



*A MAD view on the M16
elephant trunks*

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in collaboration with:

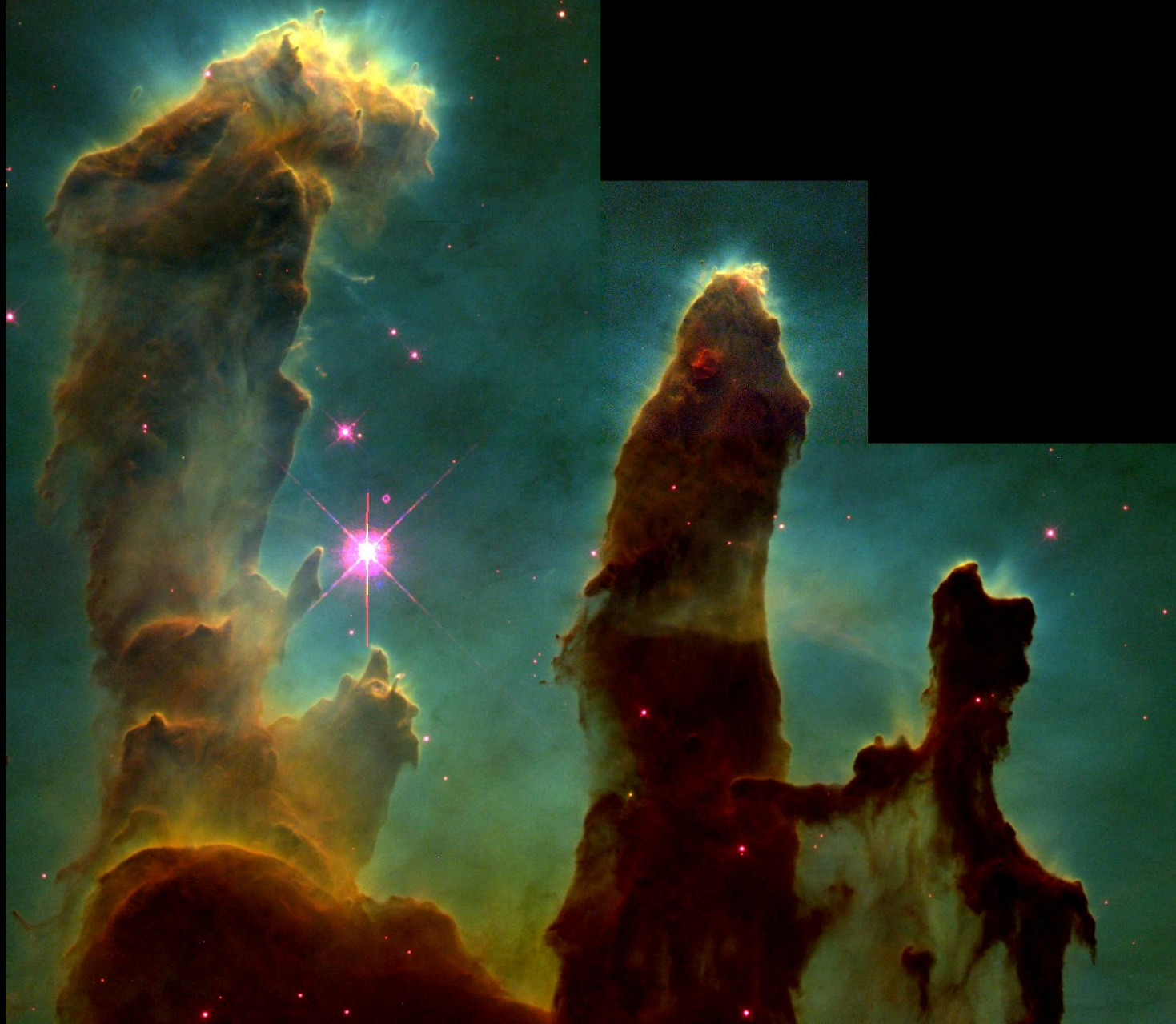
Mark McCaughrean (University of Exeter)

M16 (The Eagle Nebula) & NGC 6611



© Panther Observatory

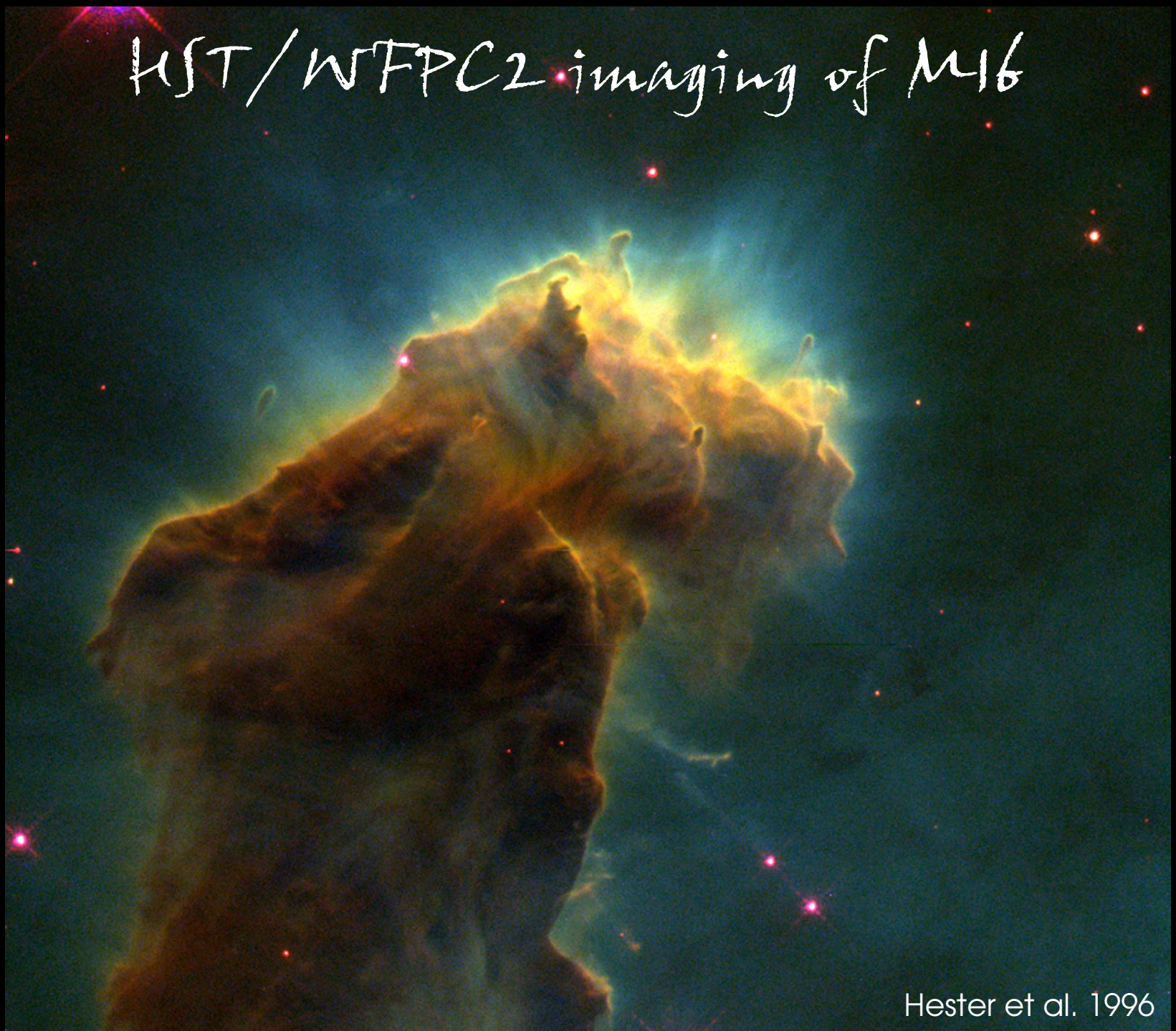
HST/WFPC2 imaging of M16



1995: HST/WFPC2
imaging in
 $H\alpha$, O(III), and S(II)

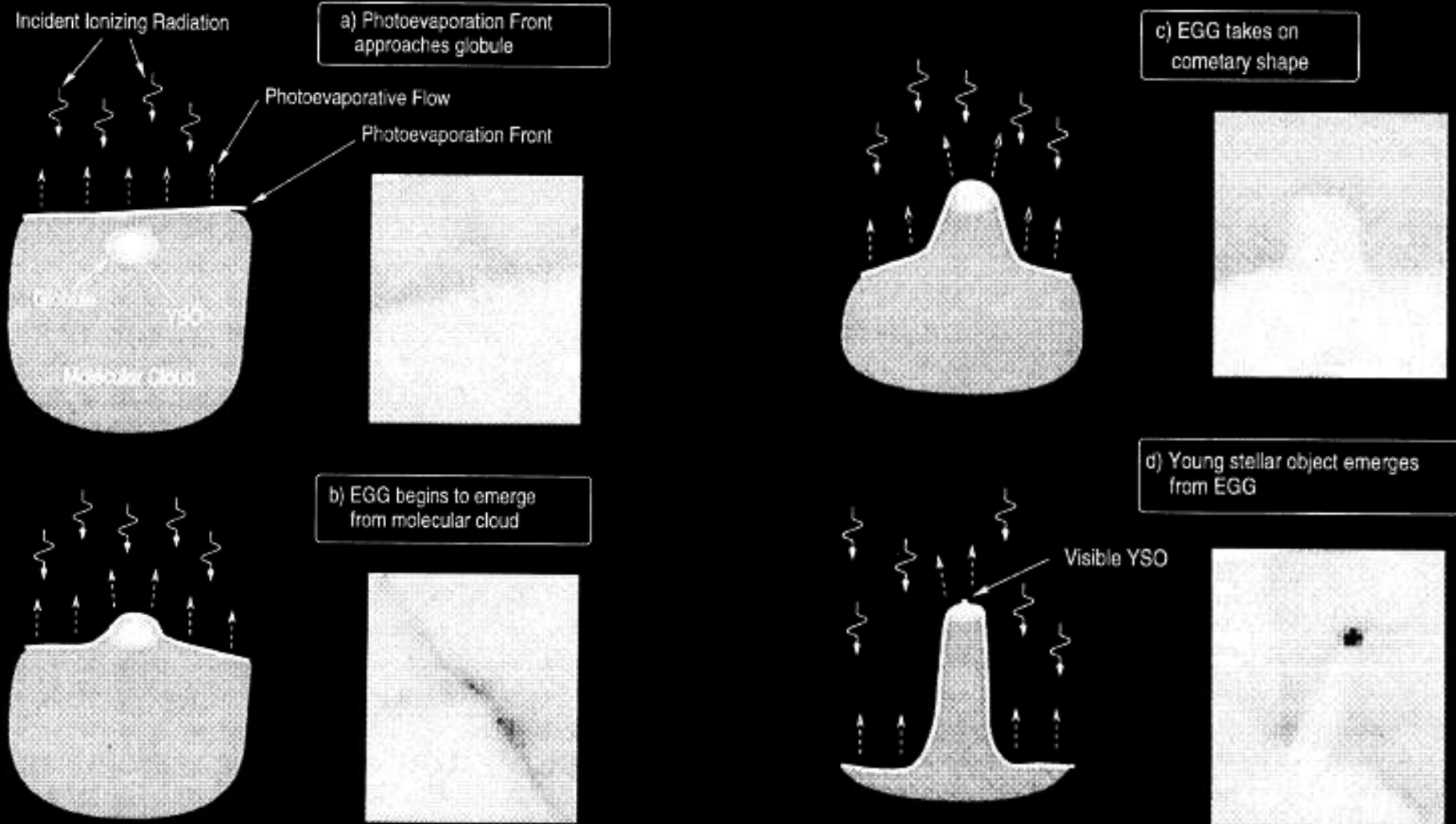
Hester et al. 1996

HST/WFPC2 imaging of M16



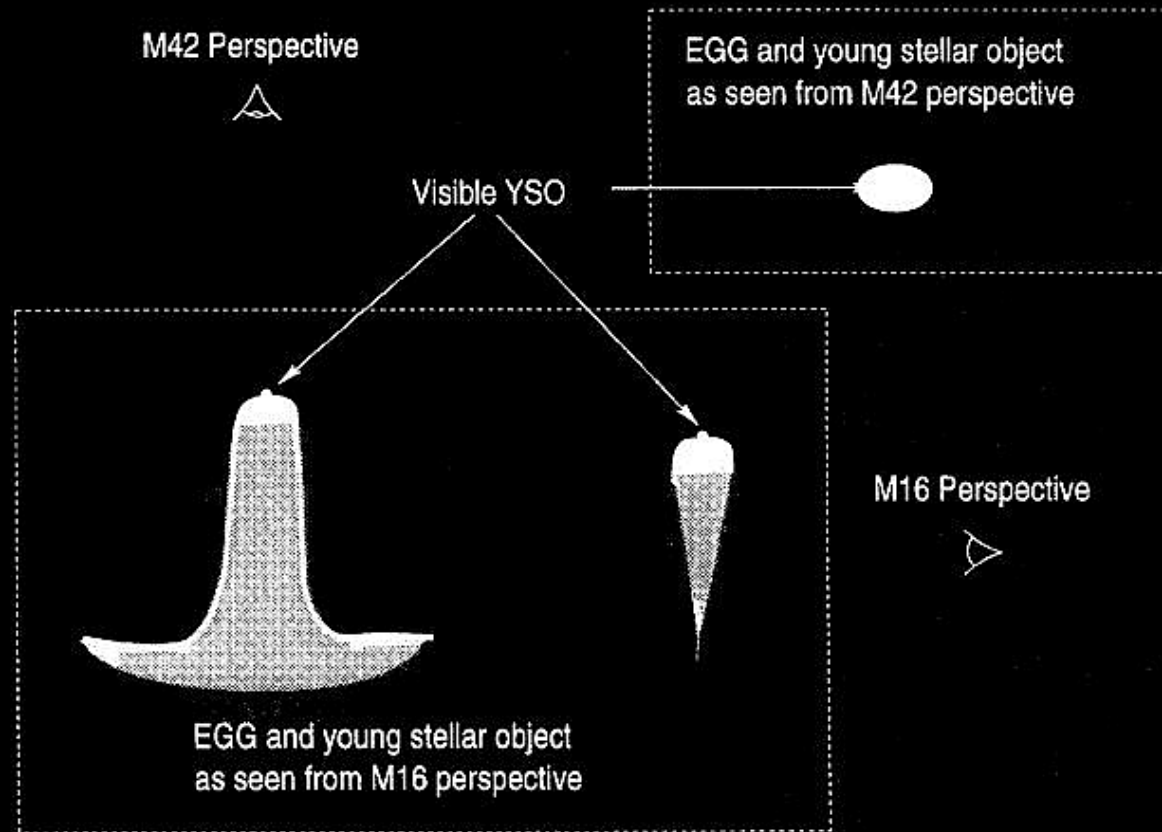
Hester et al. 1996

HST/WFPC2 imaging of M16



Hester et al. 1996

HST/WFPC2 imaging of M16



Hester et al. 1996

NIR imaging surveys in M16

Current benchmark in coverage, depth, and spatial resolution:

McCaughrean & Andersen 2002

ISAAC JHK mosaics of 9' x 9' FOV

Spatial resolution 0.3''-0.4'' FWHM

Depth (100s integration time) ~ 22.6^m – 20.4^m (JHK)



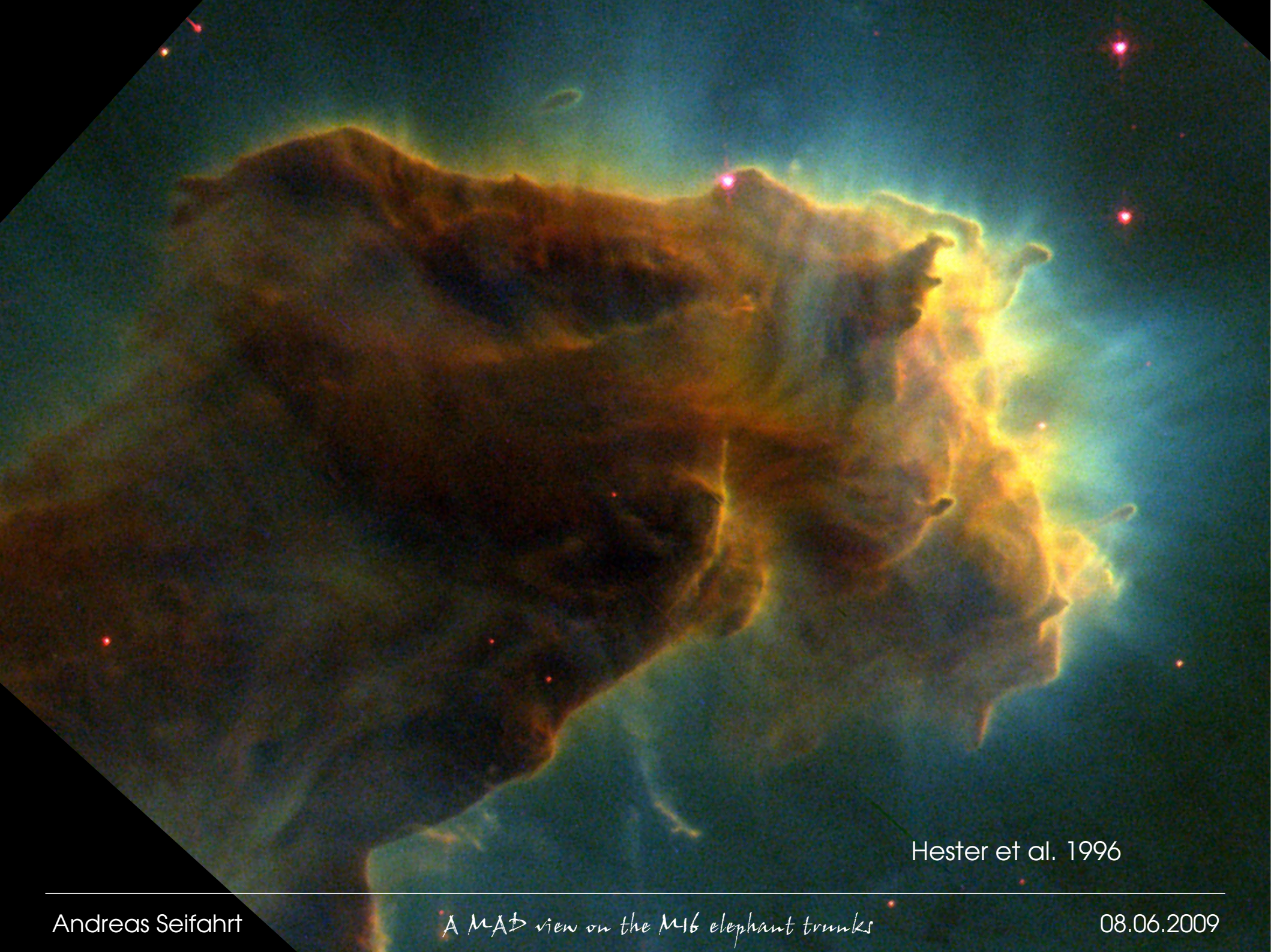
McCaughrean & Andersen 2002



McCaughrean & Andersen 2002



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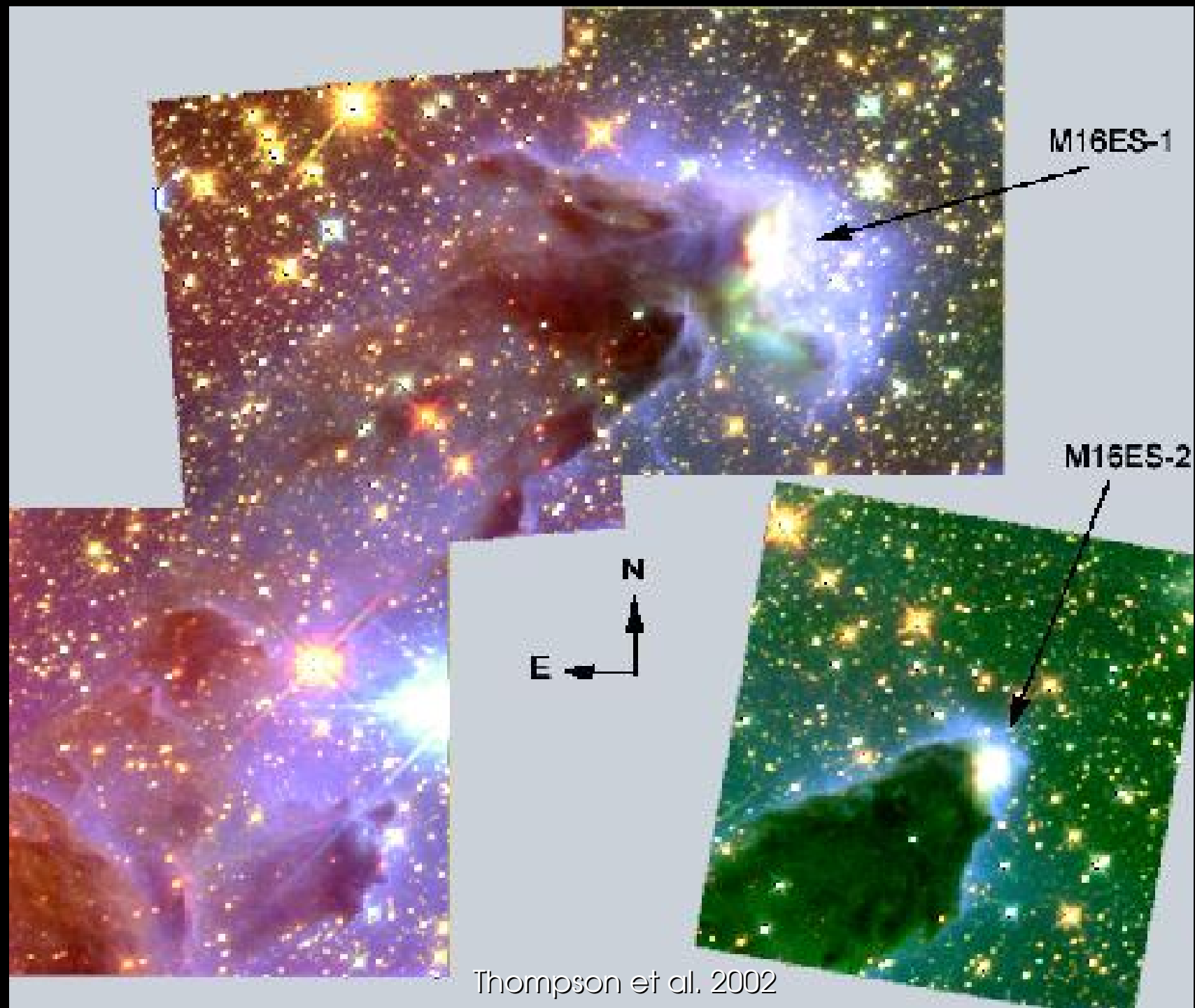
Spatial resolution 0.3''-0.4'' FWHM

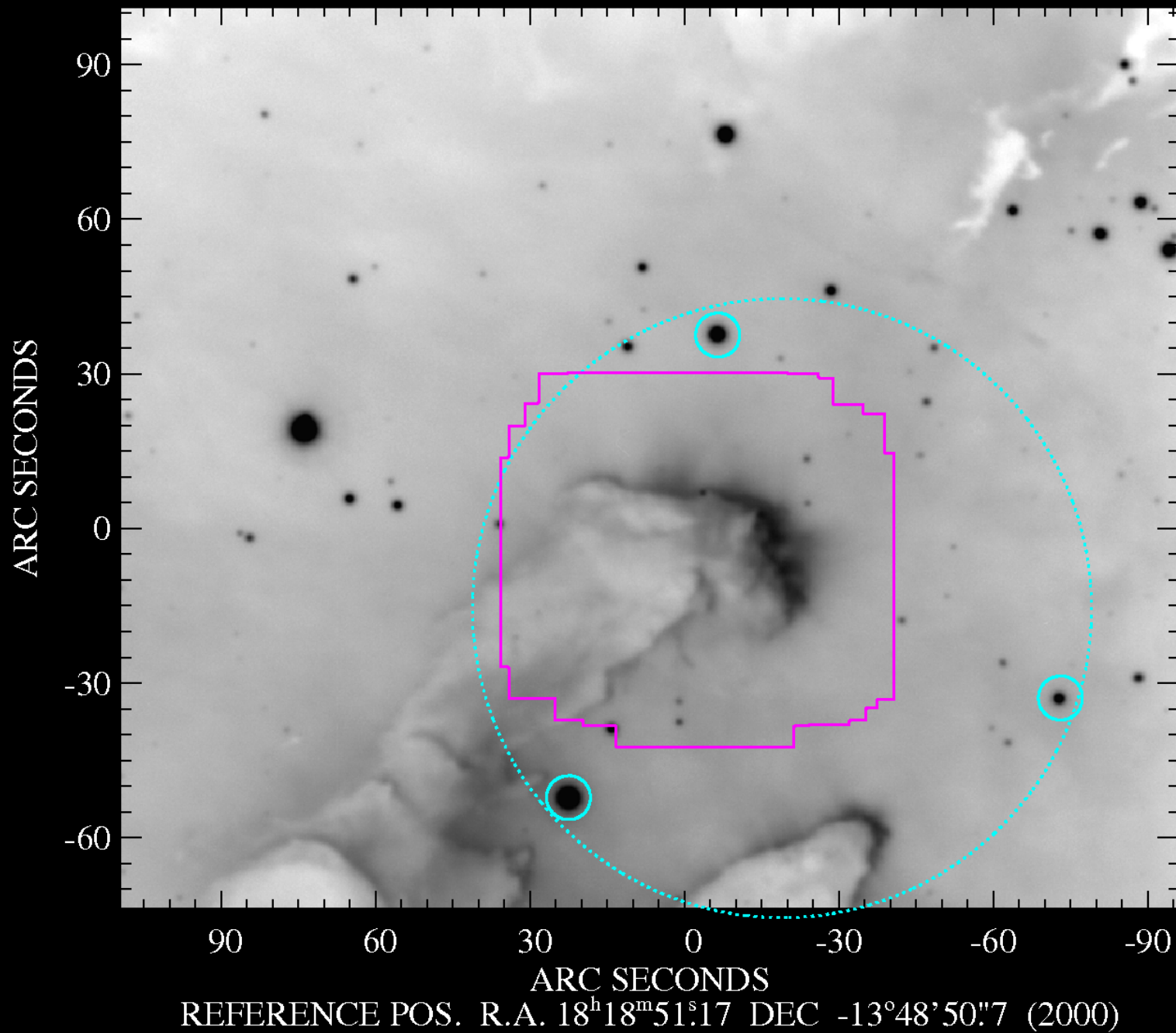
Depth (100s integration time) ~ 22.6^m – 20.4^m (JHK)

Result:

Only ~10% of all EGGs contain a protostellar source
(as seen in JHK)

Problem: Depth and Confusion with background sources





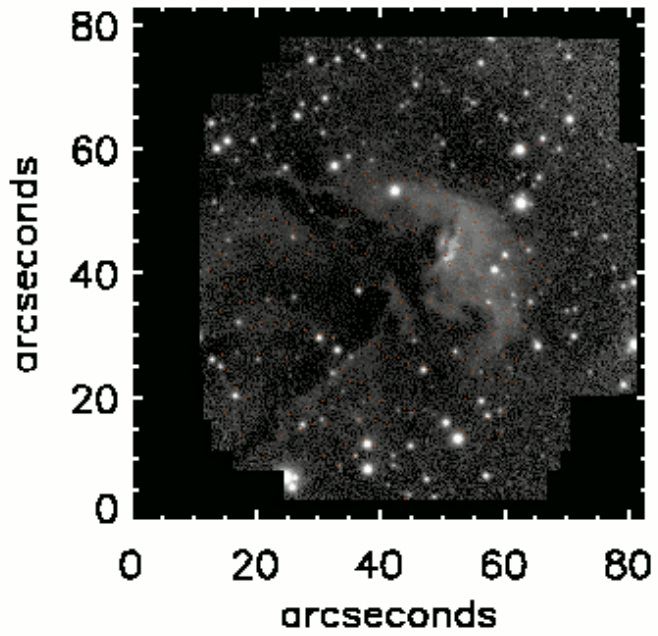


MAD sciene demonstration data
(JHK colour mosaic)

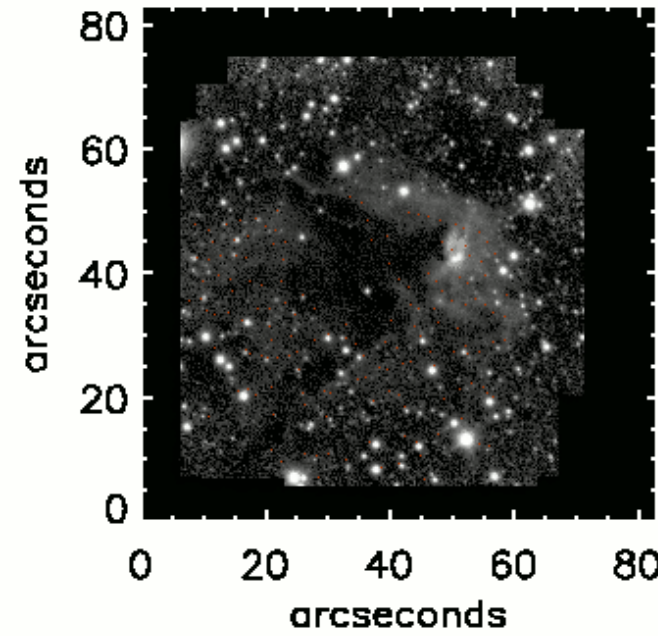


ISAAC JHK colour mosaic
McCaughrean & Andersen 2002

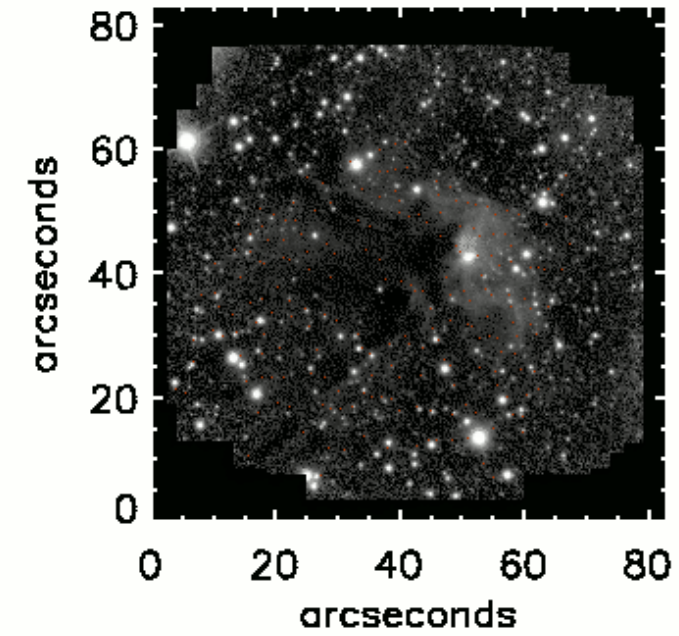
VLT MAD J-Band



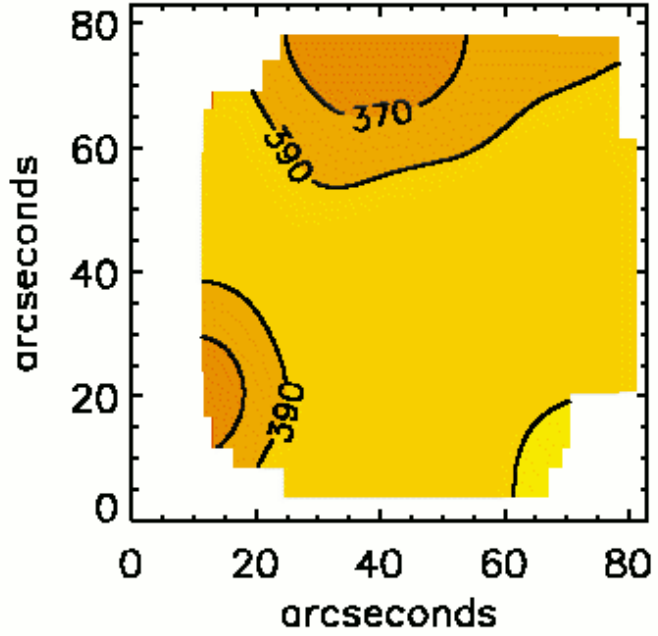
VLT MAD H-Band



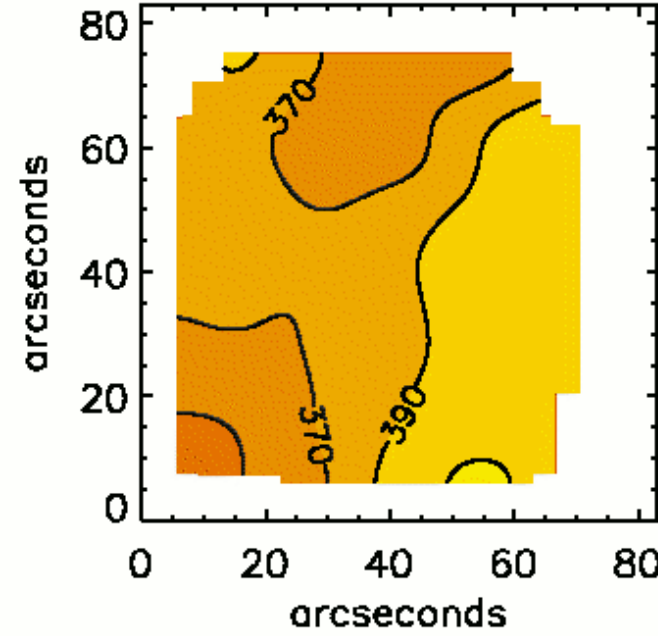
VLT MAD Ks-Band



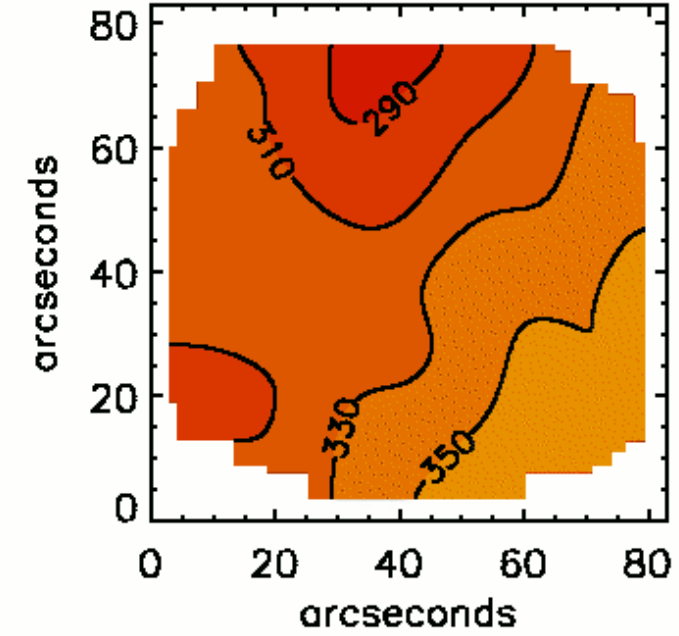
FWHM (mas)

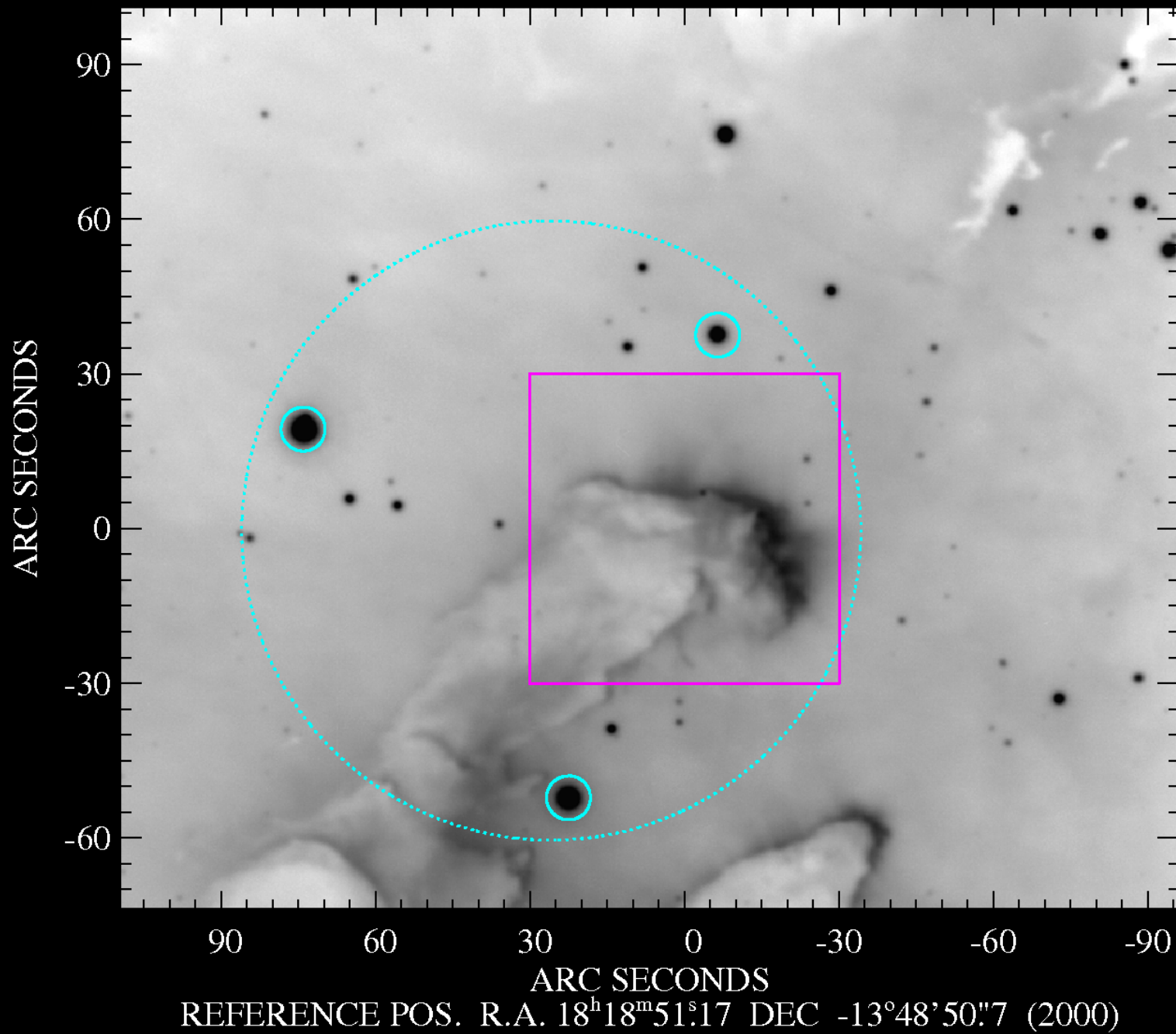


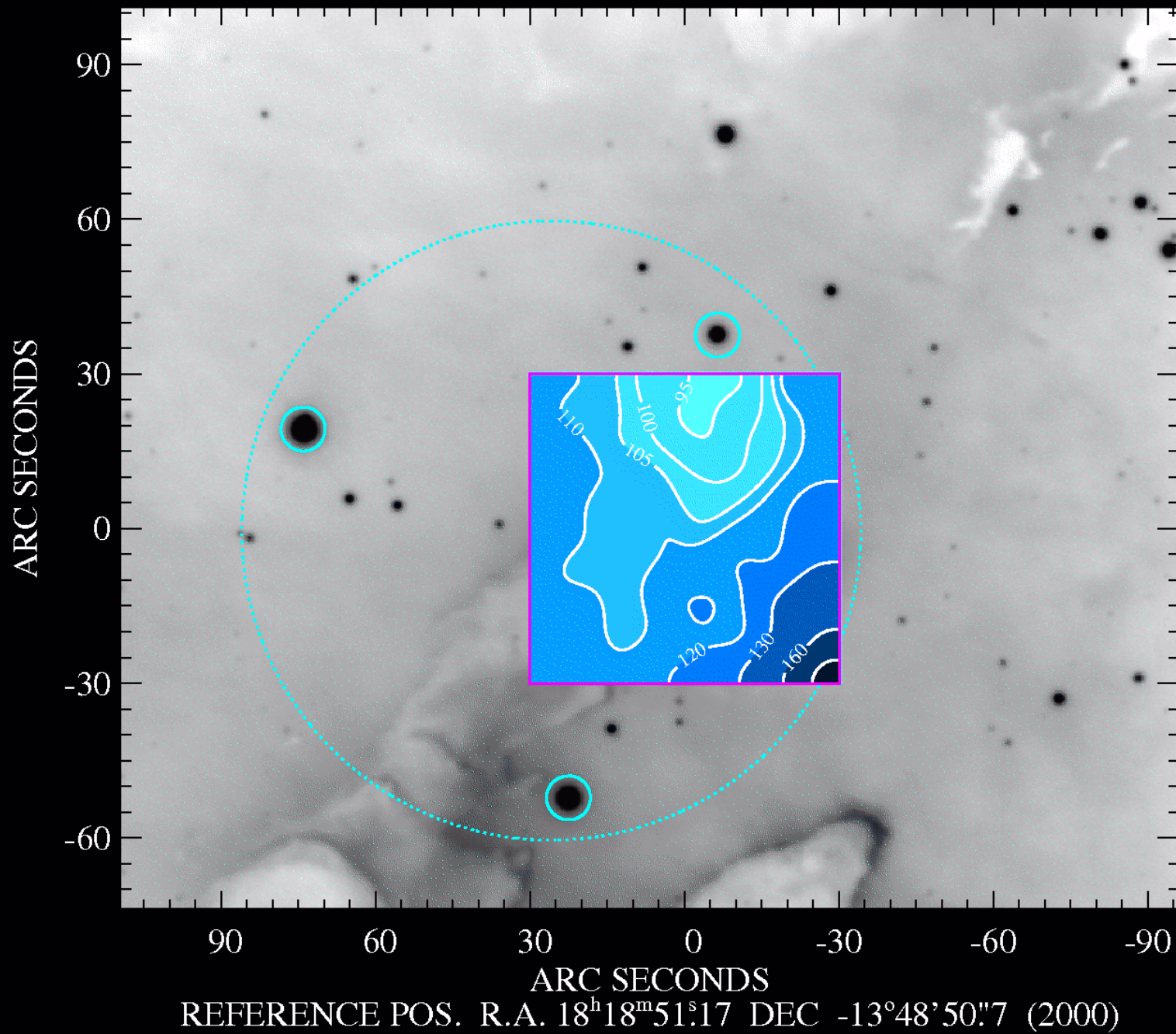
FWHM (mas)



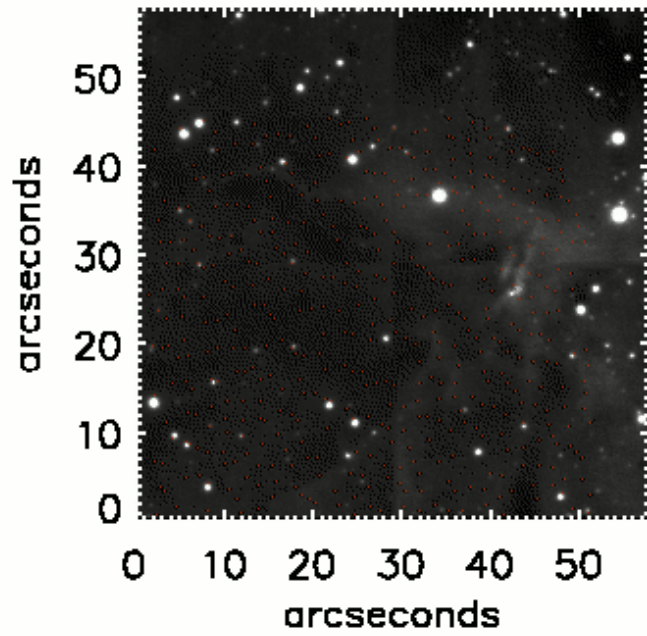
FWHM (mas)



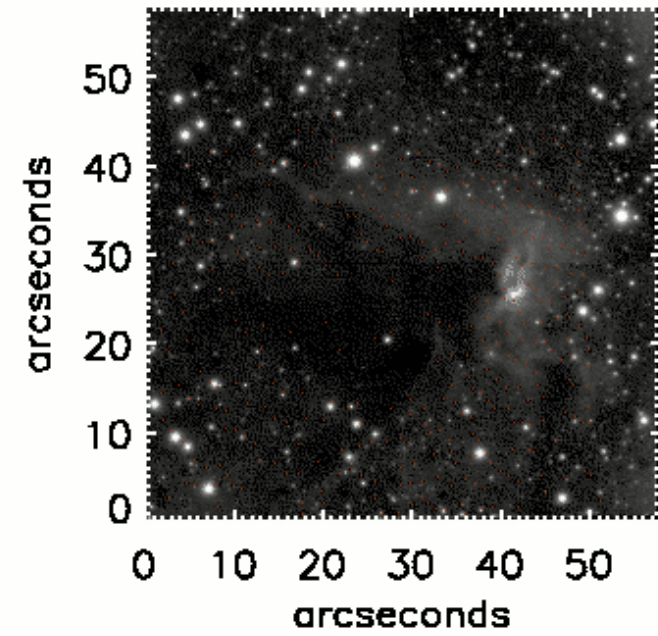




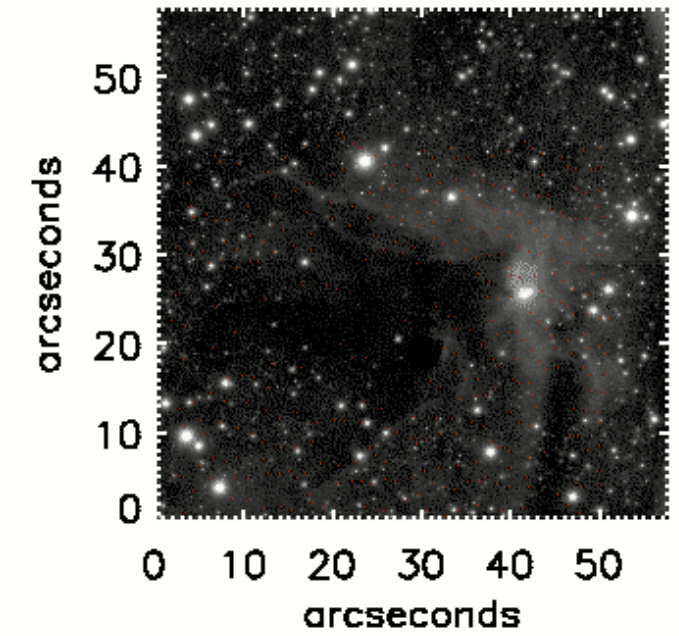
VLT MAD J-Band



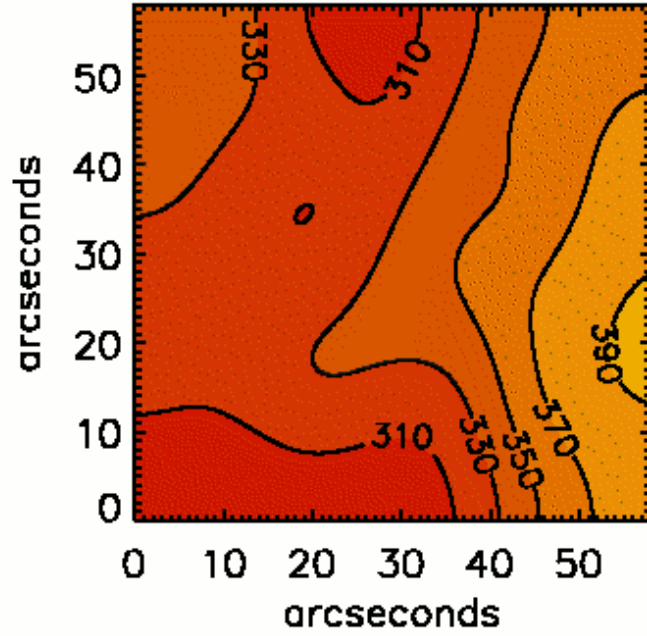
VLT MAD H-Band



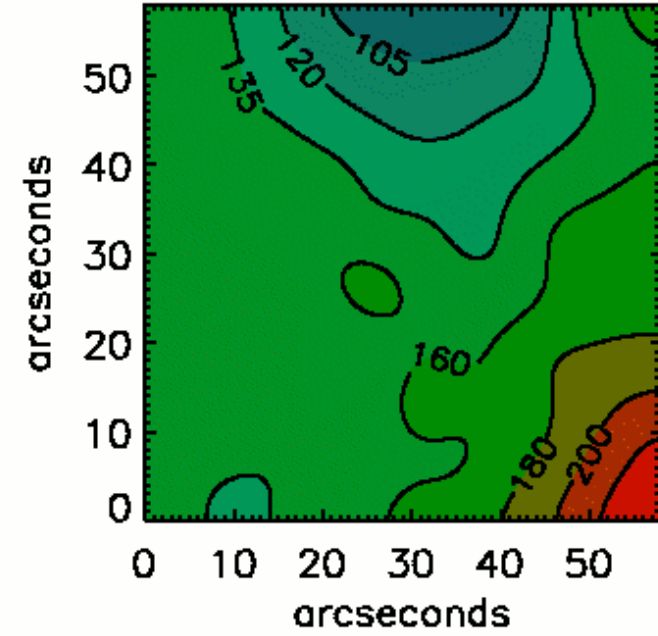
VLT MAD Ks-Band



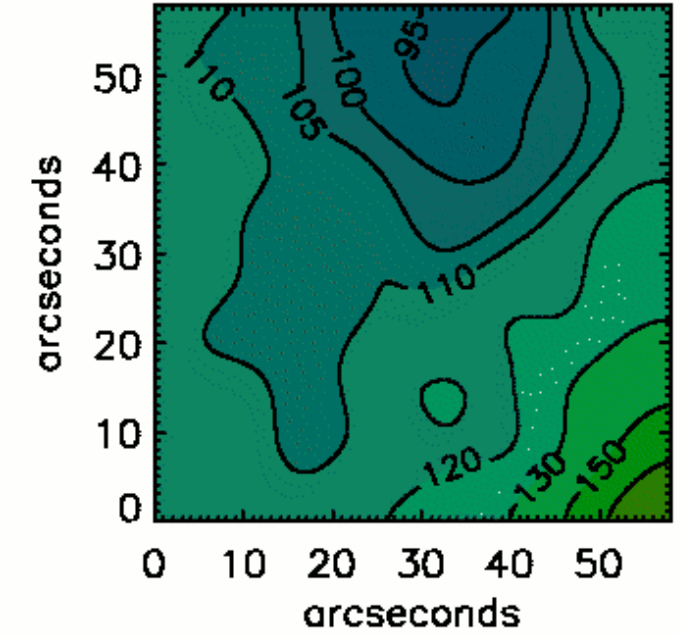
FWHM (mas)



FWHM (mas)

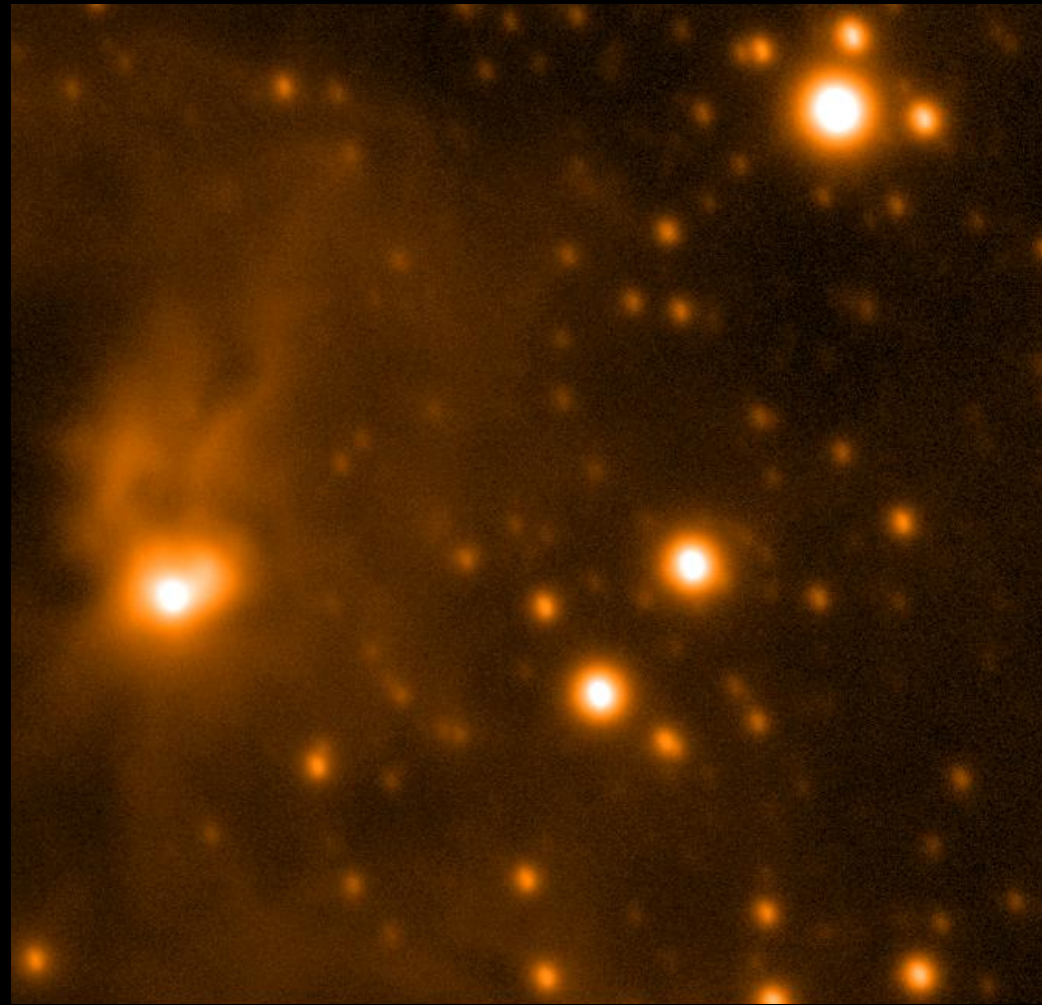
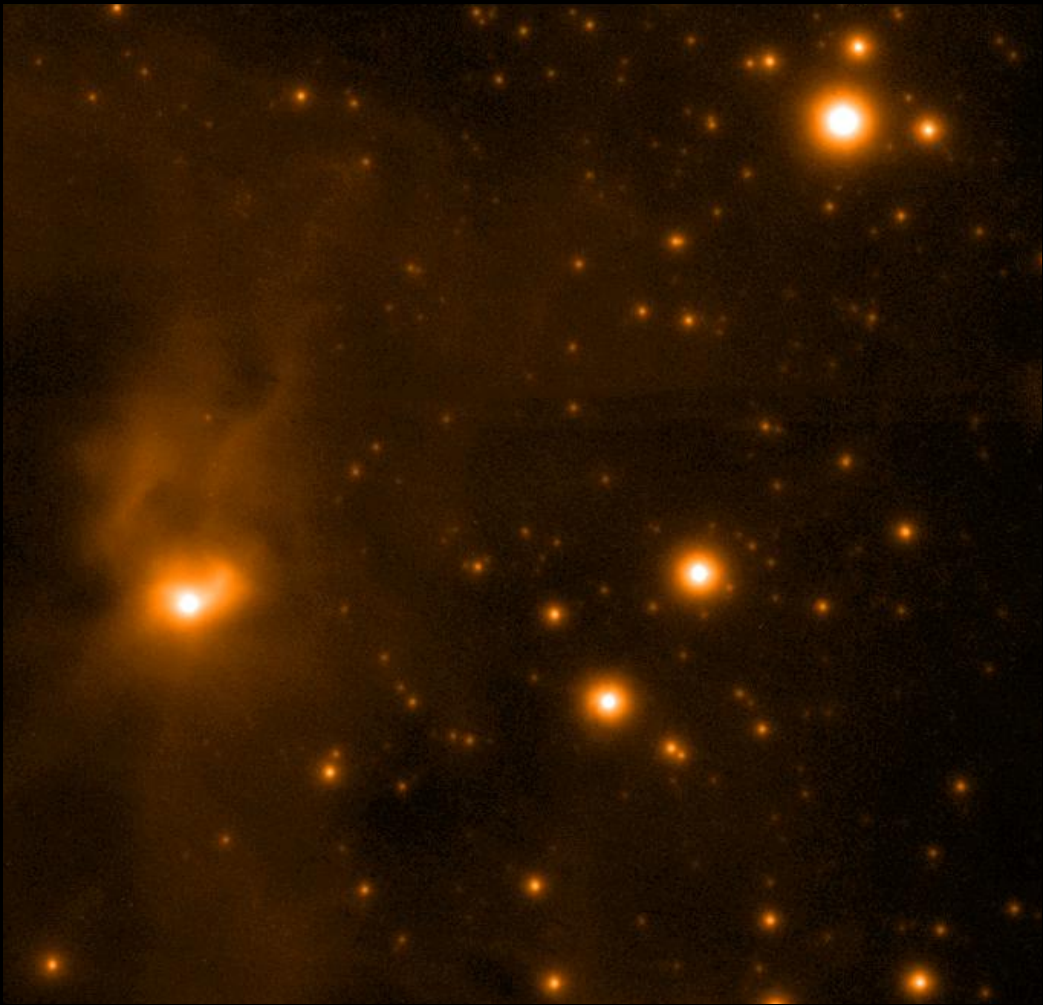


FWHM (mas)



commissioning (2007)

science demonstration (2008)

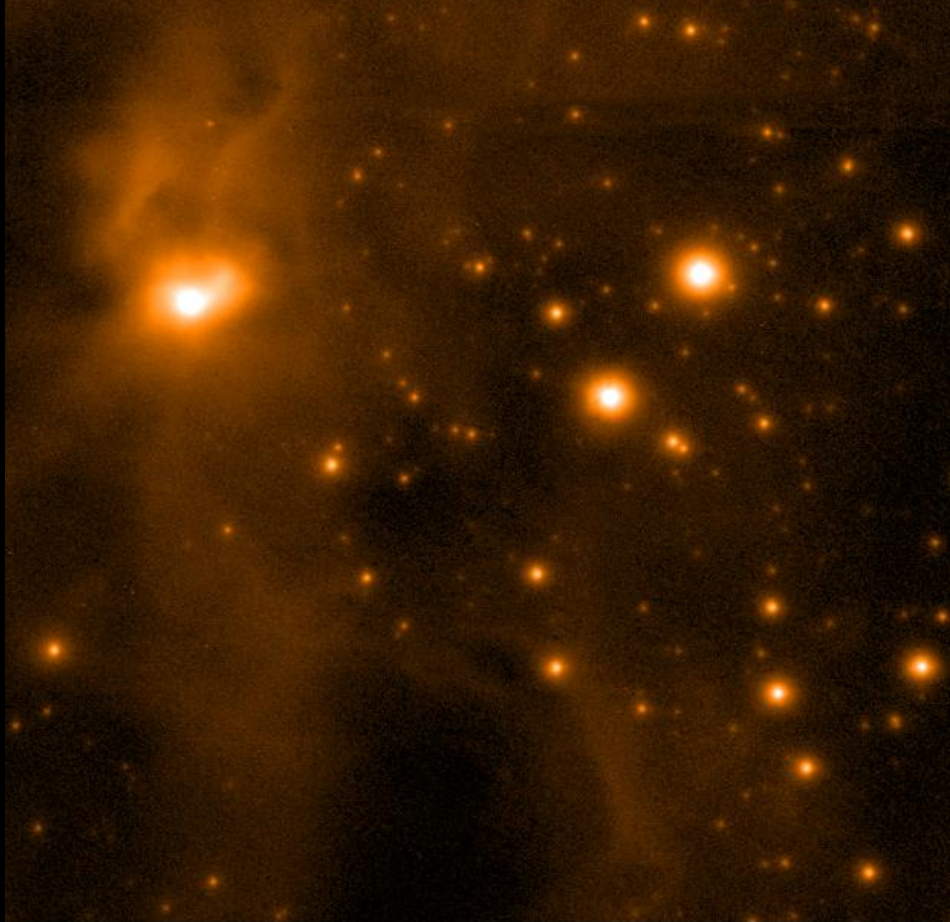


15 x 90s (1.5s DIT, 60 NDIT)
~100 mas FWHM

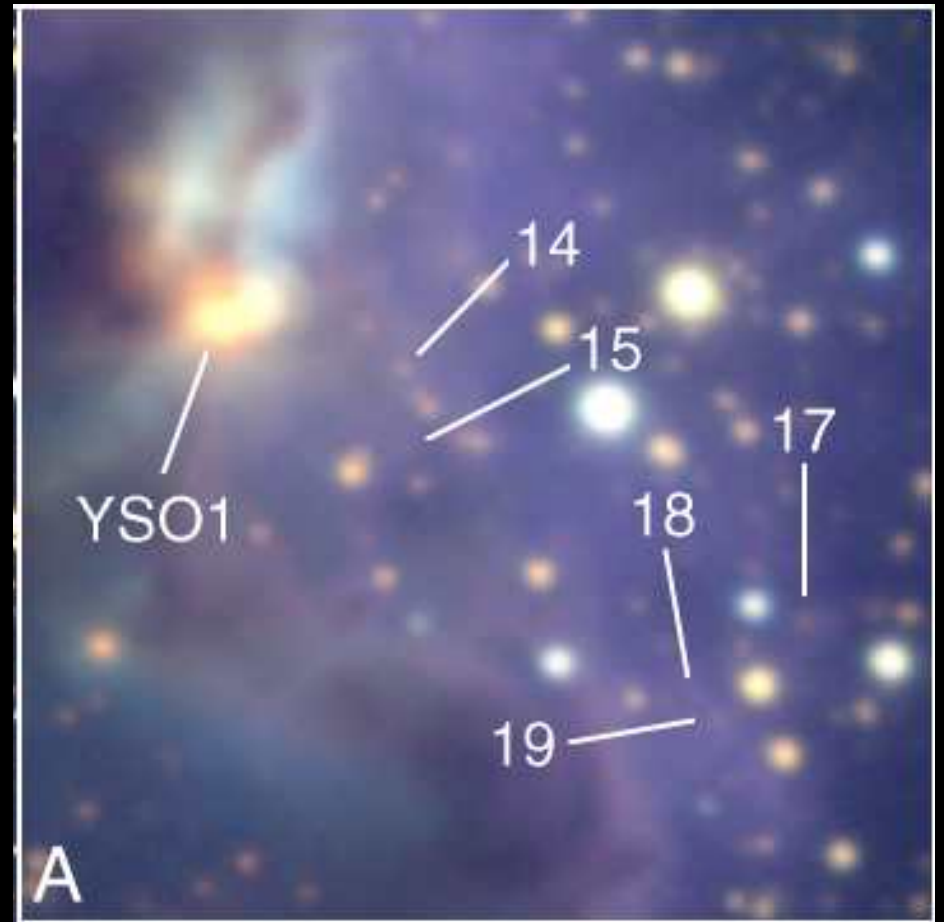
10 x 60s (2s DIT, 30 NDIT)
~350 mas FWHM

Same 'DIMM conditions' but other turbulence distribution.

MAD



ISAAC



3 times smaller FWHM in Ks
~1 mag deeper

BUT: J band not better AND smaller FOV than ISAAC



Thanks for your attention!