

# MAD

## *On sky performance testing*

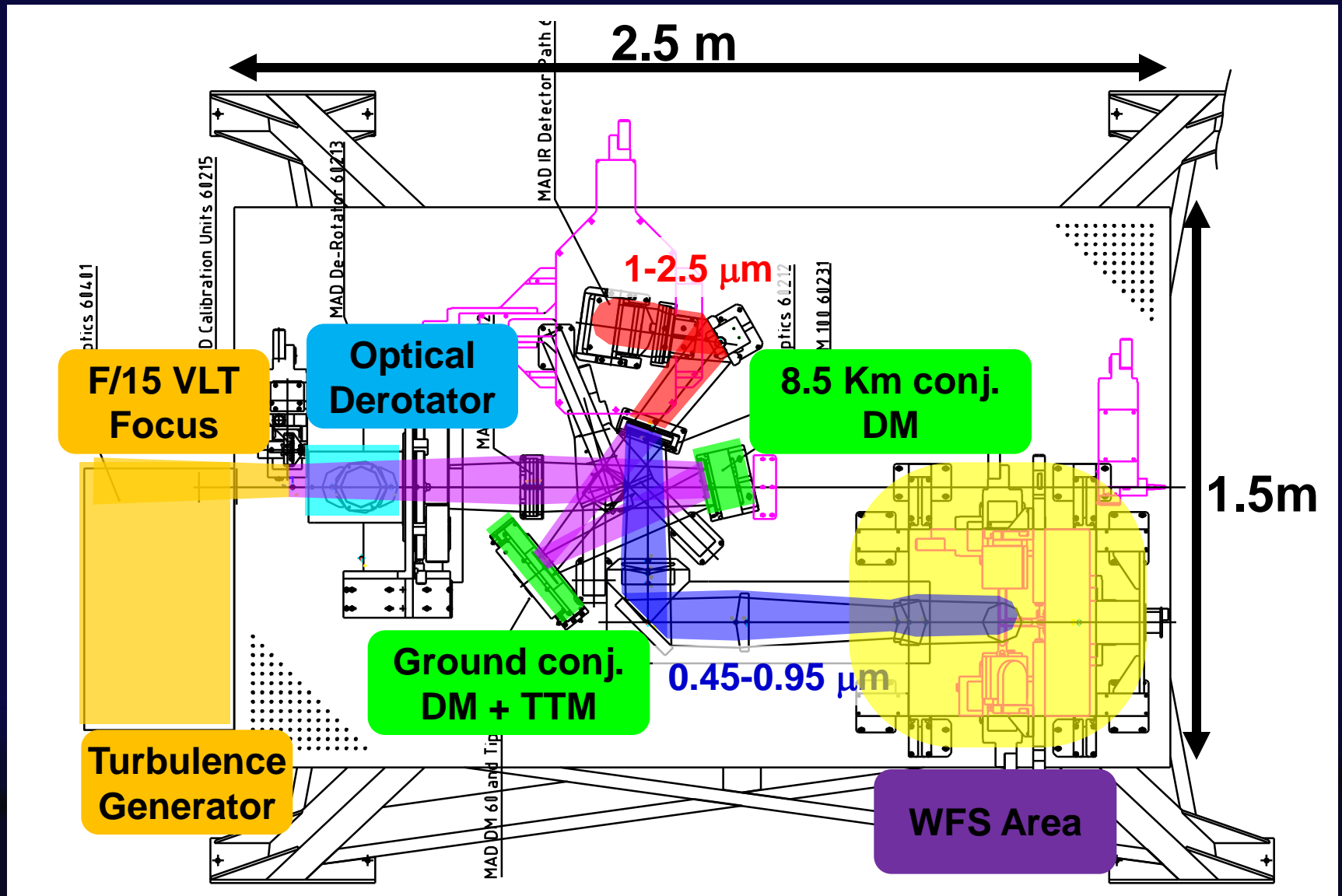
**E. Marchetti**, R. Brast, B. Delabre, R. Donaldson, E. Fedrigo, C. Frank, N. Hubin, J. Kolb, J.-L. Lizon, M. Marchesi, S. Oberti, R. Reiss, C. Soenke, S. Tordo, P. Amico, A. Baruffolo, P. Bagnara, A. Amorim, J. Lima & many others...



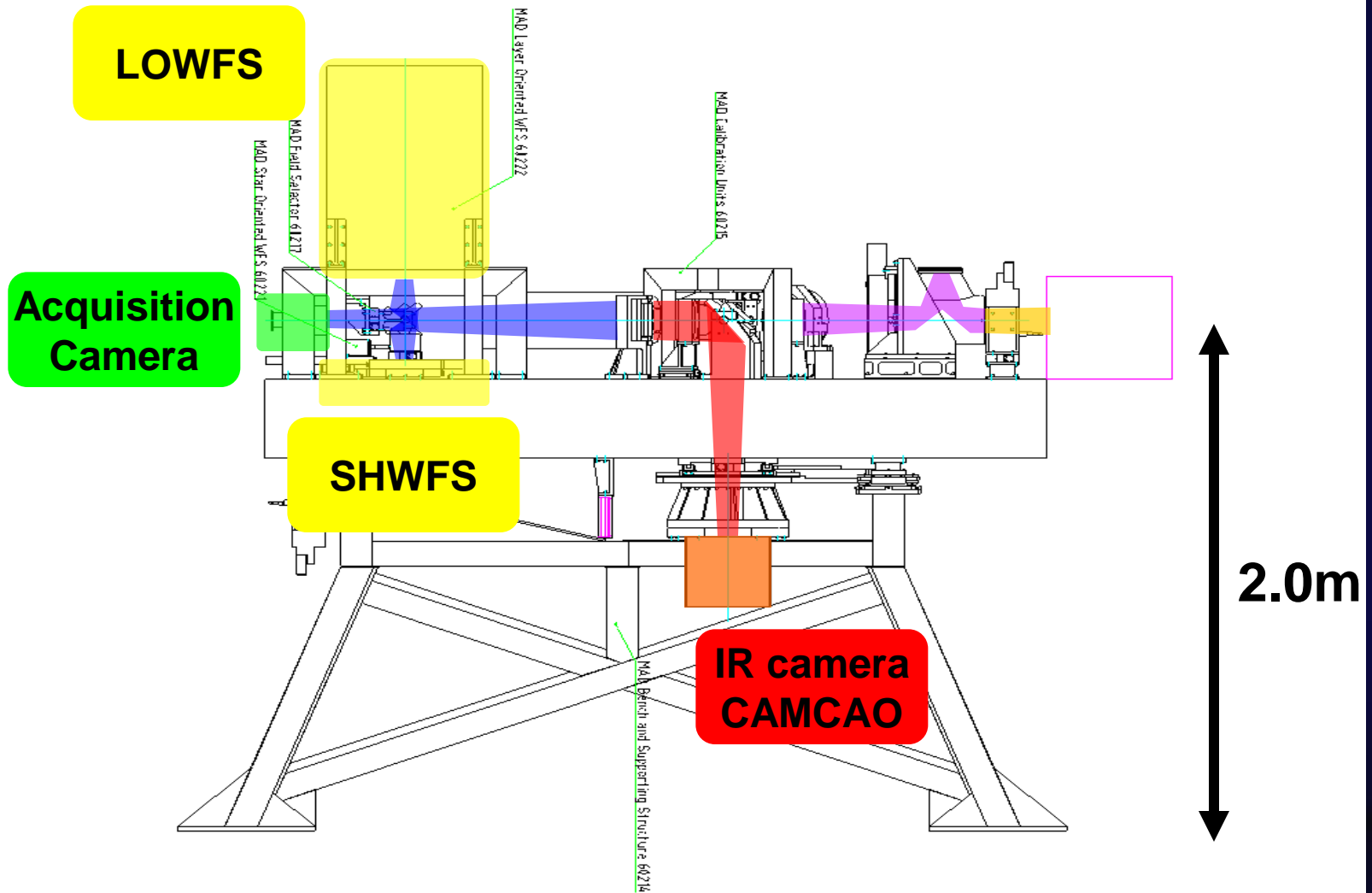
# The MAD experience

- Multi-Conjugate Adaptive Optics (MCAO) prototype demonstrator to validate of MCAO correction techniques in the framework of future VLT instrumentation and ELTs
- MAD Shack-Hartmann system: wavefront sensing on 3 Natural Guide Stars in 2 arcmin FoV (Natural Guide Stars)
- Correcting with 2 deformable mirrors (0 & 8.5 Km)
- IR camera imager CAMCAO (U.Lisboa - Portugal)
- Multi-Pyramid wavefront sensor (see R.Ragazzoni)
- Installed at Melipal's Nasmyth platform (Feb'07)
- Three science demonstration runs (Nov'07-Jan'08-Aug'08)

# MAD overview

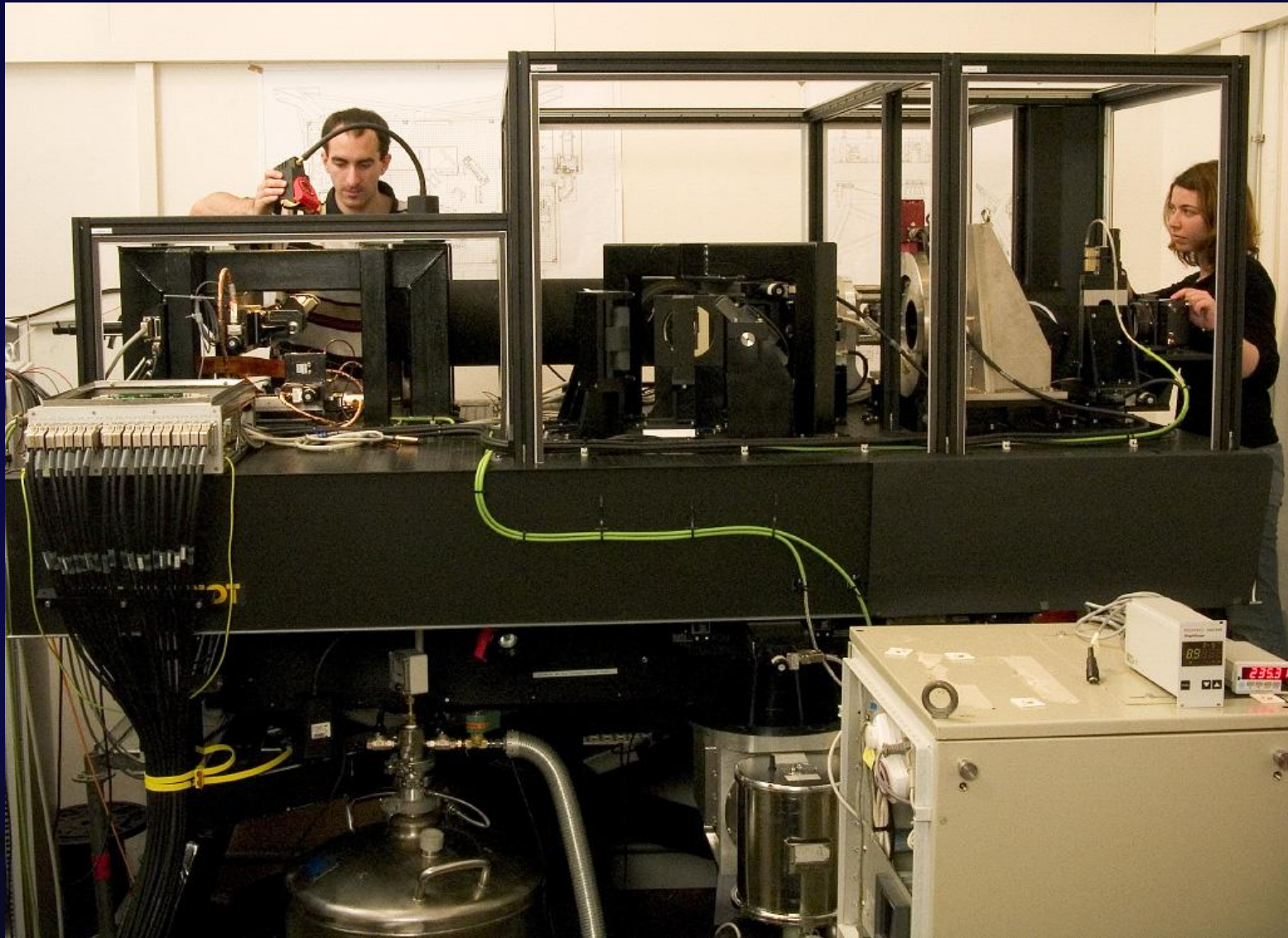


# MAD overview

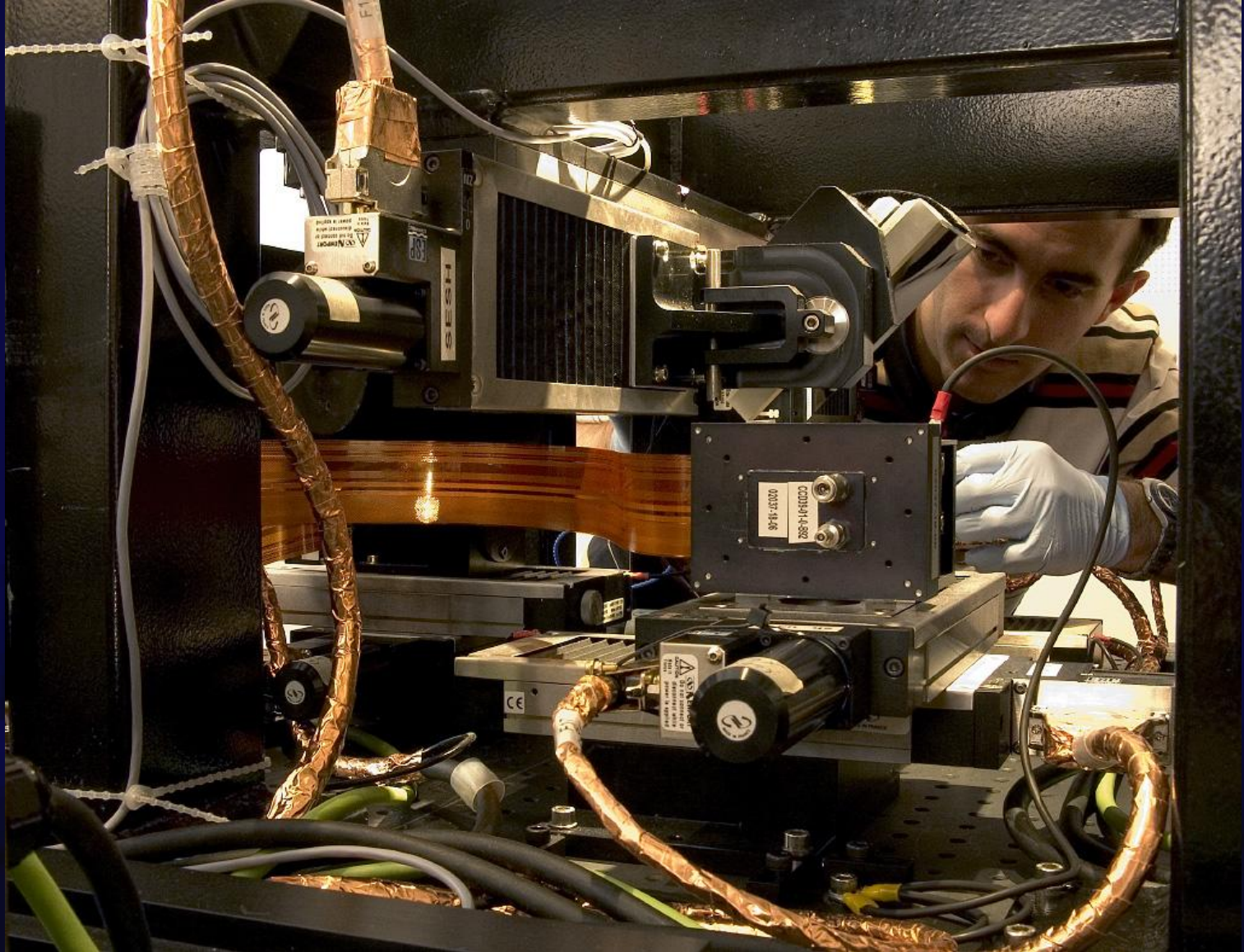




# MAD at ESO premises

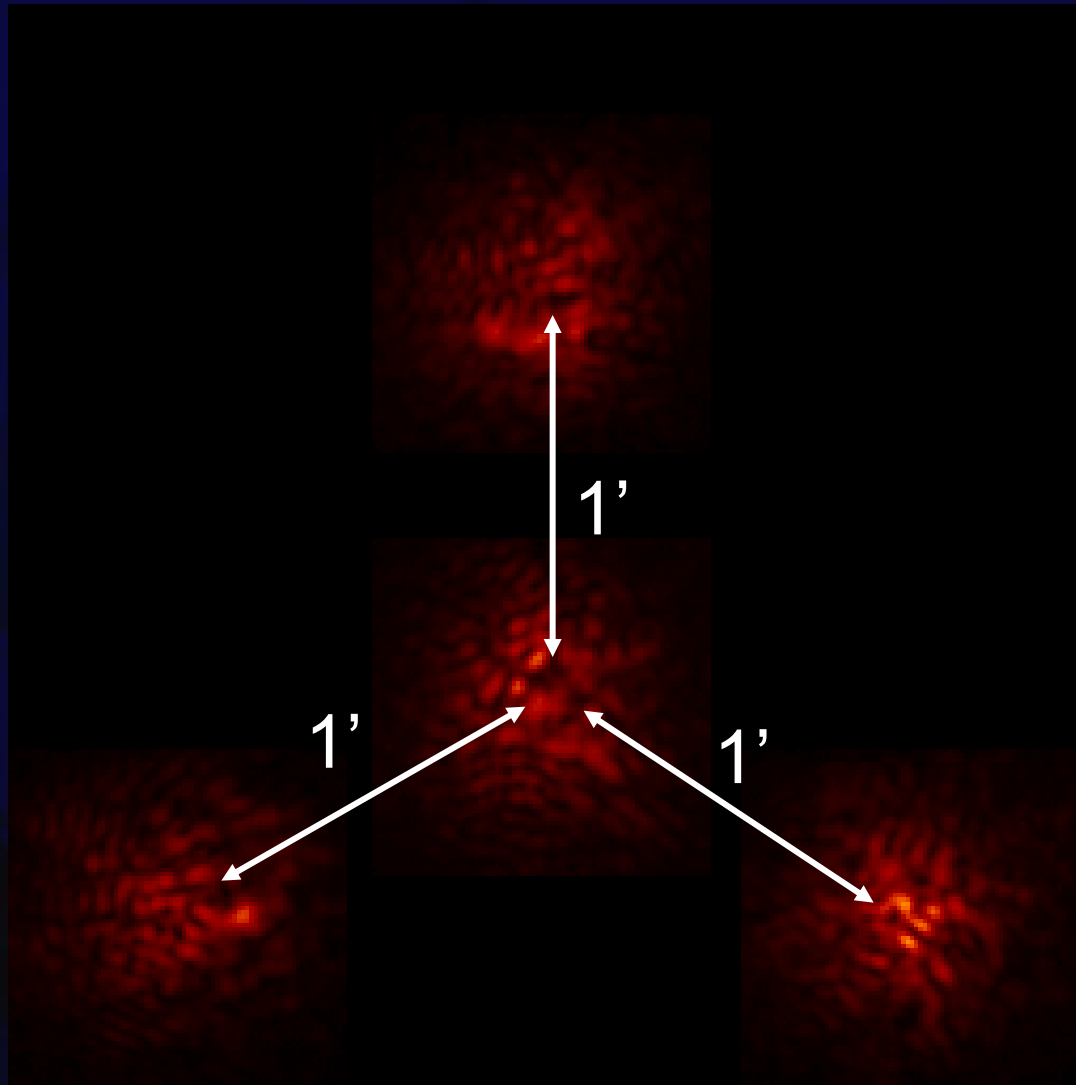


# Shack-Hartmann wavefront sensors





# MAD MCAO first closed loop in the lab



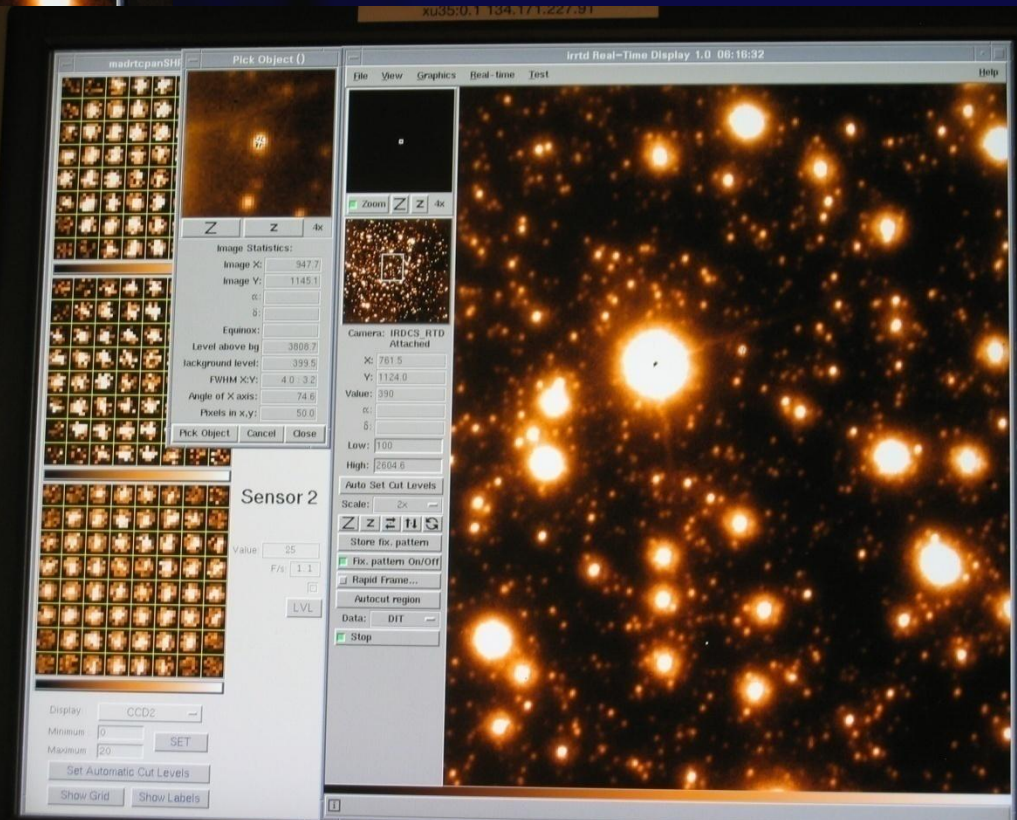
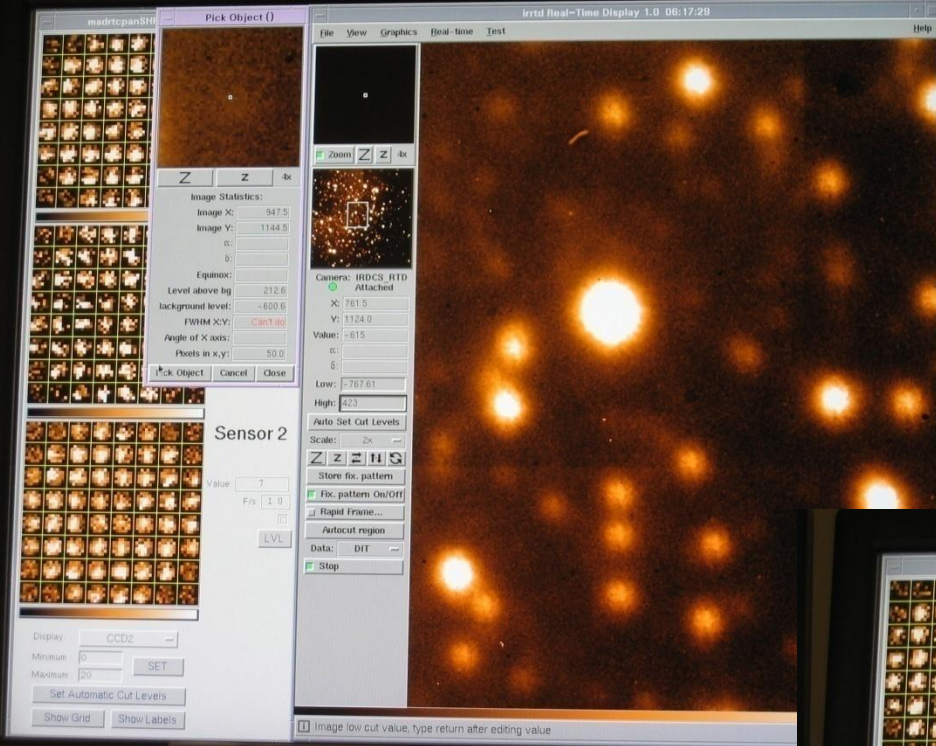


# MAD at Paranal



**Melipal (UT3) Nasmyth platform**

# Closing the MCAO loop...

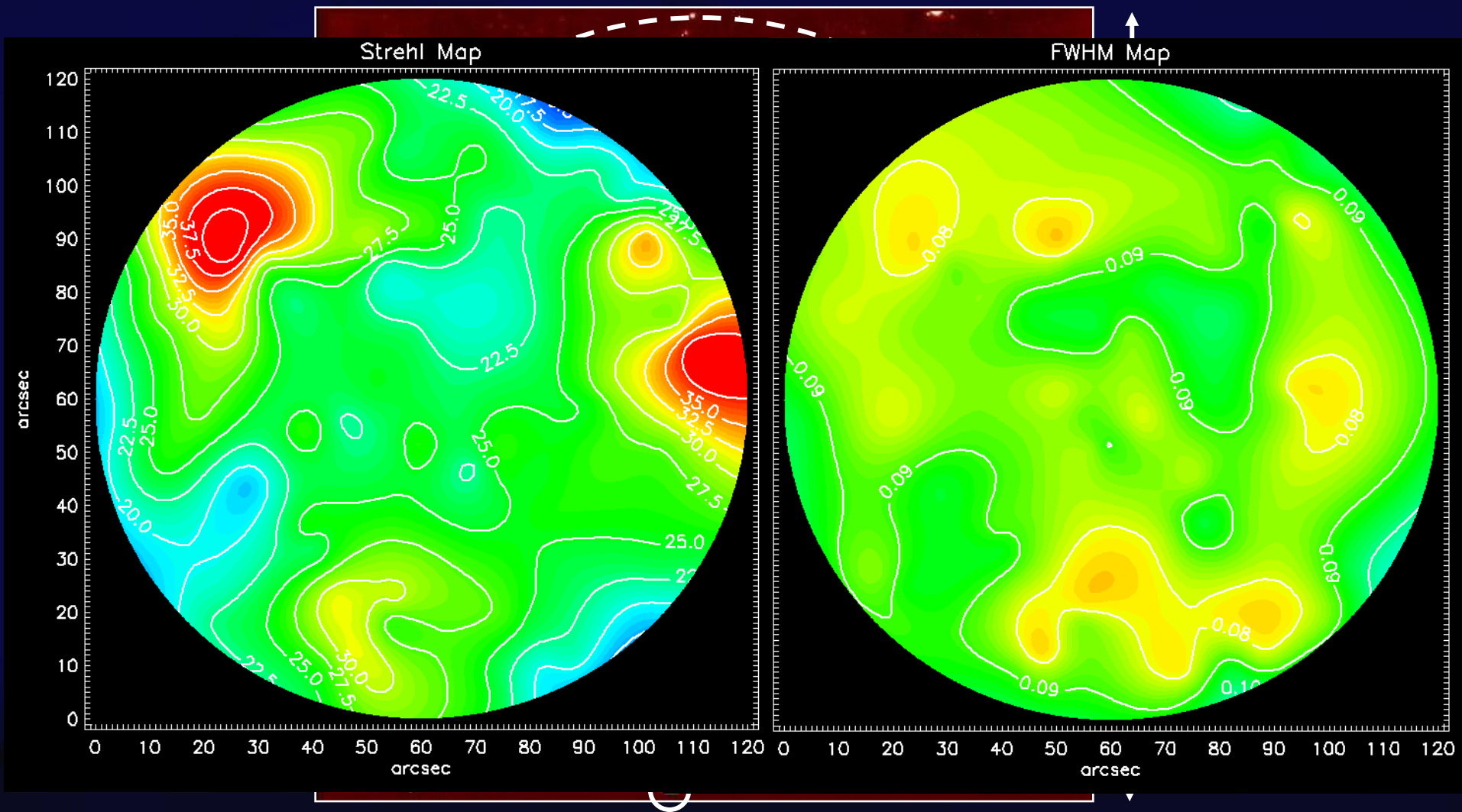


# MAD MCAO first light





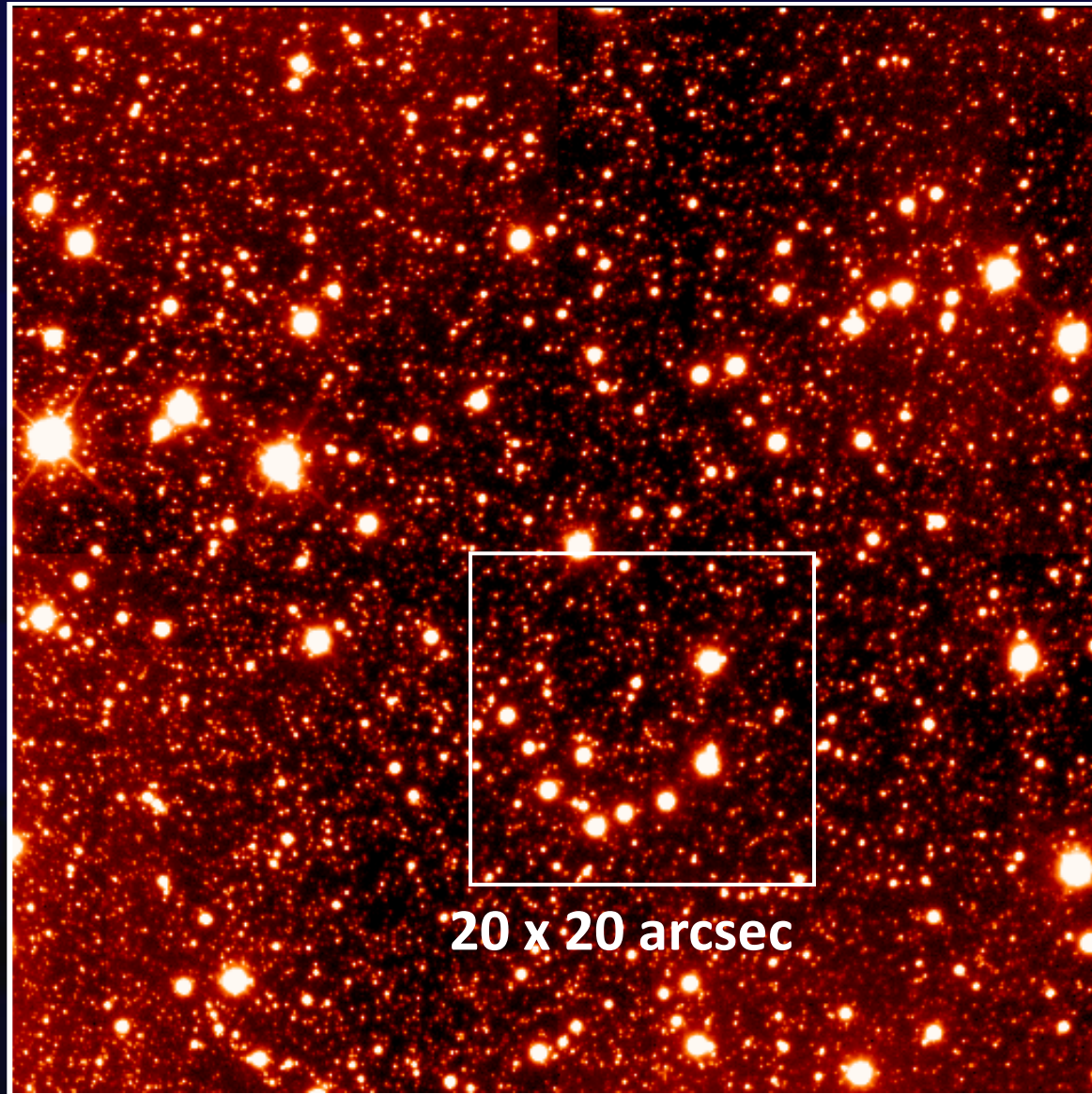
# MAD MCAO first light





# Omega Centauri

$T_{\text{exp}}$  [Ks]: 600s  
FWHM: 100mas  
 $K \sim 20.5$   
DIMM: 0.69''



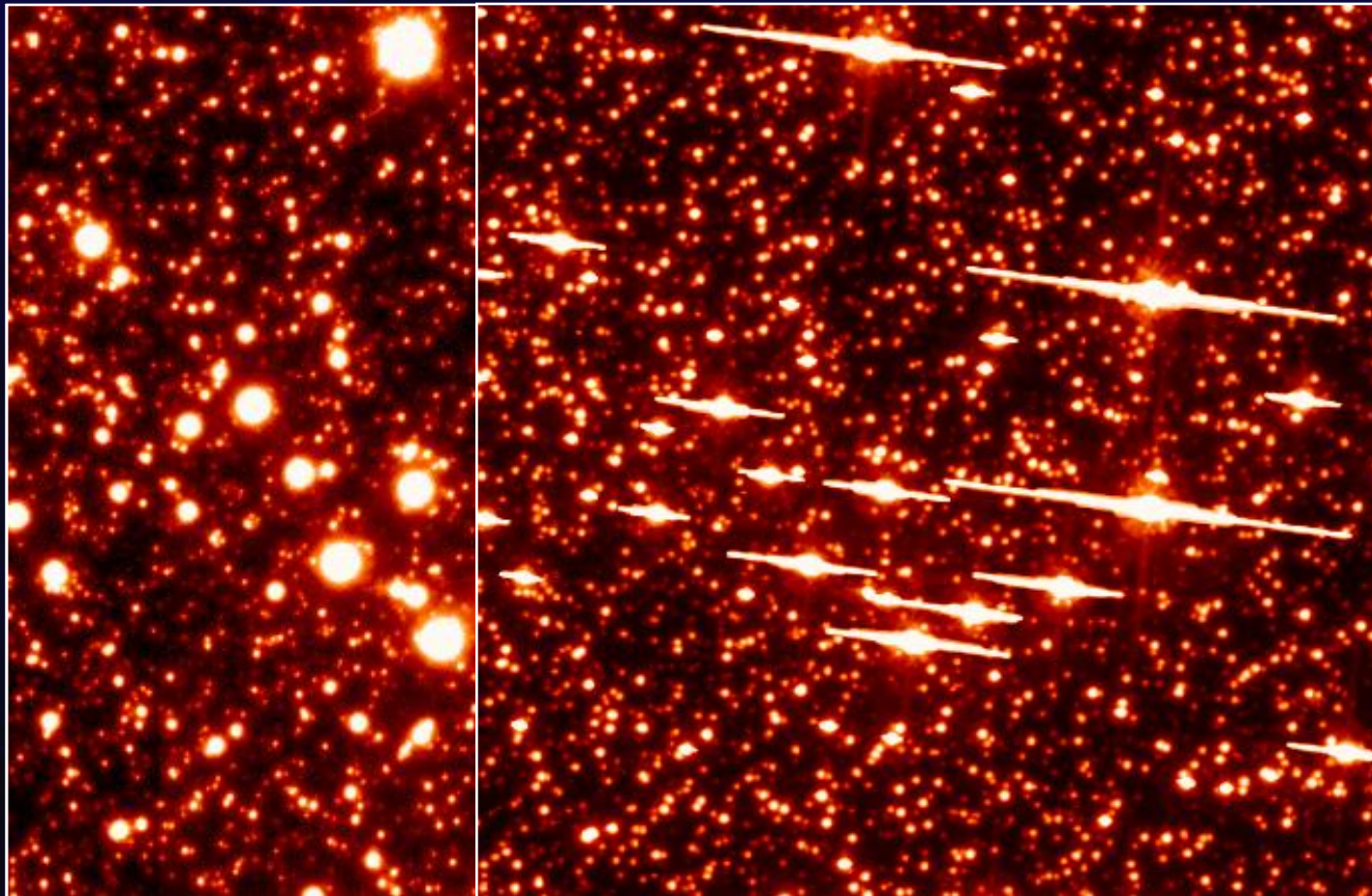
1 arcmin

20 x 20 arcsec

# Omega Centauri

Guide stars @ 60"

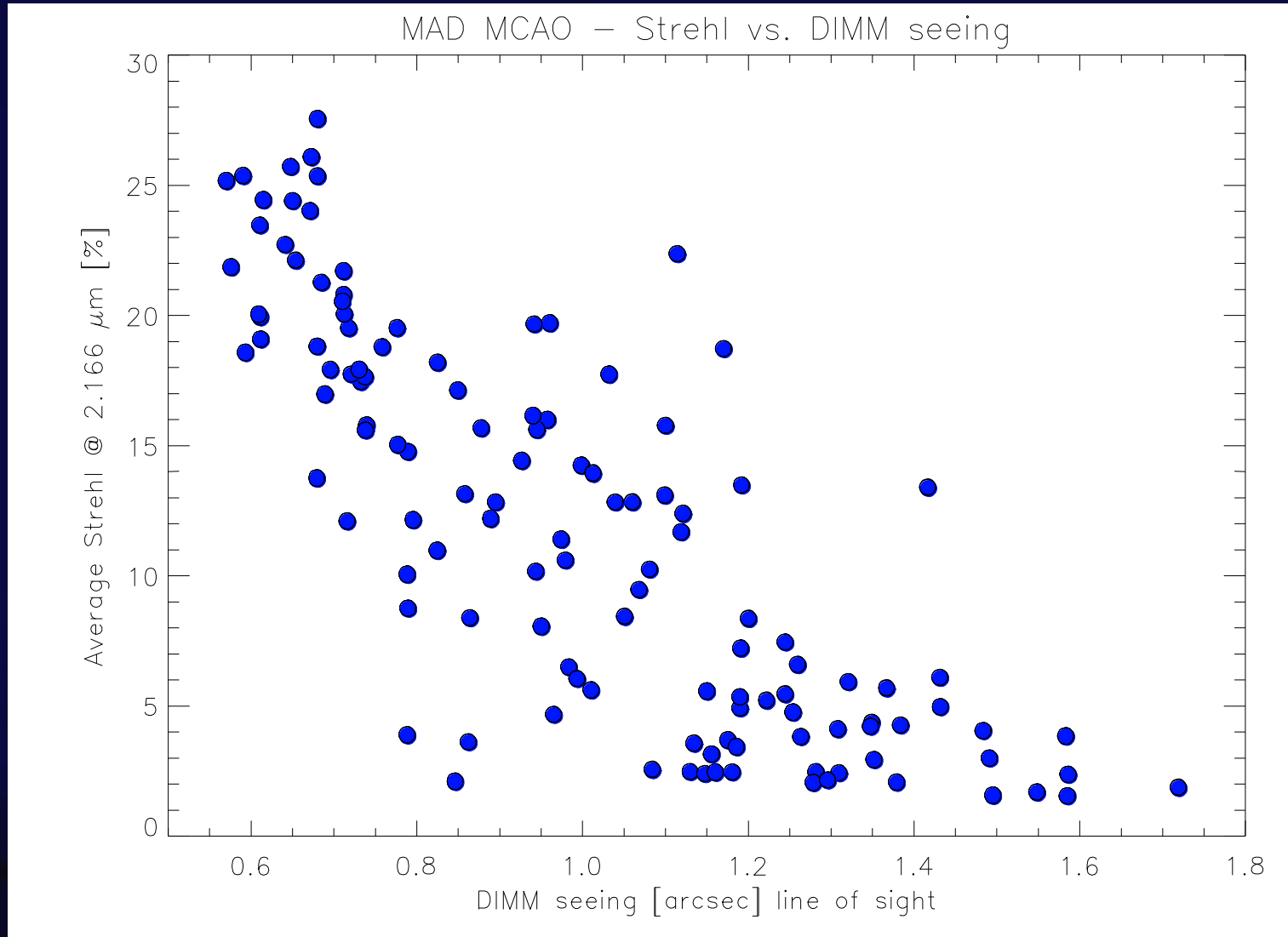
20" x 20"



MAD [BA] (0.028"/px) #51/100 SO [6935W] (0.050"/px)

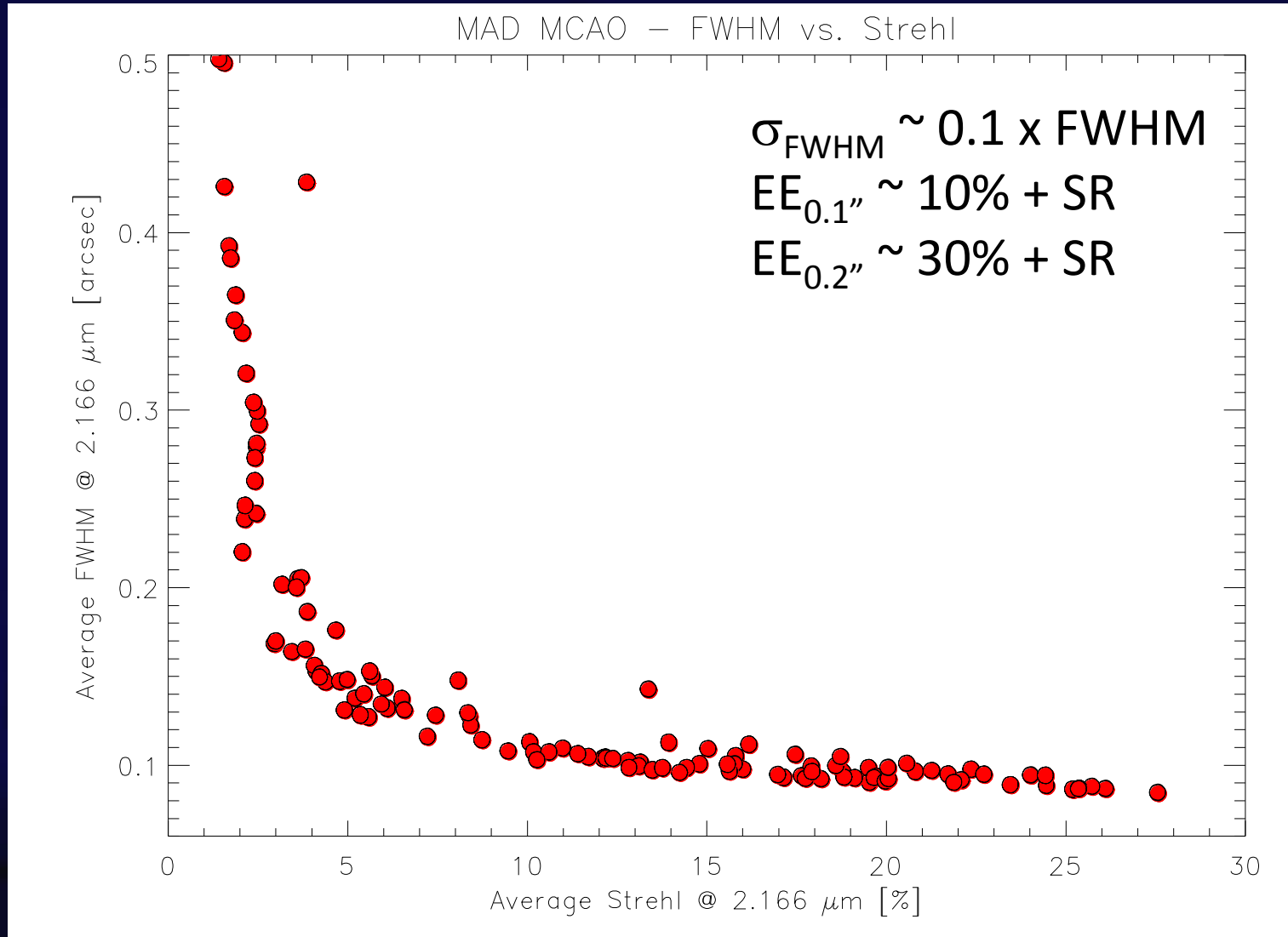
FWHM 100mas  $T_{exp} = 15000s$  100mas –  $T_{exp} : 340s$

# MAD MCAO correction





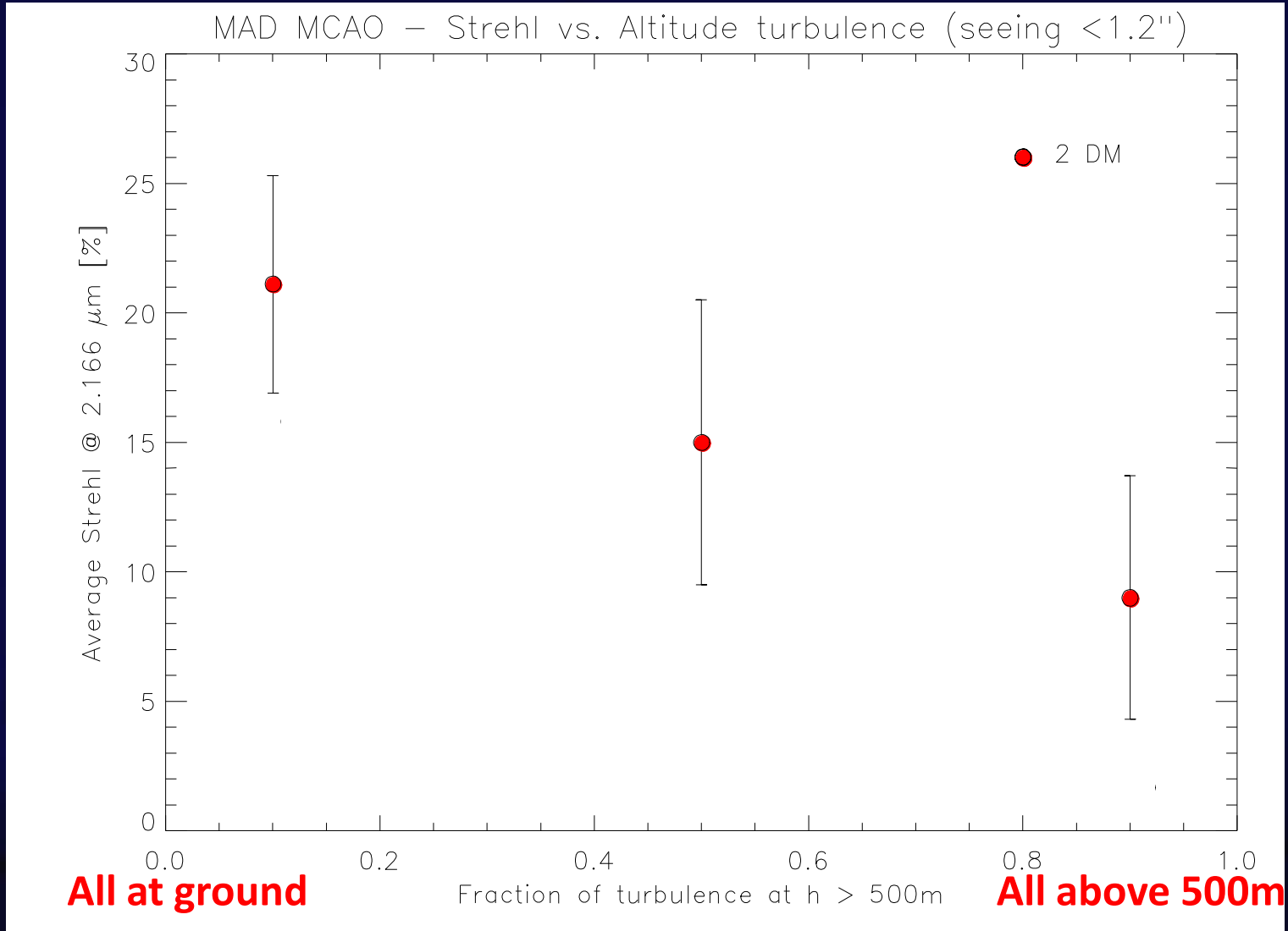
# MAD MCAO correction



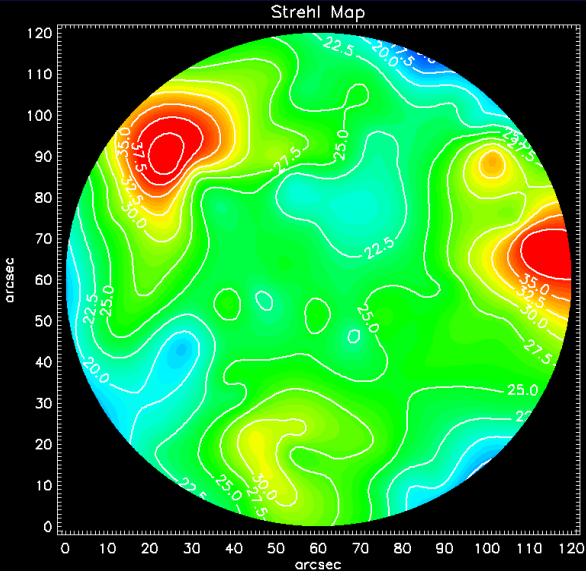




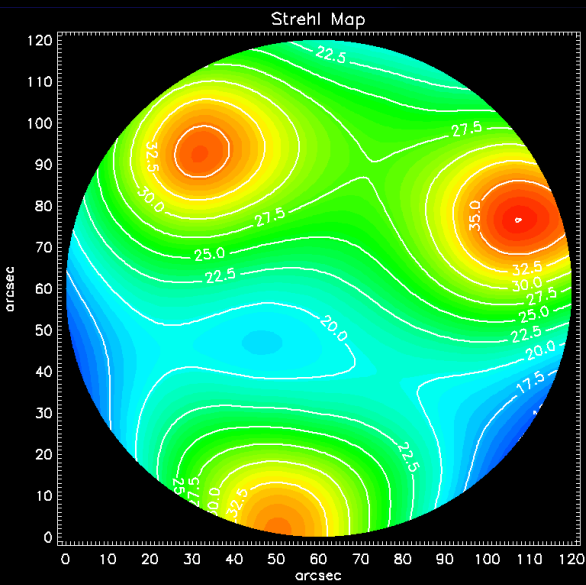
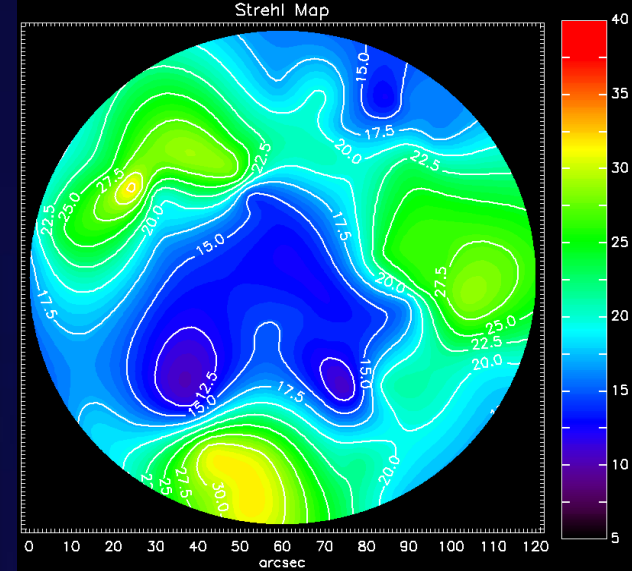
# MAD MCAO correction



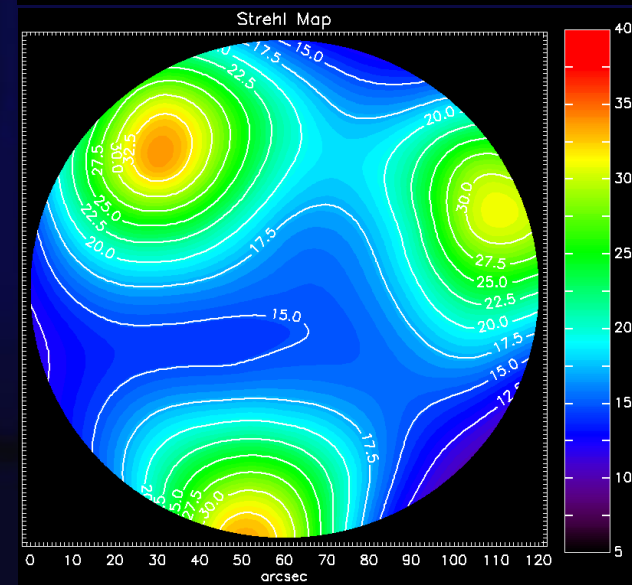
# Effects of turbulence vertical profile



0.64"	DIMM	0.64"
3.0 ms	$\tau_0$	3.0 ms
0.40	Alt	0.35
74%	$0.5 < h < 9$	55%
14%	$9 < h < 13.5$	25%
12%	$h > 13.5$	20%

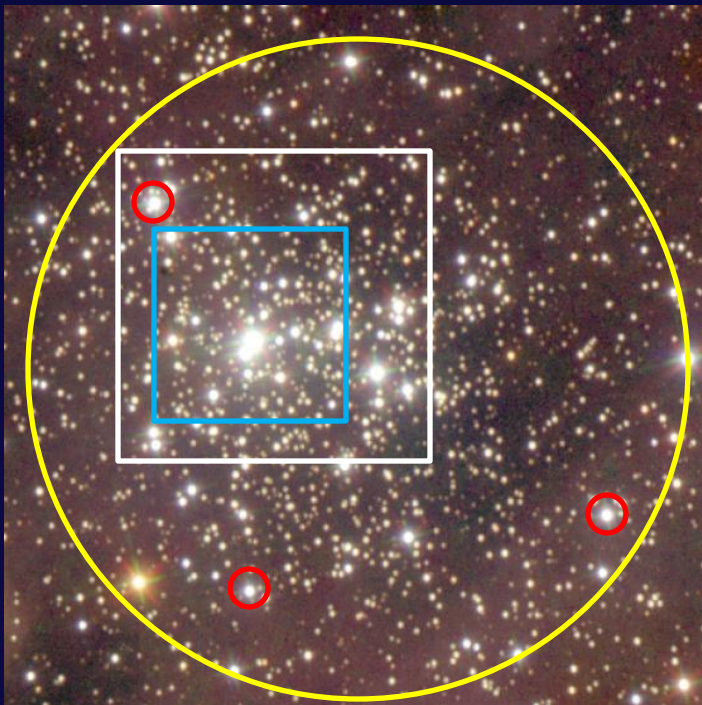


Numerical simulations  
10 layers  
measured  $C_n^2$  profile





# Classical SCAO vs. MCAO: an example



MAD [H] 0.028"/px

$T_{\text{exp}}$ : 1680s

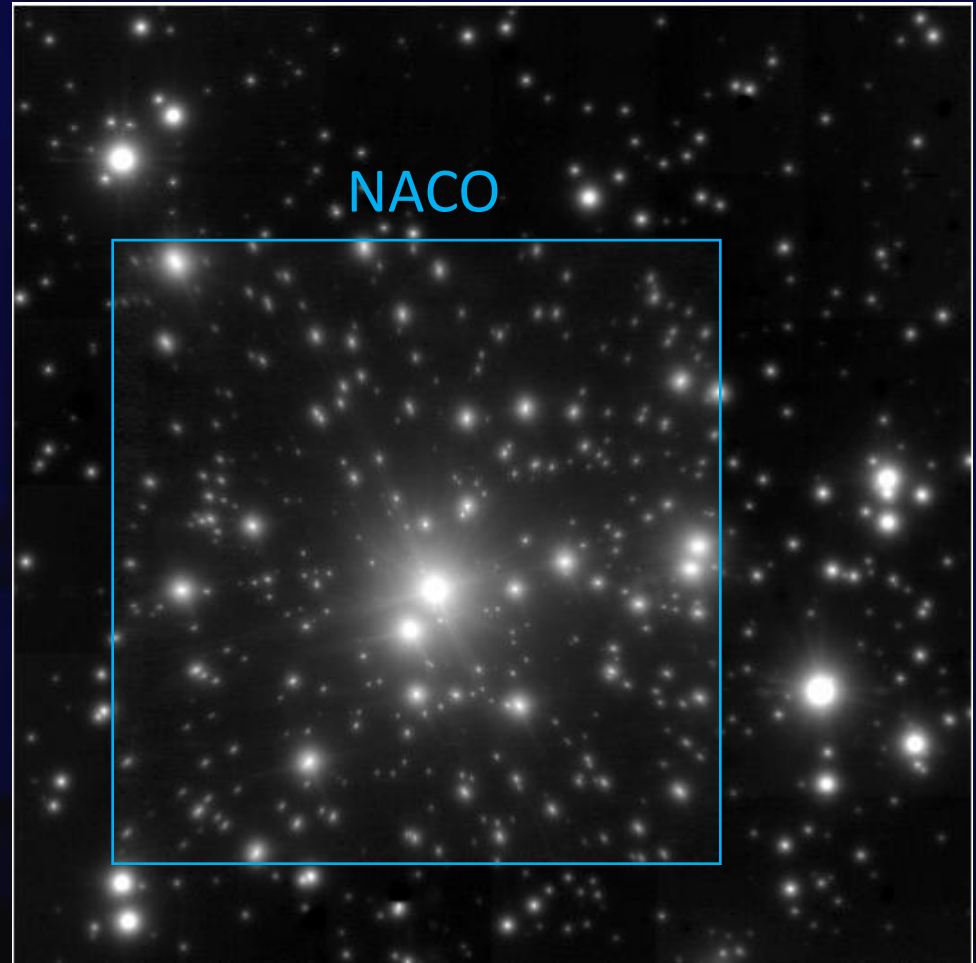
DIMM: 1.20"

NACO [H] 0.054"/px

$T_{\text{exp}}$ : 690s

DIMM: 0.85"

MAD



60" x 60"

Courtesy: J. Ascenso

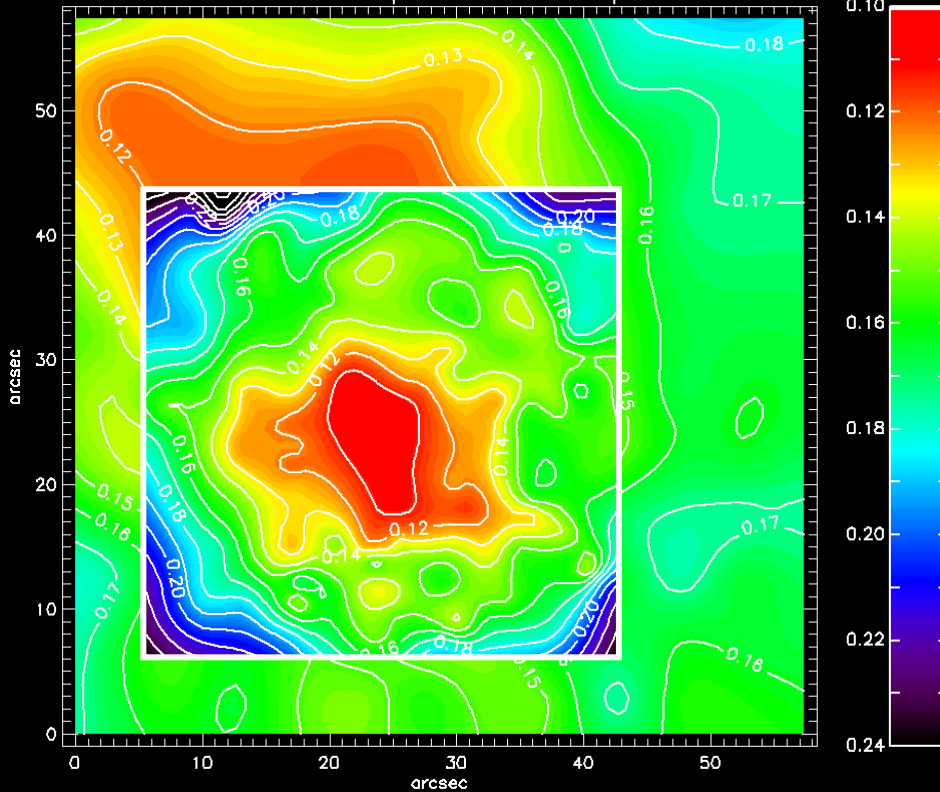


# Classical AO vs. MCAO correction uniformity

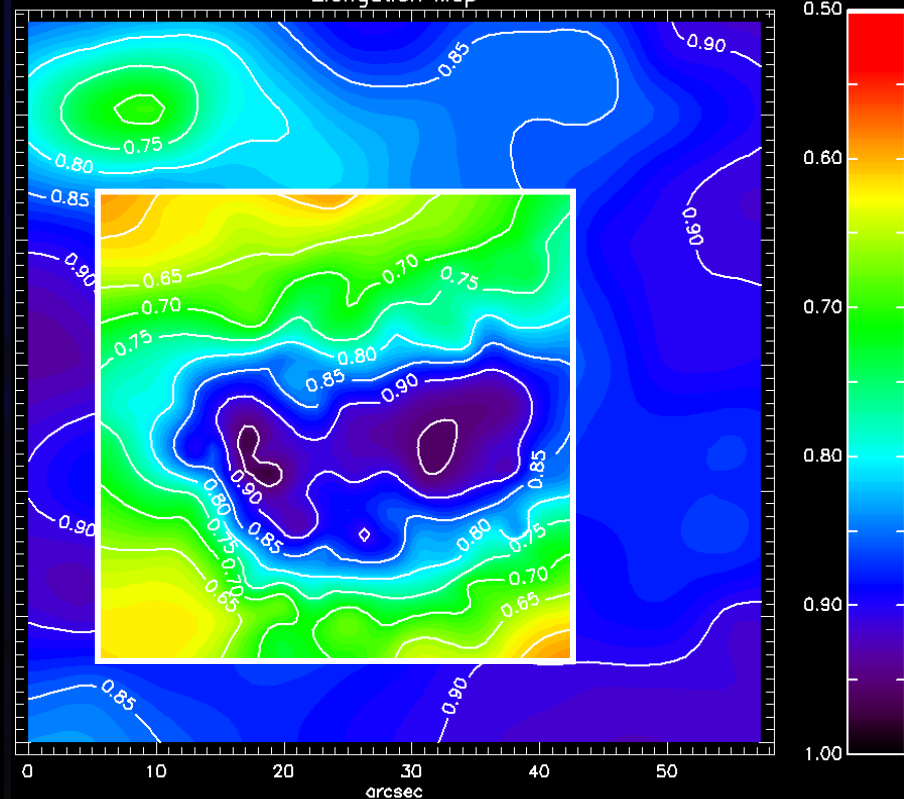
## FWHM

## Elongation

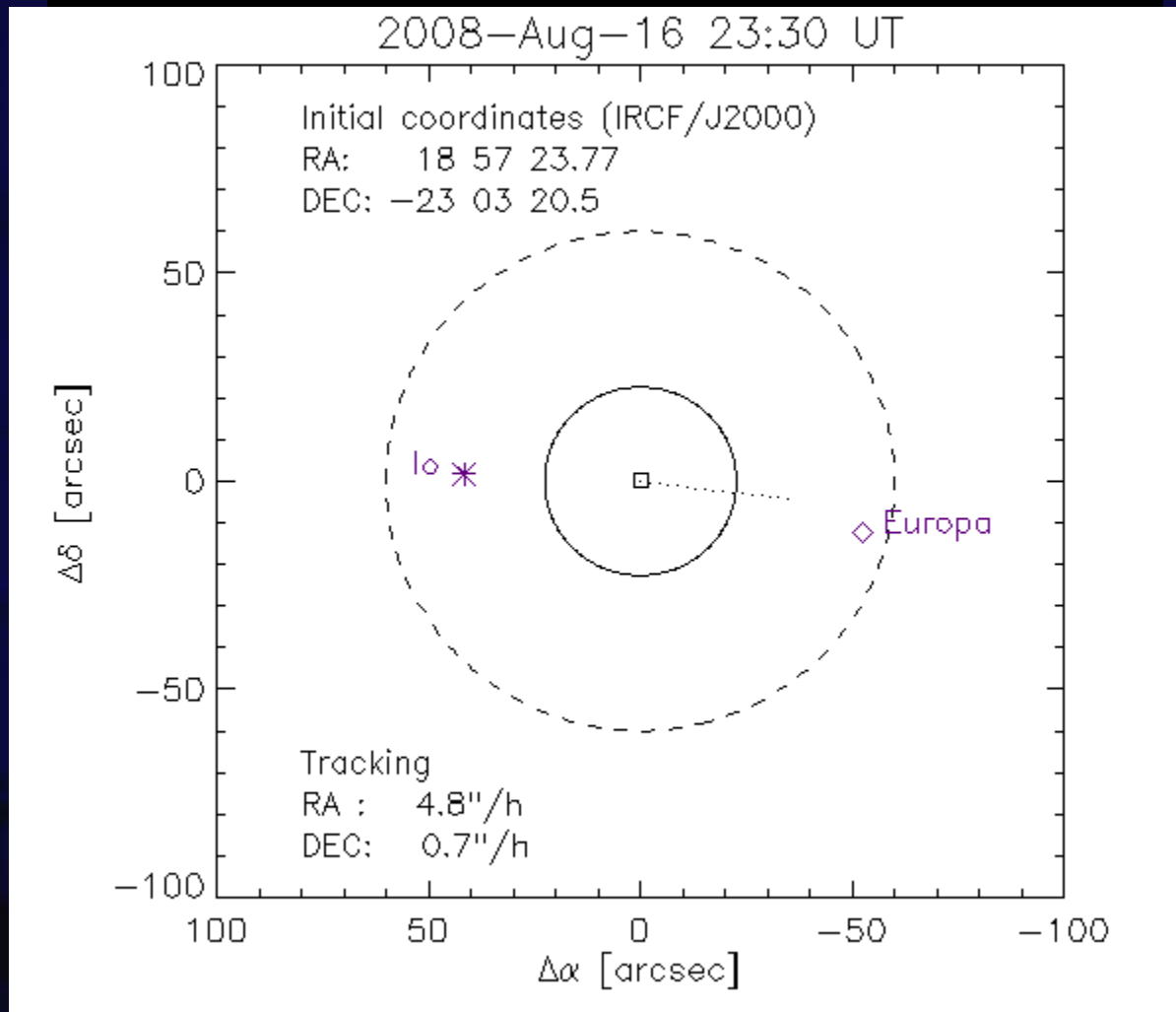
MAD & Trumpler 14 – FWHM Map



Elongation Map

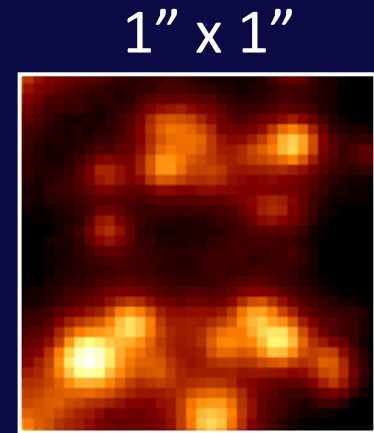
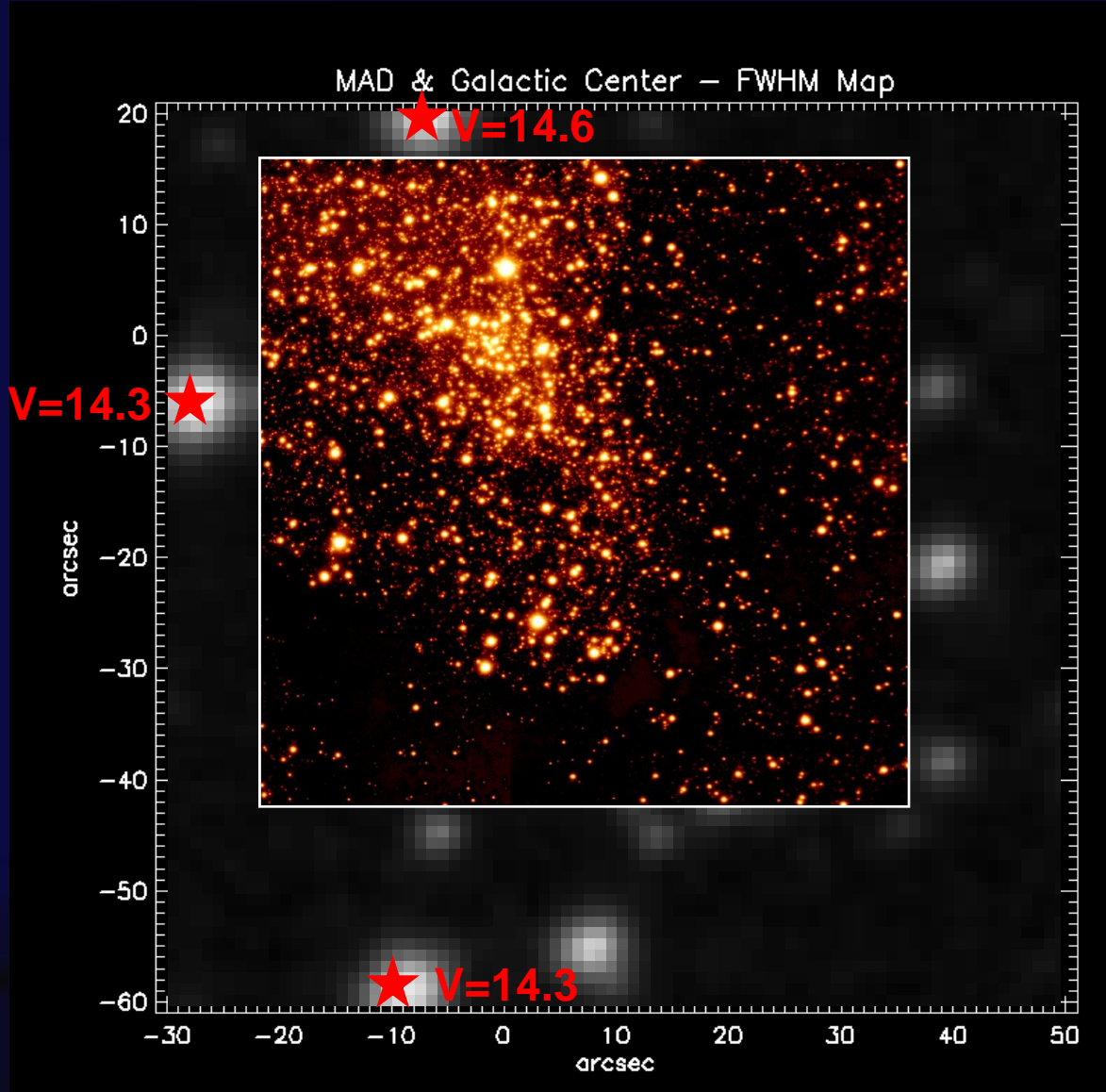


# Two guide star MCAO: Jupiter



FWHM [Ks]: 90mas      DIMM: 0.7'' – 1.0''

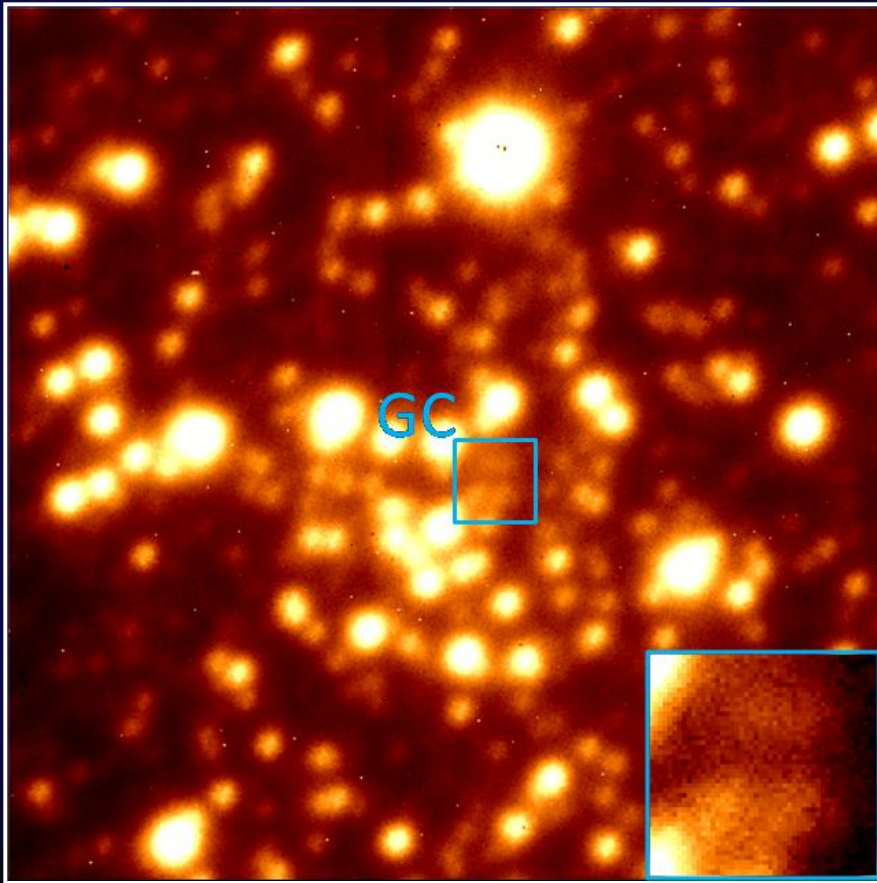
# Galactic Center



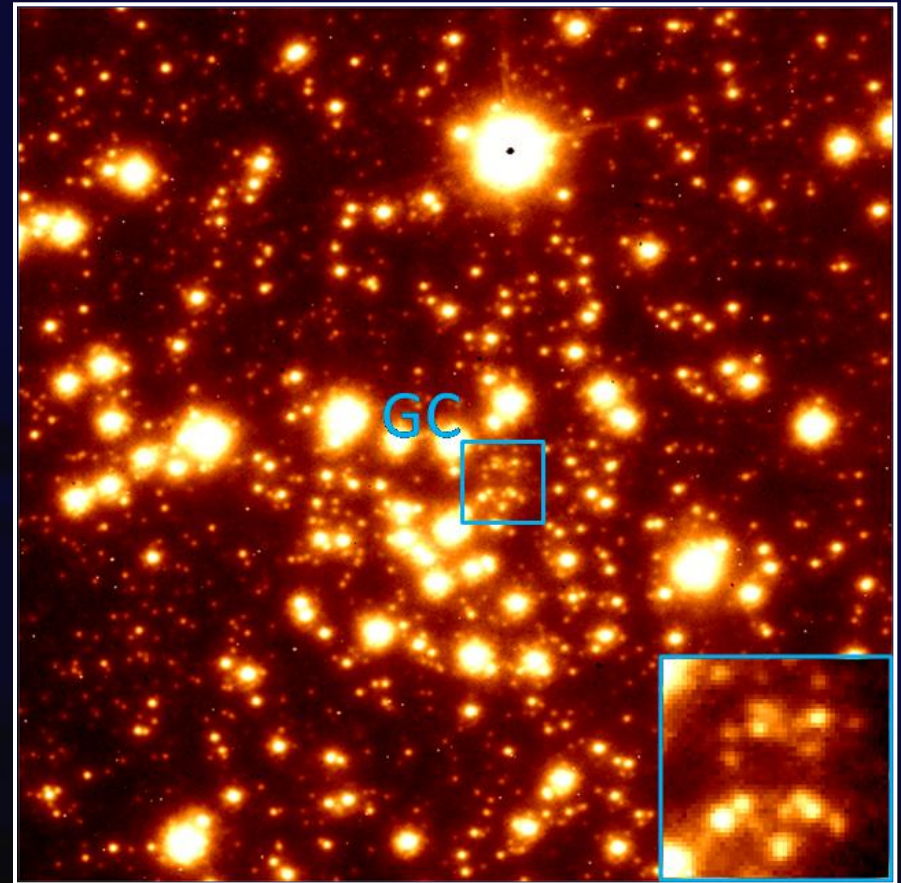
$T_{\text{exp}}$  [Ks]: 50s  
DIMM: 0.6"

# Galactic Center

$T_{\text{exp}}[\text{Ks}]$ : 10s, FoV: 15" x 15", DIMM: 0.6"



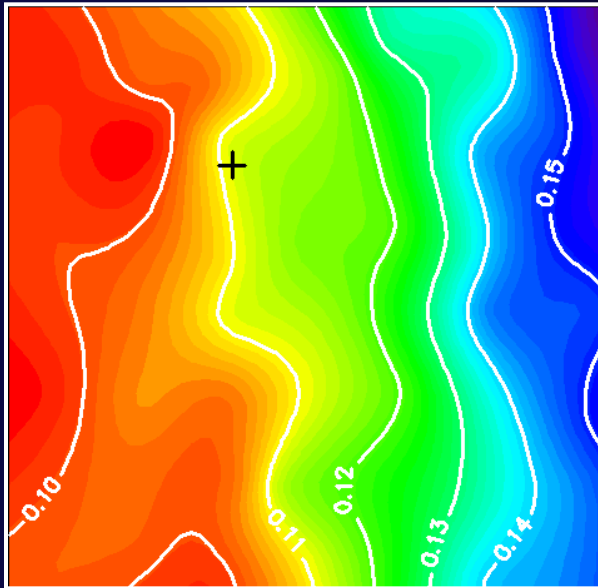
Open Loop



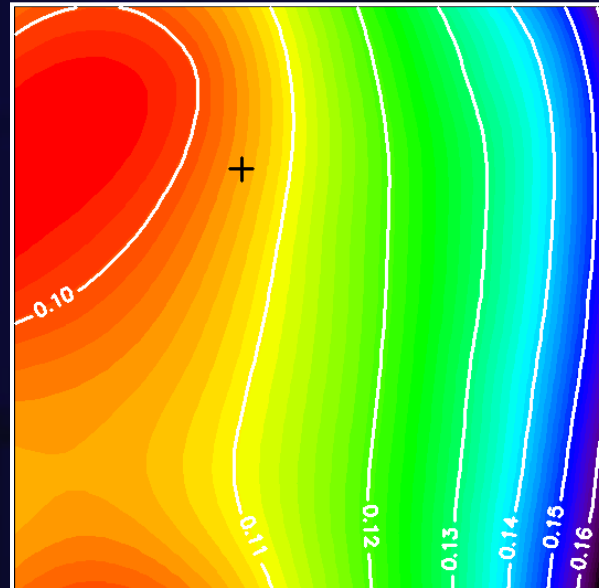
Closed Loop

# Galactic Center

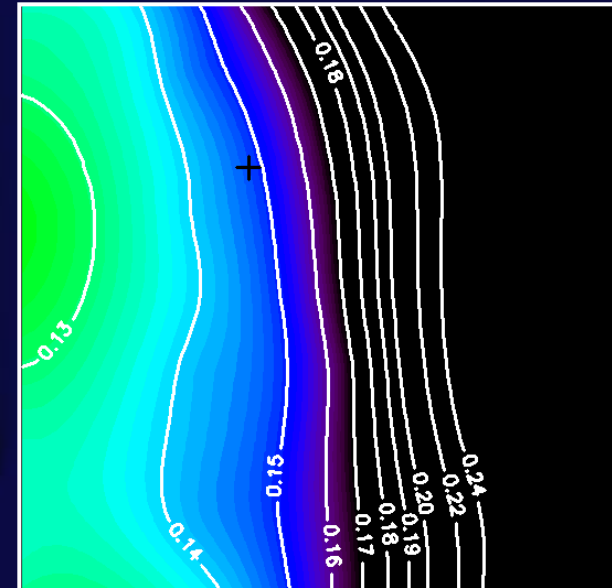
Observed



Simulated



Simulated  $m_V+1$

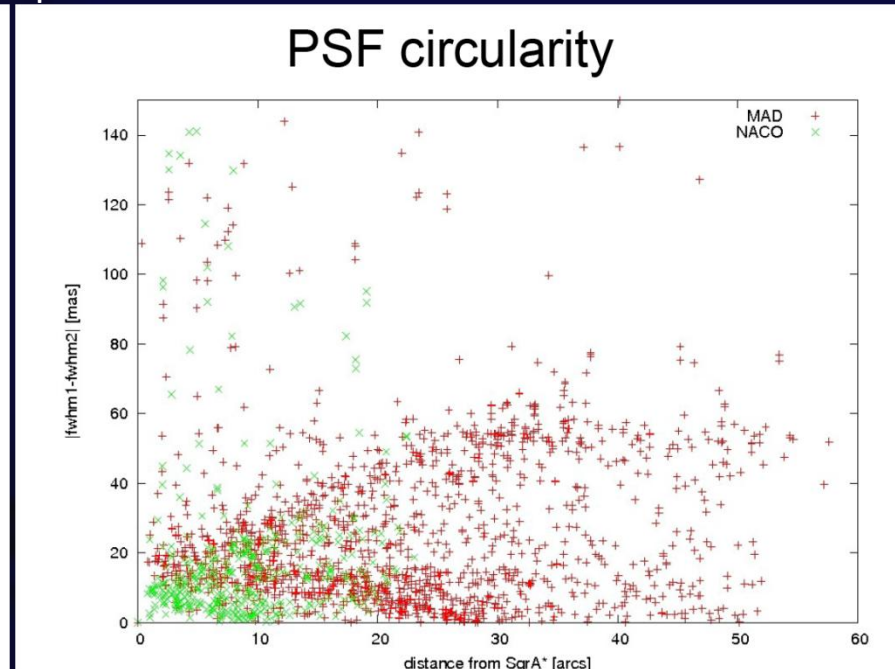
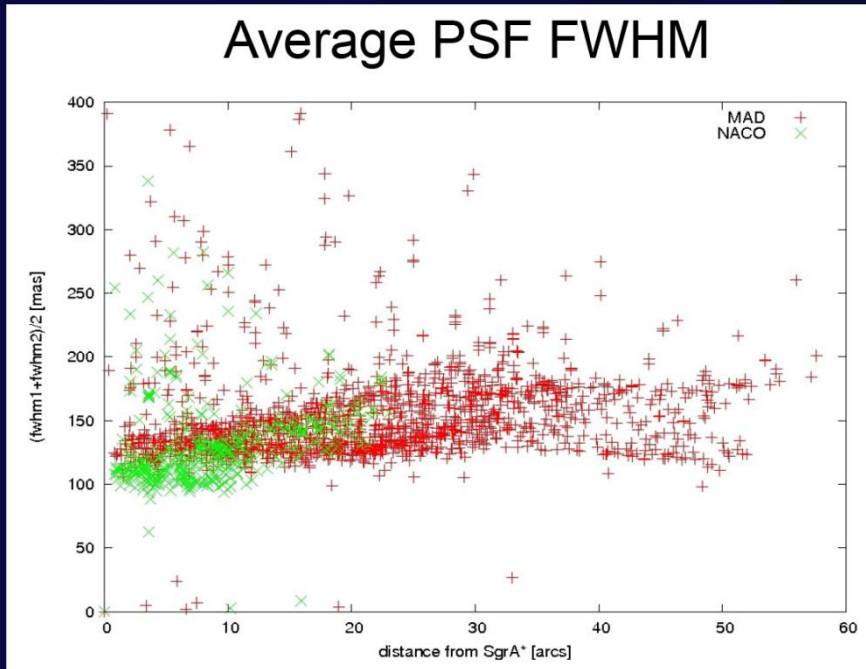






# Galactic Center: a comparison

MAD (1') vs. NACO (30'')  $T_{\text{exp}}: [\text{Ks}] 100\text{s}$ , DIMM:  $\sim 0.6''$



**Photometric accuracy (13>Ks>11):**

MAD  $\sim 0.03\text{mag}$  NACO  $\sim 0.02\text{mag}$

**Distortion**

“... just linear transformation, rotation and shift”

mag bin	completeness (%)		
	NACO	MAD-K	MAD-H
13	94	96	99
14	88	92	97
15	74	81	93
16	53	64	85
17	28	42	68
18	9	19	42

Courtesy: L. Mascetti & R. Davies



# MAD greetings from Atacama

## STC-437 (Nov'07)

*“The STC continues to be impressed with the performance of MAD, and applauds ESO and the MAD team for this remarkable demonstrator. The committee has noticed that an instrument able to further exploit multi-conjugate AO on a short timescale would be highly welcomed. We encourage ESO to investigate ways to continue to make MAD available to the community for additional observing runs to exhaust the potential observing opportunities available with this NGS system.”*