

# PIONIER : Science with a Four Telescope VLTI Instrument

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VLTI Team

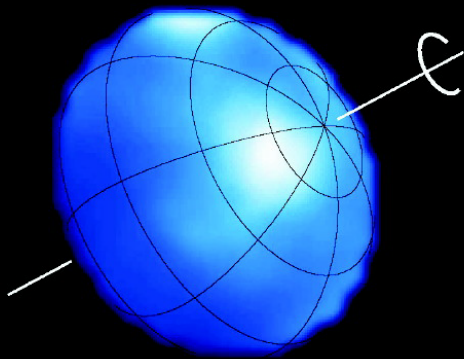
# Outline

- **Context**
- **Instrument description**
- **Commissioning**
- **Results from first year of operation**
- **Perspective**

■ **Context in 2009**

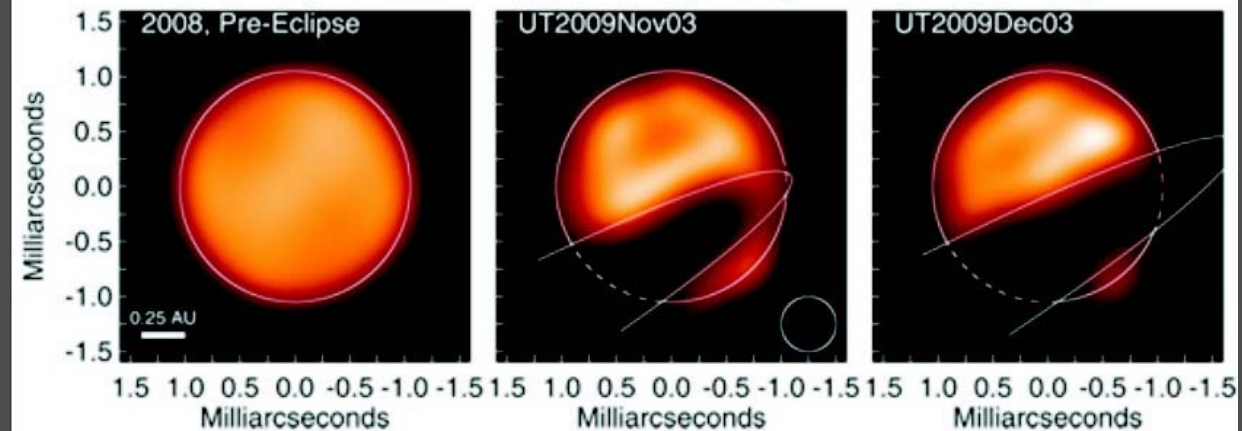
- **Optical long baseline interferometry enters era of imaging at mas scale**
- **Sensitivity limits mostly to stellar physics**
- **Game dominated mostly by US (CHARA)**
- **VLTi improving (AMBER), but time-consuming**

Actual image of Altair from the CHARA Interferometer

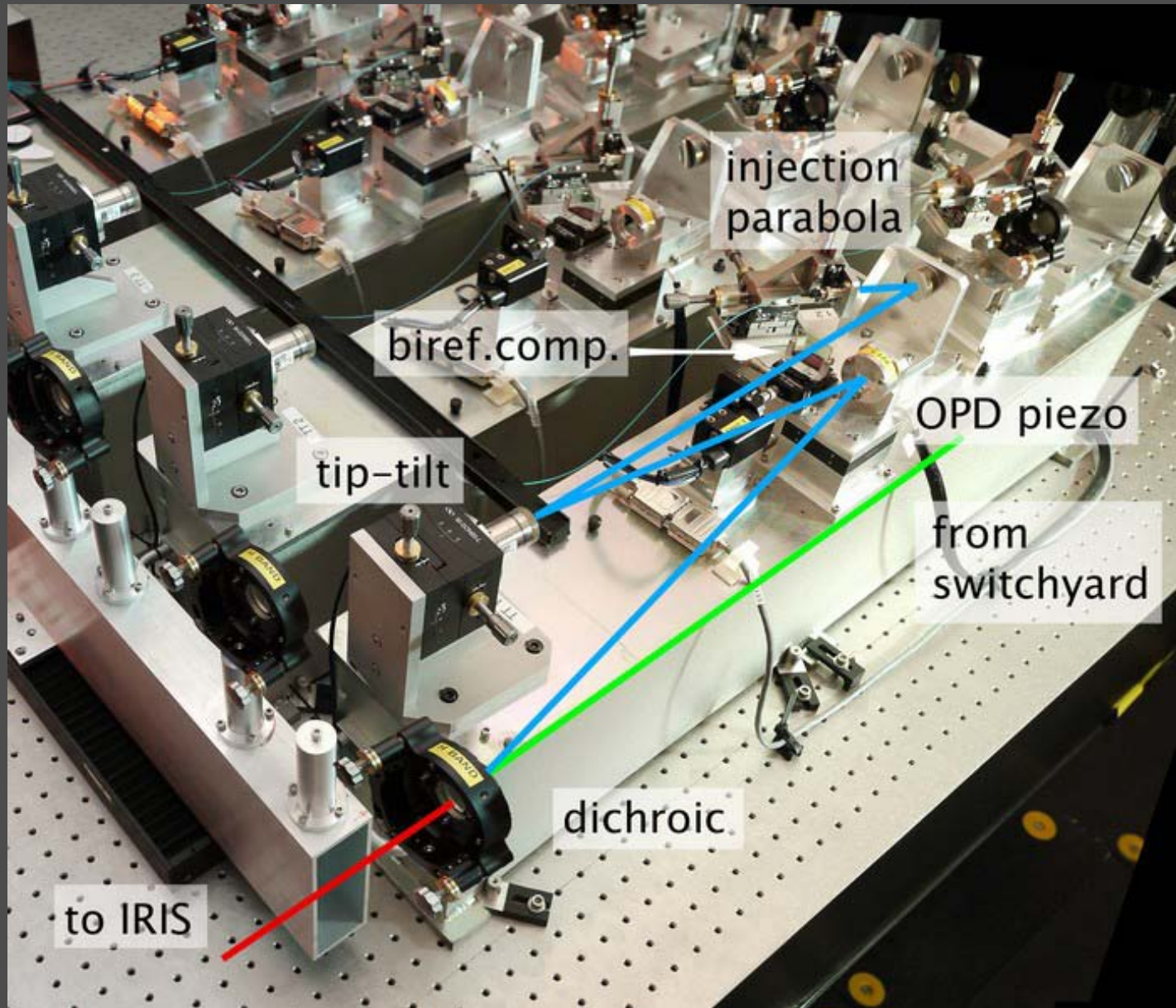


Monnier et al Science 2007

Epsilon Aurigae Eclipse (CHARA-MIRC)



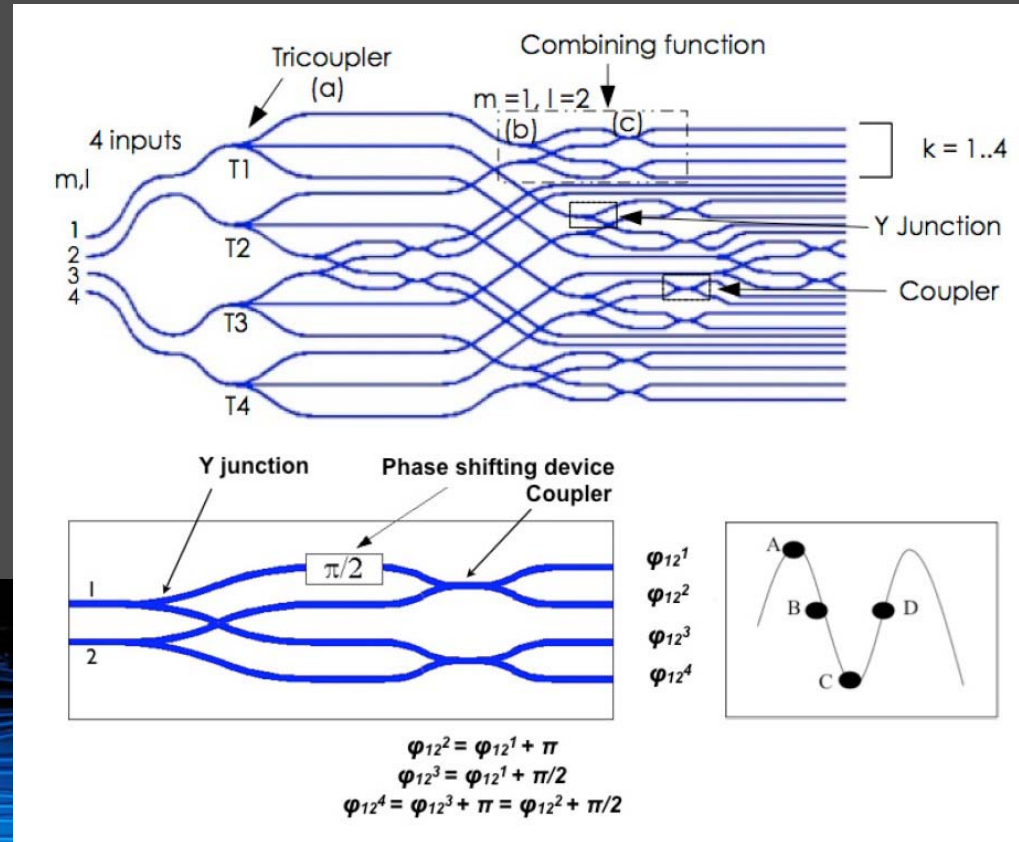
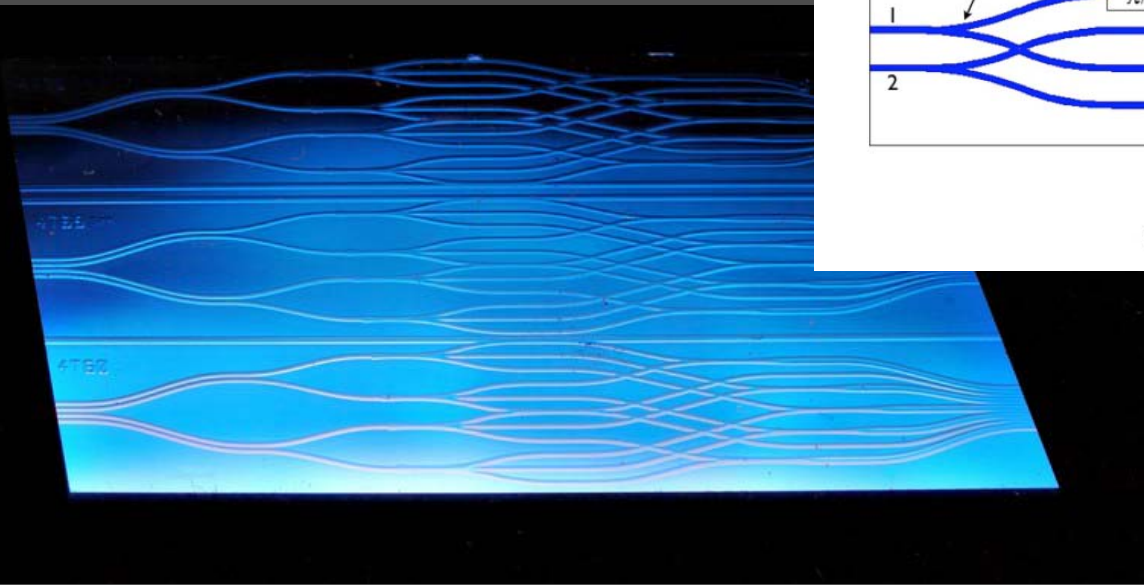
- **VLTI context in 2009**
  - 4 ATs available since 2006
  - But instruments allow for 2 or 3 telescopes only
  - Next-generation projects planned for 2014/15 (GRAVITY and MATISSE)
- **LAOG (now IPAG) context in 2009**
  - Expertise and interest in imaging interferometry
  - IONIC 4-telescope beam combiner available off-the-shelf (following 10 years of development with LETI)
- **Opportunity for a fast track project**
  - LAOG proposal to ESO
  - Visitor instrument
  - On sky end 2010
  - Approved by STC Nov 2009
- **Support**
  - Funding: Université Joseph Fourier – IPAG – INSU  
ANR "Exozodis" – ANR "2G VLTI"
  - Camera on loan from IOTA/PTI, courtesy W. Traub





■ **IONIC 4-beam combiner**

- Compactness
- Stability where it matters most
- Credit: P. Labeye (LETI) and M. Benisty (LAOG)

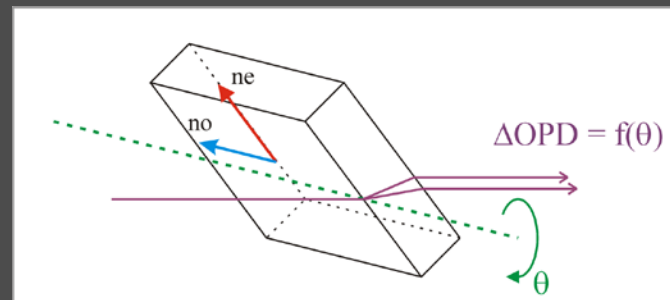
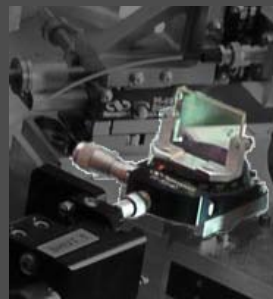
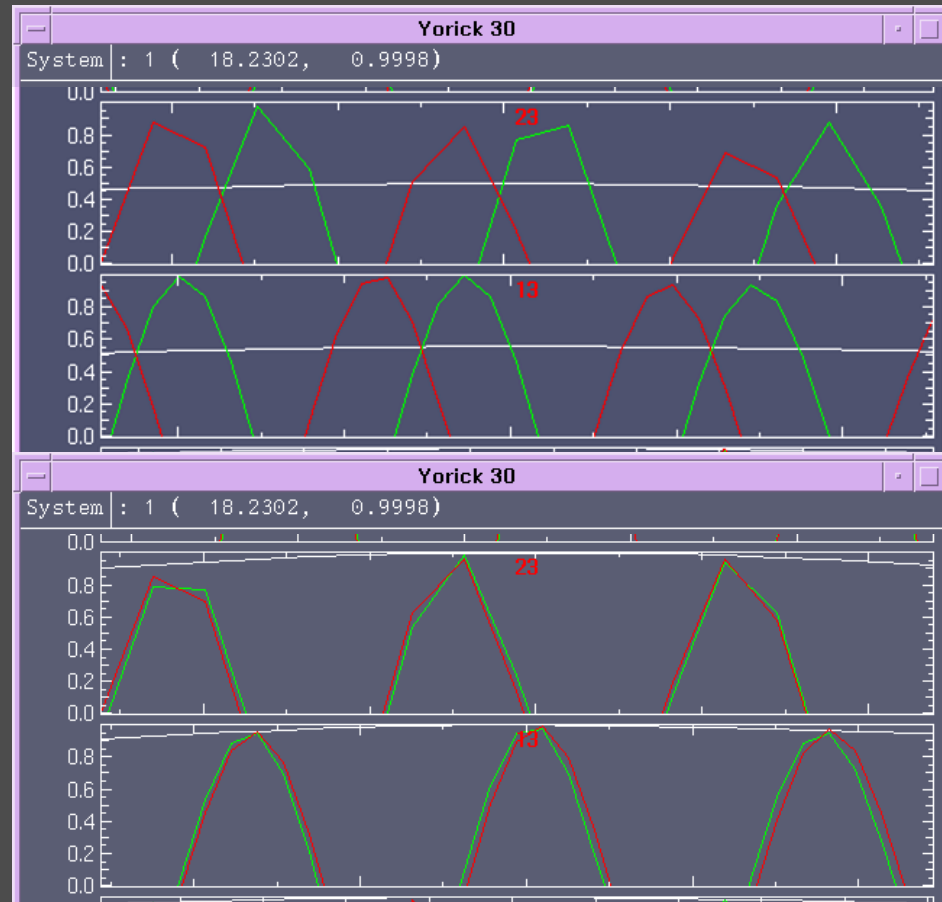


## ■ Birefringence

- Pol-maintaining fibers are birefringent
- OPDs not aligned for two polarizations
- Requires pol splitting and separate detection

## ■ Birefringence compensation

- Birefringent plates (LiNbO3) on optical path with adjustable inclination
- Phases (group and fringe) can be aligned → detected together
- Better S/N and speed



- **Operation**
- **6 baselines; 3 closures**
- **Scan the OPDs across the white fringe (150 – 400 ms typ)**
- **Three dispersion modes:**
  - **Broadband 1.55 – 1.80  $\mu\text{m}$**
  - **Small dispersion ( 3 resolution elements,  $R \approx 20$ )**
  - **"Large" dispersion (7 resolution elements,  $R \approx 40$ )**

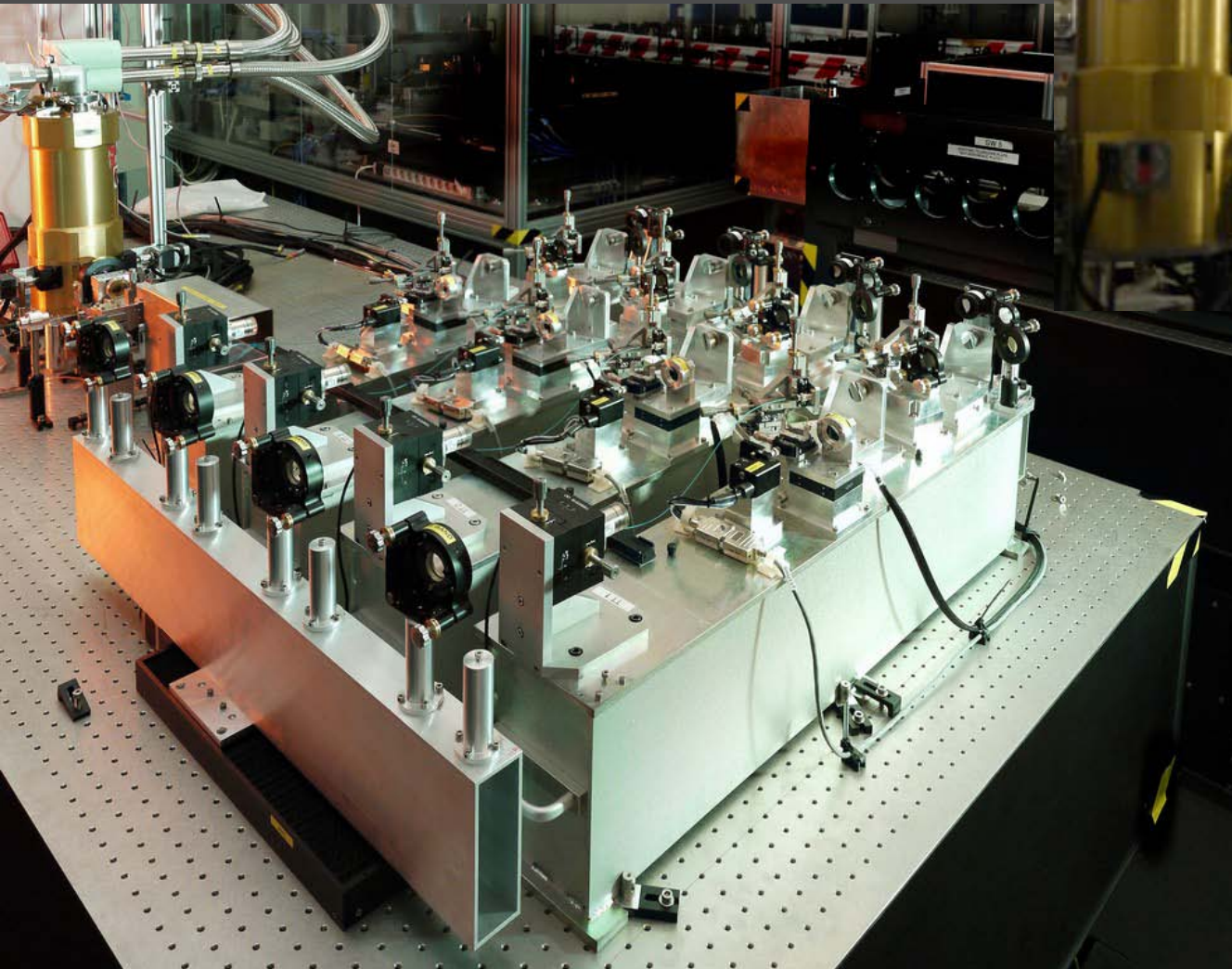


■ Packing...



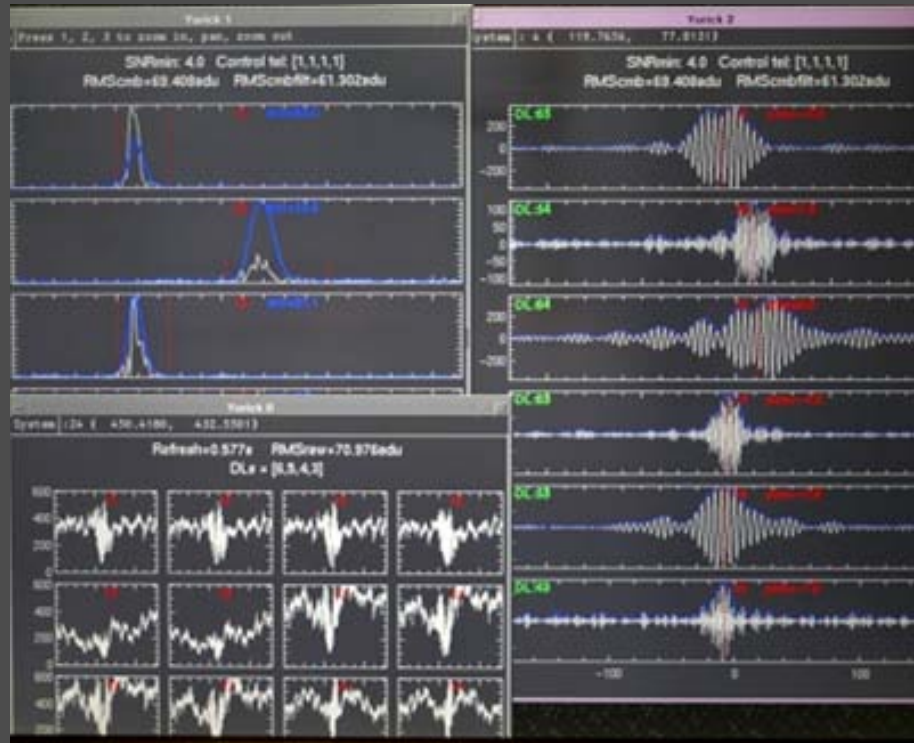


- **Unpack & Install**  
**Oct 20-24, 2010**



# Context > Instrument > **Comm** > Results > Perspective

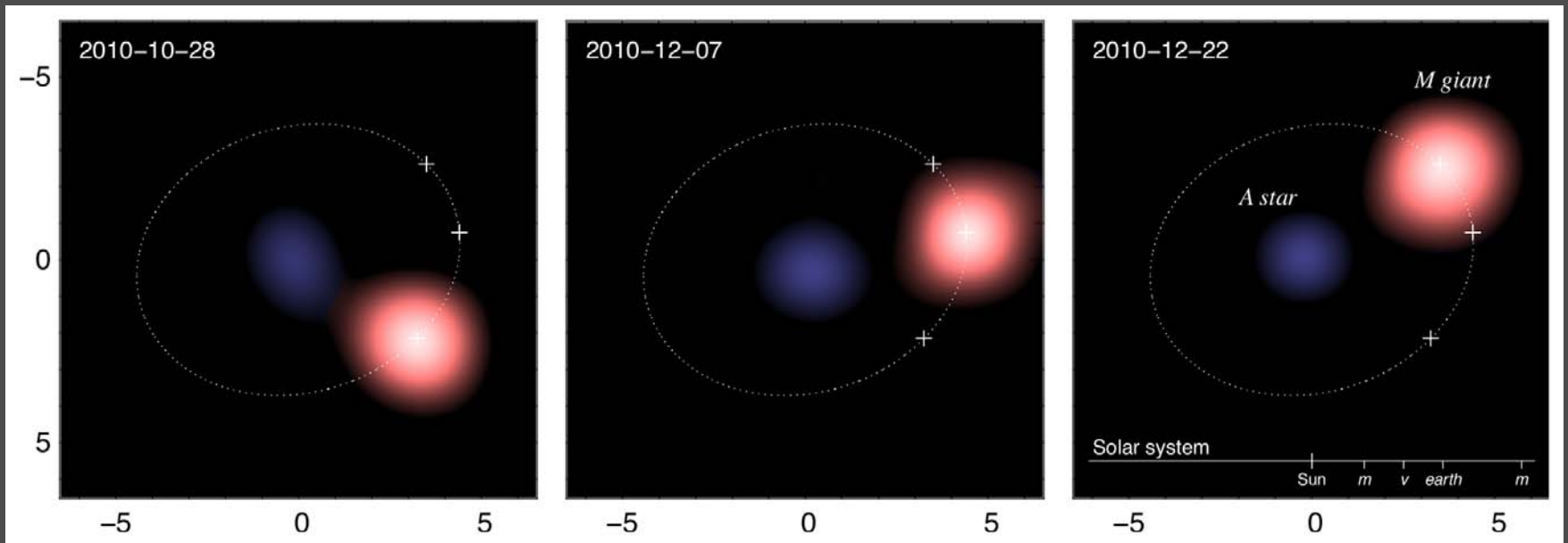
- First light + fringes  
Oct 25, 2010
- T0 : VLTI preset
- T0+10 min: fringes on 3 baselines
- T0+30min: fringes on 6 baselines
- T0+1day: science demonstration





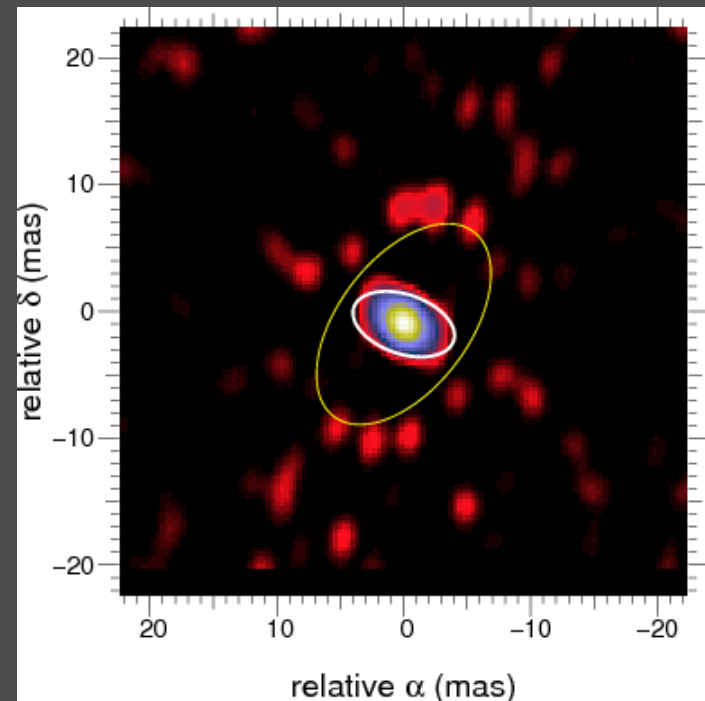
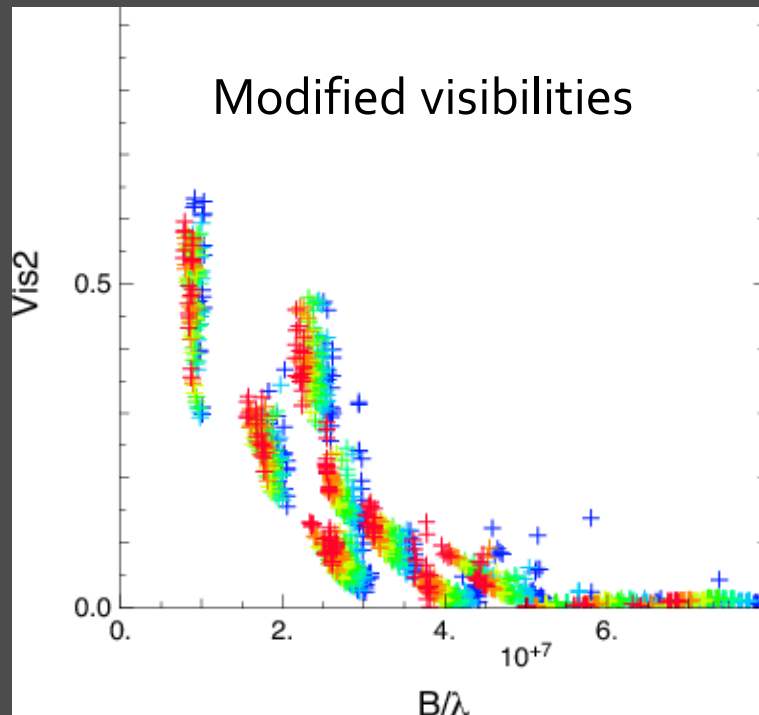
# Results : SS Lep

- An incisive look at the symbiotic star SS Leporis
  - N. Blind, H.M.J. Boffin, J.P. Berger, J.B. Le Bouquin, A. Mérand, B. Lazareff, G.Zins. **Submitted.**
  - Orbit, masses, radius (M giant), temperatures, envelope
  - Constraints on co-evolution of Algol-like system



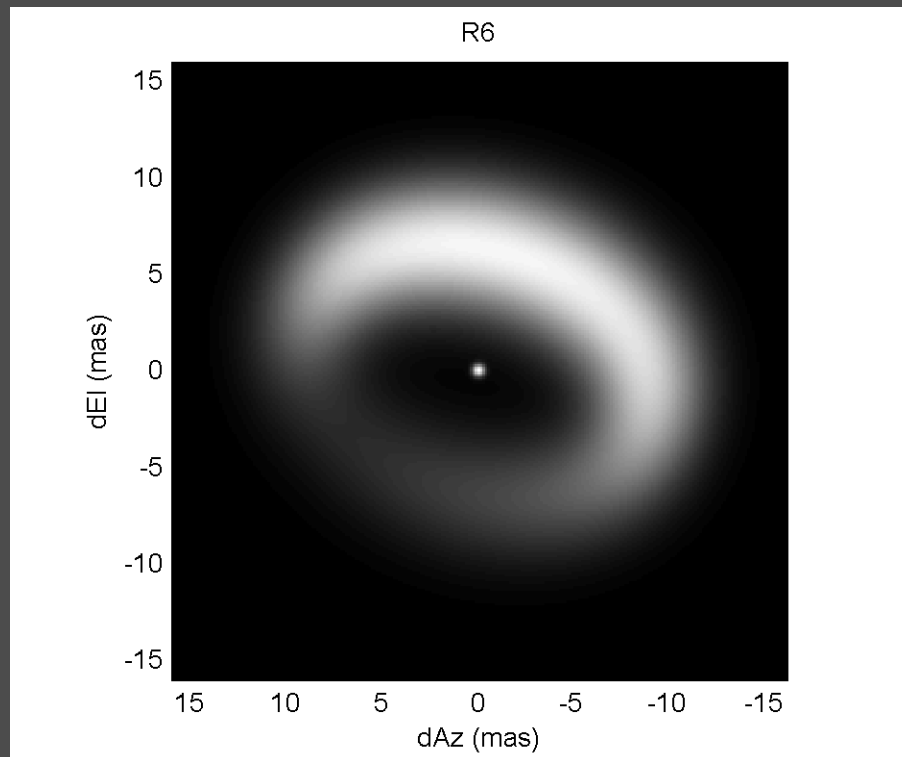
# Results : MWC158

- Imaging the inner disk region of MWC158
  - F. Malbet, J. Kluska, J.-P. Berger, F. Ménard, C. Pinte, J.-B. Le Bouquin, B. Lazareff, R. Millan-Gabet, W. Traub, M. Benisty
  - Image reconstruction with MIRA (polychromatic object)



# Results : HD45677

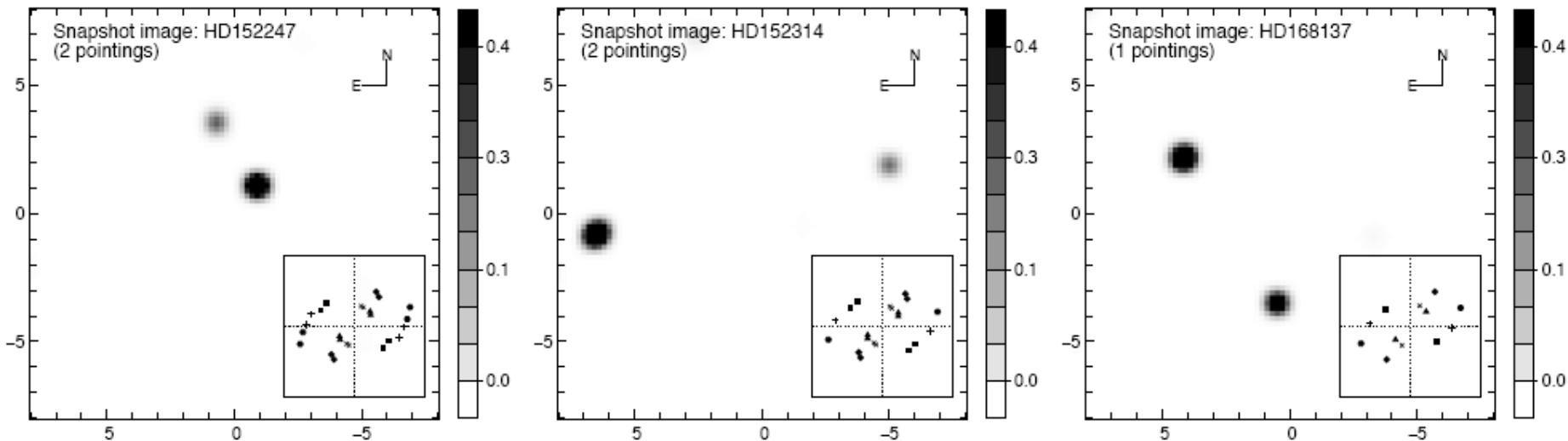
- The sublimation ring of HD45677
  - B. Lazareff, J.P. Berger, M. Benisty, W.F. Thi, J.B. Le Bouquin, G. Zins, W. Traub, R. Millan-Gabet
  - Parametric image of star and bright inner ring at sublimation radius
  - Radius (9.5mas) and temperature (1300K) of emitting ring in agreement with *puffed up ring* model





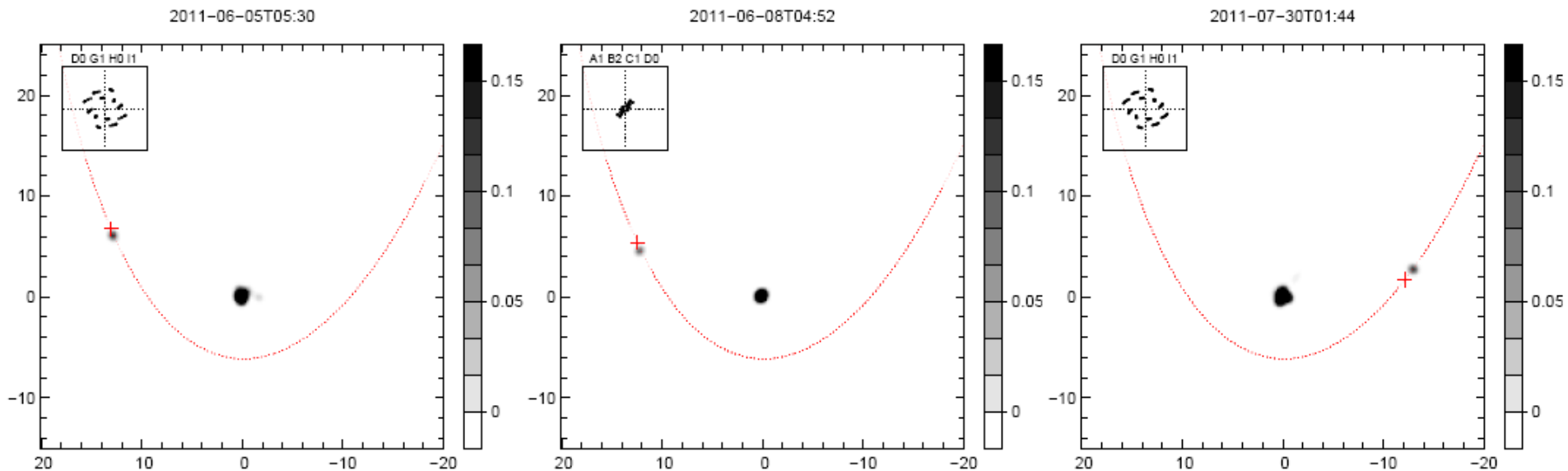
# Results : Massive star binaries

- H. Sana, J.B. Le Bouquin, M. De Becker, O. Absil, J.P. Berger  
B. Lazareff, R. Millan-Gabet, A. Merand, C. Martayan
  - Absolute mass determination of high mass stars
  - Three binaries with  $H=6\dots7.5$  observed in 1.5 hour
  - [ Pionier + AT ] as efficient as [ Amber + UT ]



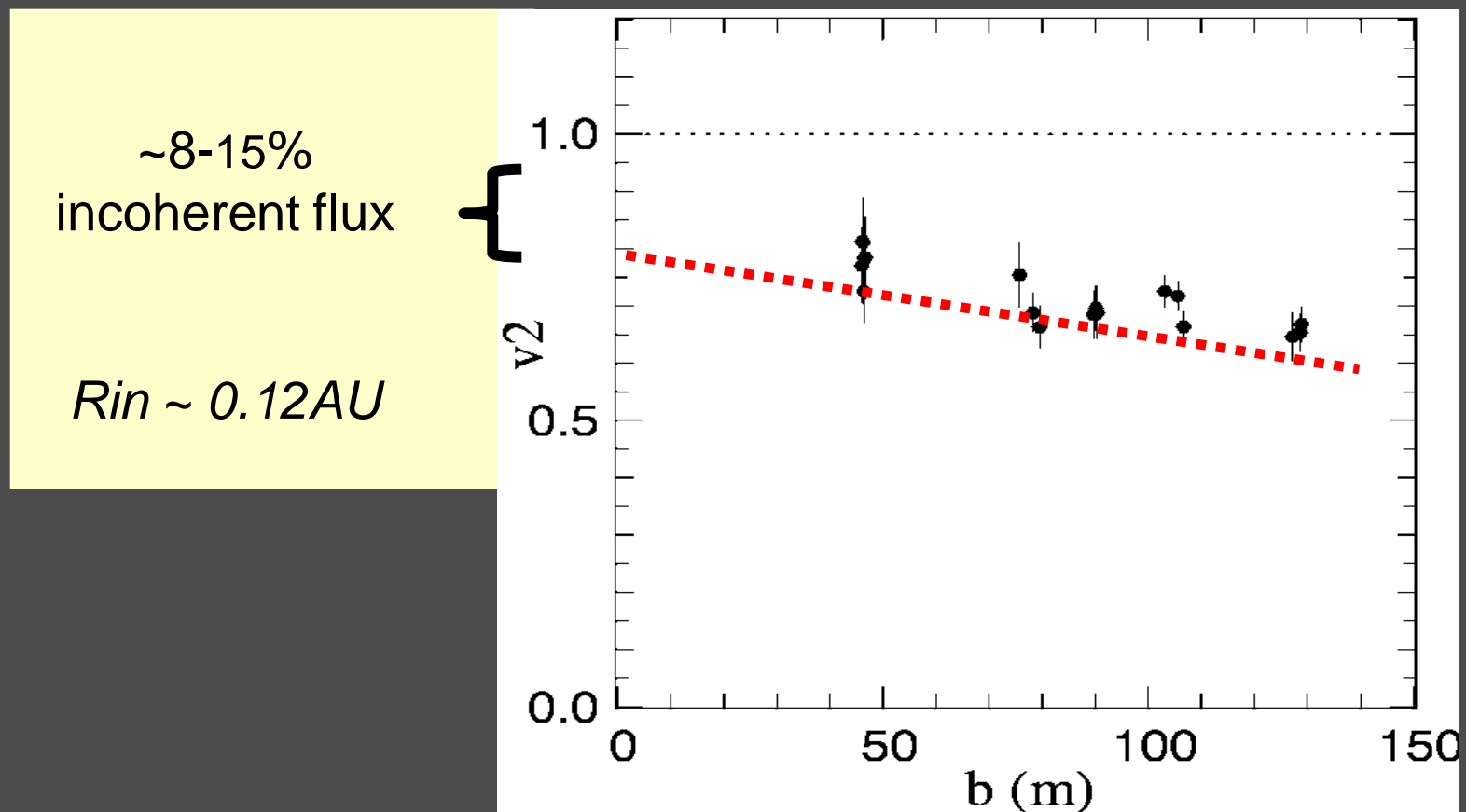
# Results : $\delta$ Sco

- A VLTI view of  $\delta$  Sco at periastron
  - J.B. Le Bouquin, S.Stefl, D.Baade, J.P.Berger, A.C.Carciofi, M.Cure, X.Haubois, B.Lazareff, R.Millan-Gabet, A.Okazaki, T.Rivinius, G.Zins
  - Model-independent imaging of central star, secondary, and disk
  - Combined with Amber Br $\gamma$  data, spectro-image of extended rotating disk
  - Perturbation of disk near periastron ?



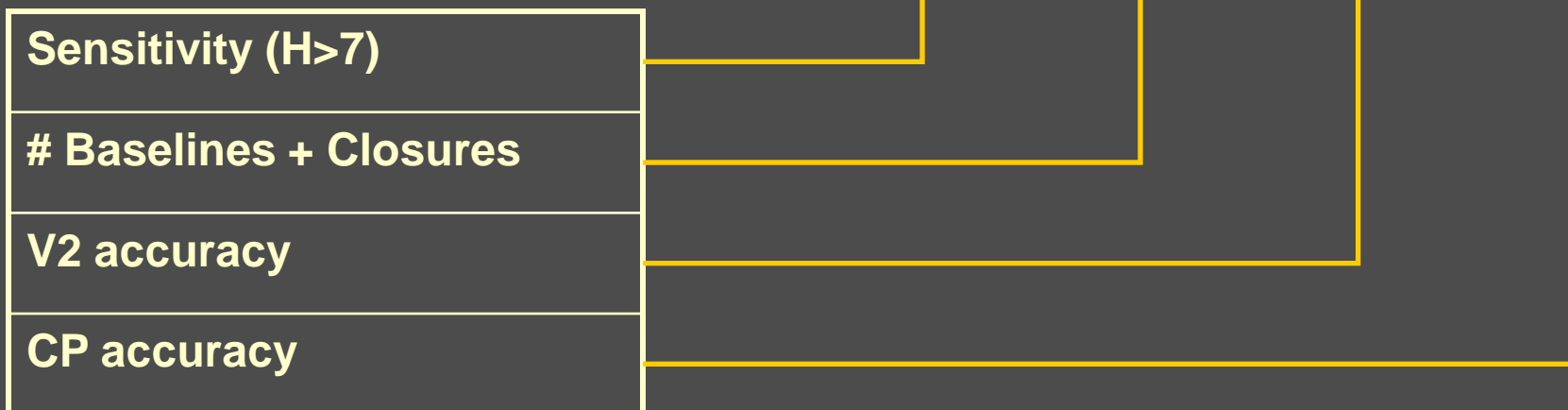
## Results : T Tauri Survey

- Extended component is frequent, different source to source
- Closure phases are very small, or null
- Example: HT Lupi, the PIONIER measurements K2, ~1Myr, class II



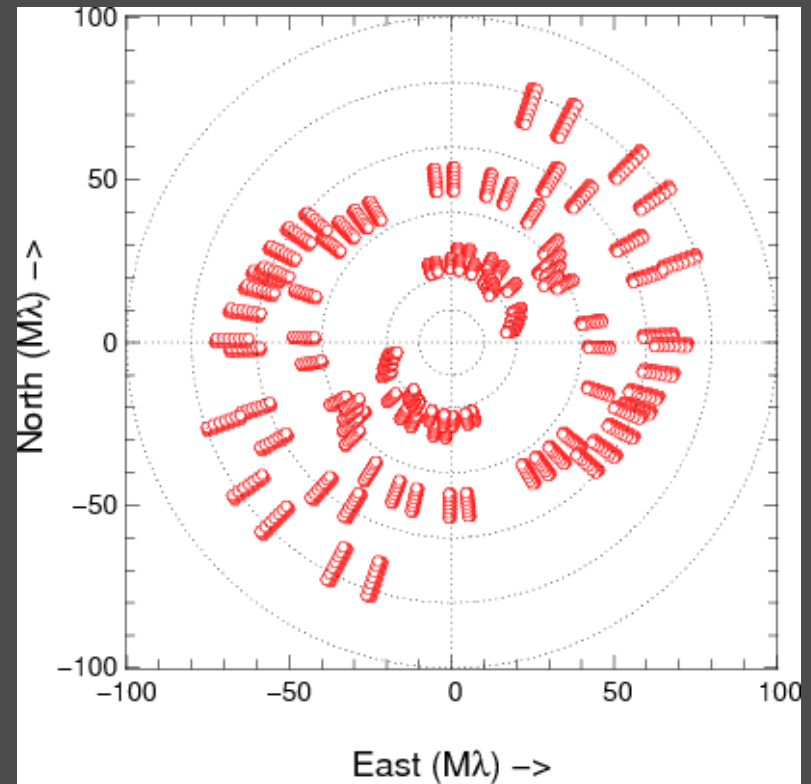
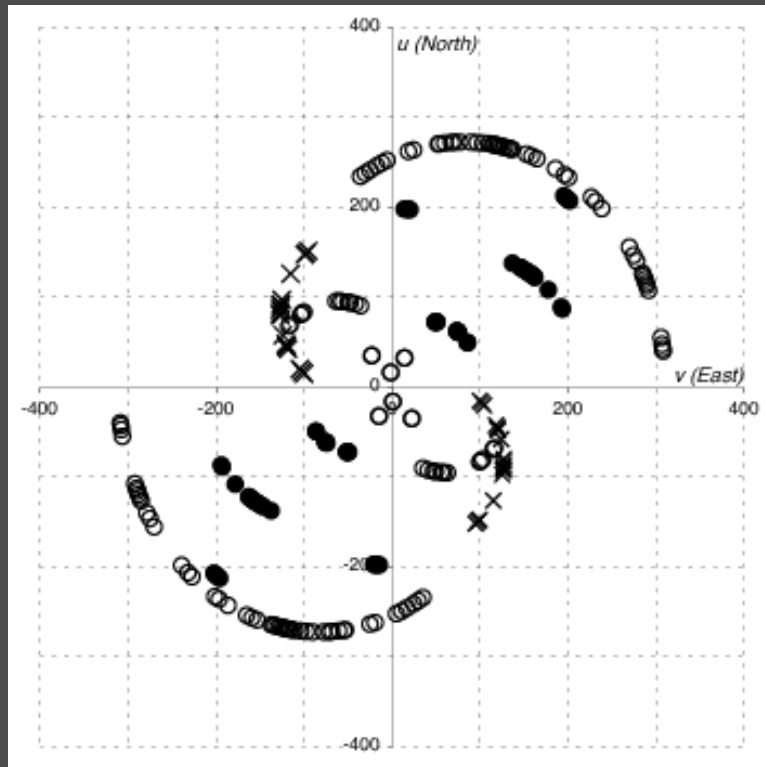
# The PIONIER benefits. Real user testimonies.

<b>T. Tauri (Menard)</b>	■	■	■	
<b>AeBe Imaging (Berger)</b>	■	■		
<b>Massive binaries (Le Bouquin)</b>	■			
<b>Transitional disks (Benisty)</b>	■	■		
<b>Faint companions (Absil)</b>		■		■
<b>Debris disks (Absil)</b>		■	■	
<b>Hot jupiters (Absil)</b>		■	■	■



# The PIONIER benefits. Achernar; uv coverage

- Vinci : Several years of observing
- Pionier: one night



# The PIONIER benefits. Massive binaries

- **Amber + UTs**      **3 – 4 targets / night**
- **Pionier + ATs**      **8 targets / night @  $H > 7\text{mag}$**

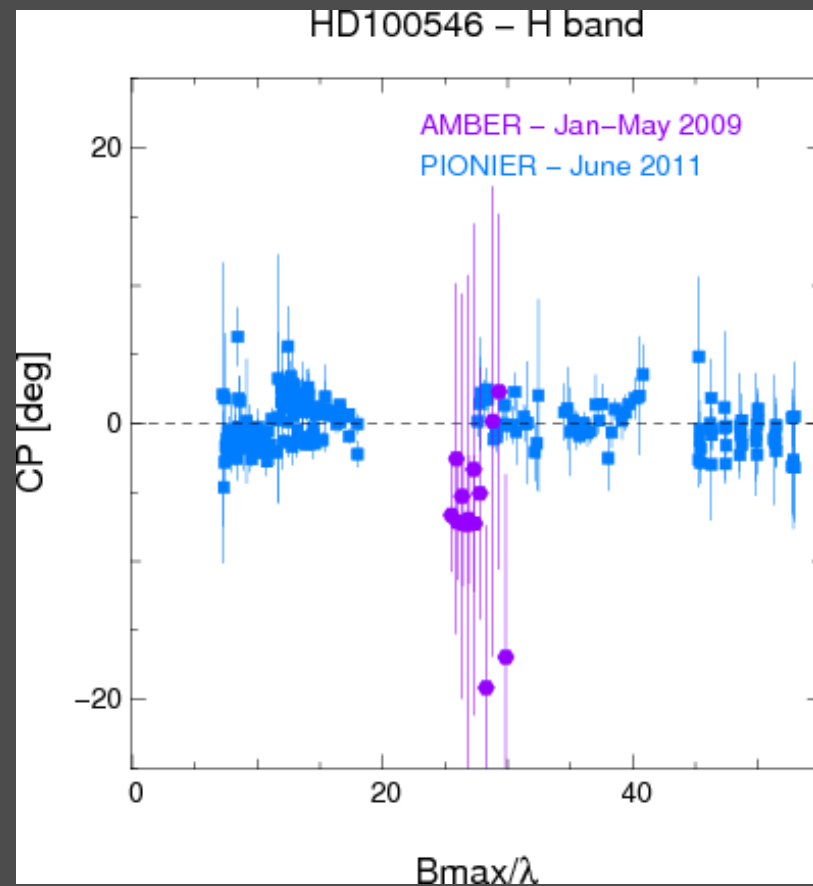
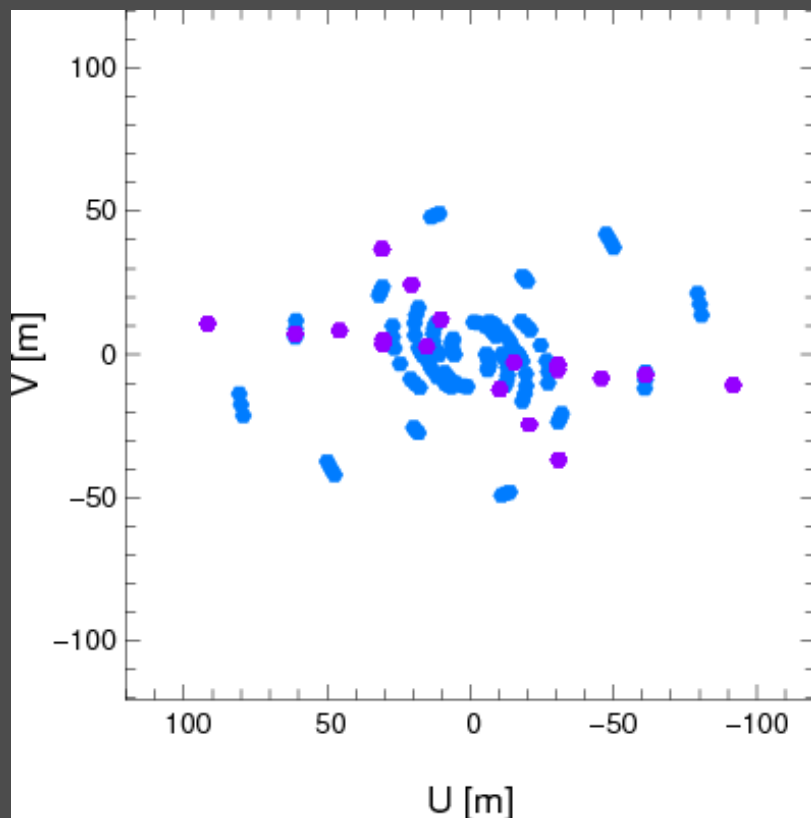


# The PIONIER benefits. Transitional disks

## ■ Observations of HD100546 with Amber and Pionier

$\lambda$  Amber

$\lambda$  Pionier



## Context > Instrument > Comm > Results > **Perspective**

- **Remote operation of disperser change – Done**
- **Identify and eliminate instrument and reduction software systematics – In progress**
- **Implement K band – Planned summer 2012**
- **Change to Rapid (EMCCD) camera; low readout noise – 2013?**
- **Most significant limitations on performance:**
  - **Phase piston (atmosphere + telescopes)**
  - **Camera electronic transients**



Acknowledgements to:

- Pionier instrument team
- Paranal VLTi team