The LINER NGC 7097: the AGN and its mirror

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Acknowledgement: Felipe A. Oliveira

Central Massive Objects, ESO Garching, 22-25 June 2010

Introduction

- Goal: Search for low luminosity AGN in early type galaxies
- Data cube observed with GMOS-IFU on Gemini South Telescope
- Apply PCA tomography on the data cube in order to detect and characterize LLAGN.

PCA tomography

- Search for spectro-spatial variance in the datacube.
- Not a set of objects; a set of spatial pixels of the same data cube. The wavelength pixels are the properties.
- Linear transformation to a new system of coordinates.
- The coordinates are orthogonal.
- Eigenspectra Show correlations between the wavelengths.
- Tomograms Show where the correlations appear in (x,y) space. Tomograms are the projections of the data on the eigenvectors; "slices" of the data in the eigenvector space.
- Most useful information in a few eigenvectors. Dimensional reduction.
- See Steiner et al 2009, MNRAS, 395, 64

NGC 7097: Eigenvalues



<u>Scree test</u> – Indicates that until Eigenvector 8 we have useful results



NGC 7097: Eigenvector 1 (99.52%):

Mainly the galaxy Bulge



NGC 7097: Eigenvector 2 (0.38%):

LINER signature. Red component correlated with the emission line features. Interstellar gas (Na I) in the line of sight to the AGN.





NGC 7097: Eigenvector 3 (0.046%):

Positive correlation is the red component of the line features. Negative correlation is the blue component of the line features. Rotating disc



NGC 7097: Eigenvector 4 (0.021%):

A second LINER, but bluer than the one found in Eigenvector 2 and displaced 0.15" eastward. Absence of Na I absorption.

Same redshift as the first one.





NGC 7097: Eigenvector 5 (0.0059%):

• The central component of the emission line features is anti-correlated with their blue and red components.

• Ionization cone (correlated) and rotating disc (anti-correlated).









CONCLUSION

- The LINER is viewed directly through the disc, reddened by dust.
- The same LINER is reflected by the clouds in the ionization cone.
- We predict that the blue component is polarized.