# Stellar population properties of nuclei in early-type galaxies:

the case of NGC 1428 (FCC277)

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in collaboration with

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## Overview of the project

- →~75 % of all types of galaxies host a NSC(e.g. Côté et al. 2006, and ref. therein )
- Their morphology and luminosity classified
- Lack of kinematics and chemical data about nuclei in early-type galaxies

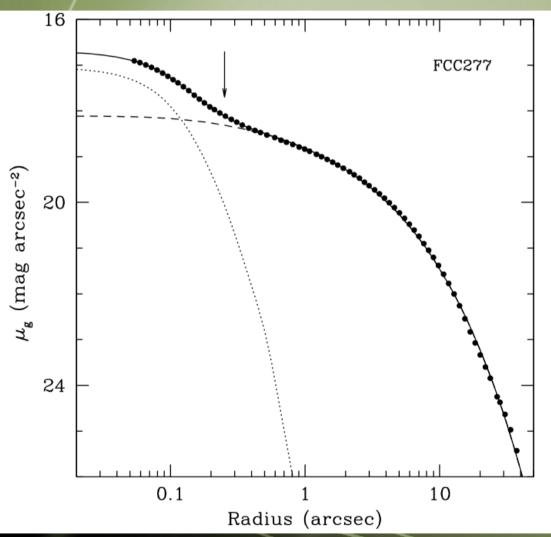
## NGC 1428 (FCC 277)

**NSC** 20"

Nucleated E5in Fornax cluster

part of ACS/FCS(Jordán et al. 2007)

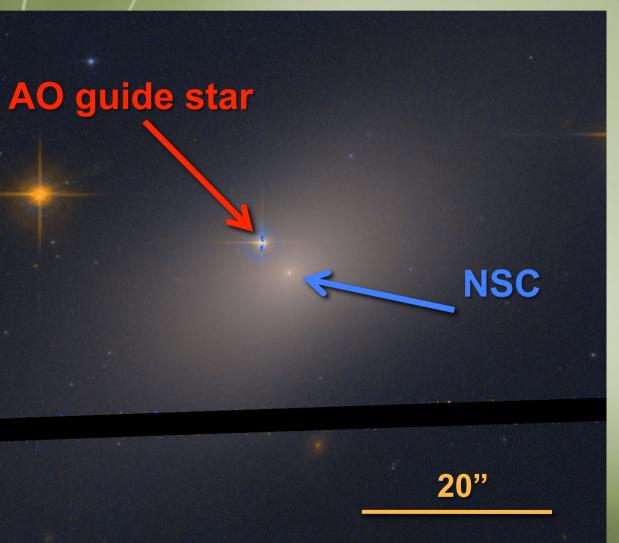
## What makes the nucleus of NGC 1428 different?



NSC r<0.25"

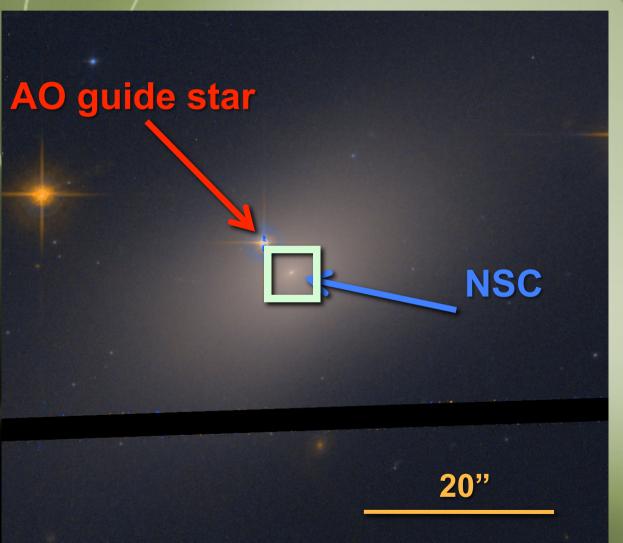
Côté et al. 2007

## NGC 1428 (FCC 277)



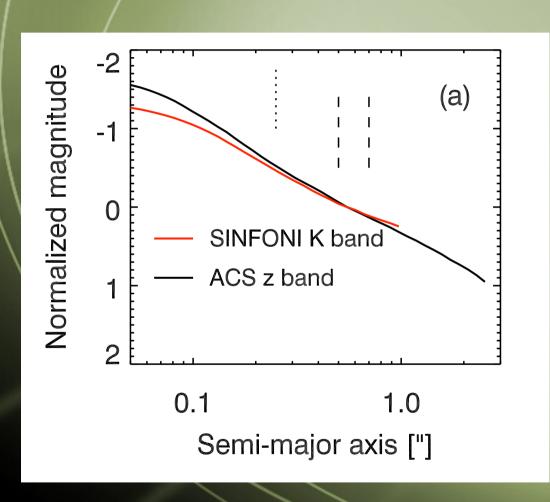
- Nucleated E5in Fornax cluster
- part of ACS/FCS(Jordán et al. 2007)

## NGC 1428 (FCC 277)

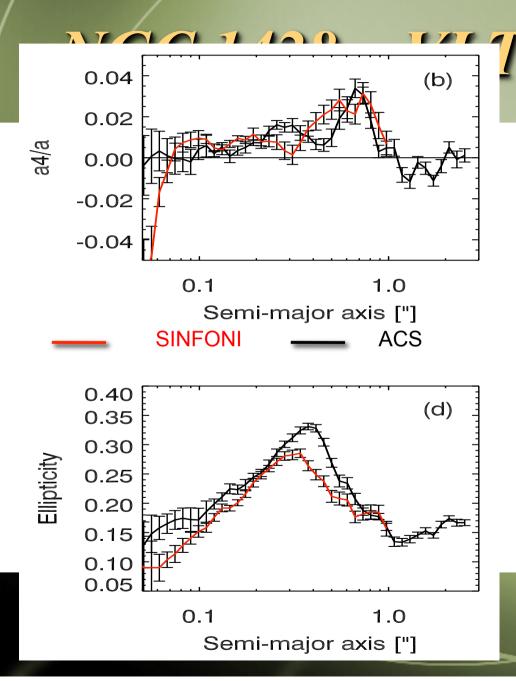


- Nucleated E5in Fornax cluster
- part of ACS/FCS(Jordán et al. 2007)

#### NGC 1428 – VLT Observations



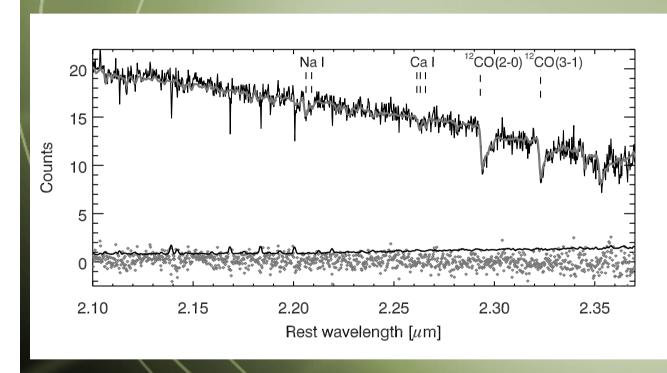
- **+ VLT/SINFONI**
- **♦NGS AO**
- **♦3"x3"** FoV
- \*0.17" (FWHM) spatial resolution



#### T Observations

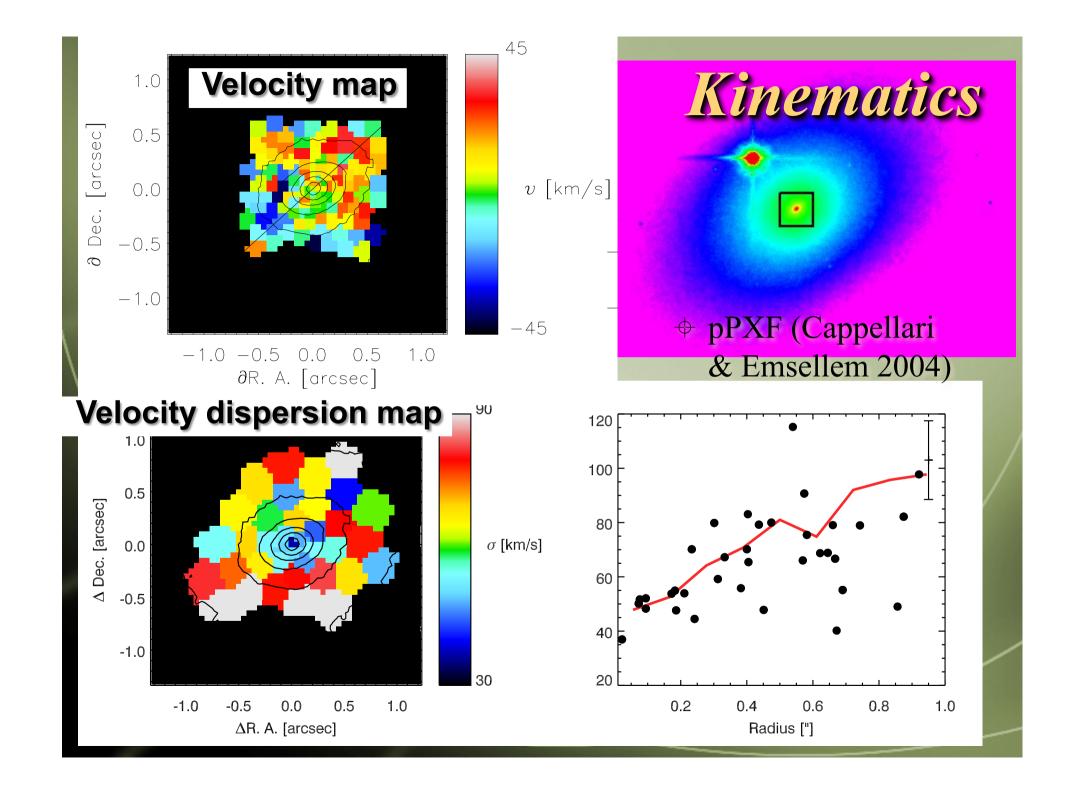
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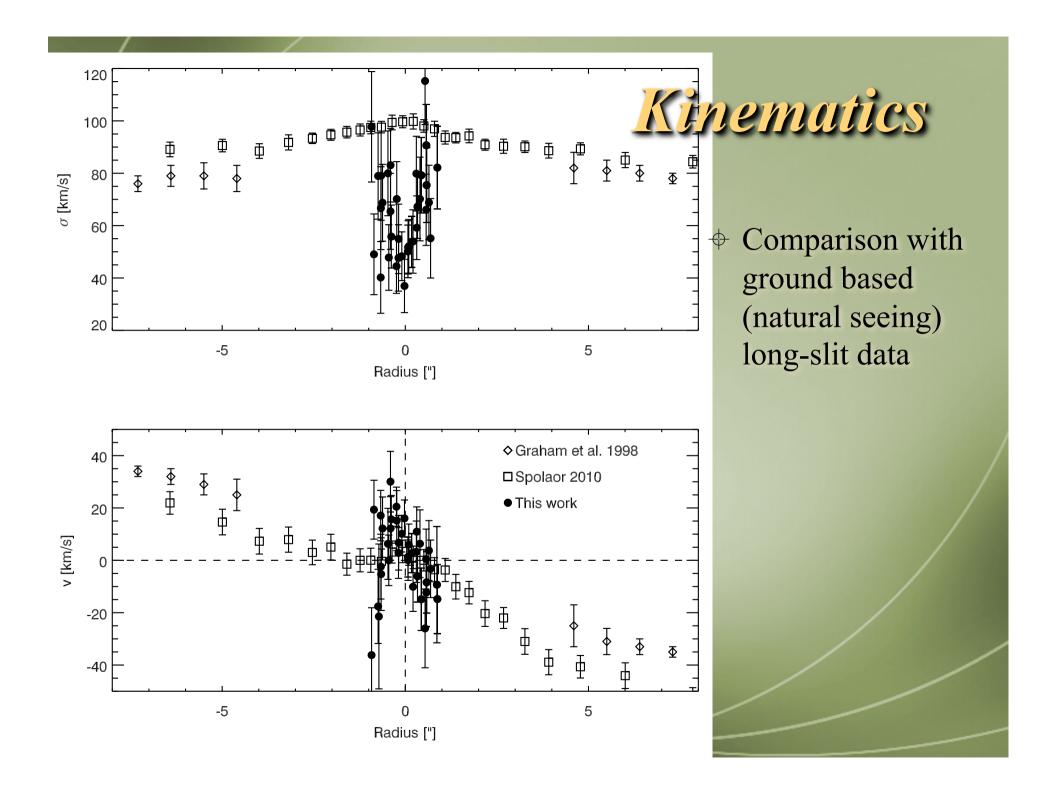
## K-band spectra



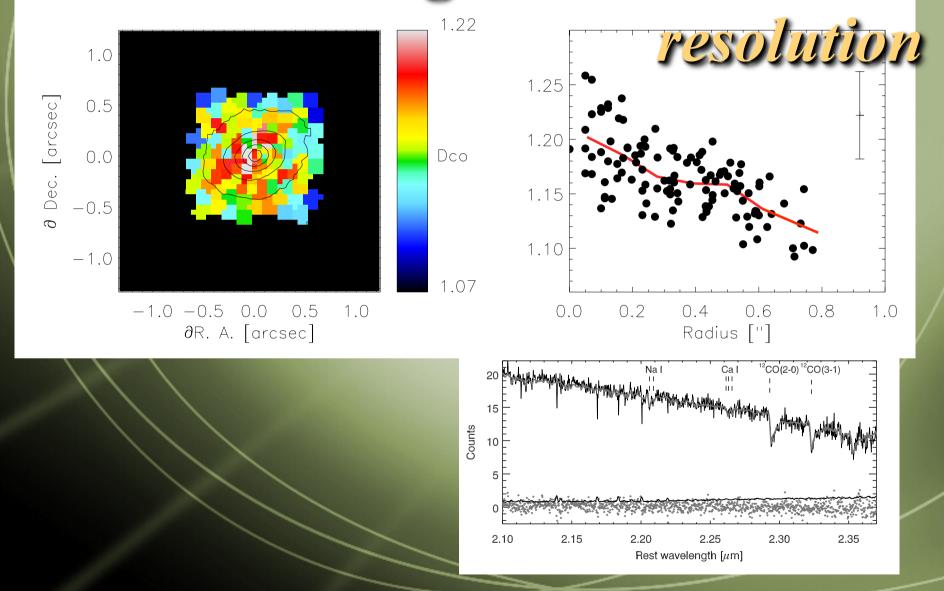
Light is dominated by old stars

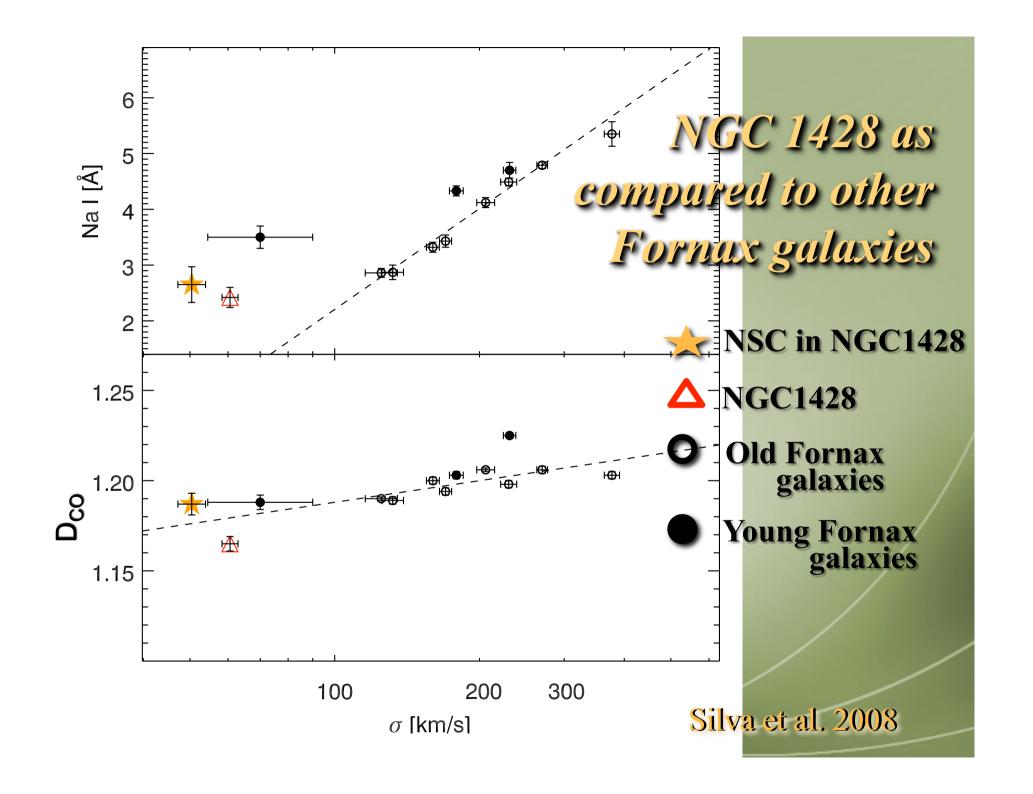
 $\Phi$  S/N  $\sim 15$ 



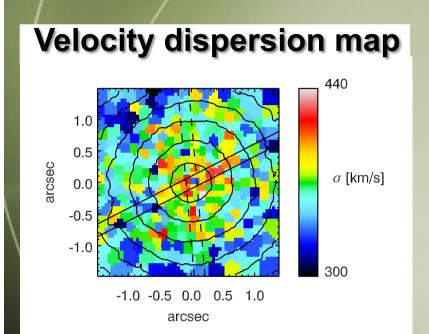


## Line strengths with HST-like

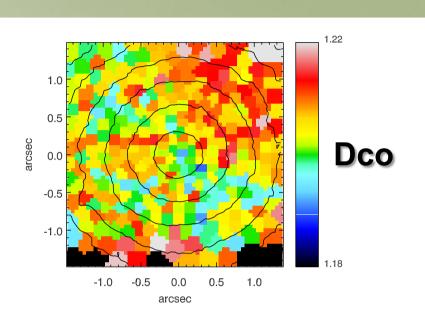


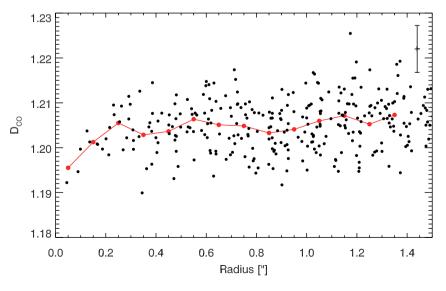


### NGC 1399 – cD in Fornax









## Concluding remarks

- ♦ We are able to resolve the nucleus of NGC1428 (FCC277) – differences in kinematics and line strengths
- Do we see several nested structures at r<1"?
- Can we determine ages and metallicities from near-IR spectra?
- Near-IR diagnostic tools highly needed: AO in Natural Guide and/or Laser Guide Star