

OmegaCAM: Key Features

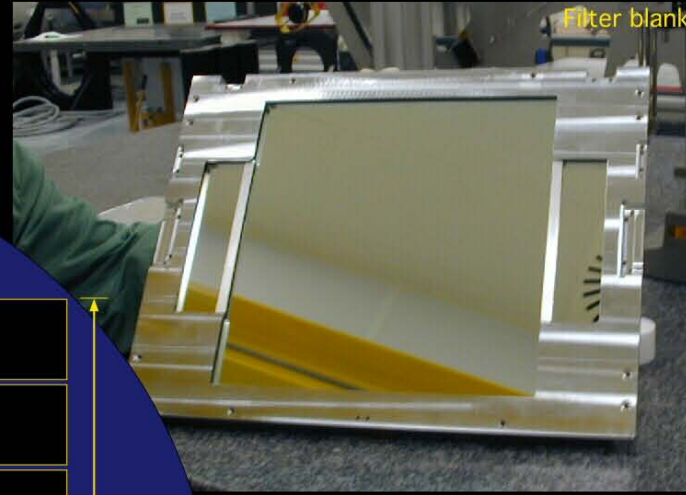
- Scope:** Sole VST instrument
- Top-level requirements:** Cover VST FoV (1 deg x 1 deg)
Critically sample best seeing on Paranal, WFS for AO
Optimize for UV sensitivity
Conform to VLT technical standards
- Filters:** SDSS ugriz, Johnson B & V, Strömgren u , H α (4 segments/ λ_0),
segmented calibration filter, public private filters.
Up to 2 x 6 filters can be mounted.
- Detector system:** 8 x 4 science mosaic of 2K x 4K e2v CCD44-82 devices
268 10⁶ 15 μ x 15 μ pixels (0.21 arcsec x 0.21 arcsec)
+ two 2K x 4K CCDs for autoguiding
+ two 2K x 4K CCDs for image analysis (AO)
- Shutter:** Dual blade, $t_{\text{exp}} \geq 0.1$ s

VST superimposed on its dome



OmegaCAM

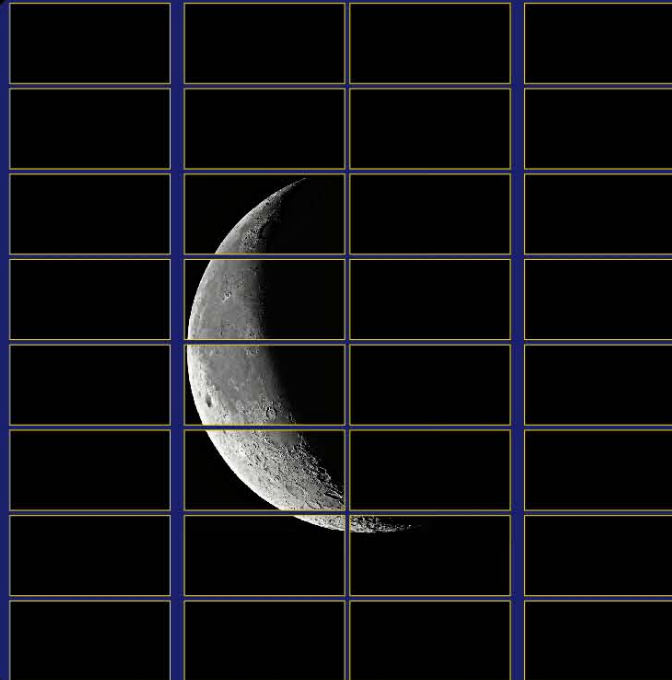
Filter blank



Guiding

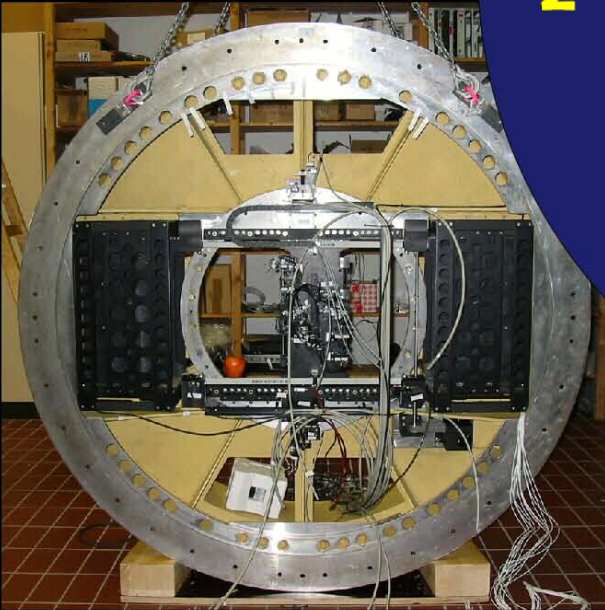
Image analysis

←
N



1°

A true scale image of **OmegaCAM** detector plane. 32 CCDs give a total imaging area of 10° with 256 Megapixels.



Camera mount in assembly.
An apple in the aperture is used for scale.

Shutter



OmegaCAM: Time Table

Announcement of Opportunity	03/1998
Letter of Intent	06/1998
MoU	01/2000
Kick-off meeting	04/2000
Conceptual Design Review	07/2000
PDR	12/2000
Specifications, SoW	11/2001
FDR	09+11/2001
Agreement	03/2002
PAE	04/2004 (DFS) + 06/2004 + 09/2005 (detector)
Packed and stored in Garching	01/2006 + 10/2006 (detector)
Re-assembly on Paranal	10/2008
Faulty CCD replaced	12/2009
Last filters shipped to Paranal	02/2010

OmegaCAM: Key Players

PI:

Co-I's:

Konrad Kuijken, Leiden

Ralf Bender, Munich

Enrico Cappellaro, Padua

Project manager:

Bernard Muschielok, Munich

**ESO Instrument
Responsible & Scientist:**

Dietrich Baade

Detector system:

Olaf Iwert, ESO

Mechanics:

Harald Nicklas, Göttingen

Electronics:

Achim Hess, Munich

Control software:

Andrea Baruffolo, Padua

Data flow software:

Edwin Valentijn, Groningen

Shutter:

Klaus Reif, Bonn

OmegaCAM: Contractual Matters

Price:

Fixed. Fully borne by OmegaCAM Consortium, incl. six person-years contributed to detector system

ESO undertakings:

Site, telescope, enclosure, infrastructure, operation

Guaranteed Time:

25% for 10 years (nominal)

20% for 10 years + 25 VLT-UT nights (contract)

?? (actual)

OmegaCAM: Design Challenges

Space between instrument flange and vertex of dewar entrance window

- 89 mm filter and shutter
- detector system deeply embedded in instrument (accessibility)

Mass limit

- in Announcement of Opportunity: 220-250 kg
- current allowances from VST: 687 kg (Cass. flange) + 300 kg (co-rotator)
- measured: 765 kg + 318 kg

Distribution of items on co-rotator

Large beam

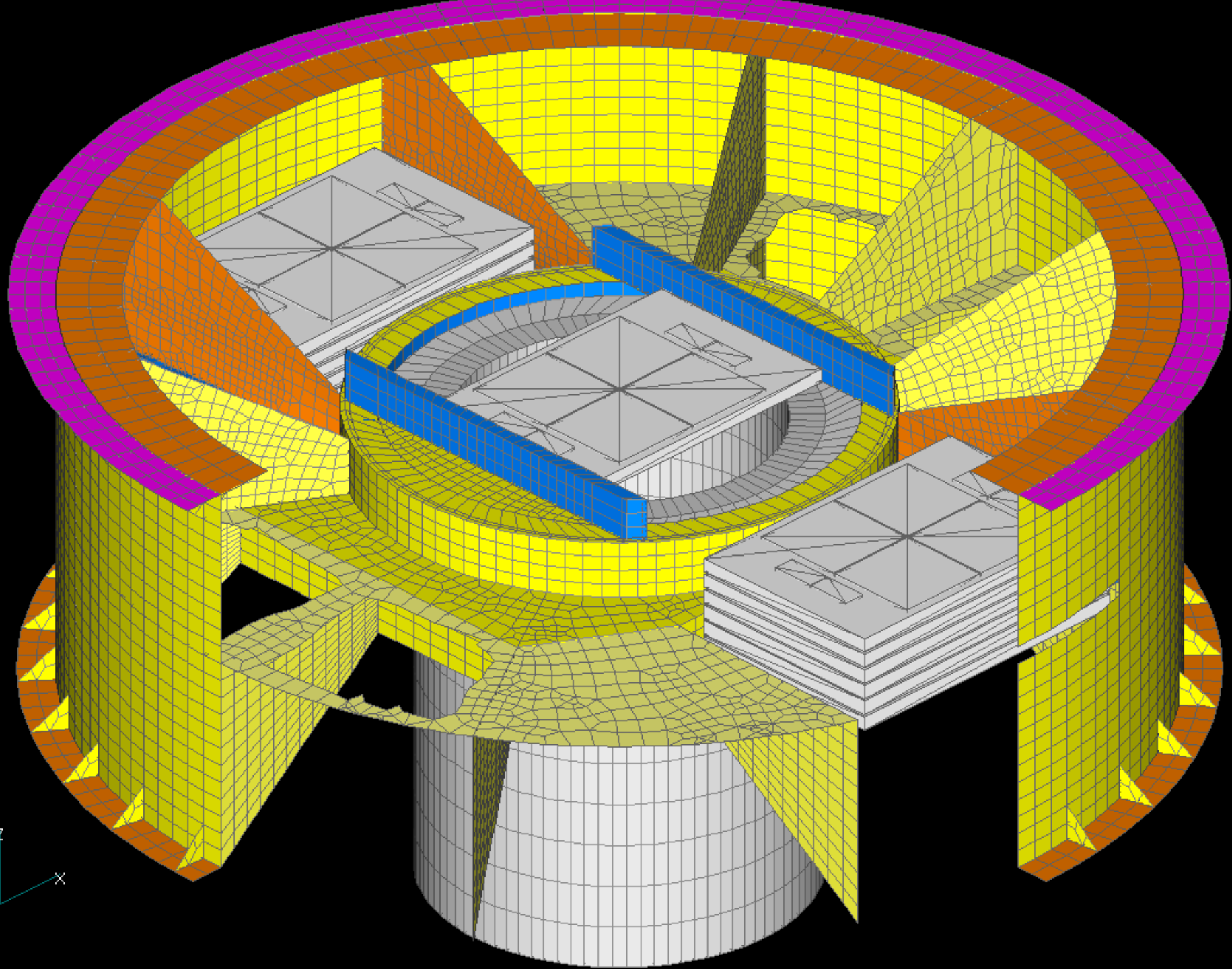
- Large shutter + large filters (274 mm x 274) mm + large mosaic

Detector system

- synchronization (<10 nsec) of 2 halves/controllers
- parallel operation of science mosaic, guider CCDs, and IA CCDs
- cooling

Long March (长征)

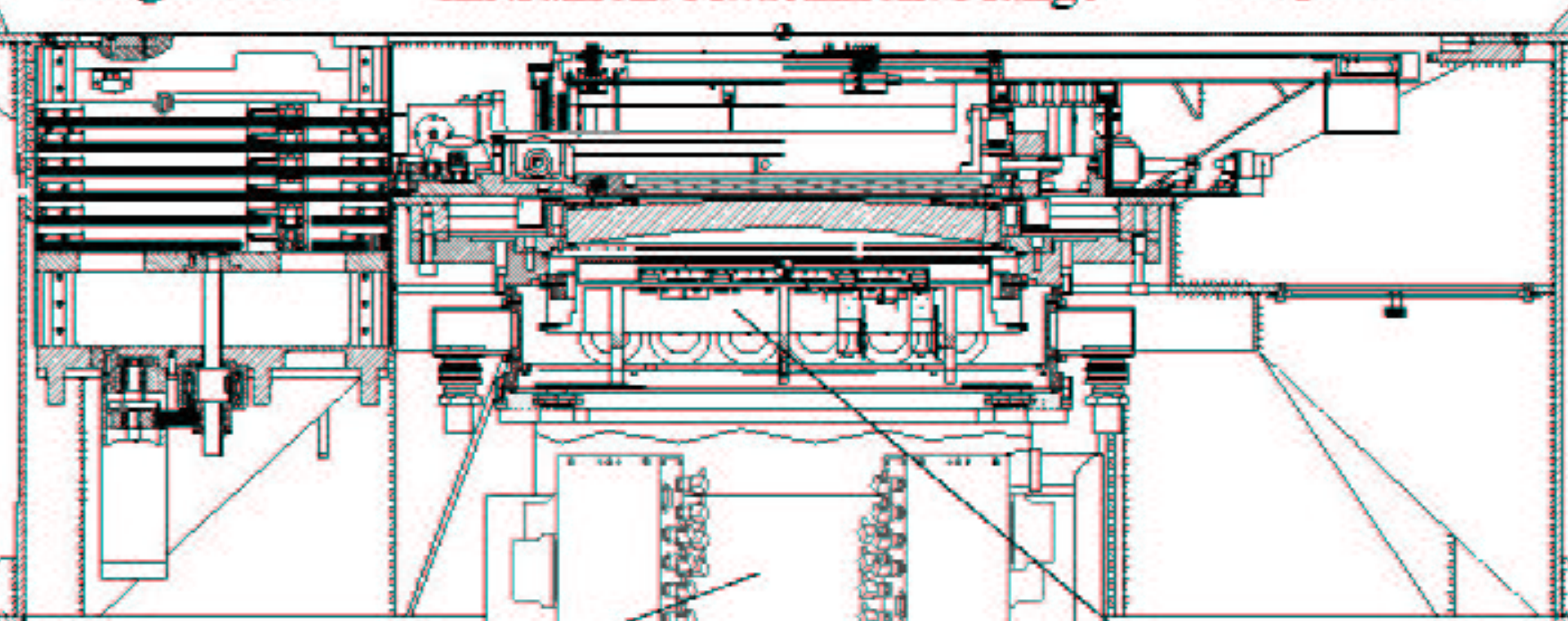
- people leaving & retiring
- fading memory



Filter Magazine A

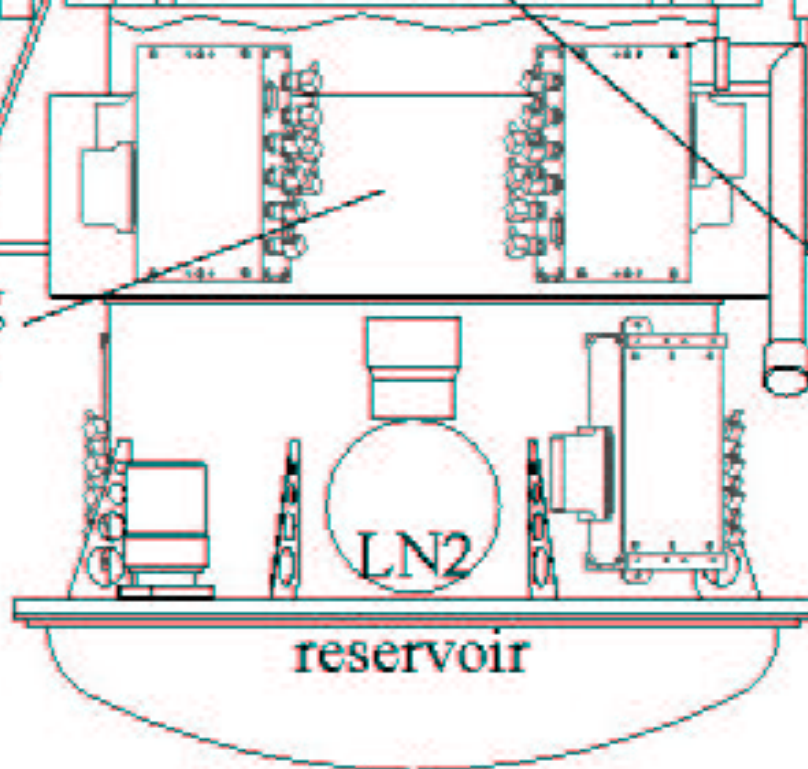
Instrument Attachment Flange

Filter Magazine B

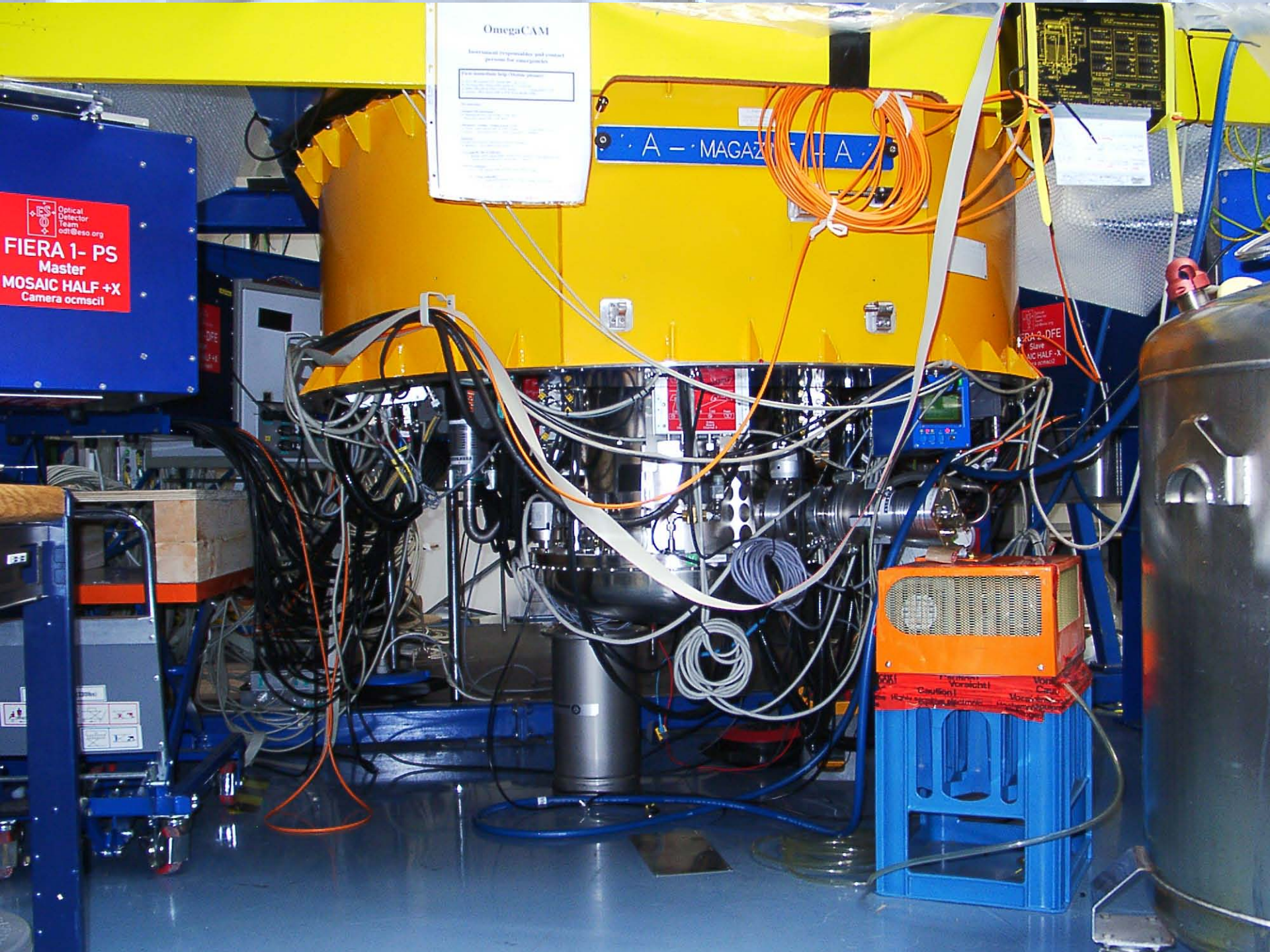


Cooling System

Detector Head



reservoir



OmegaCAM

Instrument responsible and contact persons for camera

For consultation only (Not for printing)

A - MAGAZIN - A

Optical Detector Team
odt@eso.org
FIERA 1- PS Master
MOSAIC HALF +X
Camera ocmsci

FIERA-2-DFE Slave
MOSAIC HALF -X
Camera ocmsci

Orange electronic device with a warning label:
Vorsicht!
Caution!
Wichtig!
Important!

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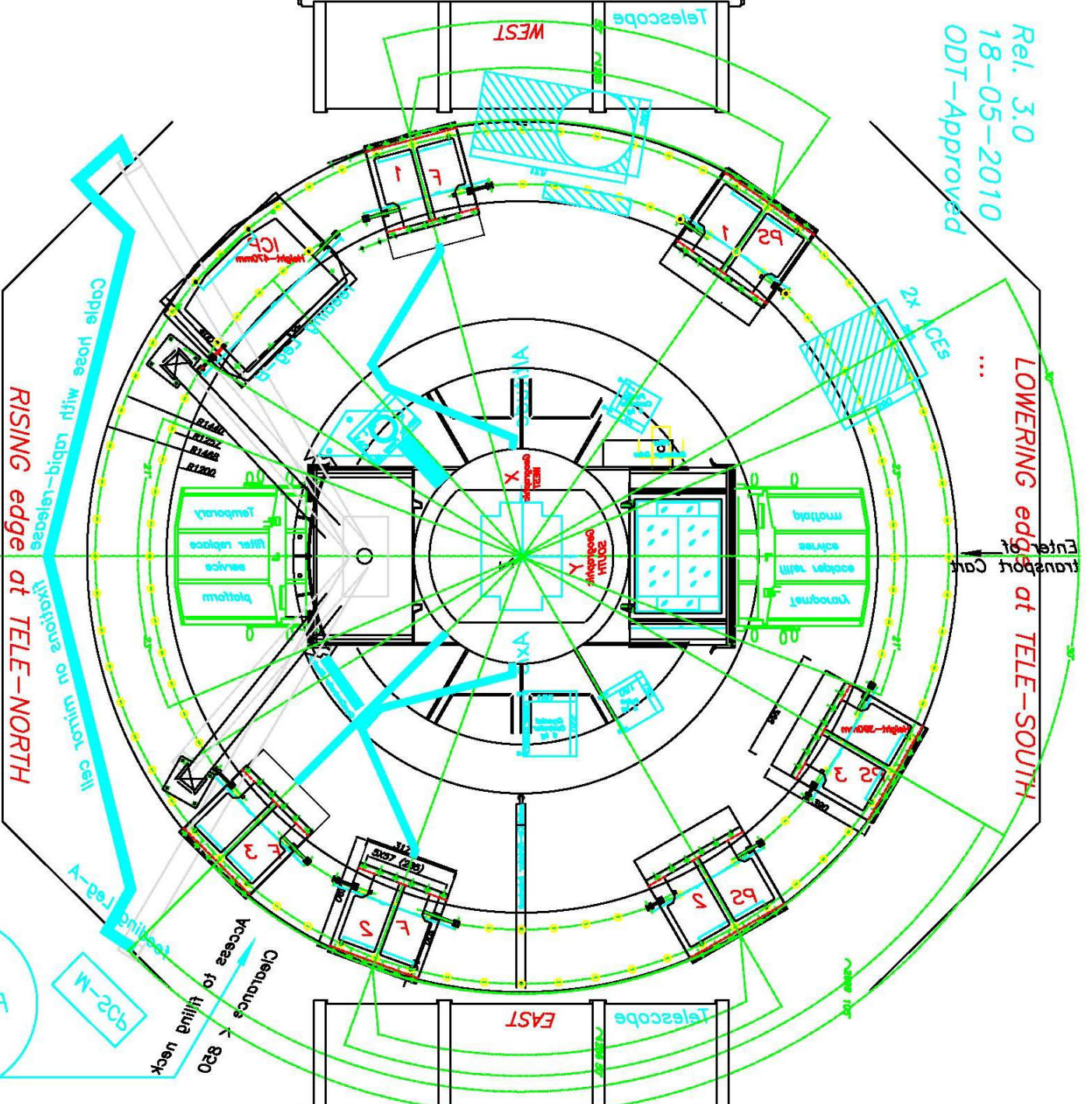
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Rel. 3.0
18-05-2010
ODT-Approved



Lowering edge at TELE-SOUTH
Rising edge at TELE-NORTH
BOTTOM VIEW from floor

Re-filling tank
LMS

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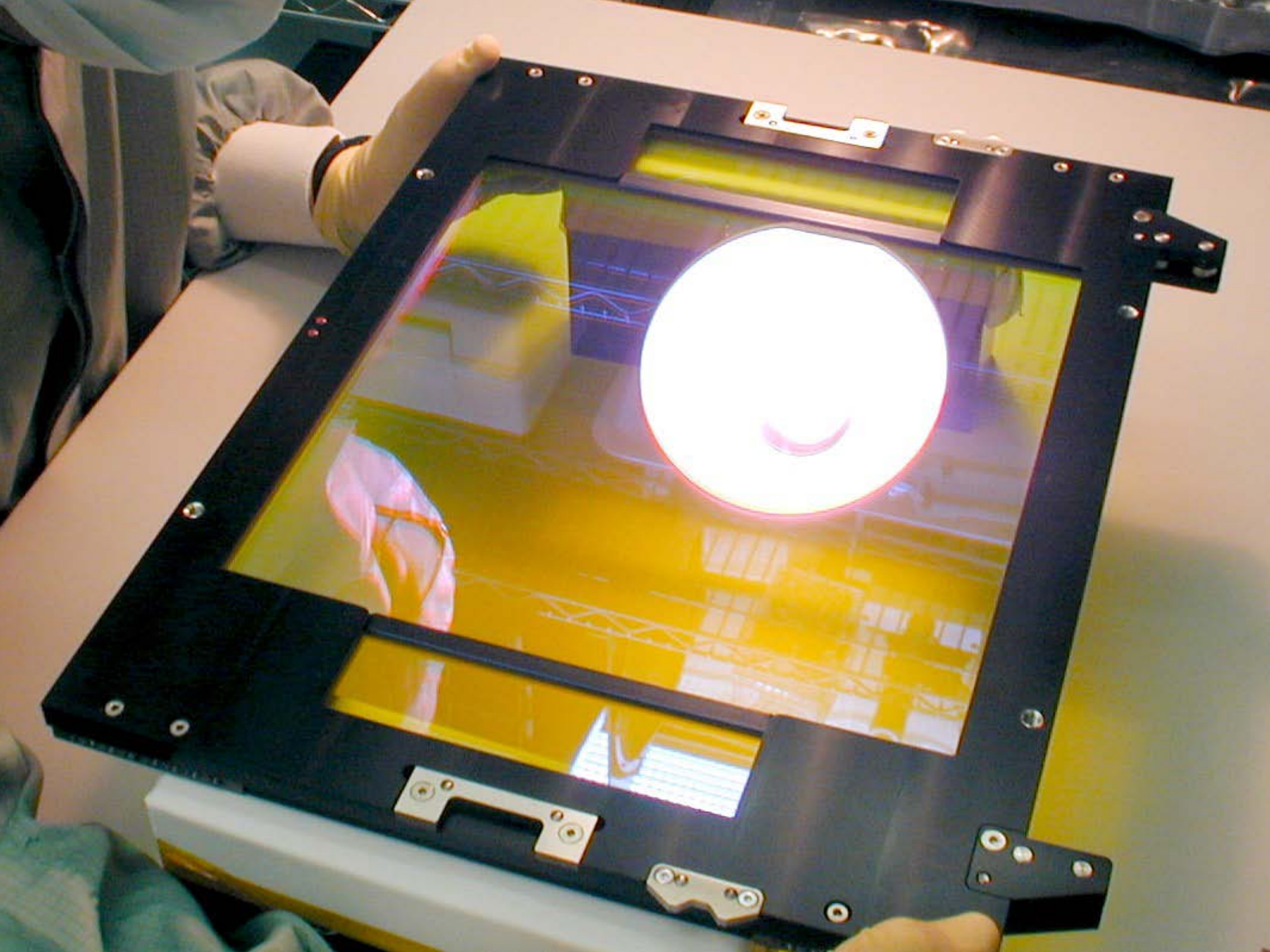
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Long March (長征)

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OmegaCAM: Performance in the Lab

Differential flexure: $< 3 \mu$

Mosaic readout: RON $< 5 \text{ e}^- @ 30 \text{ s}$

Nominal zeropoints

Filter	mag (AB)
u'	24.9
g'	26.4
r'	26.1
i'	25.6
z'	24.3

Nominal detection limits

Filter	V mag of A0 star
u'	22.6
g'	24.4
r'	24.2
i'	23.5
z'	22.2

Filter change:

Calibration unit:

Control S/W I/F to VST:

Pipeline throughput:

Assumed conditions:

Exp. Time	300 s (single)
S/N	5
Seeing	1.0 arcsec
Age of moon	7 days
ADC	used

no longer an issue

OmegaCAM: Status

Overall:	pre-assembled, tested OK, dismantled, and safely stored on Paranal, <u>ready for commissioning</u>
Filters:	received, tested, accepted
Detector mosaic:	faulty CCD replaced on Paranal, new one tested OK
On-line Data Flow:	not tested with instrument but reported OK
Electromechanics:	excellent reliability
Handling tools:	successfully tested
Documentation:	mostly very good
Standard stars:	TBD within one year after first light

OmegaCAM: Where does it stand relative to others?

Growth of CCD mosaics

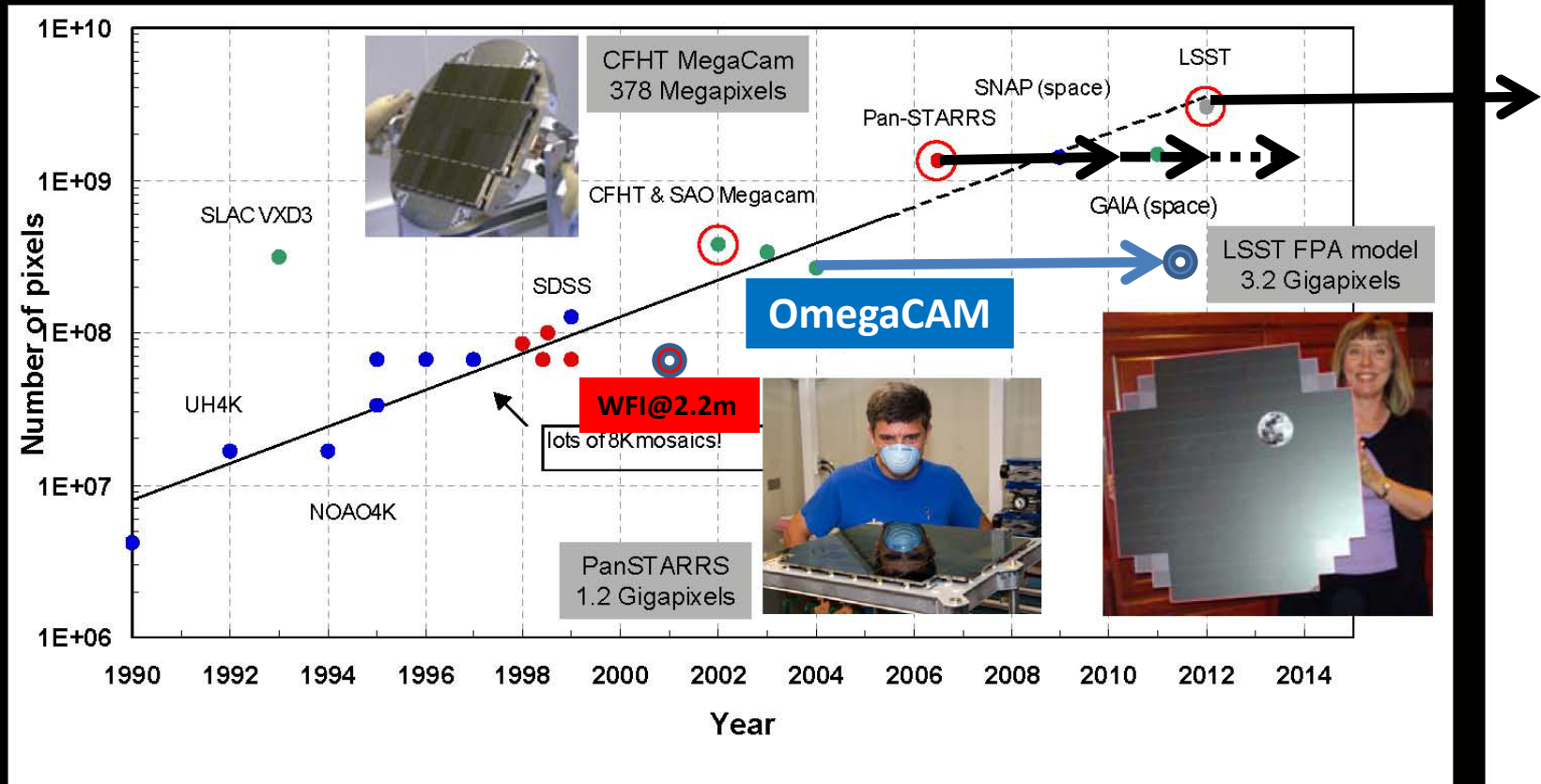


Illustration of large focal plane sizes, from Luppino 'Moore's' law

Focal plane size doubles every 2.5 years

OmegaCAM: Commissioning Challenges

- **Interfaces VST/OmegaCAM**
- **Combination of new telescope and new instrument**
- **Optical alignment**
- **Image analysis and active optics (open/closed loop)**
- **Near-time data processing (bandwidth to Europe)**
- **Detector safety**
- **On-site data flow system**
- **Correction for sky concentration**
- **Secondary standard stars**
- **Overweight**