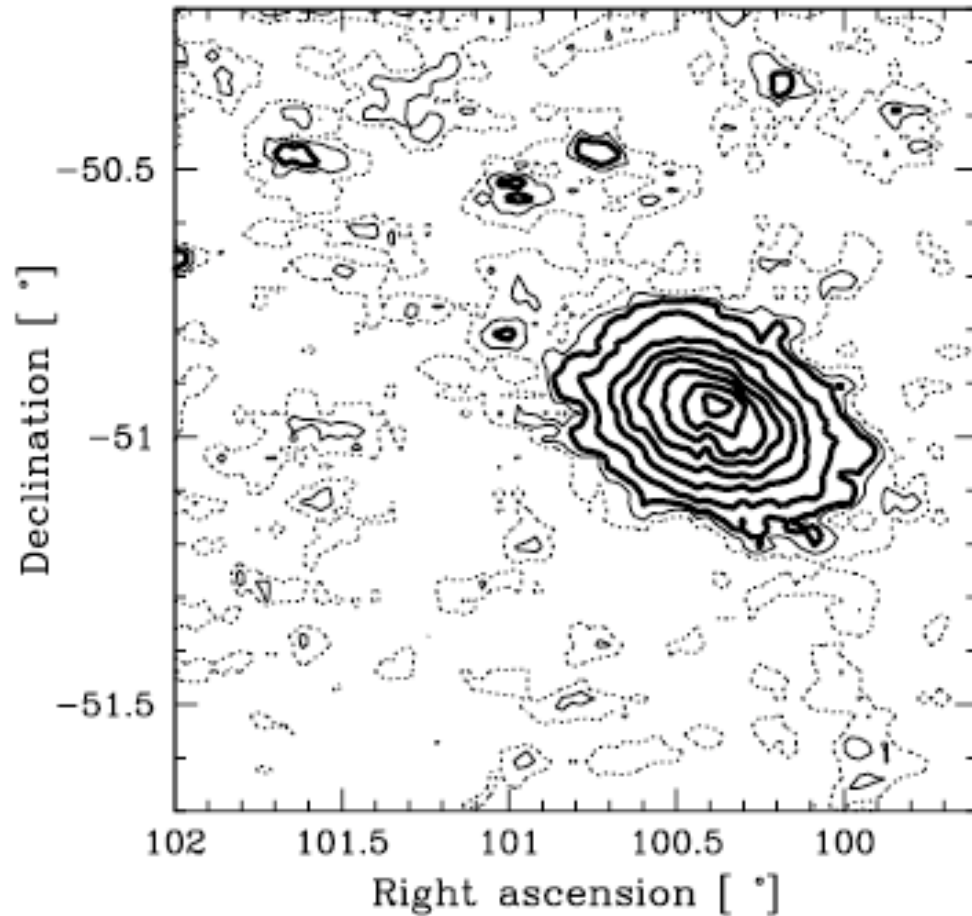


Fig. 1. Contour plot of the Carina dwarf spheroidal. The density levels correspond to background value (dotted line), 1σ above that (thin solid line), 2σ , 5σ , 10σ and so on (thick solid lines). No significant departure from the spheroidal shape can be seen. A galactic gradient can be seen from the northeastern to the southwestern corner.

Carina

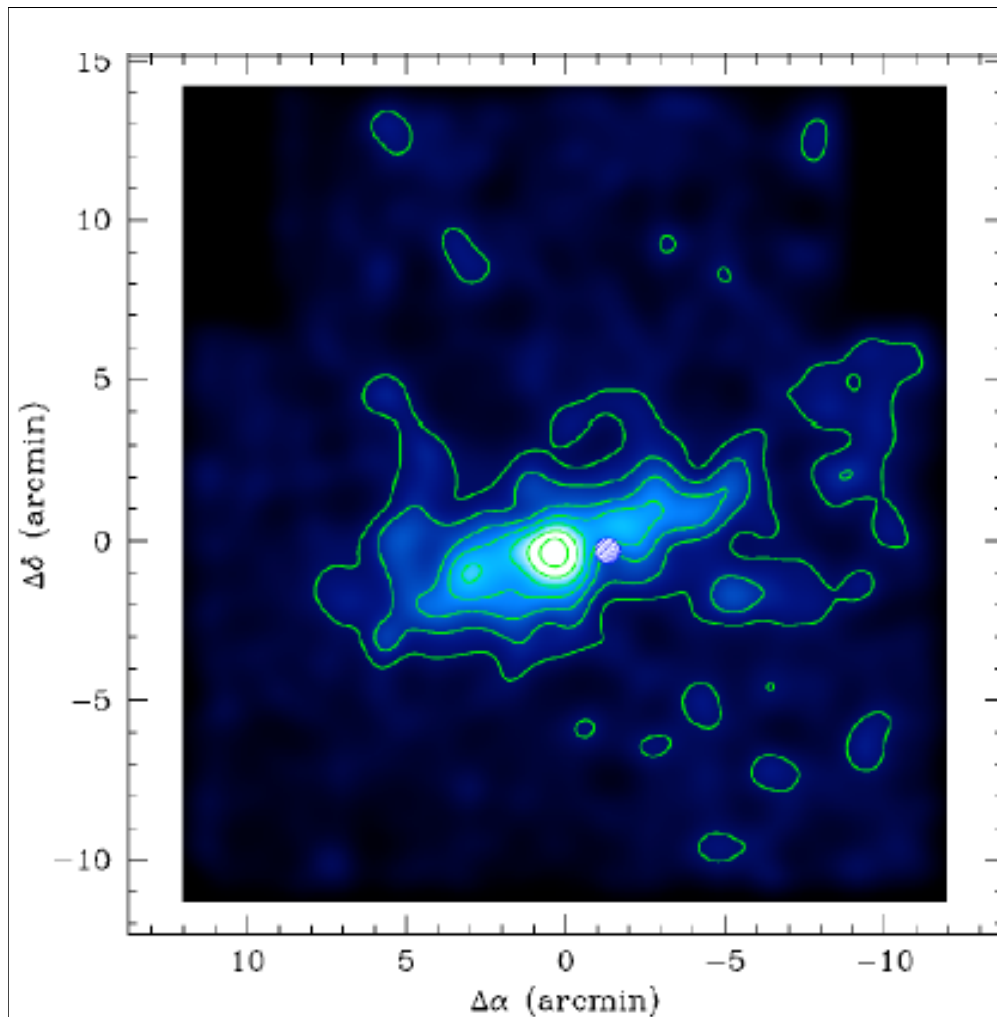
$D=93\text{kpc}$

(Walcher et al. 2003)

$$\left(\frac{M}{L}\right)_{0,V} = 30$$

$$\left(\frac{M}{L}\right)_{0,V} = 31$$

(Mateo 1998)



Hercules

$D=130\text{kpc}$

(Coleman et al. 2007)

TABLE 1
PROPERTIES OF THE HERCULES DSPH

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$c = \log(r_t/r_c)$	0.74 ± 0.25

The distortions apparent in many of the dSph satellites do not support the notion that they are shielded by $10^9 M_{\odot}$ dark-matter halos.

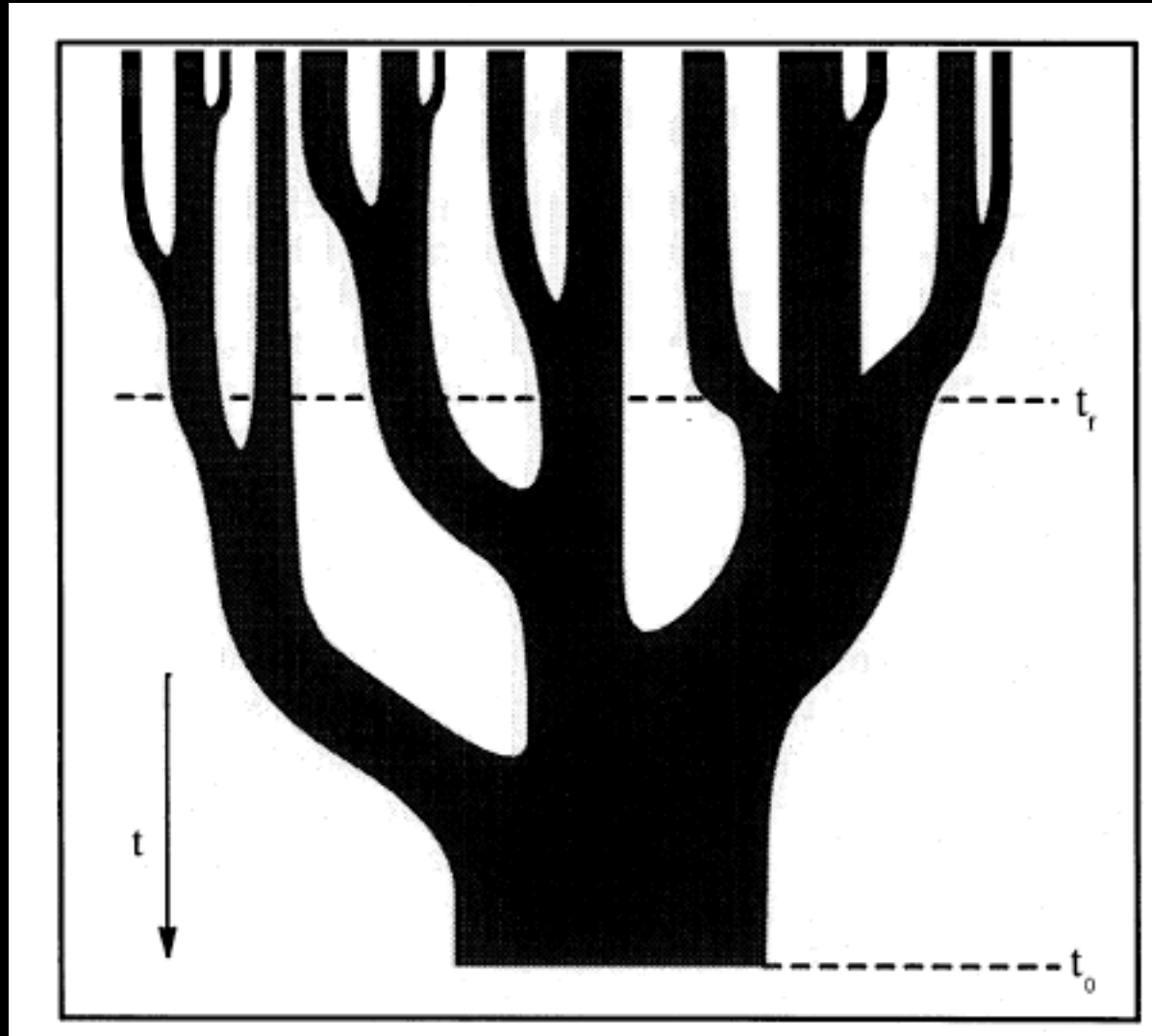
~~dark-matter halos~~

The
spatial distribution
of the
MW satellites

... further clues

Structures form according to the cosmological merger tree

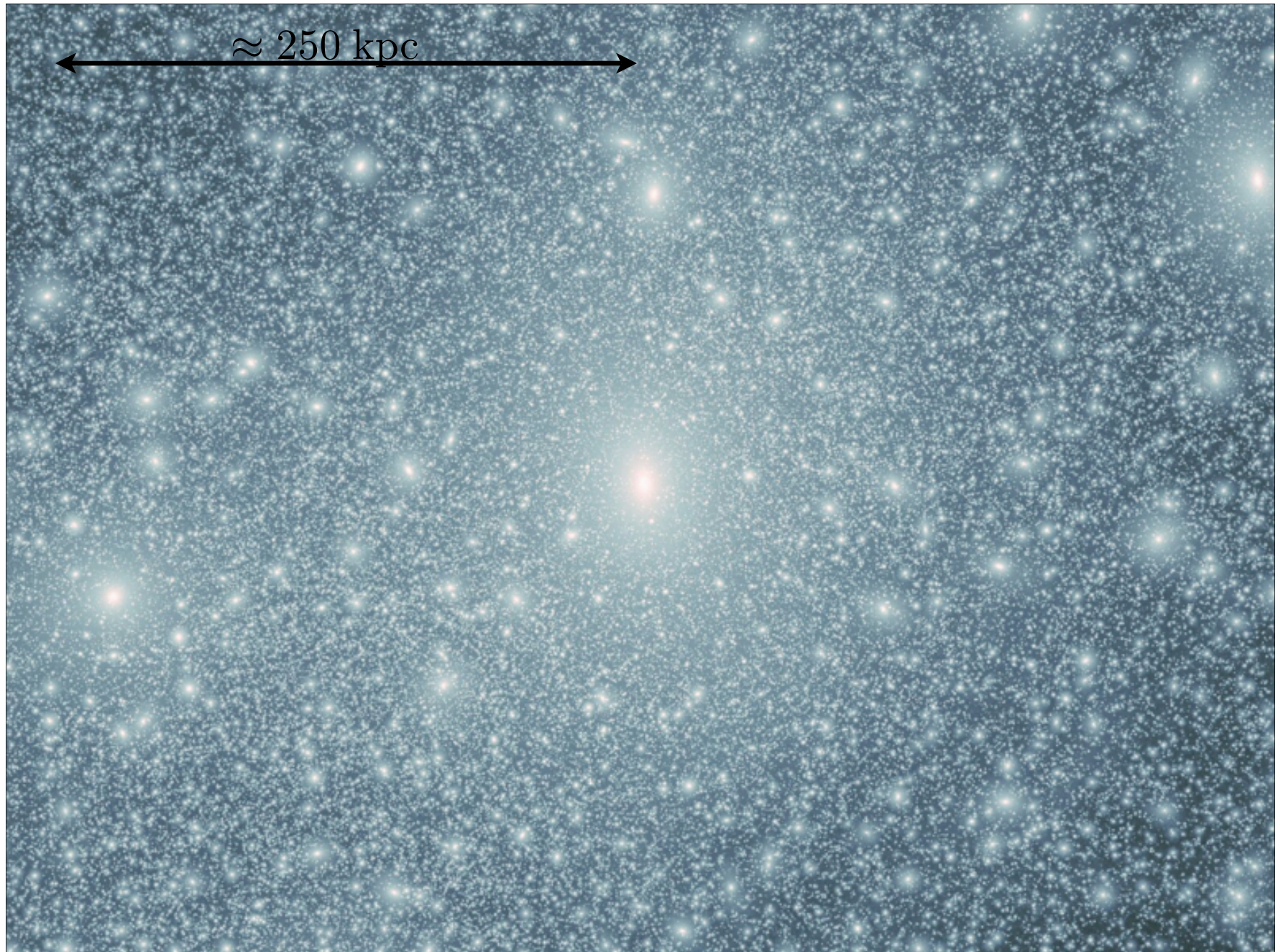
Lacey & Cole
(1993)



the
beginning

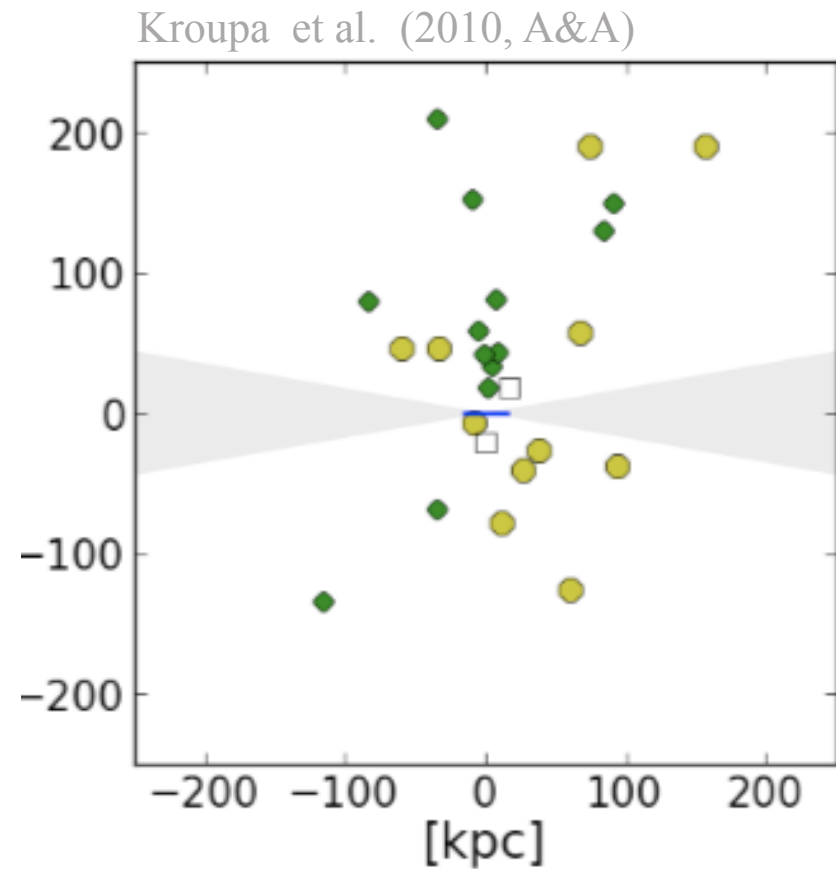
today

Pavel Kroupa: ALFA, University of Bonn



MW satellites are in a disk-like configuration:

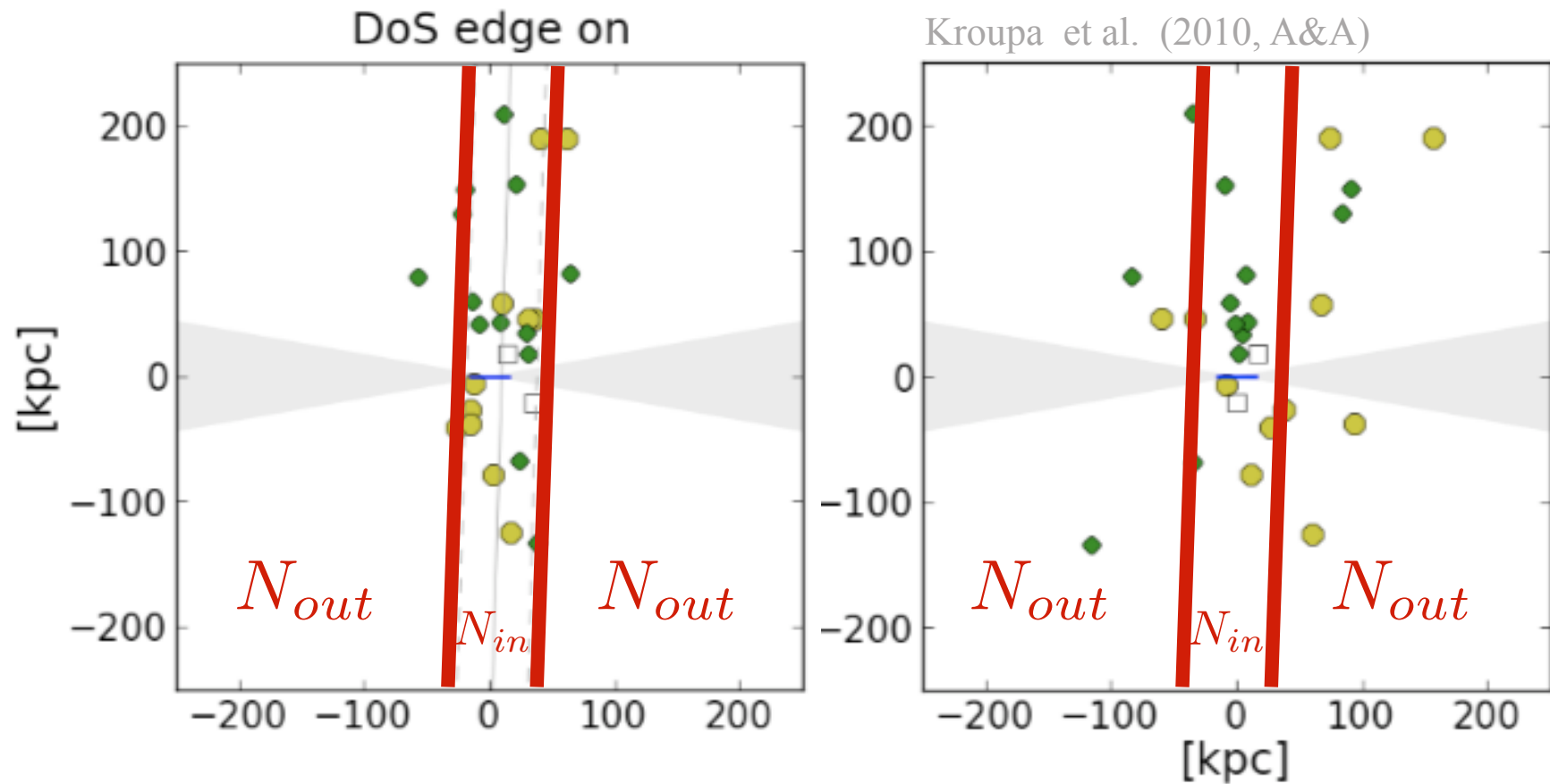
- the 11 “classical” (brightest) satellites
- new satellites



Pavel Kroupa: AITA, University of Bonn

MW satellites are in a disk-like configuration:

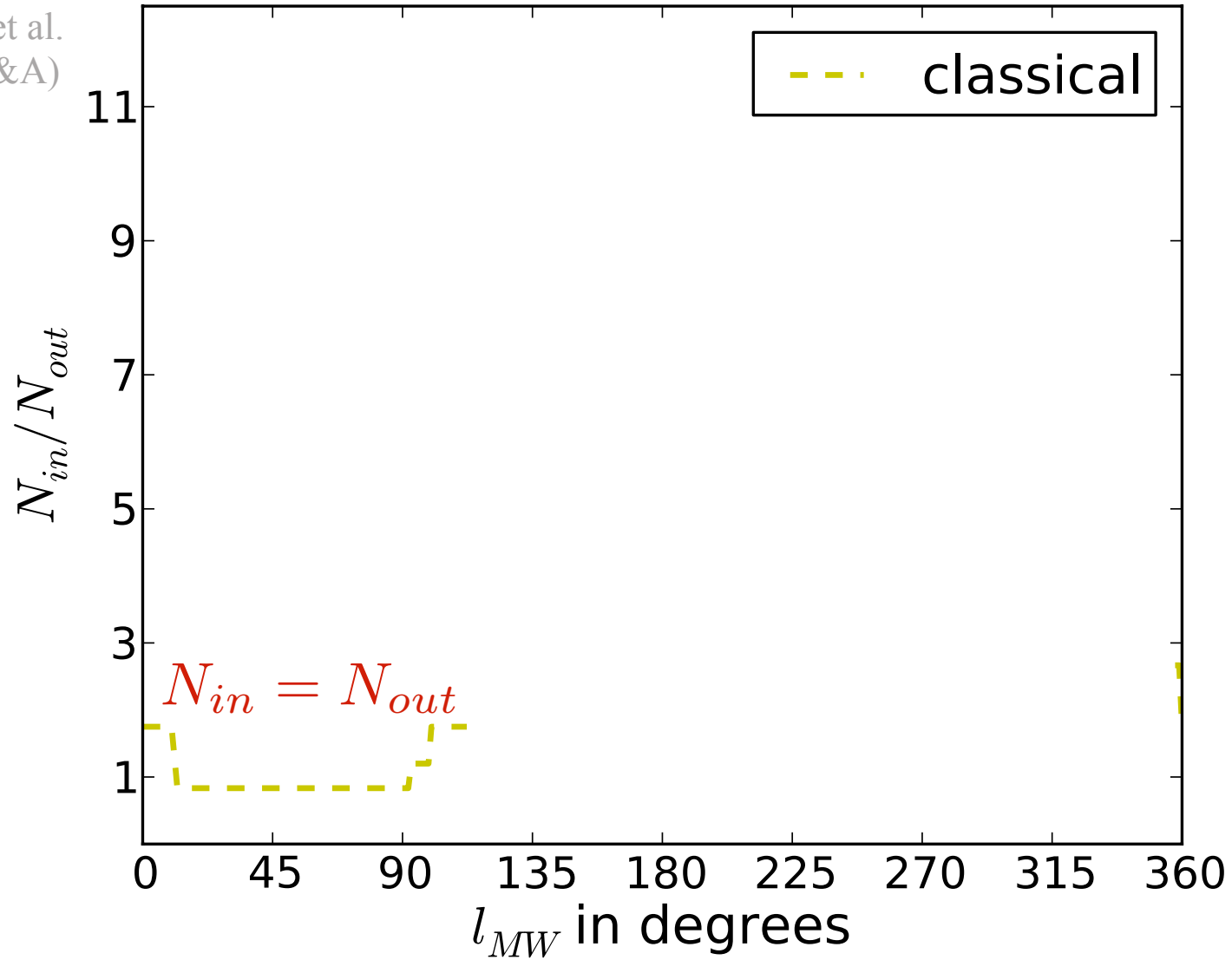
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Pavel Kroupa: AITA, University of Bonn

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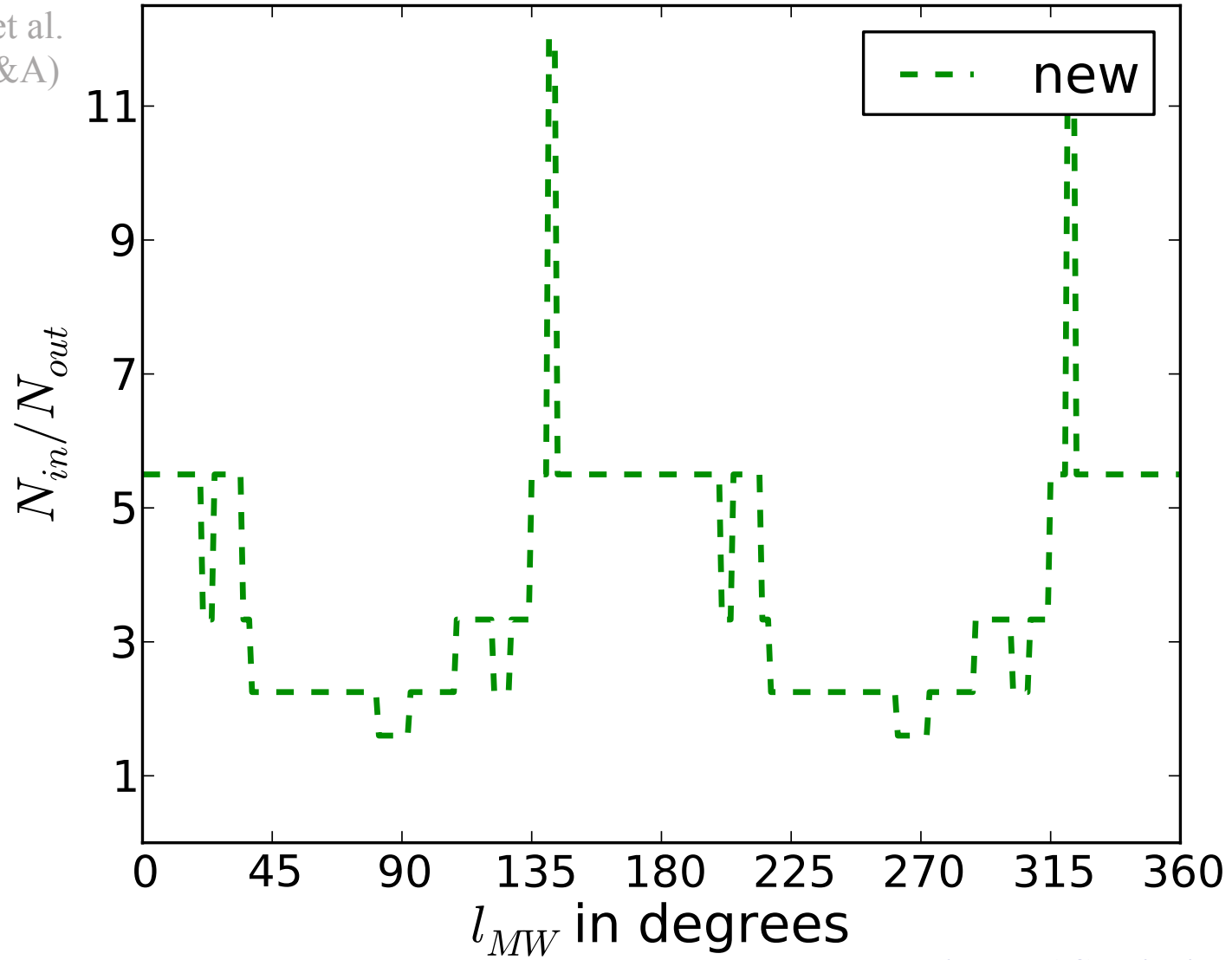
Kroupa et al.
(2010, A&A)



Pavel Kroupa: AlfA, University of Bonn

MW satellites are in a disk-like configuration:

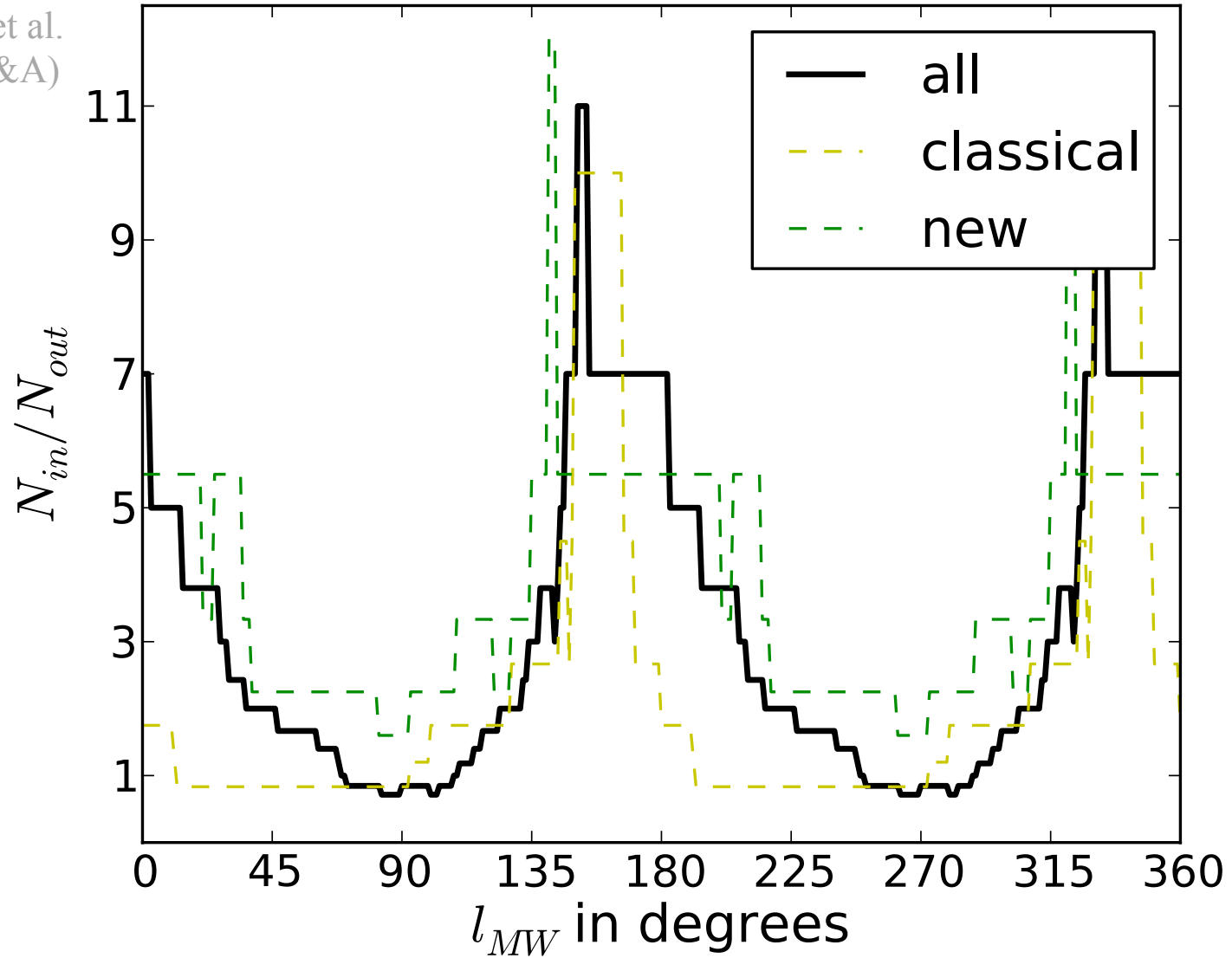
Kroupa et al.
(2010, A&A)



Pavel Kroupa: AIfA, University of Bonn

MW satellites are in a disk-like configuration:

Kroupa et al.
(2010, A&A)

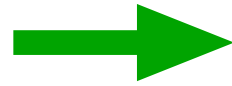


Pavel Kroupa: AIfA, University of Bonn

Thus,

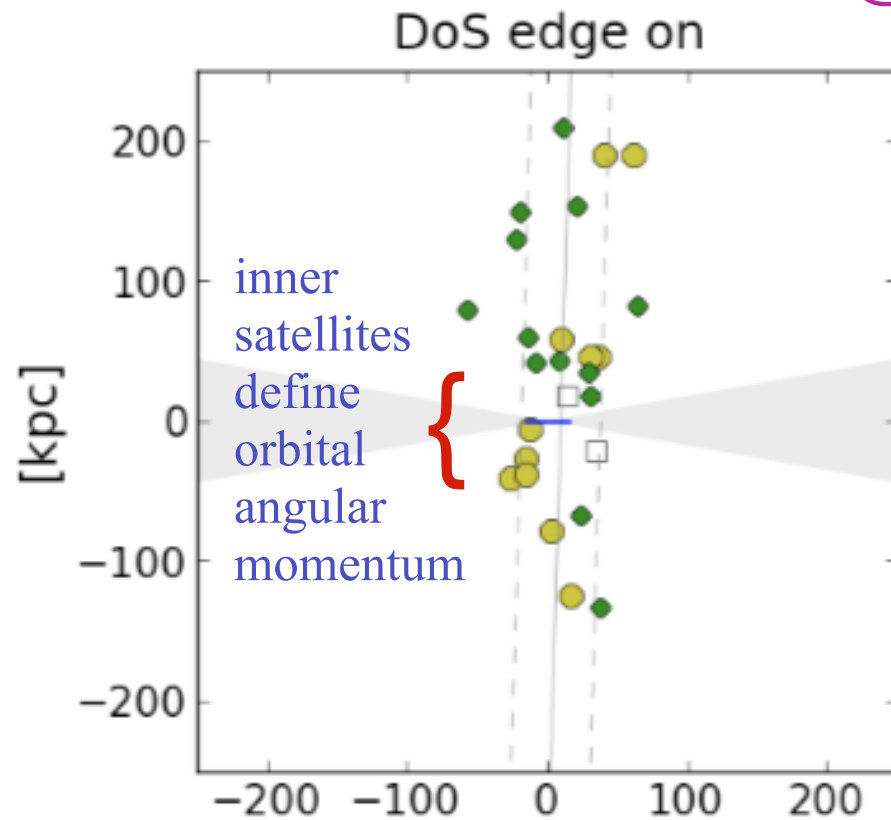
the
13 new ultra-faint
and the
11 classical satellites
independently
define the
same Disk of Satellites
(DoS).

Disk of Satellites



a rotational structure ?

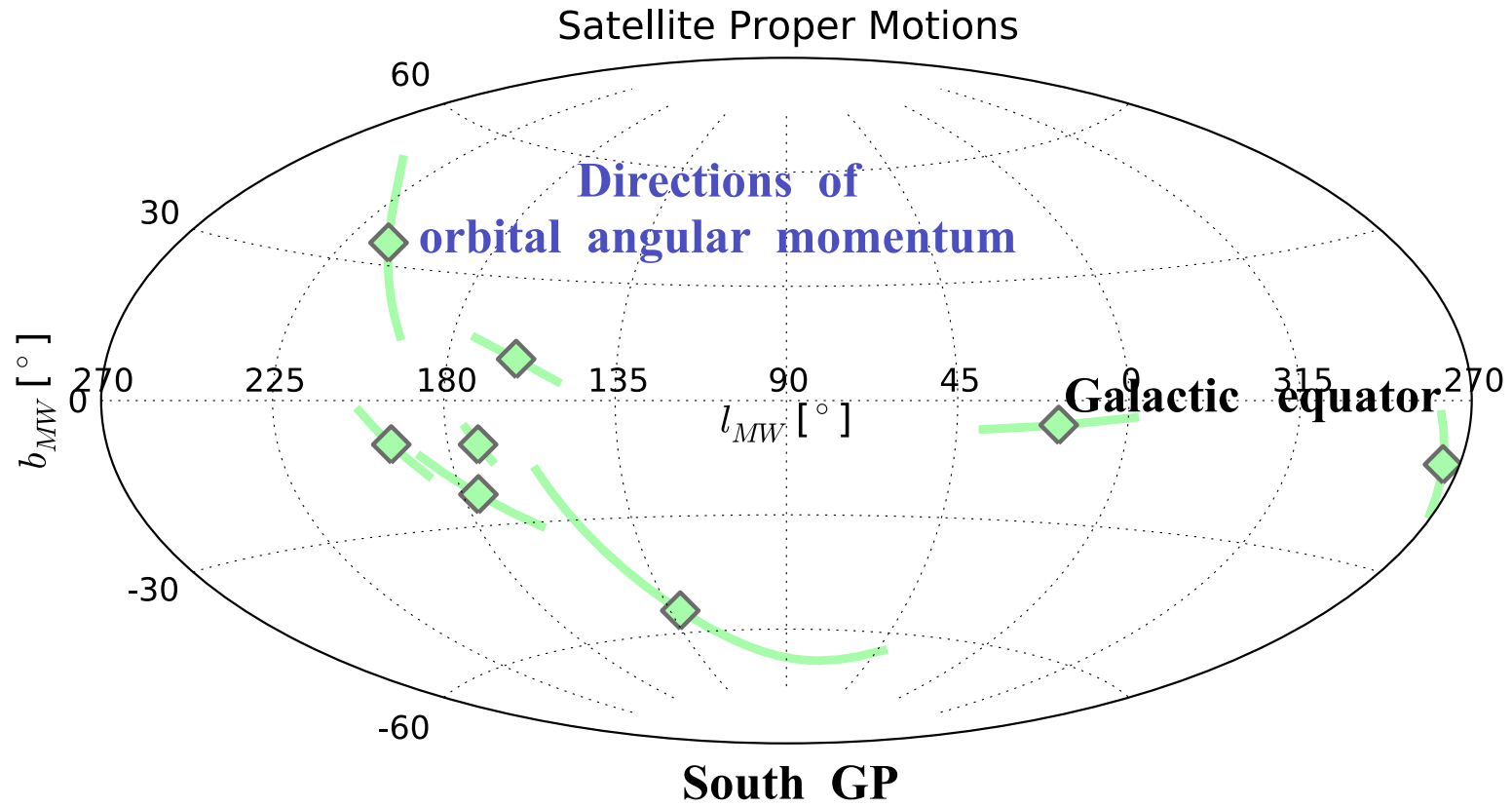
- the 11 “classical” (brightest) satellites
- new satellites



Kroupa et al. (2010, A&A)

the Galactic sky

(Galactic spherical coordinates)

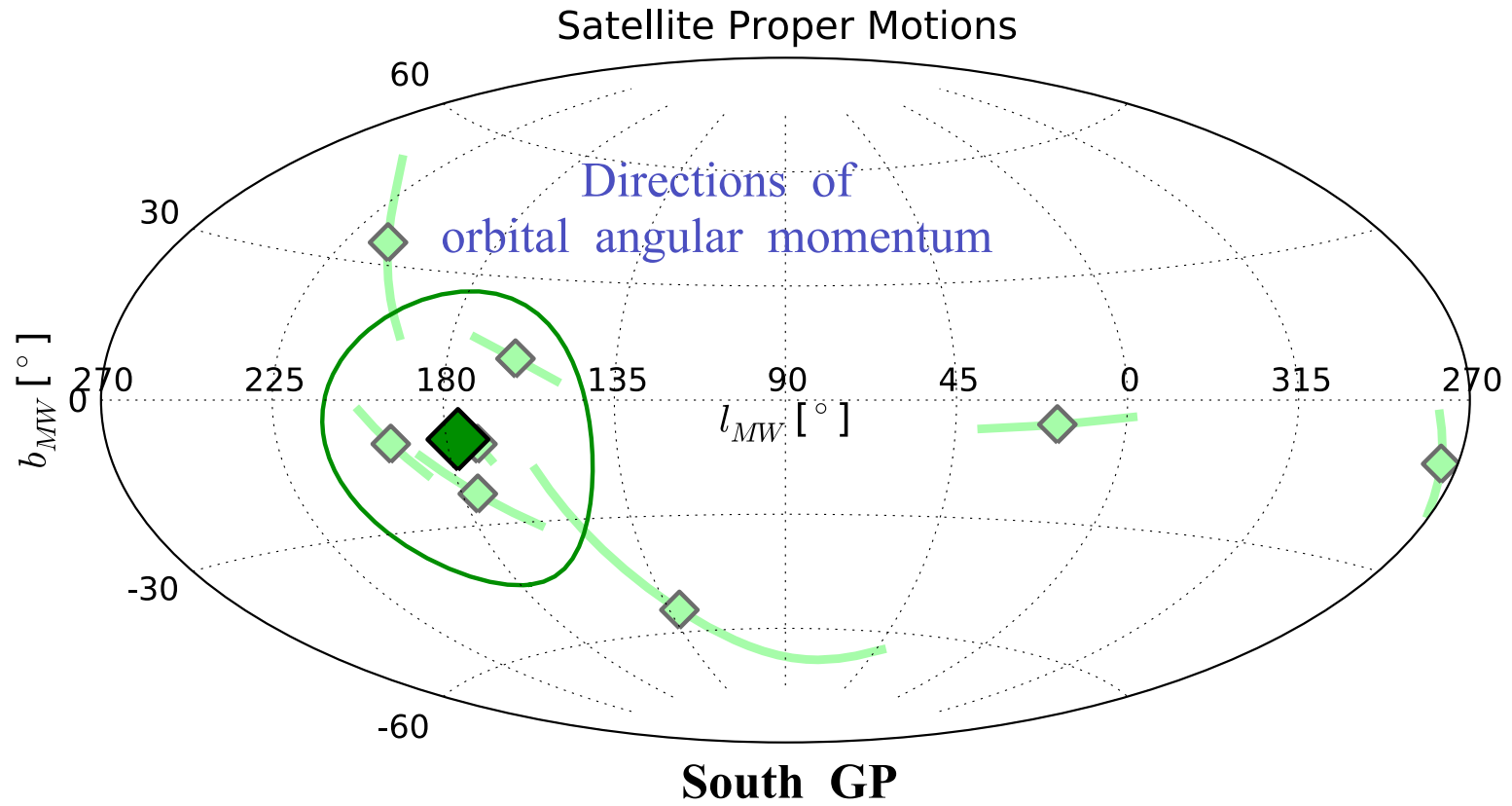


Pawlowski et al. 2011

Pavel Kroupa: AIfA, University of Bonn

the Galactic sky

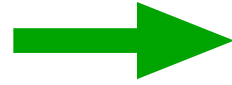
(Galactic spherical coordinates)



Pawlowski et al. 2011

Pavel Kroupa: AIfA, University of Bonn

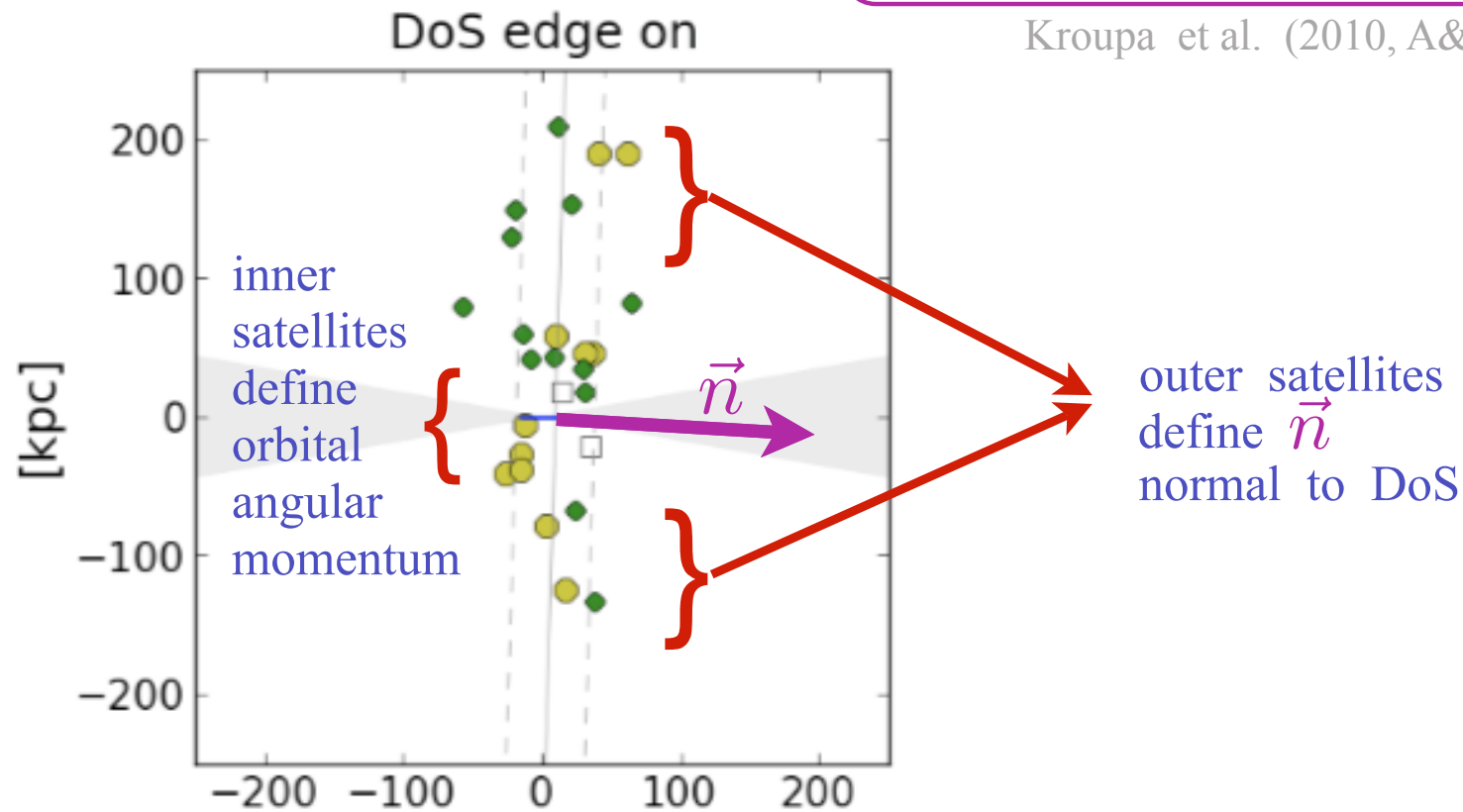
Disk of Satellites



a rotational structure ?

- the 11 “classical” (brightest) satellites
- new satellites

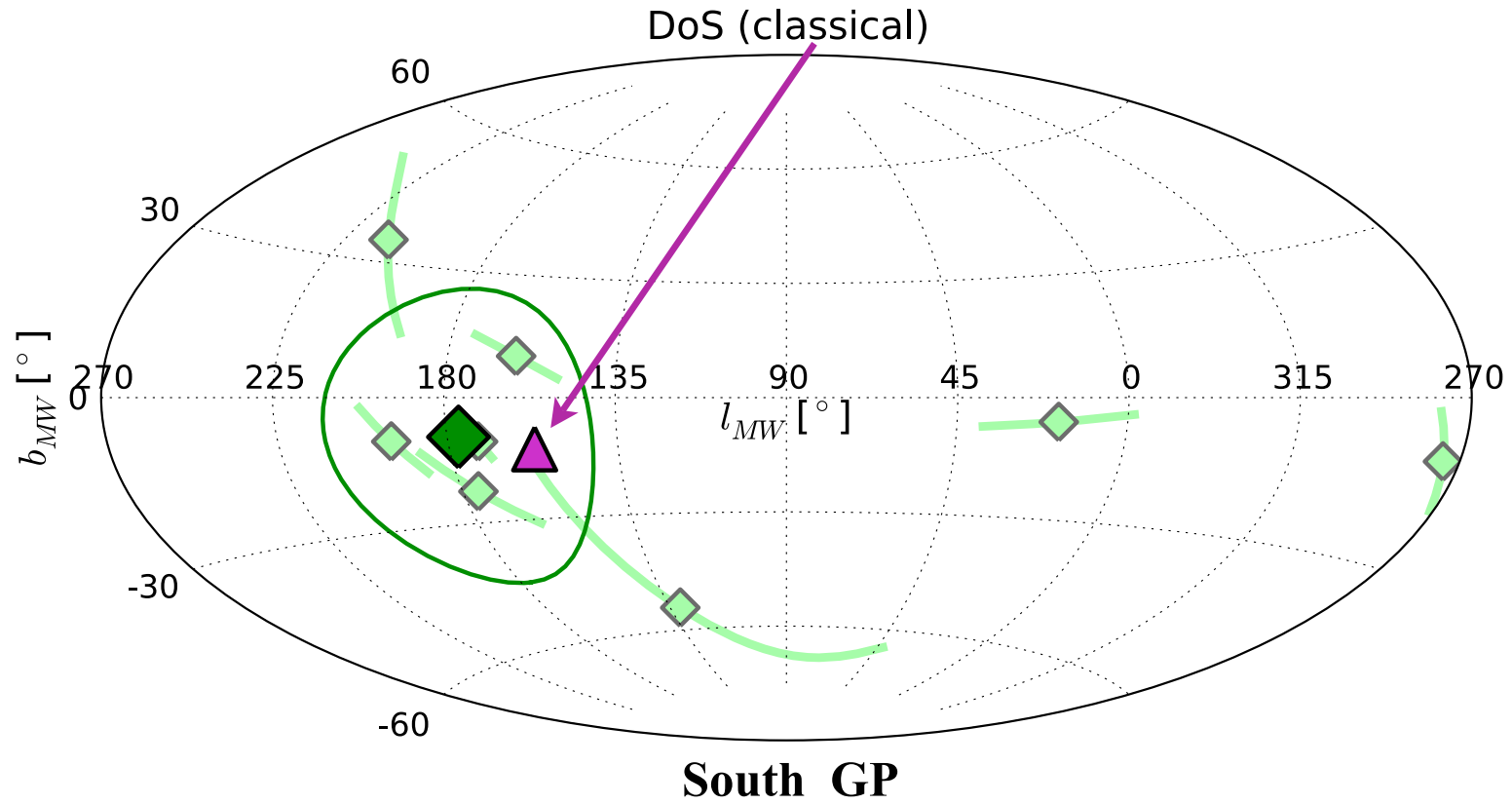
Kroupa et al. (2010, A&A)



Pavel Kroupa: AIfA, University of Bonn

the Galactic sky

(Galactic spherical coordinates)

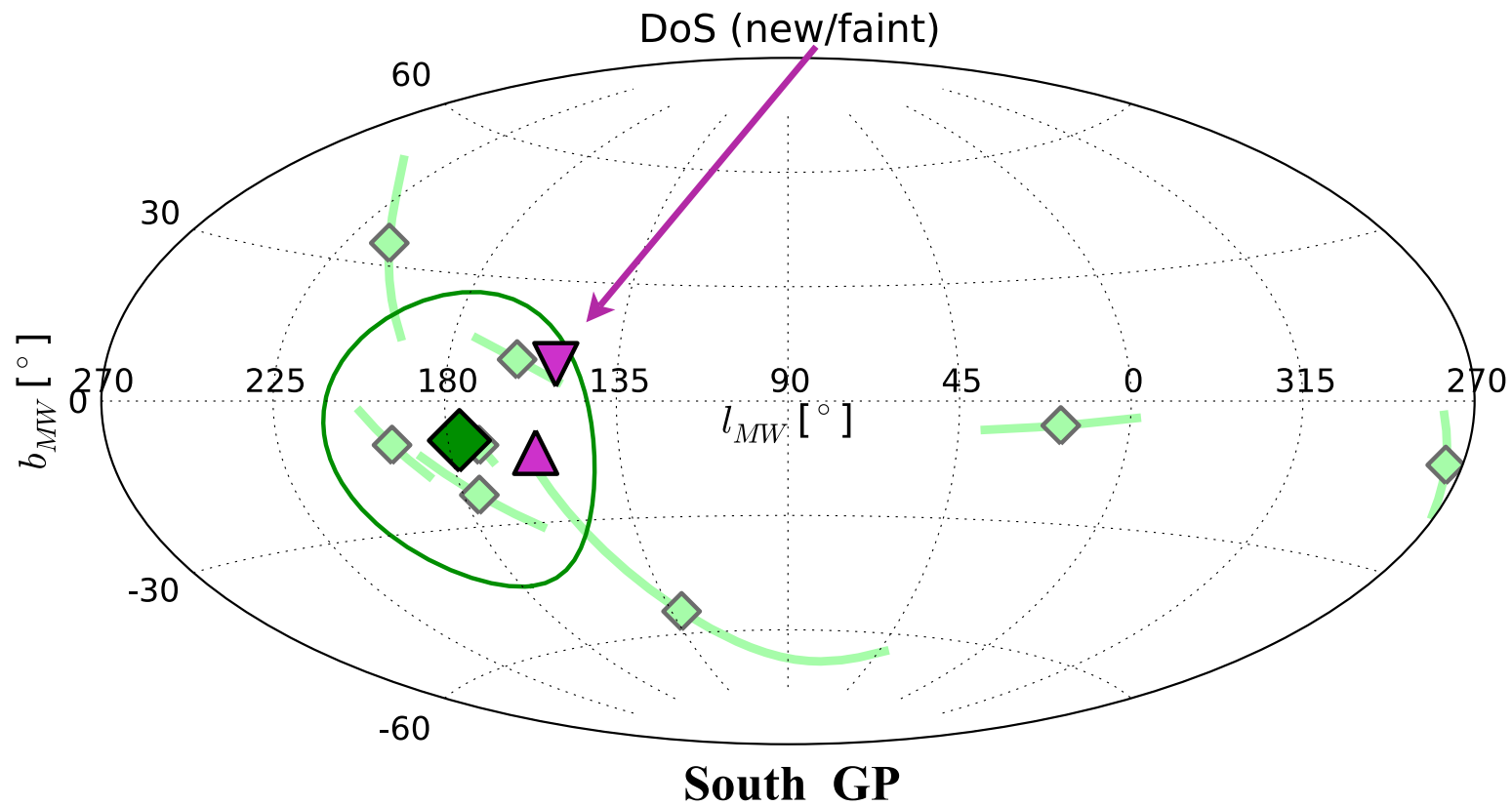


Pawlowski et al. 2011

Pavel Kroupa: AlfA, University of Bonn

the Galactic sky

(Galactic spherical coordinates)



Pawlowski et al. 2011

Pavel Kroupa: AIfA, University of Bonn

This
correlated phase-space
population
is *inconsistent* with
the satellites being
dark-matter sub-haloes
that fell into the MW halo
in a group or individually.

~~dark-matter halos~~

Logical inconsistencies
within
LCDM framework

Deason et al. (2011, MNRAS) (abstract):

To get the DoS:

"The satellite galaxies have been accreted relatively recently"

Nichols & Bland-Hawthorn (2011, ApJ) ignore the DoS (abstract):

Get gas-poor dSph around the Galaxy and M31:

"if the dwarfs fell in at high redshifts ($z \sim 3-10$)."

All these independent arguments



The dark-matter ansatz fails.
(no consistent natural solution
has come forth)

Remember the two implications of Fritz Zwicky's two conjectures :

~~1. There exist large numbers of dark-matter dominated satellite galaxies~~

2. There exist large numbers of newly formed (tidal-dwarf) satellite galaxies.

No, there is *only one type* of dwarf galaxy !

But, *which* one ? And *why* only one ?

Lets consider now the
tidal-dwarf type satellite
dwarf galaxy :

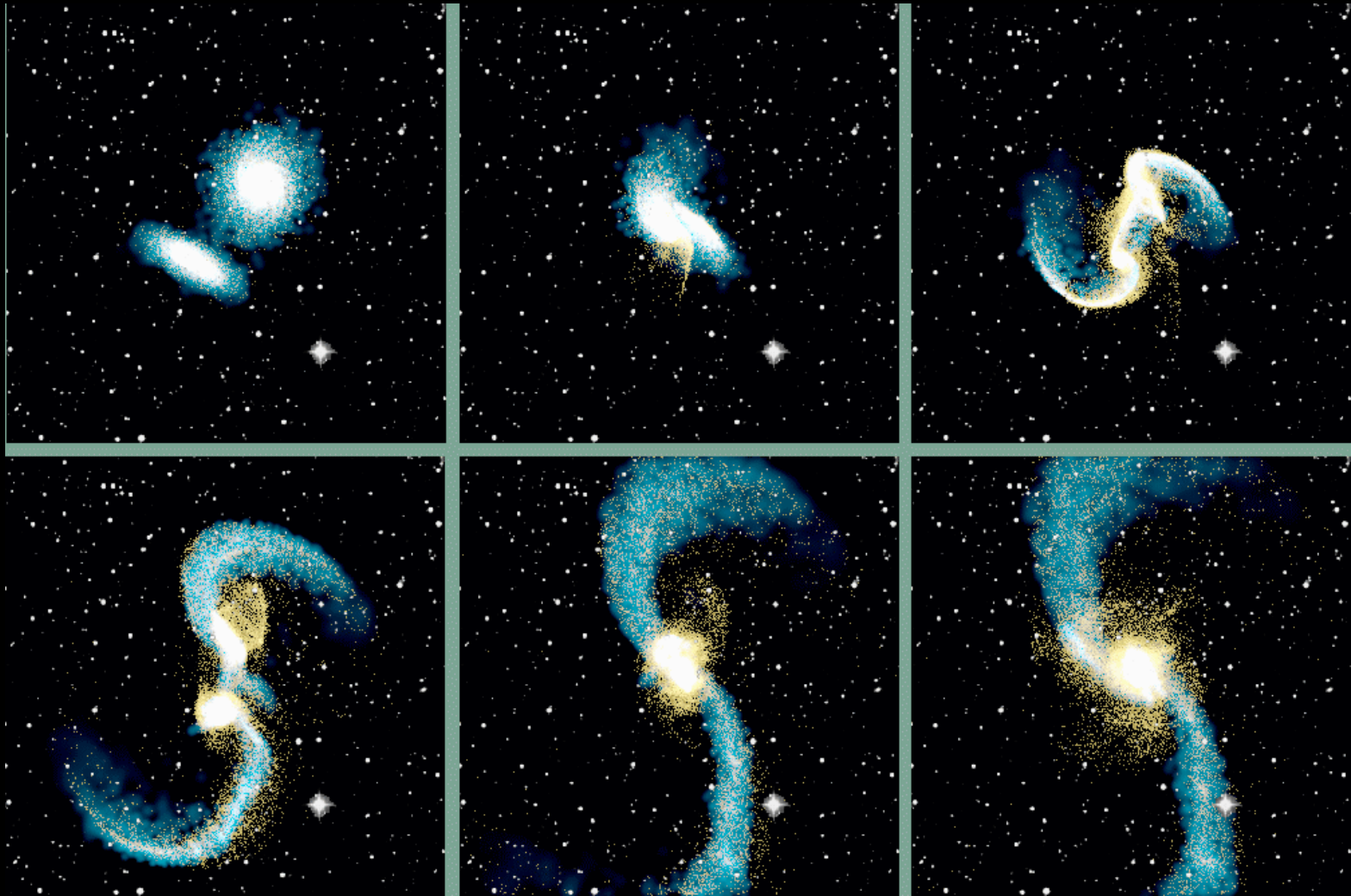
Tidal
dwarf
galaxies.

Pioneering
work by

Felix Mirabel,
Pierre-Alain Duc,
Frederic Bournaud,
Francoise Combes,
Olivier Tiret
=
"The French"

And see also
Barnes & Hernquist;
Elmegreen;
Wetzstein

Tidal tails



Miho & Maxwell, web

Pavel Kroupa: AIfA, University of Bonn



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(Weilbacher et al. 2000)

$$N_{\text{TDG}} \approx 14$$

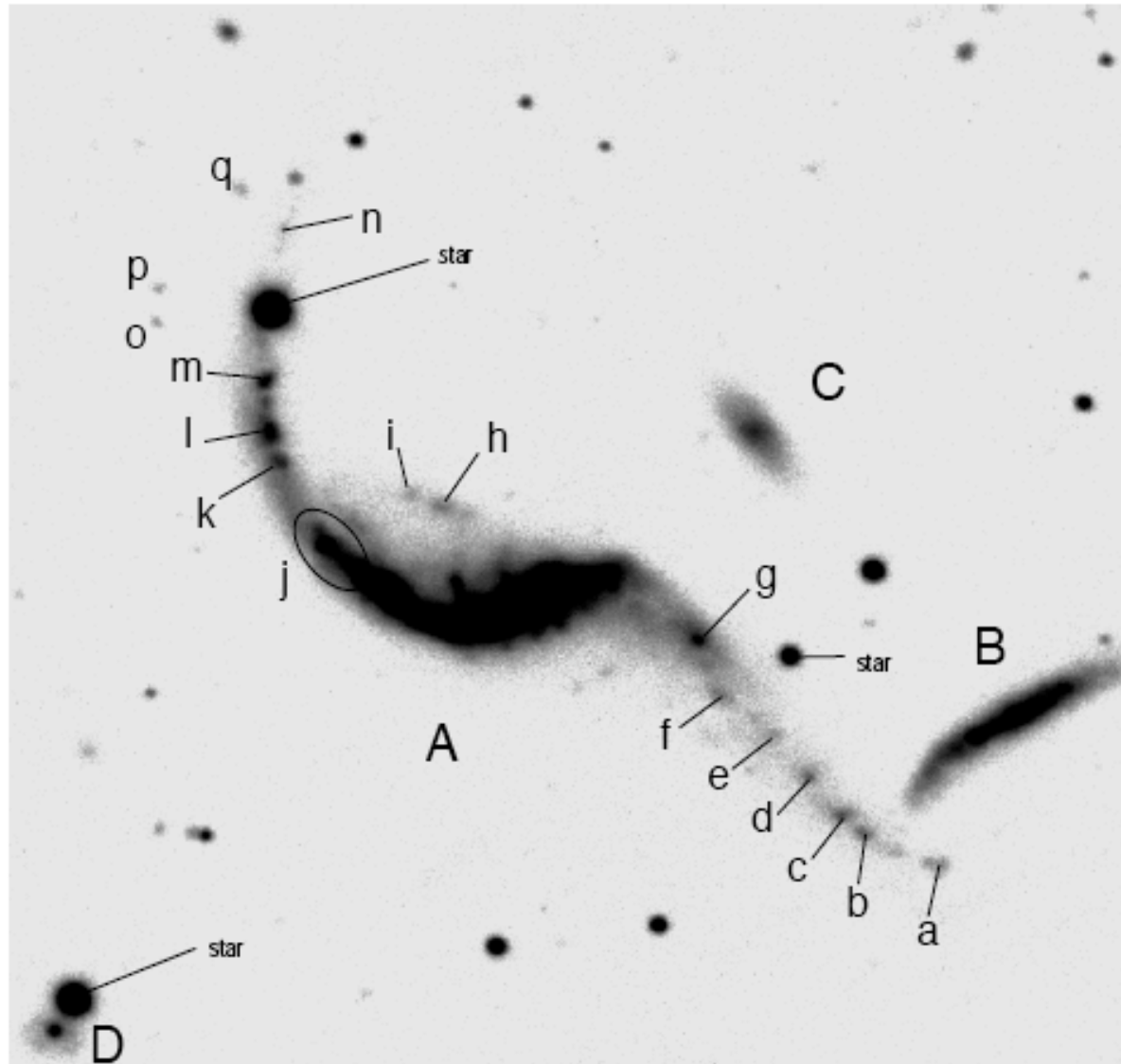
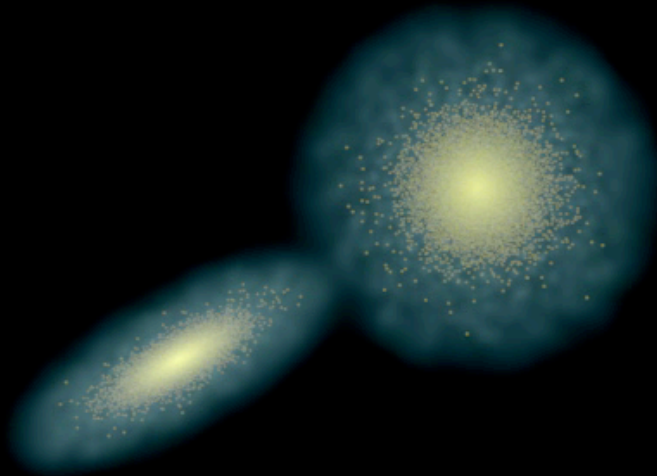


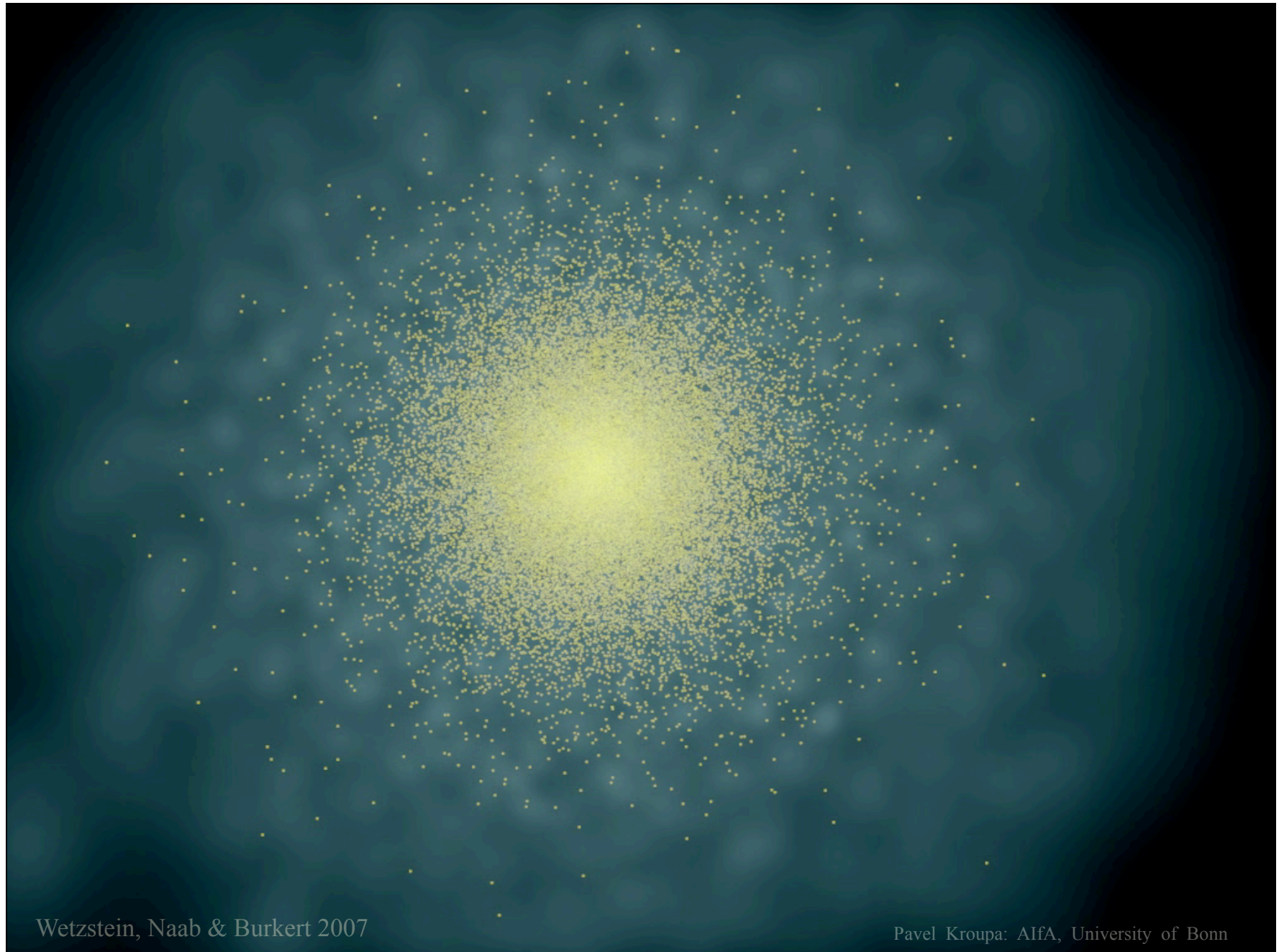
Fig. 21. Identification chart of field 10 around AM 1353-272.

Relevance : The collision of two disks at high redshift



Wetzstein, Naab & Burkert 2007

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Wetzstein, Naab & Burkert 2007

Pavel Kroupa: AIfA, University of Bonn

Thus, by direct observation *new dwarf galaxies* with masses comparable to dE/dSph galaxies form like *shrapnel*.

They are baryon dominated

(Barnes & Hernquist 1992).

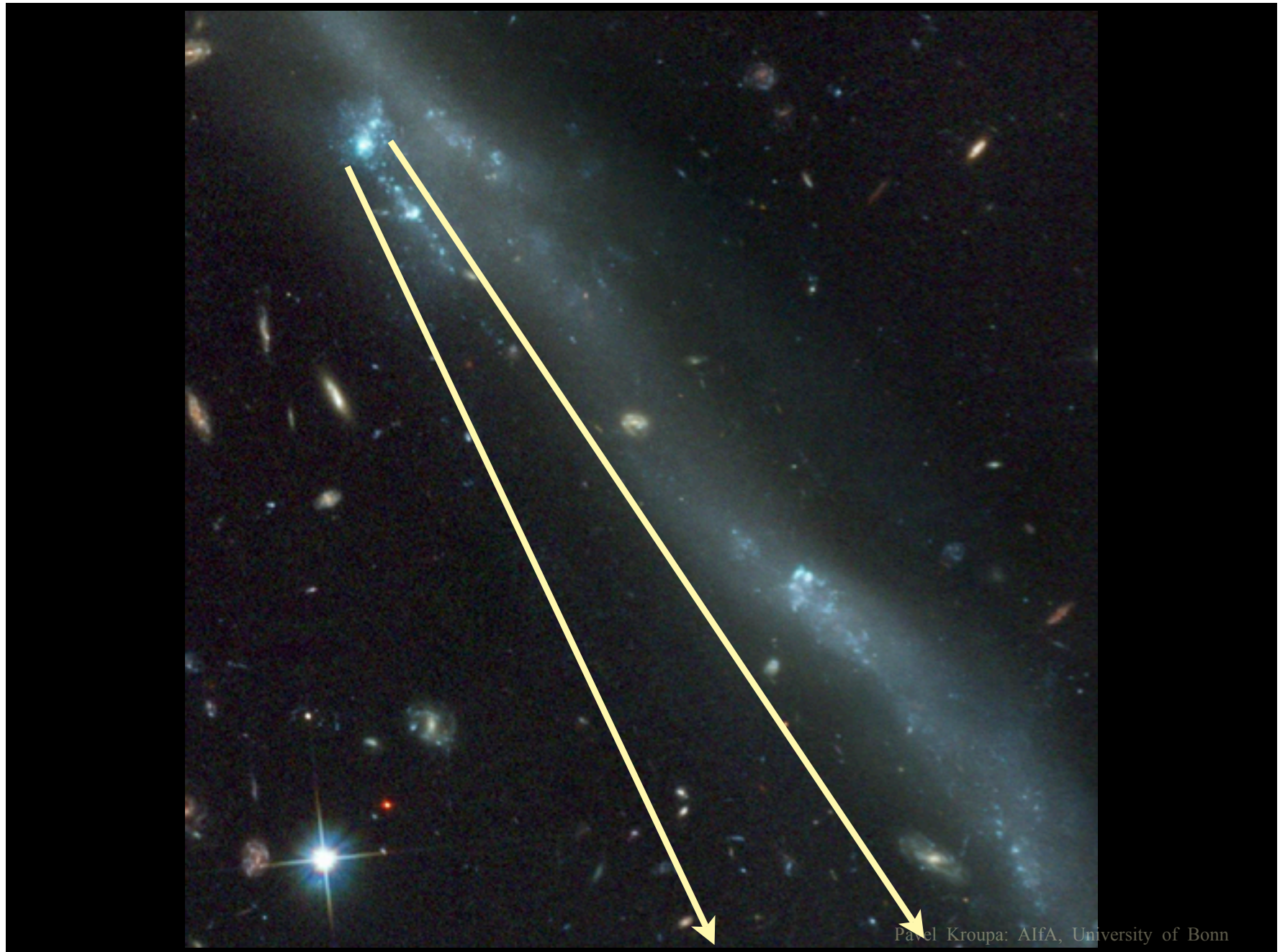
Evolution of TDGs

The Time 



Evolve dwarf galaxies
w/o dark matter
in a computer

Pavel Kroupa: AIfA, University of Bonn

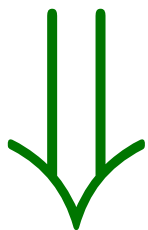


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Star-cluster complex
(cf *Tadpole*)

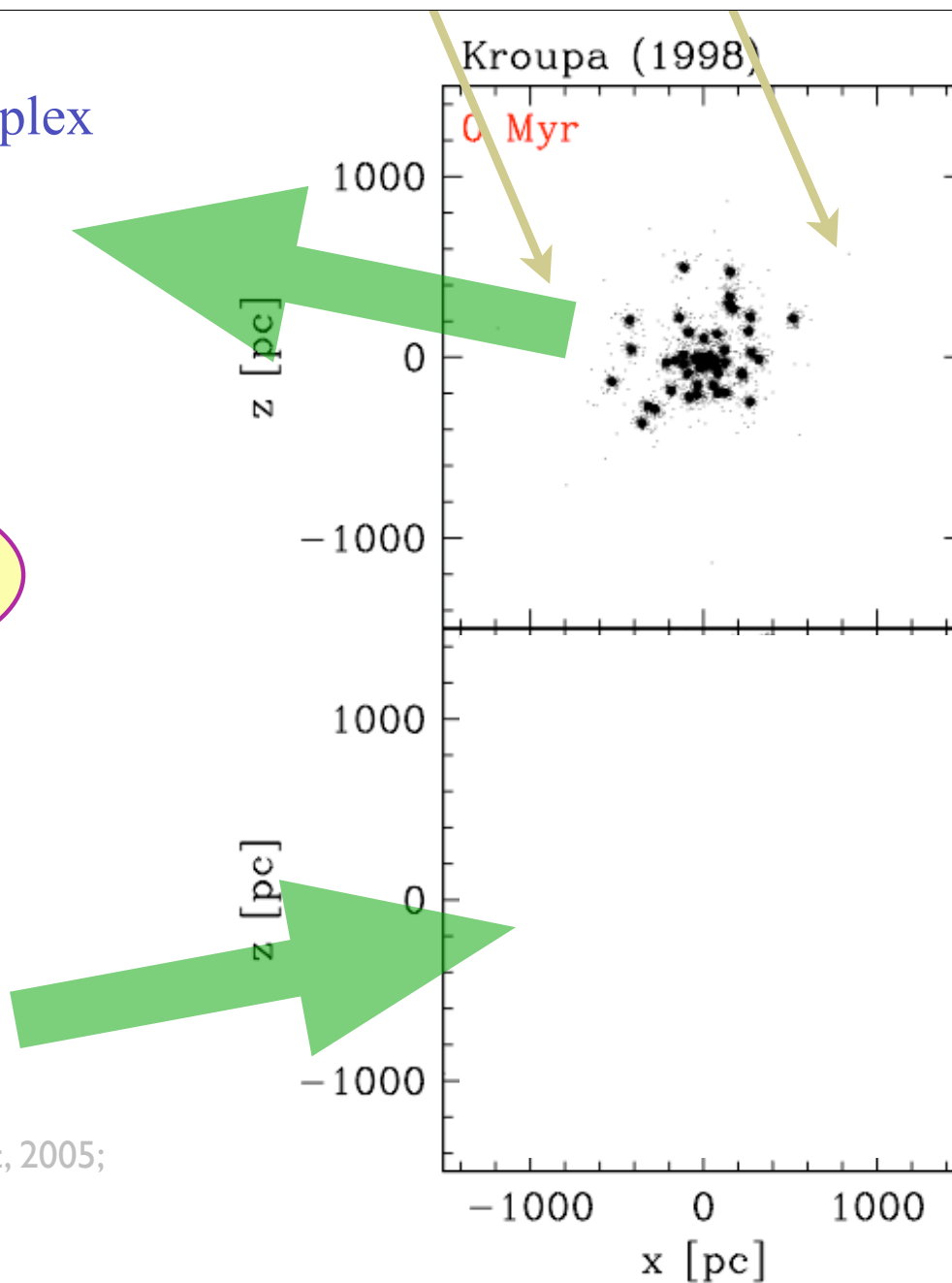
50 clusters
each $10^6 M_{\odot}$

clusters as
*fundamental galactic
building blocks*



Spheroidal dwarf
galaxy !

(Fellhauer et al. 2001, 2002a,b,c, 2005;
Bekki et al. 2004)



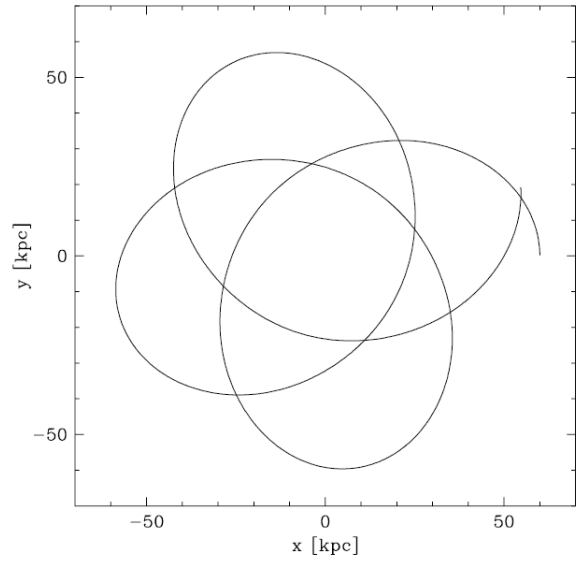
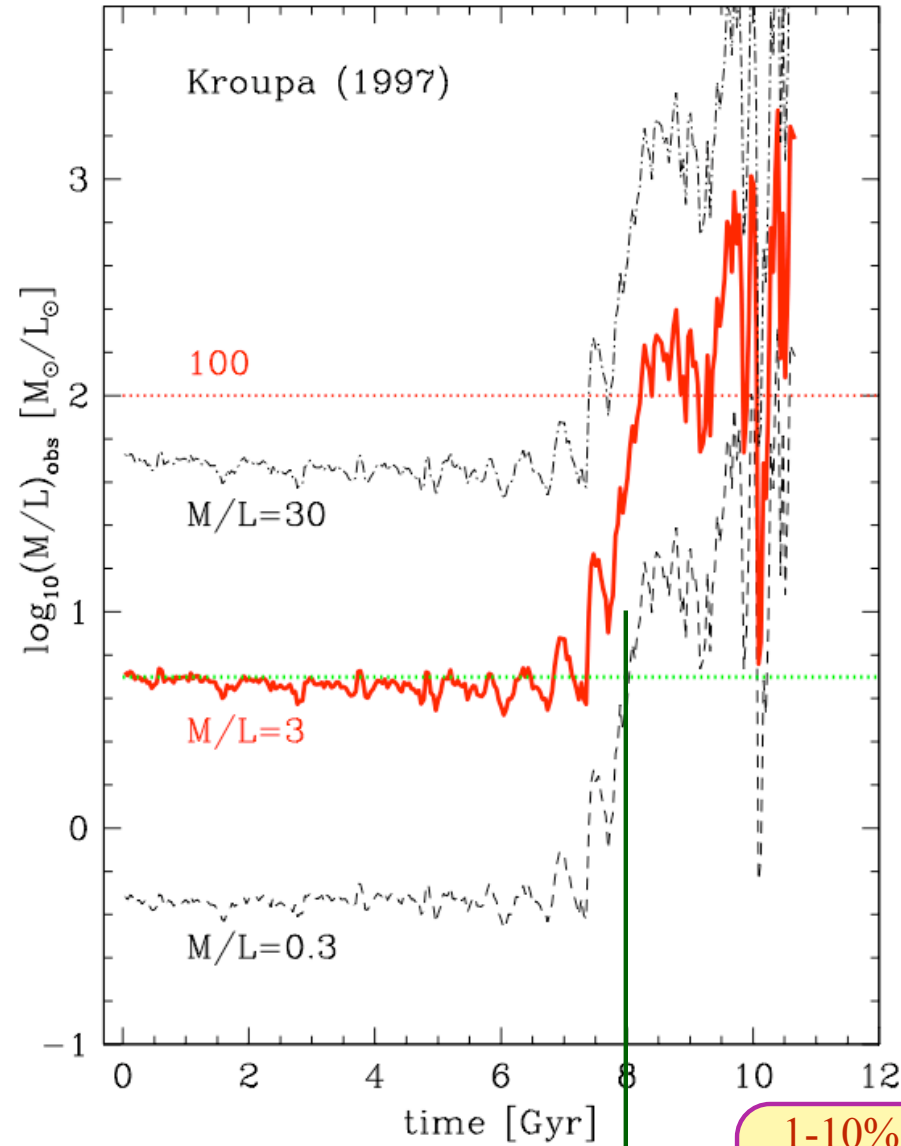


FIG. 3.—Orbital path of the satellite in simulations RS1-113 and Sat-M2.



1-10% of population in remnant

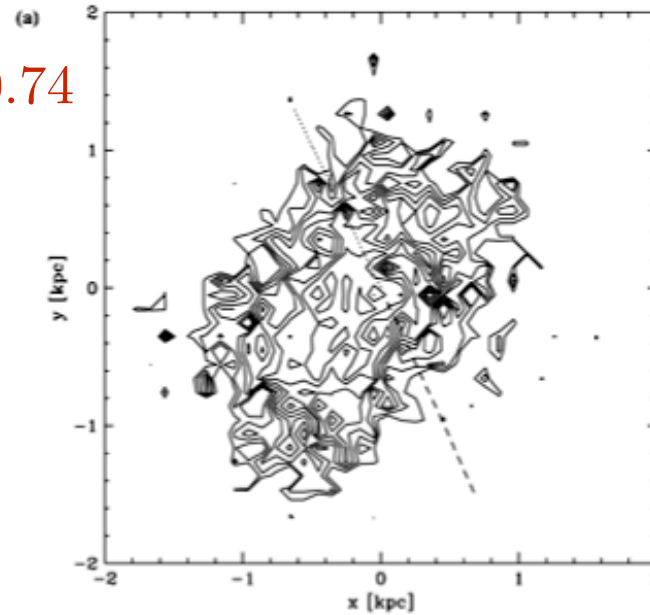
(Kroupa 1997)

Remnants have a highly
anisotropic $f(\mathbf{R}, V)$
and mass $\approx 10^5 M_{\odot}$

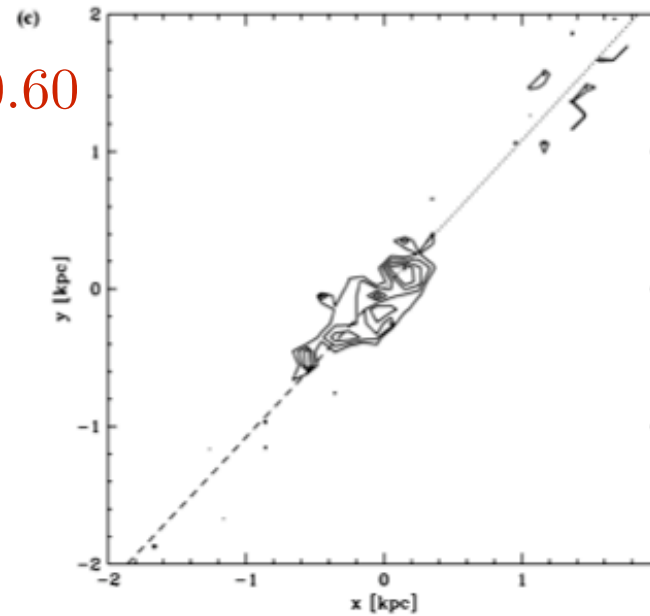
$R \approx$ few 100 pc

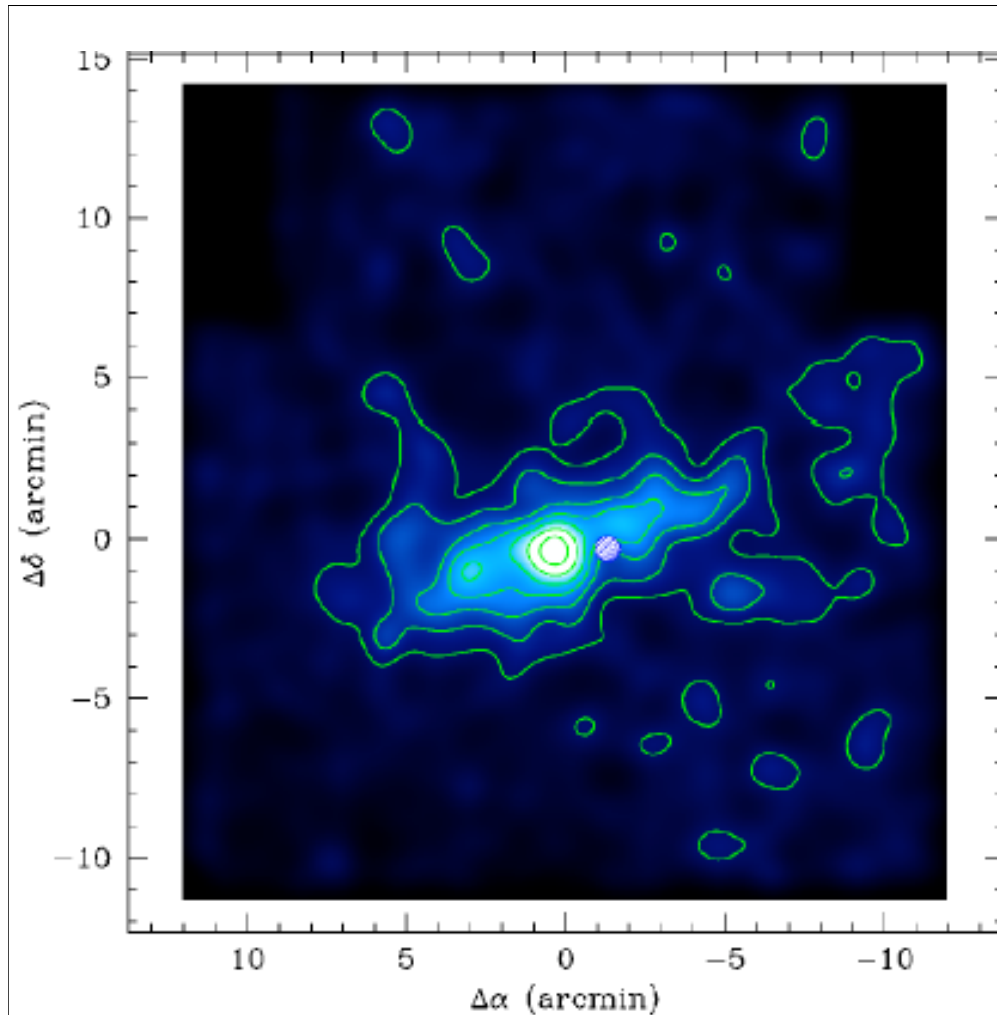
and $\frac{M}{L} \approx 10^{2-3}!$

$e = 0.74$



$e = 0.60$





Hercules

$D=130\text{kpc}$

(Coleman et al. 2007)

TABLE 1
PROPERTIES OF THE HERCULES DSPH

Parameter	Value
R.A. (J2000)	16:31:02.0
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Distance (kpc)	132 ± 12
[Fe/H]	-2.1 ± 0.2
Age (Gyr)	13 ± 3
King r_h	$4.37' \pm 0.29'$ (168 ± 11 pc)
King r_c	$4.74' \pm 0.57'$ (182 ± 22 pc)
King r_t	$25.9' \pm 11.1'$ (994 ± 426 pc)
$c = \log(r_t/r_c)$	0.74 ± 0.25

Hercules

D=130kpc

(Coleman et al. 2007)

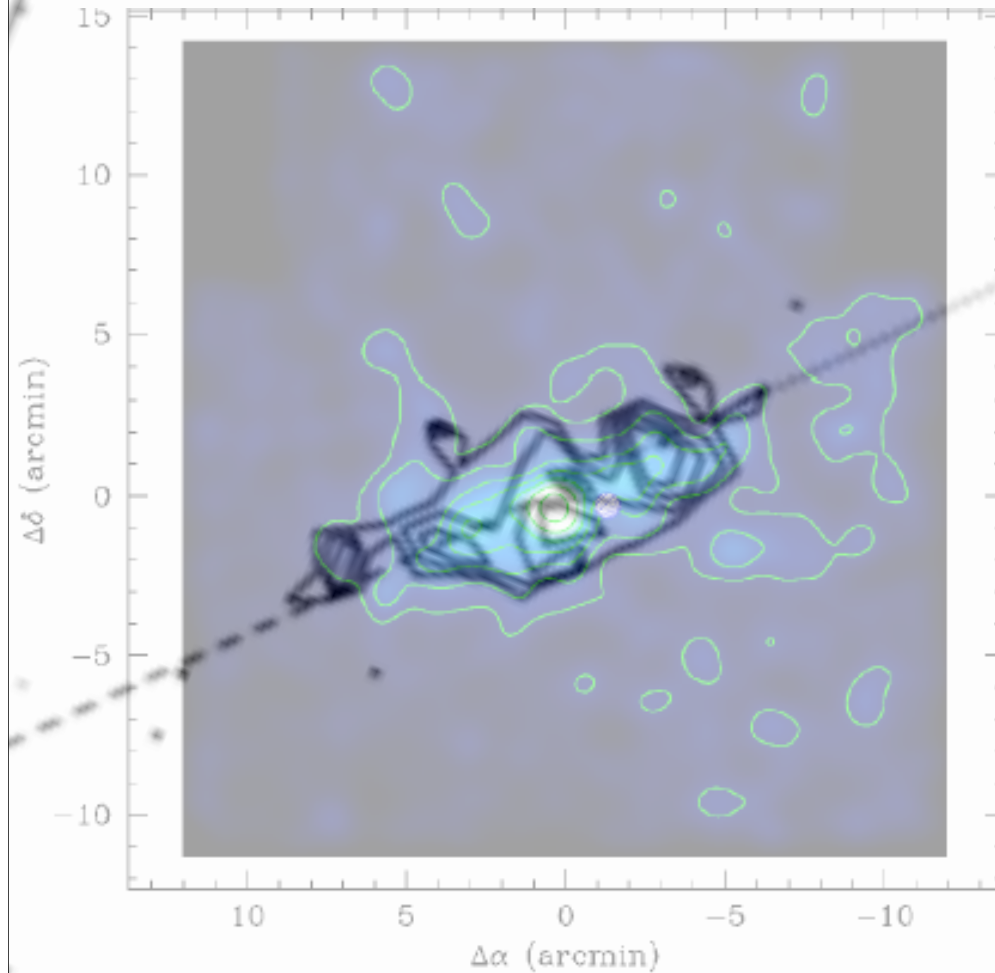
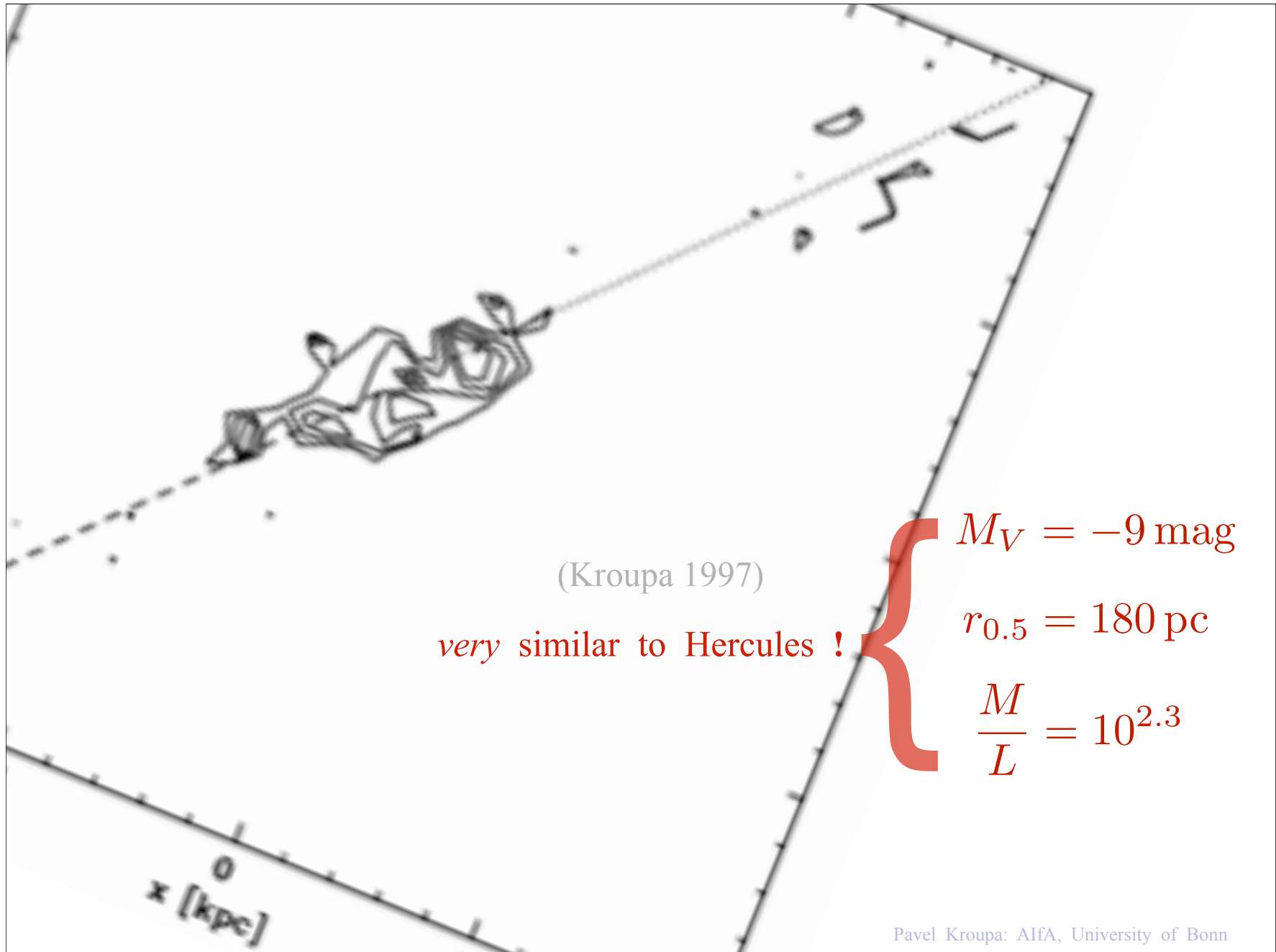


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Pavel Kroupa: AlfA, University of Bonn



This is a
real prediction
10 years before
the discovery
of this
type of celestial object !

For TDGs we know today that

The early (<100Myr) star-formation and chemical enrichment evolution is similar to the observed dSph satellites.

(Recchi et al. 2007)

Later dynamical evolution does not destroy the satellites.

(Kroupa 1997)

The number of old TDGs amounts to the dE population observed.

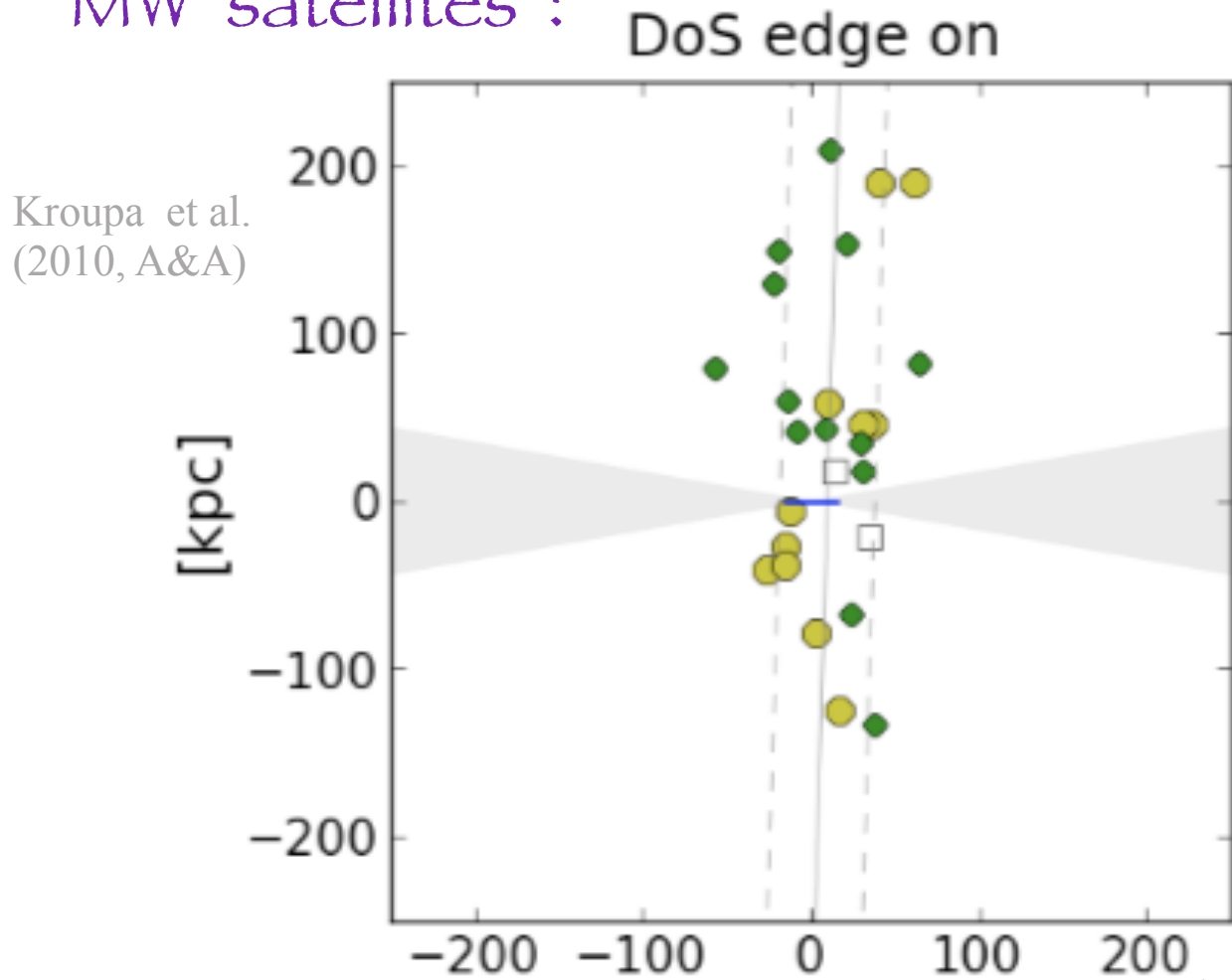
(Okazaki & Taniguchi 2000)

dE galaxies are observed to contain *no Dark Matter*, consistent with them being *TDGs*.

(Toloba et al. 2010, arXiv:1011.2198v1)

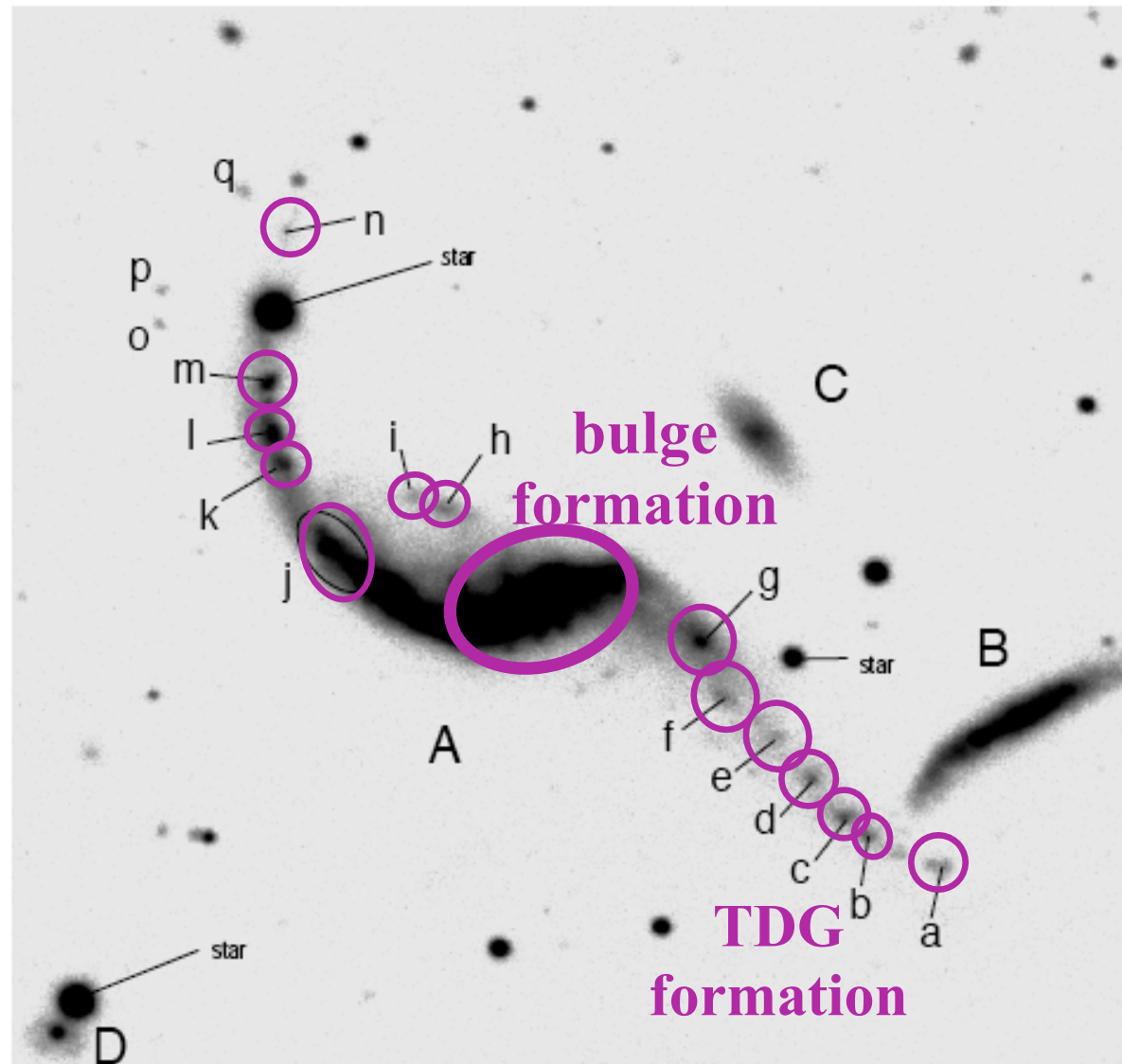
What about the disk-like configuration of MW satellites?

- the 11 “classical” (brightest) satellites
- new satellites



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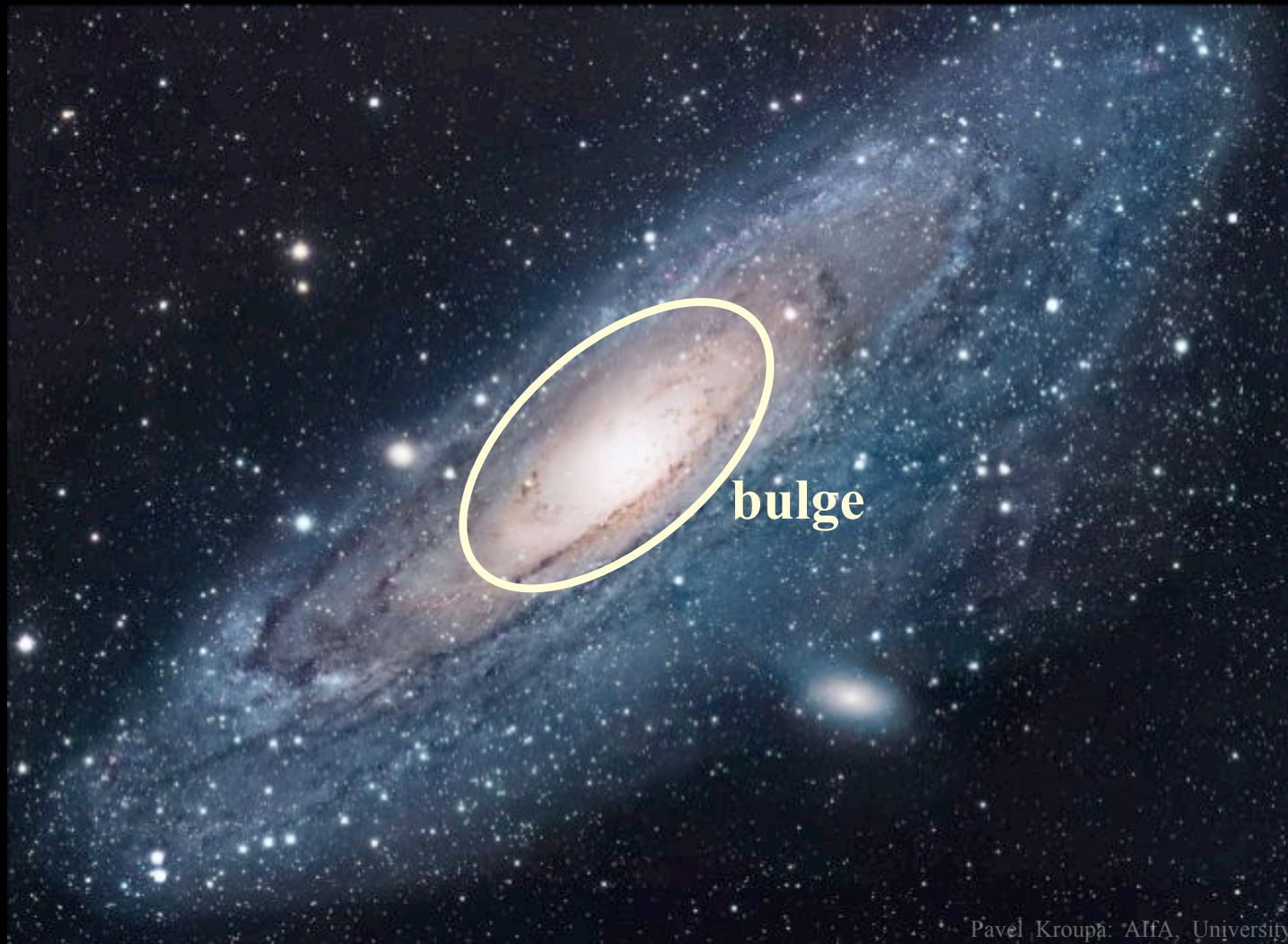
(Weilbacher et al. 2000)



*Phase-space
correlated
satellites form
naturally in the
same event
as a bulge does.*

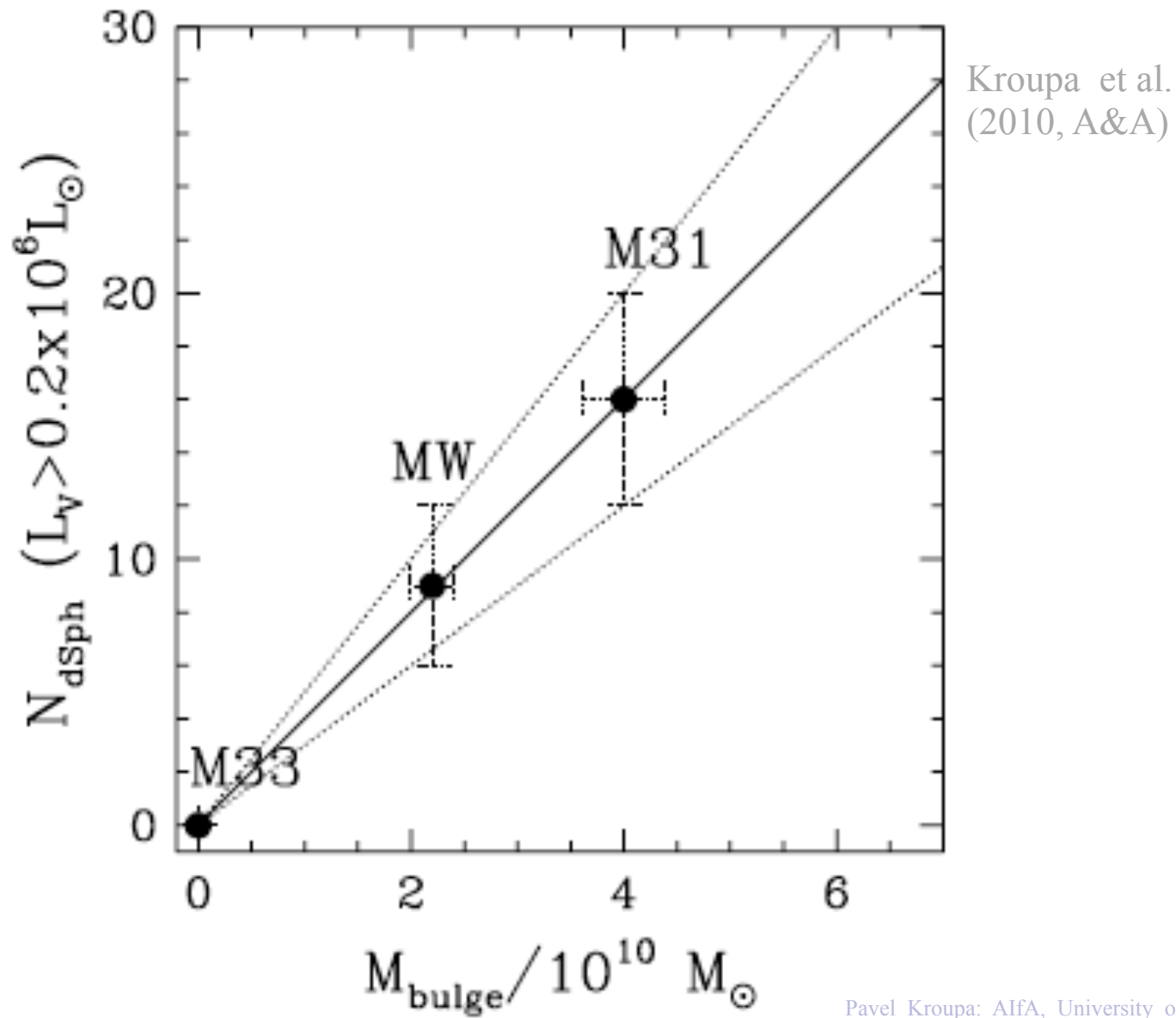
Fig. 21. Identification chart of field 10 around AM 1353-272.

... and, a bulge mass vs number of satellites correlation?



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A bulge - satellite correlation



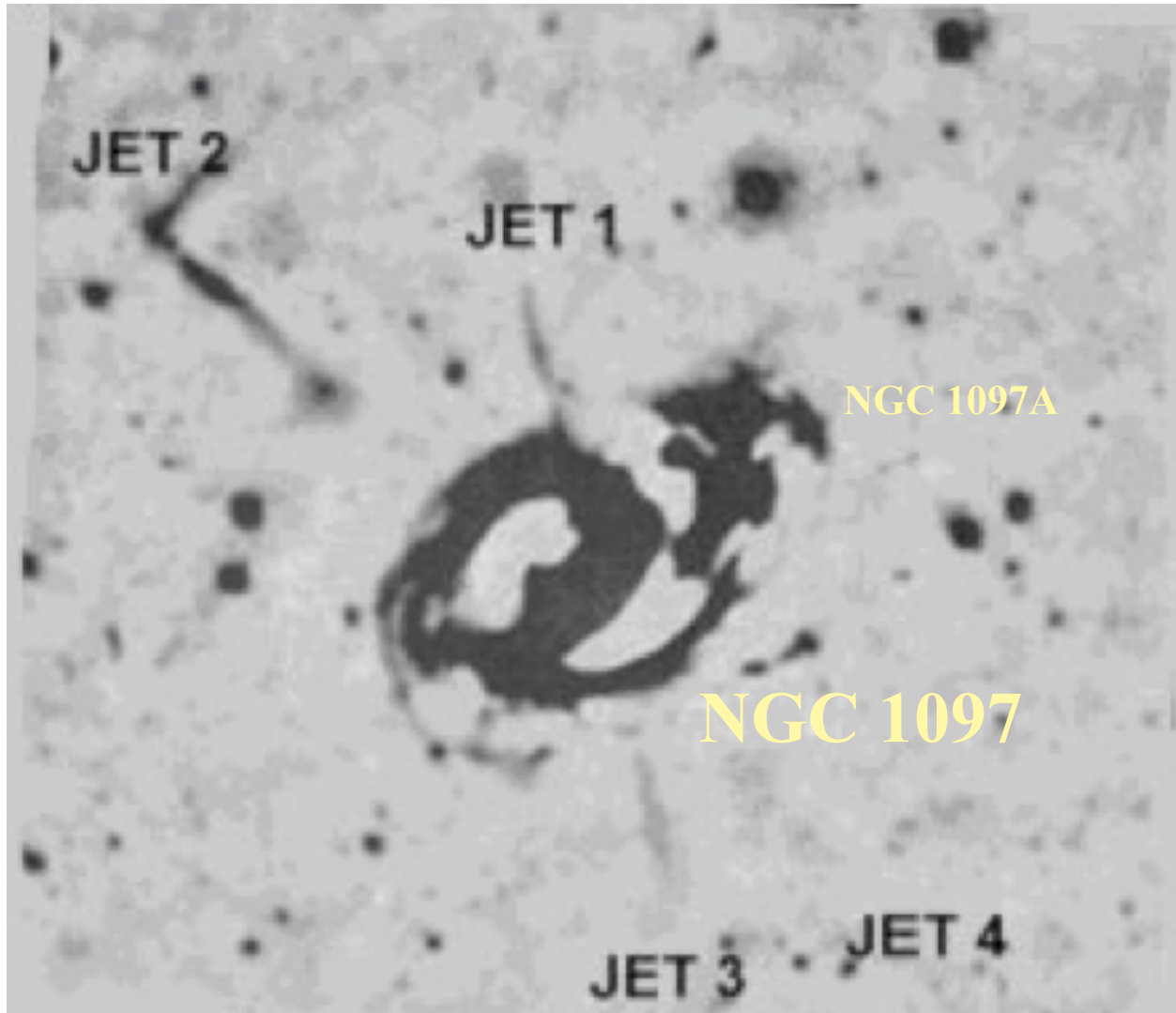
Both,
the *Disk of Satellites*
and
the *bulge--satellite correlation*
are naturally understandable
if the MW satellites are *ancient TDGs*.

(Kroupa et al. 2010)

Other extragalactic
correlated
dSph satellite systems

NGC 1097's "dog-leg" tidal stream

Galianni, Patat et al.
(2010 A&A)

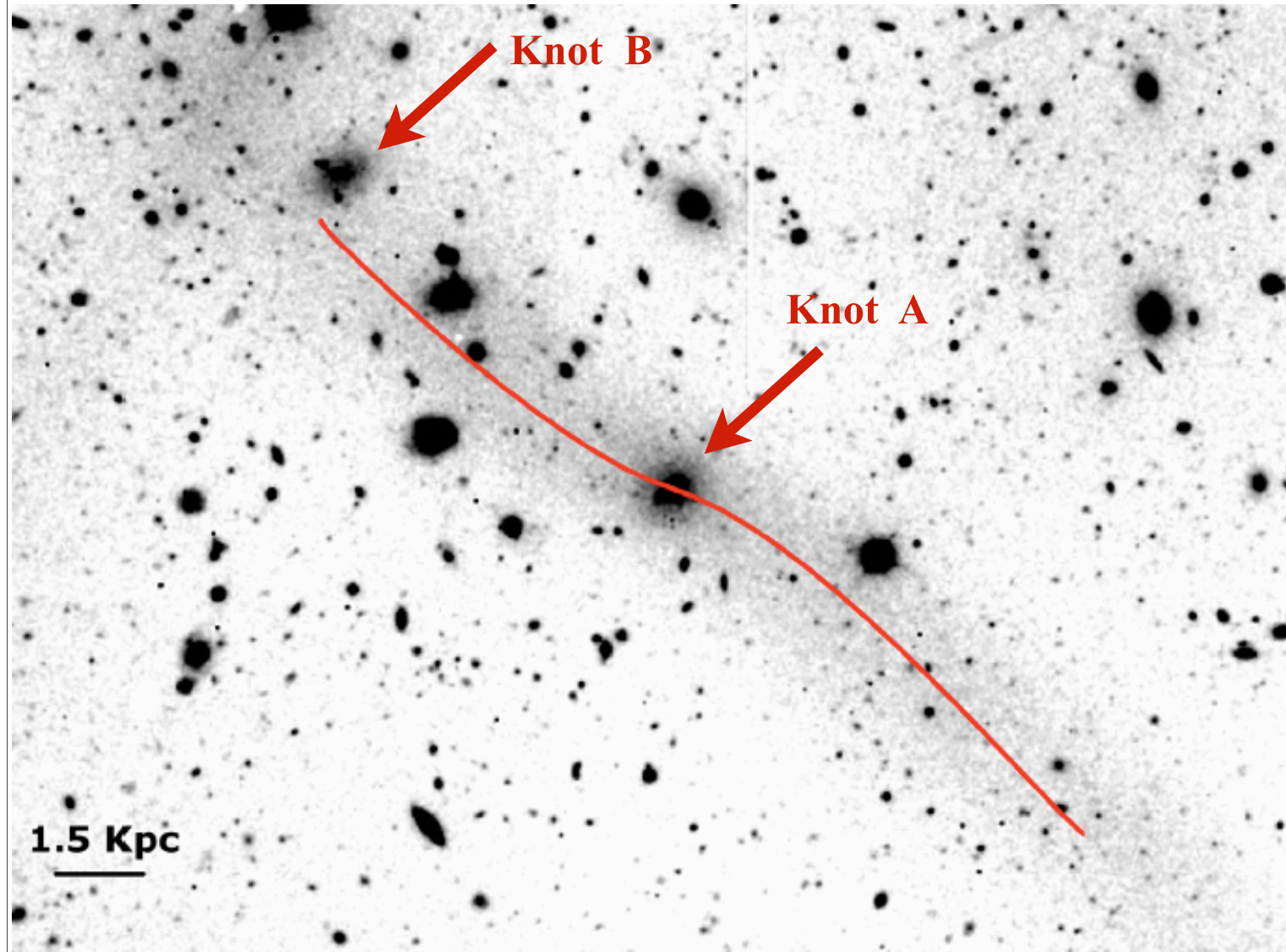


"Jet2" = dog leg
is a stellar stream

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NGC 1097's “dog-leg” tidal stream

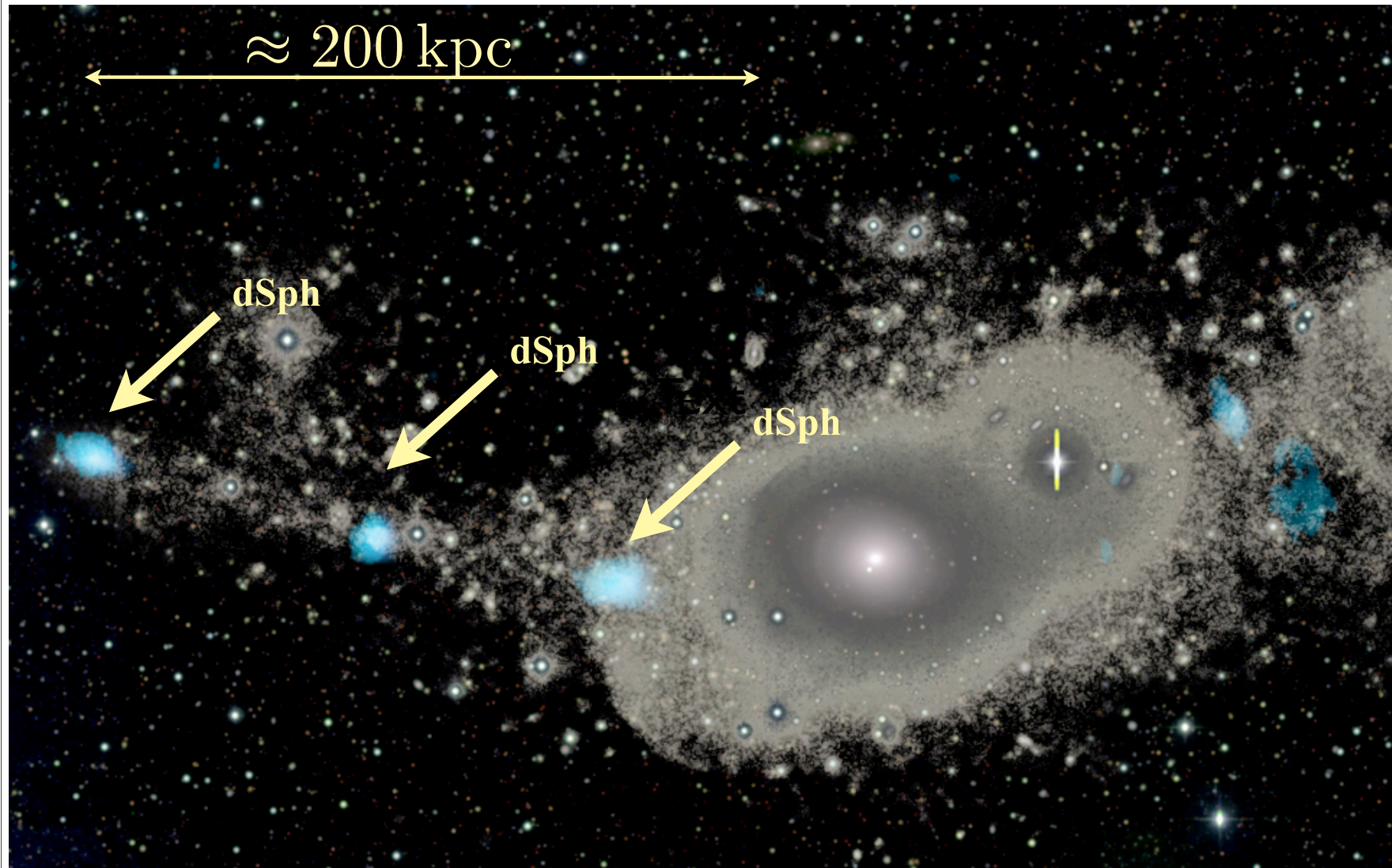
Galianni, Patat et al.
(2010 A&A)



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NGC 5557 (post-merger 2-3 Gyr)

Duc et al. (2011
MNRAS)



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The
observed and theoretical
TDGs
coincide
with dE galaxies !

Thus,

(1) a *fully self-consistent TDG scenario* thus emerges which very naturally accounts for the properties of dE and satellite galaxies;

(2) *no consistent*, and in fact a *contradictory picture* emerges in the *dark-matter framework*;

(3) there is simply *no evidence for* the existence of *DM satellites*.

Dark matter galaxies

vs

baryonic ones :

why so similar ?

The Tully-Fisher Relation :

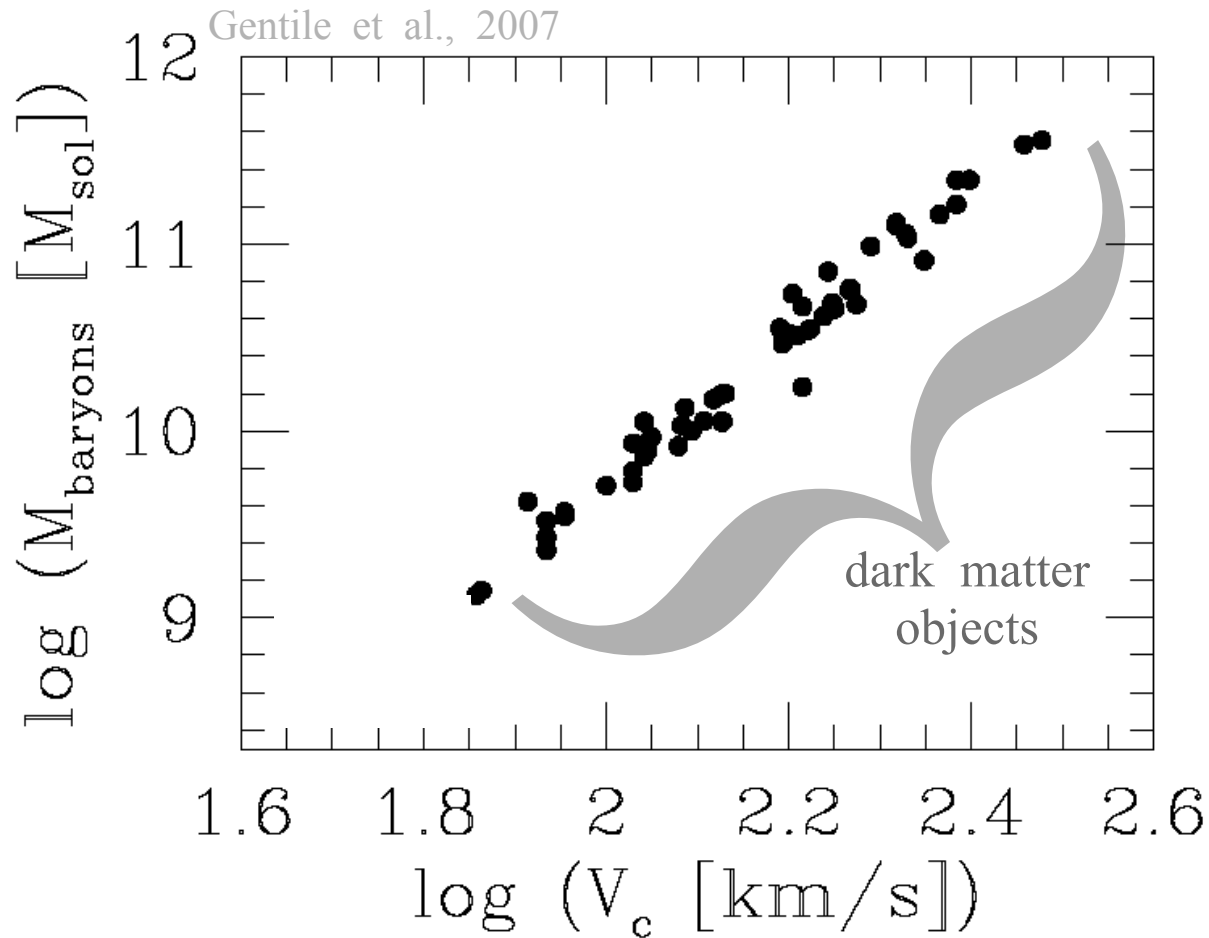
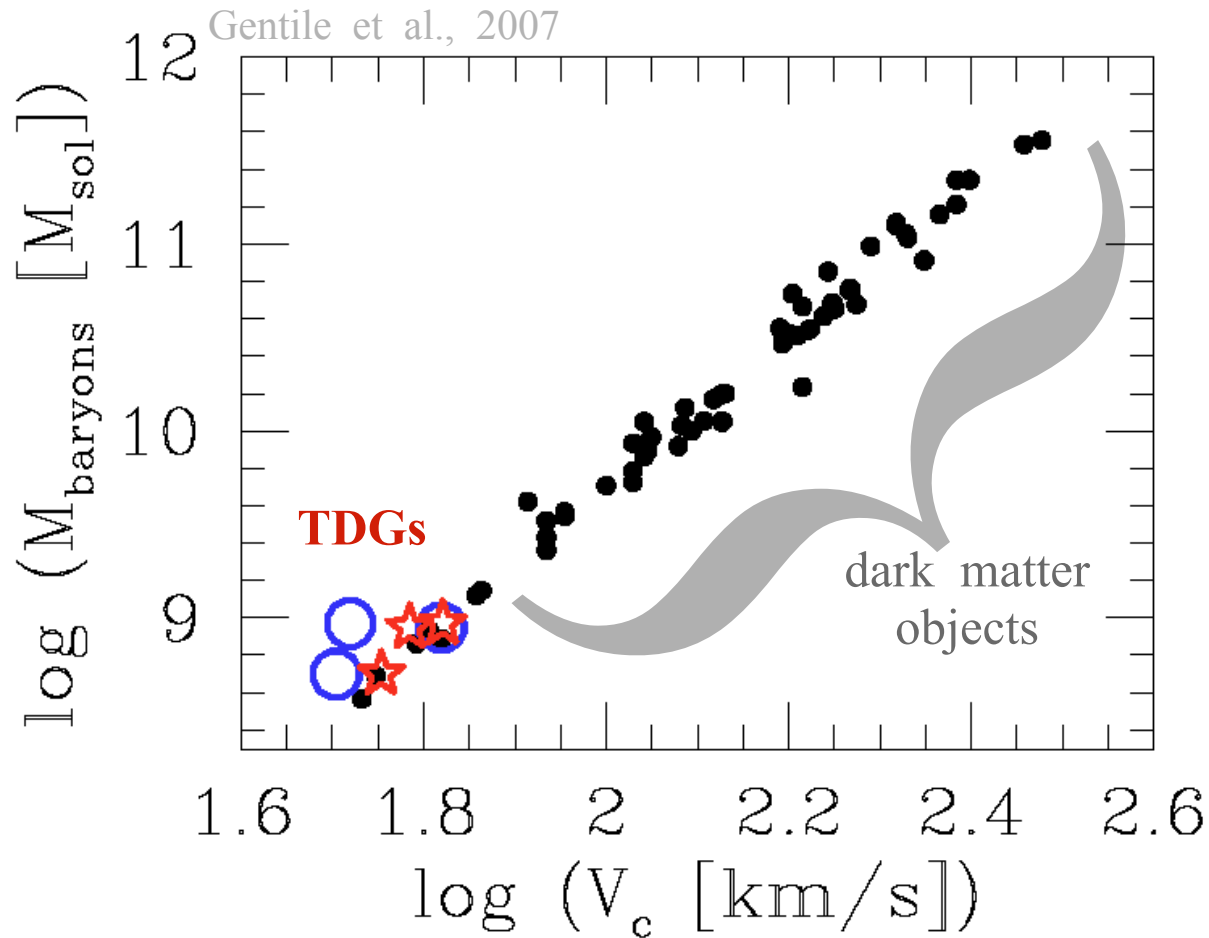


Fig. 2. Baryonic Tully-Fisher relation (baryonic mass vs. circular velocity). The small full circles are the disk galaxy data from McGaugh (2005). The 3 TDGs studied here are shown as blue empty circles ($i = 45^\circ$) and red stars (i as a free parameter).

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The Tully -Fisher Relation and internal dynamics of young TDGs :

But TDGs are purely baryonic objects.

So, why are they so similar to the other galaxies ?

Fig. 2. Baryonic Tully-Fisher relation (baryonic mass vs. circular velocity). The small full circles are the disk galaxy data from McGaugh (2005). The 3 TDGs studied here are shown as blue empty circles ($i = 45^\circ$) and red stars (i as a free parameter).

...

Summary & Conclusions

...

The two Fritz Zwicky Conjectures :

A. DM dominates large galaxies
(*postulate* from kinematics).

B. New dwarf galaxies form in tidal
debris (*direct observational fact*).



CONTRADICTION

because

1. There *must be* many satellite
galaxies with massive dark matter
halos (*type A dwarfs*).

2. there *must be* many (tidal -
dwarf) satellite galaxies without dark
matter (*type B dwarfs*).

But

only one type of dE / dSph galaxy
is observed to exist.

This can only be *type B* because this type is a
consequence of well-known physical laws.

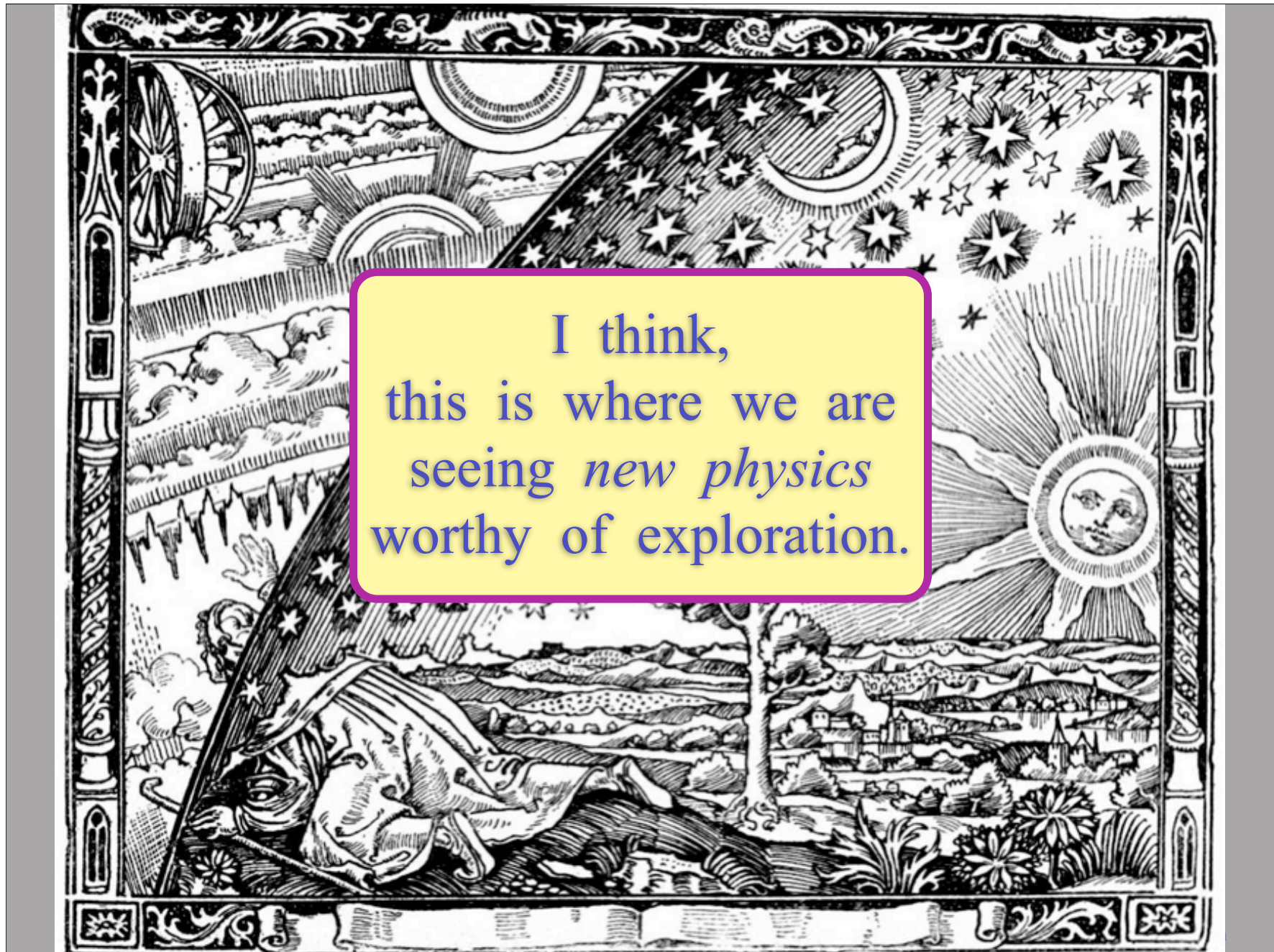
The DoS and bulge-satellite correlation are
additional, independent, supporting evidence.



DISCARD

Conjecture A.

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I think,
this is where we are
seeing *new physics*
worthy of exploration.

dSph = TDG \Rightarrow ~~LCDM~~

The END