



The Globular Cluster - UCD - Dwarf Galaxy Connection

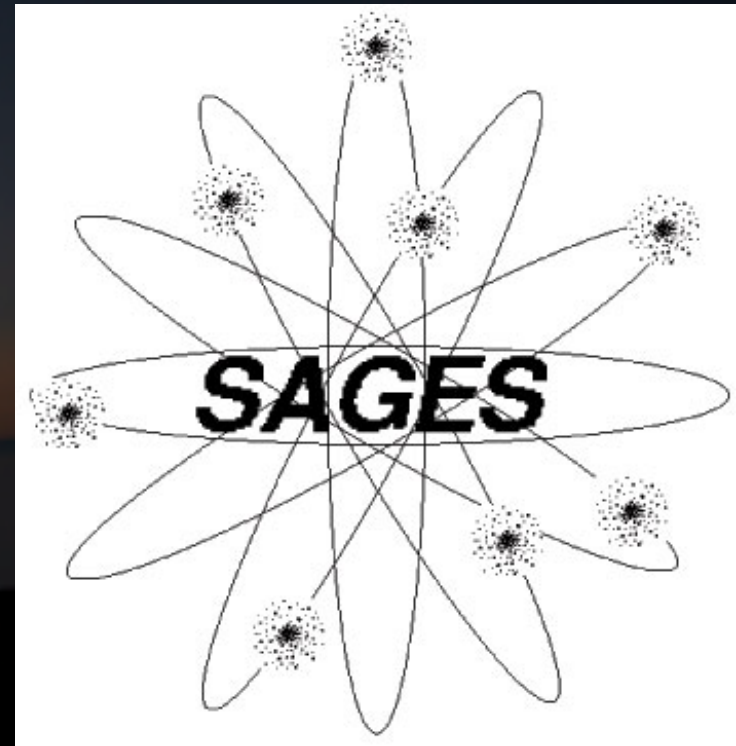
What is a Galaxy?
Voting Results

DUNCAN FORBES

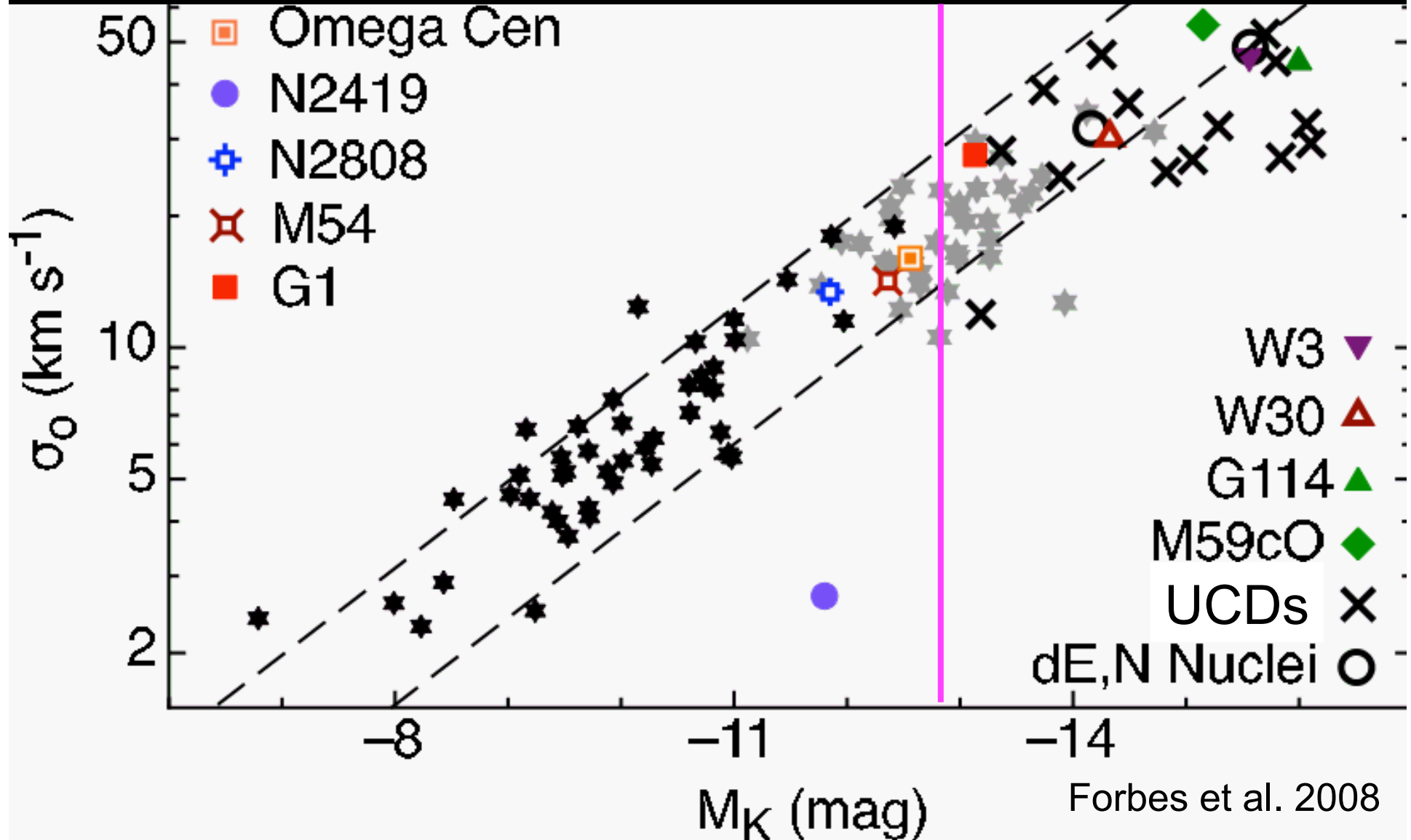
SWINBURNE UNIVERSITY

Collaborators

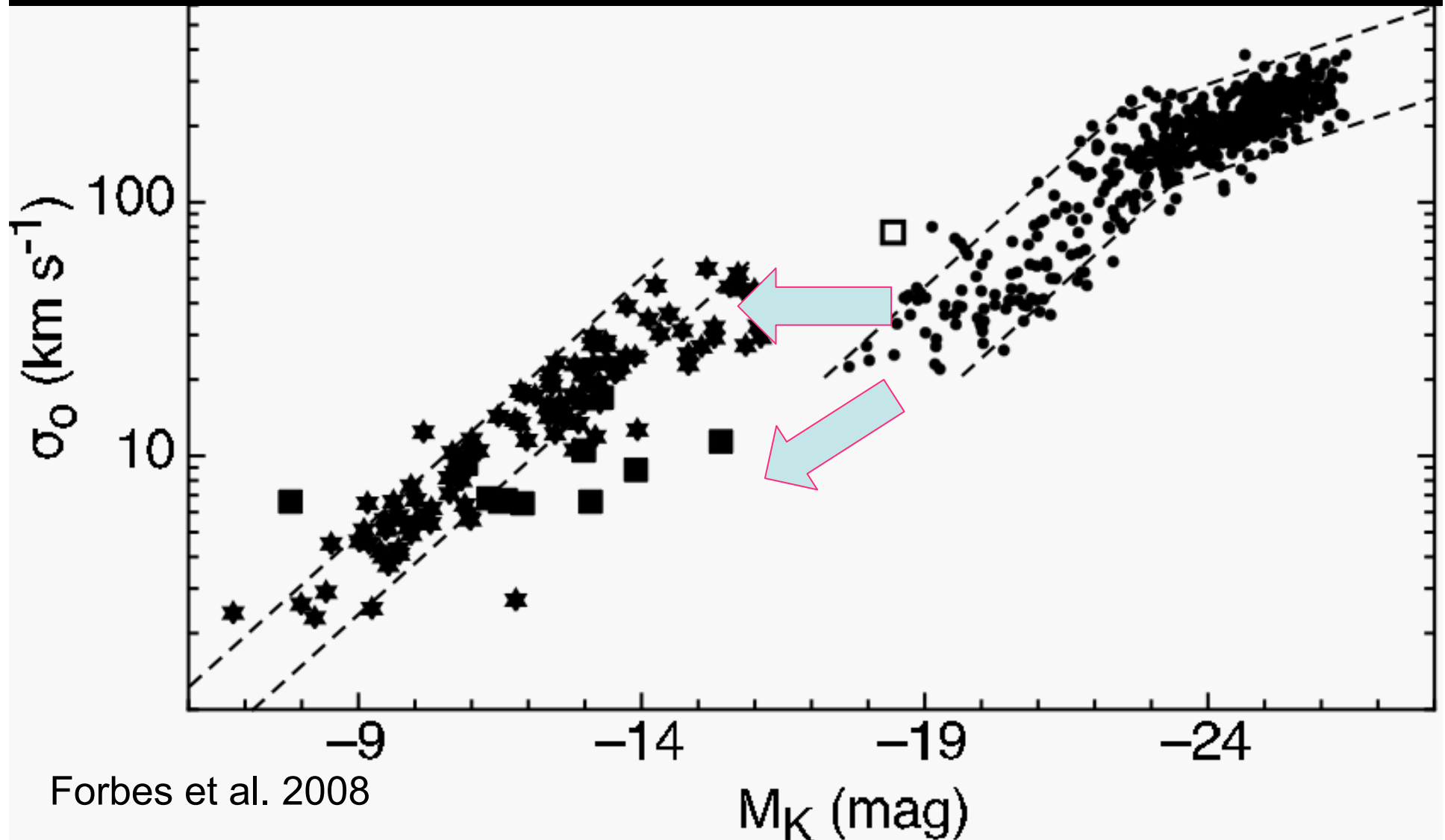
- Jean Brodie (UCSC)
 - Aaron Romanowsky (UCSC)
 - Jay Strader (Harvard)
 - Soeren Larsen (Utrecht)
 - Vincenzo Pota (Swinburne)
 - Chris Usher (Swinburne)
-
- Caroline Foster (Swinburne)
 - Lee Spitler (Swinburne)
 - Alister Graham (Swinburne)
 - Andrew Benson (Caltech), George Hau (ESO)



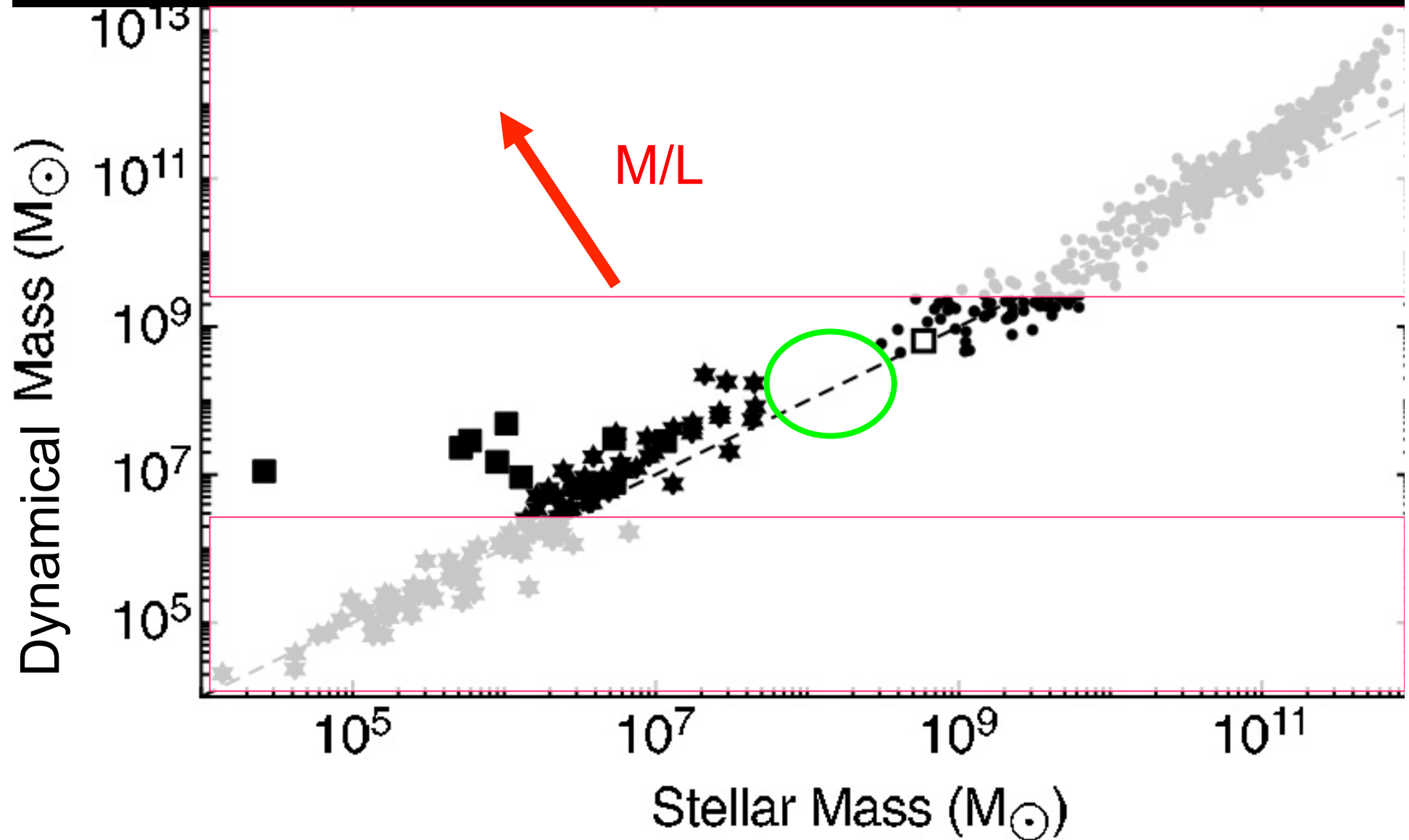
Velocity dispersion vs M_K



Velocity dispersion vs M_K



Dynamical mass vs stellar mass



Dwarf elliptical velocity dispersions

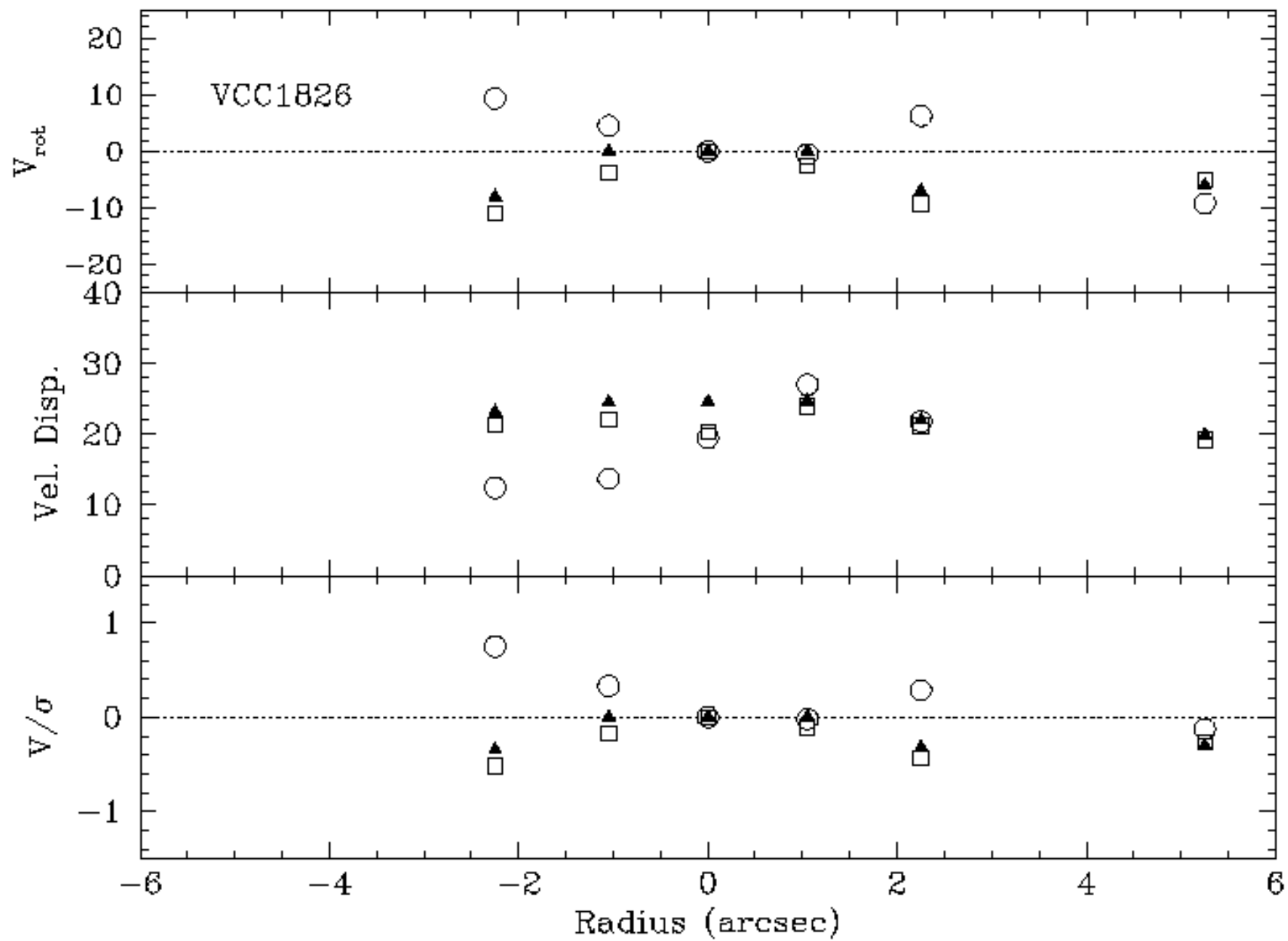
“...it is still challenging (or even impossible) to obtain accurate velocity dispersions for such low surface brightness [dwarf elliptical] galaxies.”

“This will be a promising science case for the ...ELT or the JWST.”

Misgeld & Hilker 2011 MNRAS in press

Keck/ESI observations
Echelle spectra (res ~ 15 km/s)
5 dE (no nucleus), $M_V \sim -16$
In Leo, Eridanus and Virgo
 ~ 2 hrs per galaxy, S/N ~ 20



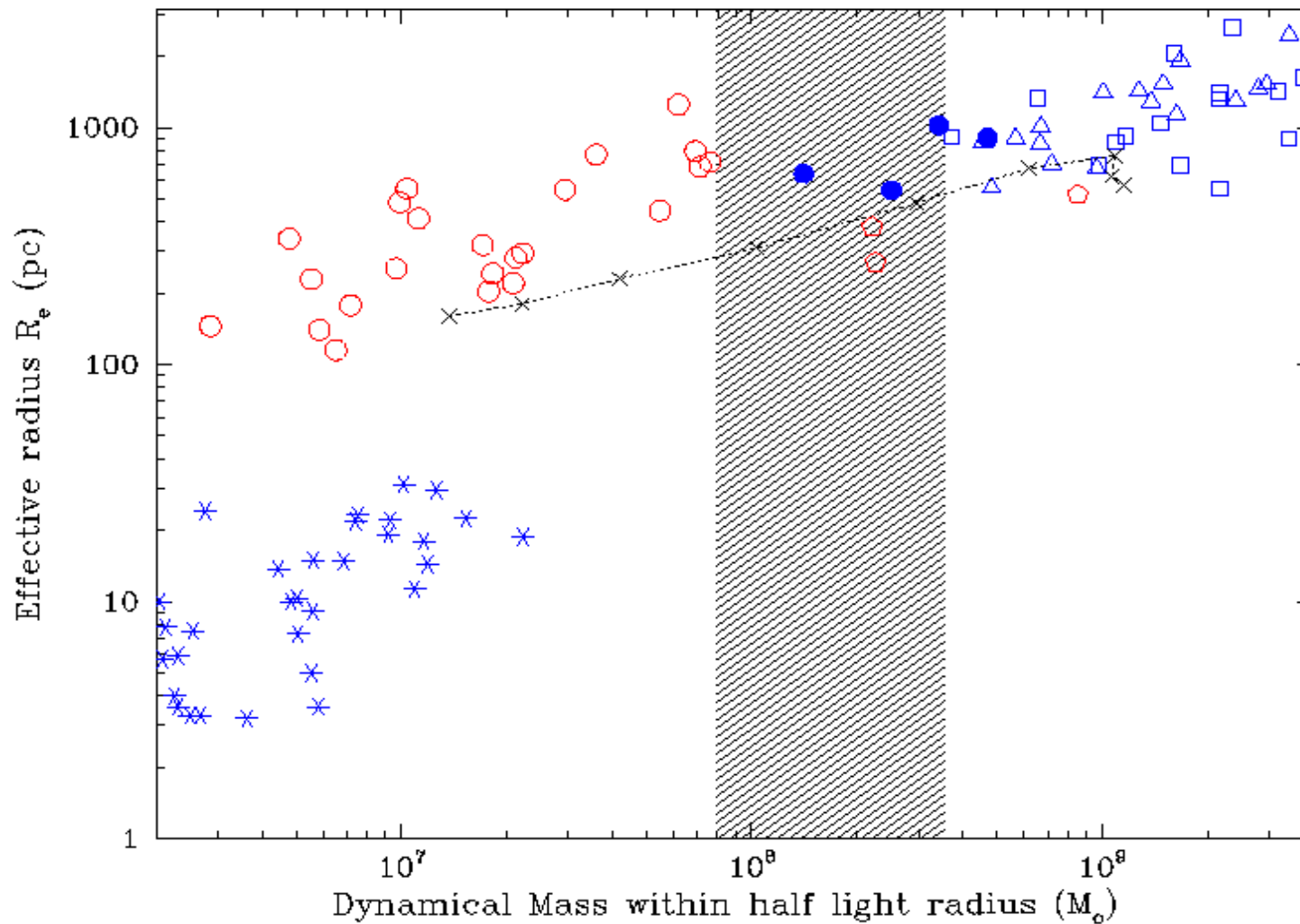


dE galaxy measurements

- Central galaxy velocity dispersions from Ca Triplet and Mg lines using ESI (galaxies have little or no nucleus)
- Effective radii (R_e) from Sersic fits
- Dynamical mass within deprojected R_e from Wolf et al. (2010) formula, ie $4 \sigma^2 R_e / G$ which is insensitive to orbital anisotropy
- Plus data from literature for Virgo/Fornax UCDs, Virgo dE, LG dSph, LG dEp (NGC 147, 185, 205)

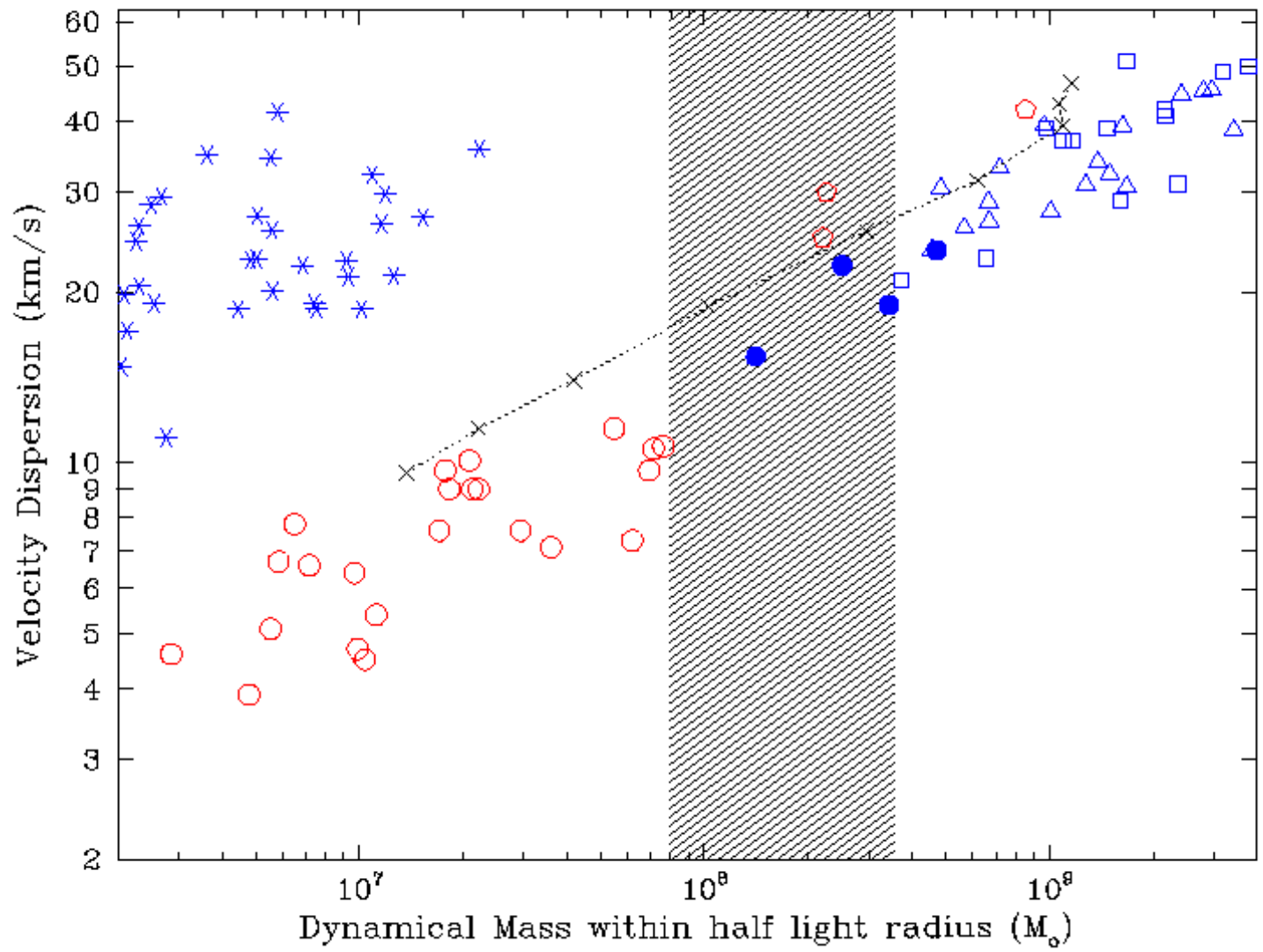
Size vs Dyn. Mass

within deprojected half light radius



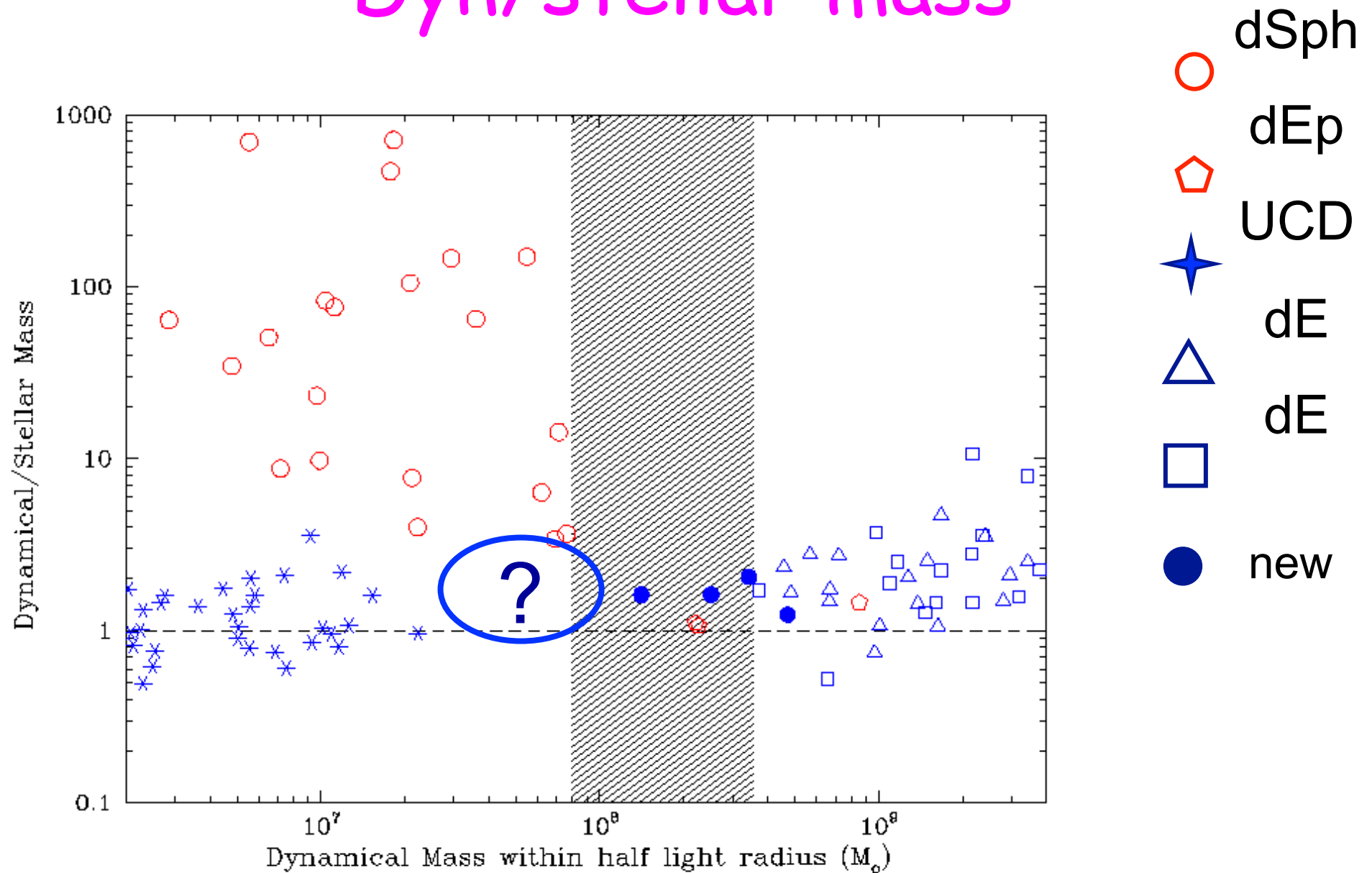
- dSph
- ⬠ dEp
- ★ UCD
- △ dE
- dE
- new
- LG
- Virgo

Sigma vs Dyn. Mass within deprojected half light radius

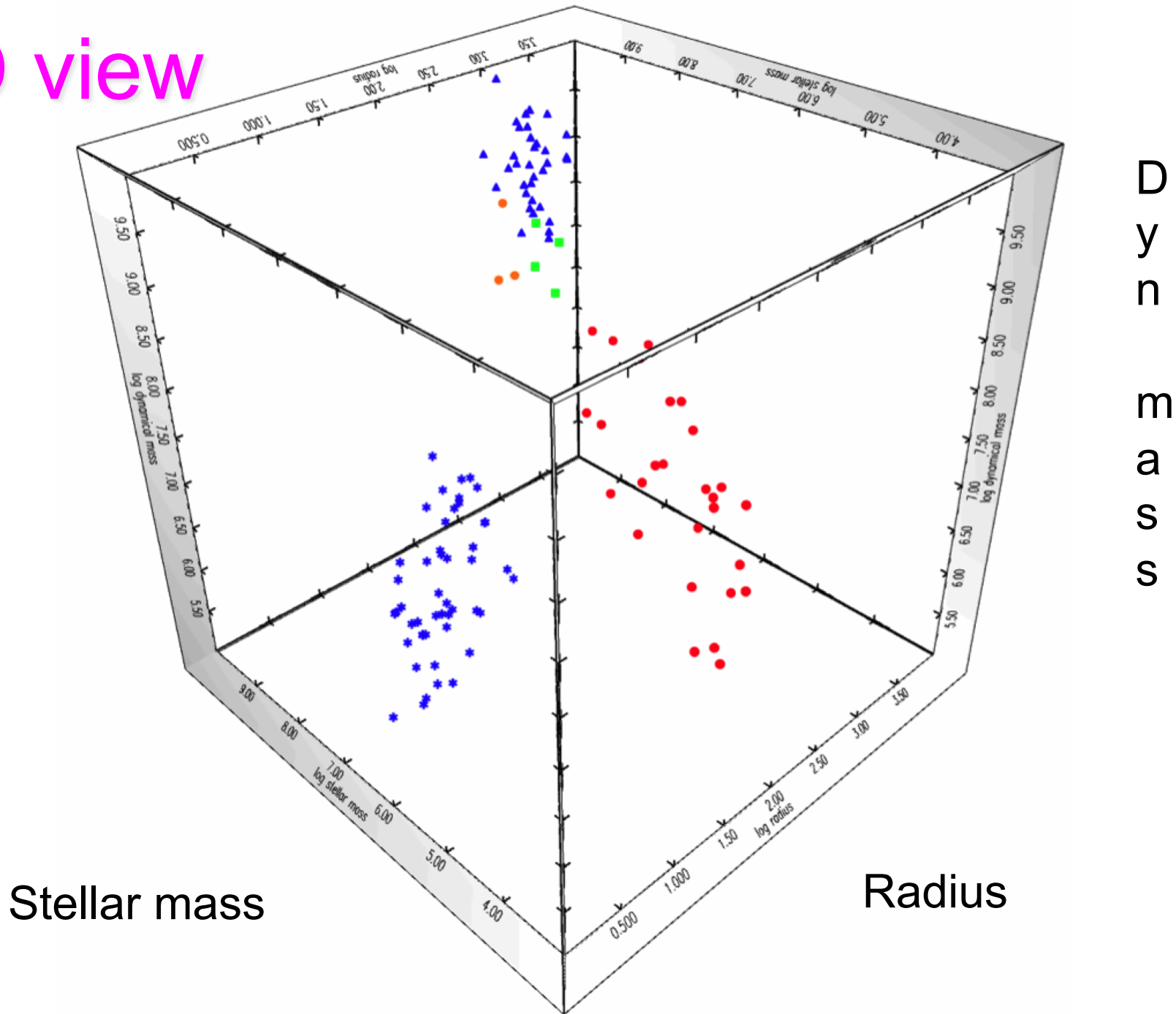


- dSph
 - ⬠ dEp
 - ★ UCD
 - △ dE
 - dE
 - new
- LG
- Virgo

Dyn/stellar mass



3D view



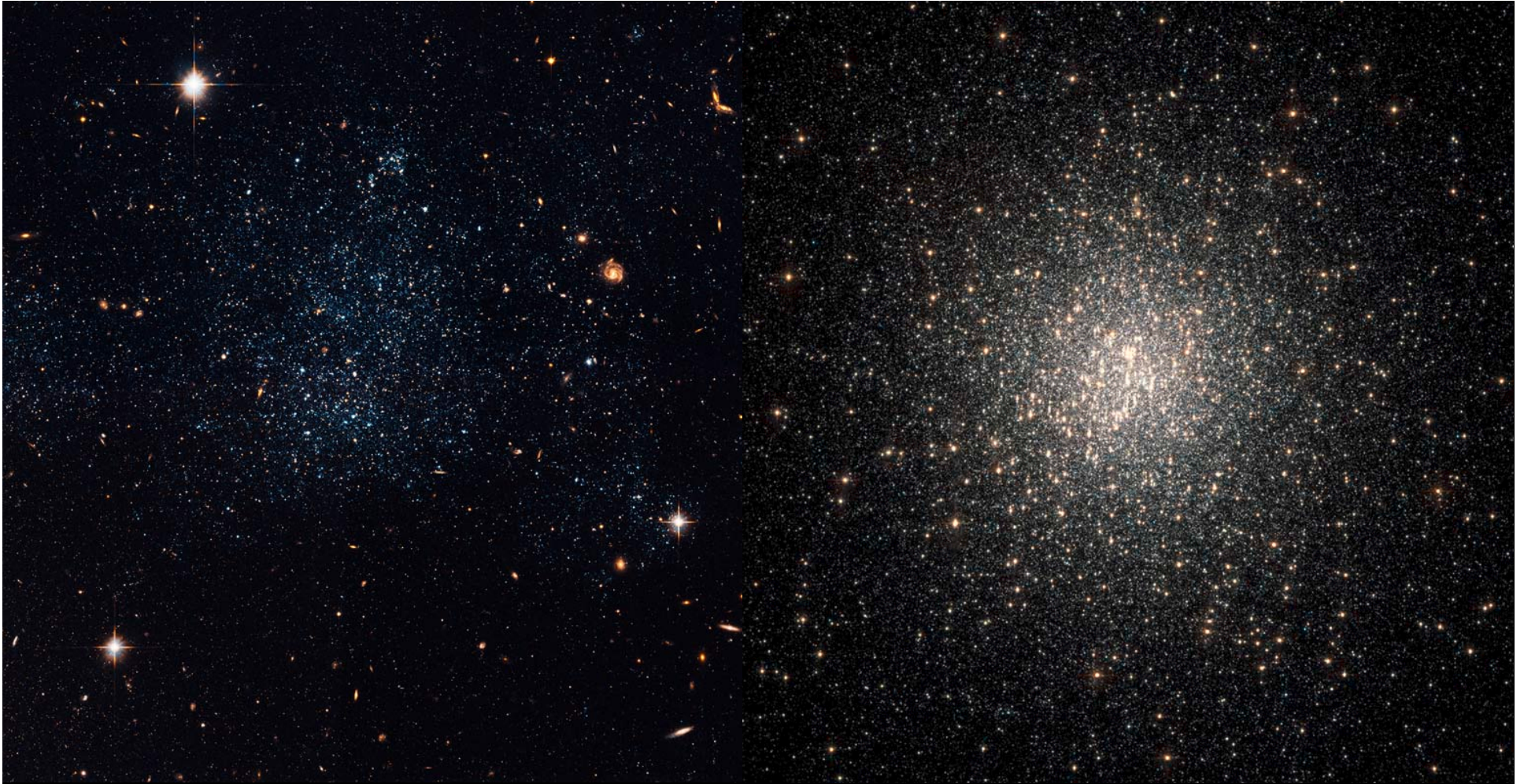
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Stellar mass

Radius

Interpretation

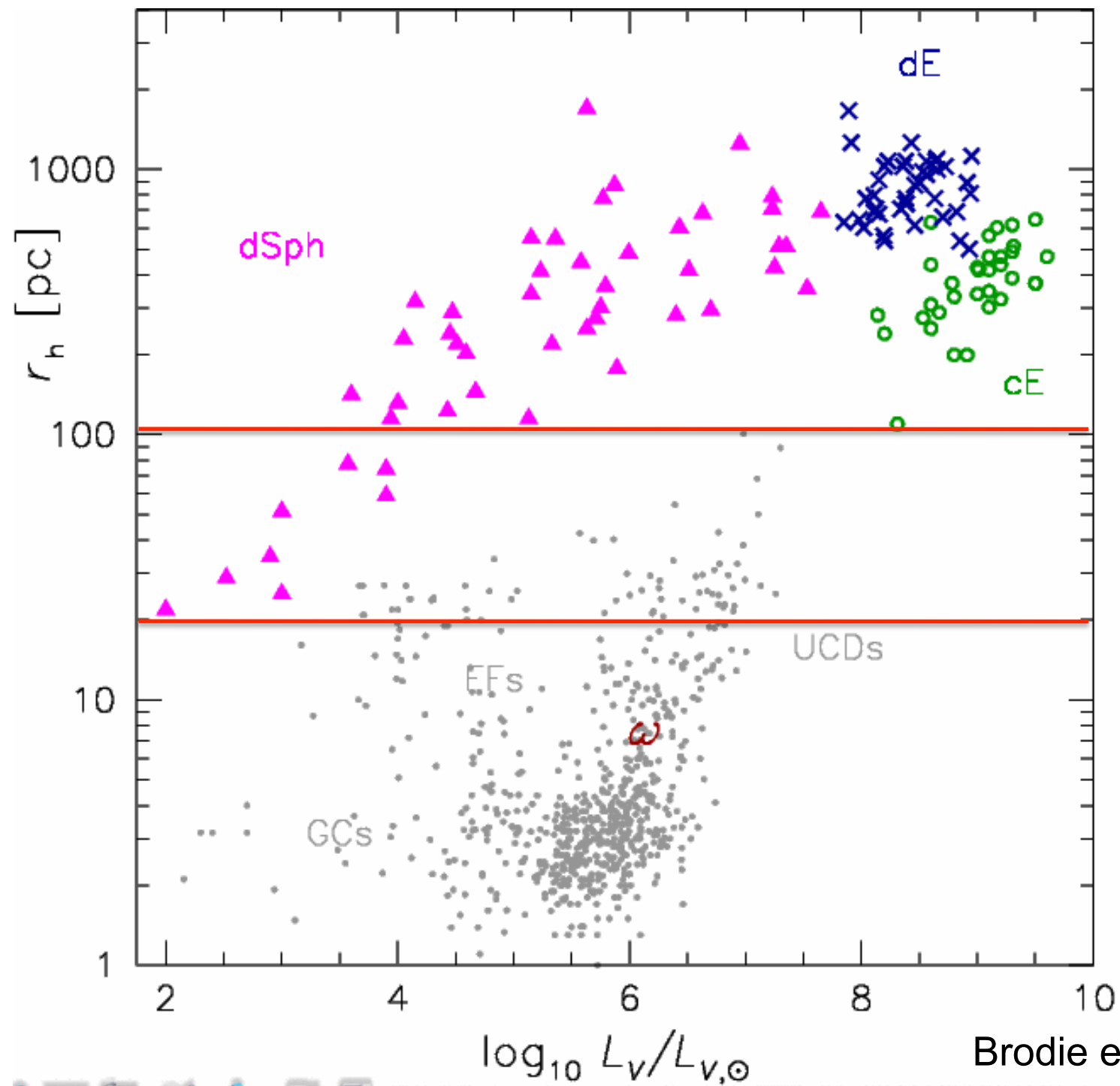
- Ultra Compact Dwarfs appear to be dark matter free star clusters, like globular clusters.
- Low luminosity dE galaxies are relatively dark matter free within R_e . SN feedback \rightarrow lower DM fractions in dEs
- dSph galaxies appear to be dark matter dominated. Tidal and ram pressure stripping \rightarrow higher DM fractions in dSph
- Dabringhausen, Zaritsky, Tollerund, Misgeld



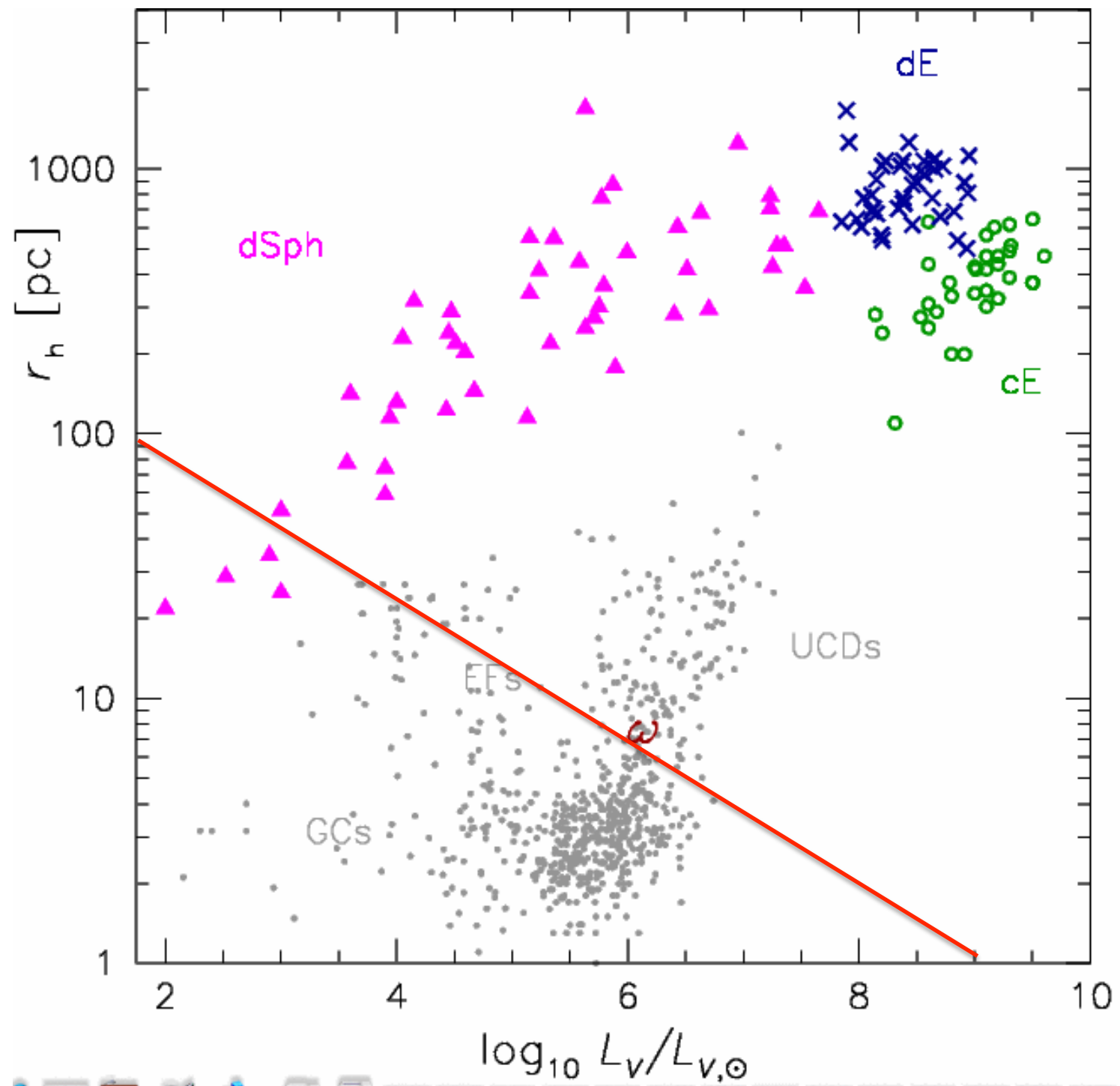
What is a galaxy ?

What is a Galaxy ?

- **Gilmore et al. 2007**: Galaxies have dark matter and effective radii > 100 pc. GCs and UCDs are dark matter free star clusters.
- **Kroupa 2007**: Galaxies have Relaxation times $>$ Hubble time. UCDs are galaxies, GCs are star clusters. Both GCs and UCDs are dark matter free.



Brodie et al. 2011



Galaxy Definition

- **Gravitationally bound**

Excludes: tidal material

- **Contains stars**

Excludes: dark galaxies

Includes globular clusters...so perhaps we need additional criteria - Forbes & Kroupa 2011

Galaxy Definition

- Half light radius > 100 pc

Excludes: GCs and UCDs

- Relaxation time $>$ age of Universe

Includes: UCDs and tidal dwarf galaxies

- Hosts satellites

Excludes: most dwarf galaxies and UCDs

Galaxy Definition

- Presence of dark matter

Difficult to measure for individual objects

Excludes: GCs, tidal dwarf galaxies, probably UCDs, perhaps some dEs

- Presence of complex stellar pops

Includes: some massive GCs

What is a Galaxy? - Crowd Wisdom

- Write paper on “What is a Galaxy?”
 - Submit it to PASA, get it accepted
 - Setup surveymonkey voting site
- 1) Have you read the paper?
 - 2) Pick the best definition(s) for a galaxy
 - 3) Comments
- Issue Media Release
 - Watch the votes and comments come in

What is a Galaxy? - voting results

- 1638 votes in 2 months
- 56% or 920 people have read the paper
- Multiple choices allowed
- 68% voted for Complex Stellar Populations
- 28% voted for Long Two-body Relaxation
- 31% voted for Sizes > 100 pc
- 31% voted for Presence of Dark Matter
- 29% voted for Presence of Satellites