# High angular resolution science with the E-ELT



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## **E-ELT Overview and Status**

## The European Extremely Large Telescope

- Most ambitious optical telescope ever built
- Primary mirror of ~40m diameter
- Working at optical / infrared wavelength
- Adaptive optics with 6 lasers diffraction limited
- Status: Handed in Construction Proposal after the detailed design phase (B)
- Start construction 2012 first light in 2021



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#### Main structure: Alt-Az mount, 2800 t

# **Dome:** 80m x 90m $\emptyset$ , fully air conditioned



![](_page_7_Figure_0.jpeg)

Three mirror anastigmat + two flat mirrors (M4, M5) Strehl > 99% out to 3' radius for  $\lambda$  > 360nm

![](_page_8_Picture_0.jpeg)

Three mirror anastigmat + two flat mirrors (M4, M5) Strehl > 99% out to 3' radius for  $\lambda$  > 360nm

![](_page_9_Picture_0.jpeg)

Three mirror anastigmat + two flat mi Strehl > 99% out to 3' radius for  $\lambda > 30$ 

![](_page_10_Picture_0.jpeg)

Strehl > 99% out to 3' radius for  $\lambda$  > 3(

![](_page_11_Picture_0.jpeg)

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![](_page_12_Picture_0.jpeg)

Strehl > 99% out to 3' radius for  $\lambda$  > 3(

![](_page_13_Picture_0.jpeg)

Strehl > 99% out to 3' radius for  $\lambda$  > 3(

![](_page_14_Figure_0.jpeg)

#### M4: flat, 2.4-m diameter

- Zerodur 2mm thin shell
- 5000 contact-less actuators
- 198 removable actuator bricks
- Fitting error 145nm rms @ 0.85" seeing
- 10 tons, 8.4 kW

#### M1: F/0.93, 39-m diameter

- 798 Segments 1.44-m average size (corner to corner)
- 6 sectors of 133 segments
  + 1 spare set 931 total
- Each segment controlled in Piston / Tip / Tilt position
- Each segment controlled in shape

![](_page_14_Figure_12.jpeg)

# E-ELT instruments at the diffraction limit

# Eight instruments and two adaptive optics modules were studied in phase A (2007-2010)

All instruments working in the near- or thermal infrared intend to exploit the diffraction limit of the telescope

> AT baseline ~200m UT baseline ~130m

![](_page_16_Figure_3.jpeg)

# Eight instruments and two adaptive optics modules were studied in phase A (2007-2010)

All instruments working in the near- or thermal infrared intend to exploit the diffraction limit of the telescope

> AT baseline ~200m UT baseline ~130m

![](_page_17_Figure_3.jpeg)

# Eight instruments and two adaptive optics modules were studied in phase A (2007-2010)

All instruments working in the near- or thermal infrared intend to exploit the diffraction limit of the telescope

> AT baseline ~200m UT baseline ~130m

![](_page_18_Figure_3.jpeg)

# Implementation of the instrument plan

![](_page_19_Figure_1.jpeg)

![](_page_19_Figure_2.jpeg)

![](_page_19_Figure_3.jpeg)

year	ELT-IFU	ELT-CAM	ELT-MIR	ELT-4	ELT-5	ELT-6	ELT-PCS
				HIRES)	HIRES		
2012	Decide science requirements, AO architecture.		VISIR start on- sky	Develop science requirements for MOS/HIRES			Call for proposals for ETD
2013			TRL Review	Call for proposals for MOS/HIRES			
2014							
2015				Selection ELT-MOS/HIRES		Call for proposals	
2016							
2017							TRL check
2018							TRL check
2019						Selection	TRL check
2020							TRL check
2021 Tel technical 1 <sup>st</sup> light							TRL check
2022 Inst Comm starts							
	Pre-studies taking the form of Phase-A or ∆-Phase-A work and/or ESO-funded enabling technology development (ETD)						
	Decision point						
	Development of Technical Specifications, Statement of Work, Agreement, Instrument Start.						

## High Angular Resolution Science Cases

#### Long term astrometric precision of 50-100 µarcsec

#### Galactic Centre

![](_page_21_Figure_3.jpeg)

## Proper motions

10 µarcsec/year

![](_page_21_Picture_6.jpeg)

at 10kpc: 0.5 km/s

at 50kpc: 2.5 km/s

#### Photometry and spectroscopy of individual star beyond the Local Group

![](_page_22_Figure_2.jpeg)

![](_page_22_Figure_4.jpeg)

Simulations taken from the Design Reference Mission report: <u>http://www.eso.org/sci/facilities/eelt/science/drm/drm\_report.pdf</u>

#### mid-IR: combining highest spectral with highest spatial resolution

![](_page_23_Figure_2.jpeg)

Simulation of a METIS image cube of the CO P(8) line from SR 21 for an assumed distance of 125 pc (Pontoppidan et al. 2009).

Resolve proto-planetary disks to a few AU at 150 pc distance, with a spectral resolution of 100.000

#### Contrast: reaching Earth-mass planets in habitable zones

![](_page_24_Figure_2.jpeg)

Simulation for the E-ELT project by S.Gladysz, analysed by J.Ascenso

For stars at < 10 pc, contrasts allowing to detect Earth-like planets can be reached, also inside the respective habitable zones

## Conclusions

![](_page_26_Figure_0.jpeg)

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![](_page_27_Picture_0.jpeg)

# The End

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