



Lessons from MAD: From the Lab to the Sky

- Testing and optimization
- Problems... solved before the nights
- Observations

Beyond MAD
ESO Garching, June 8-10 2009

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Testing and optimization..



Once upon a time...

June 2004 - December 2006:
MAD in the optical laboratory at ESO

- Mounting, optical alignment
- Test of components
- Real-Time Computer development
- Functionalities (motions...)
- Performance verification and optimization
- Preliminary Acceptance Europe

MELLES GRIOU

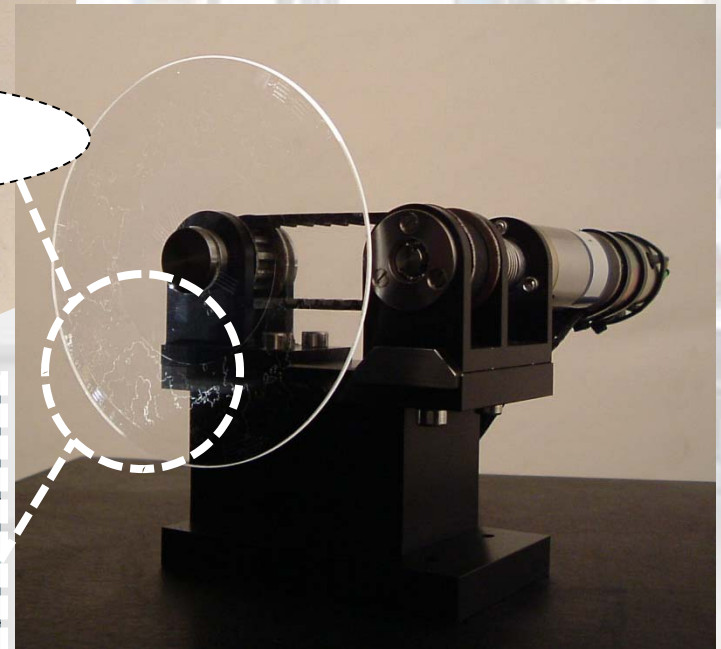
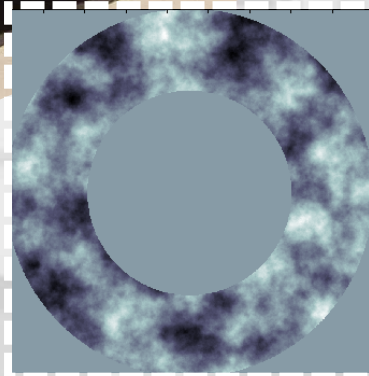
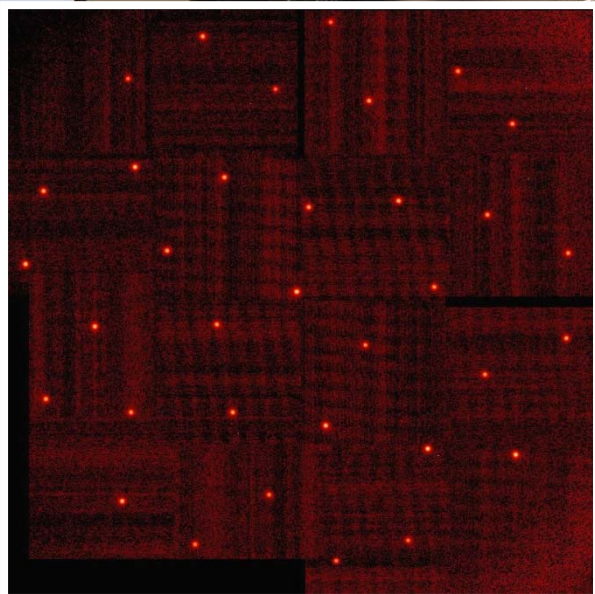
The turbulence generator MAPS

“Stars”

“VLT”

For MAD: equivalent
to being at the VLT

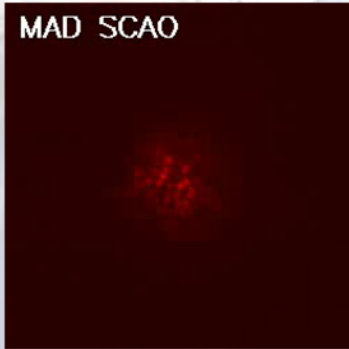
“Turbulence”





MAD performance in the lab

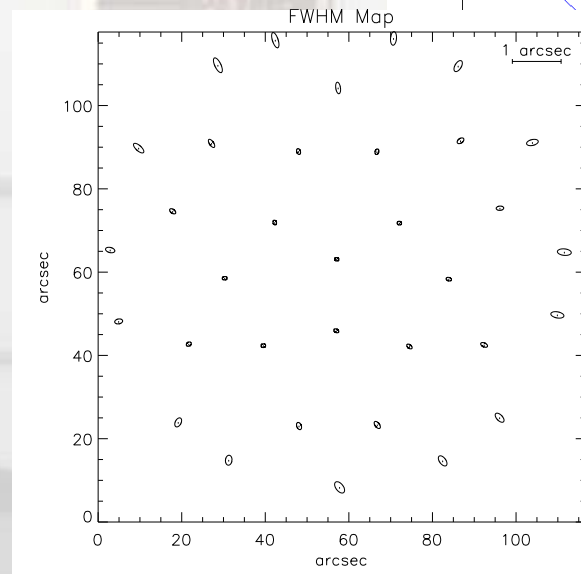
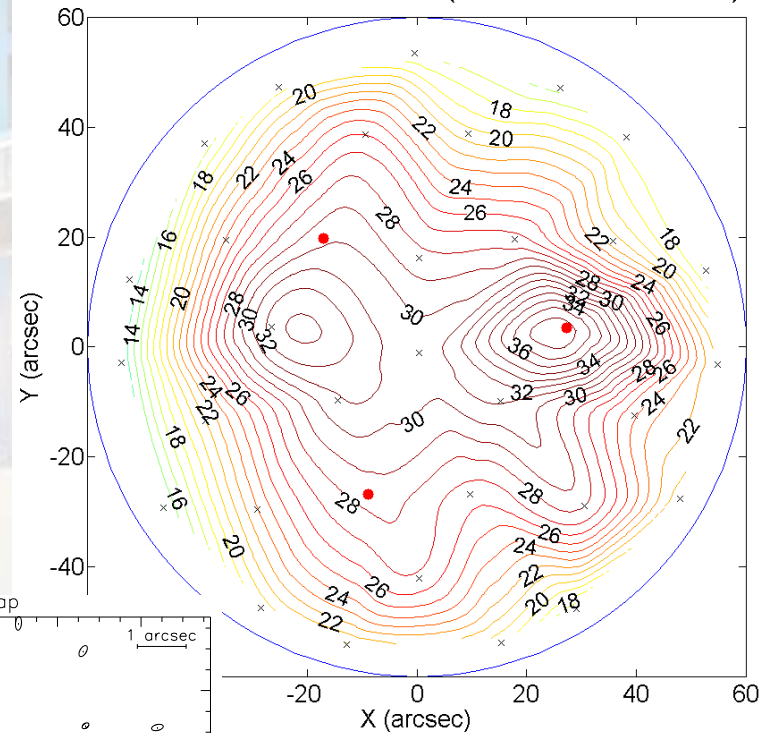
MAD SCAO



37 Technical reports,
including one on performance
evaluation (and optimization)
under different:

- Guide Stars asterisms
- Guide Stars magnitudes
- Seeing conditions

MAD MCAO 1' correction (SR in % at 2166 nm)





Installation at Paranal: Feb-Mar '07

<http://www.eso.org/projects/aot/mad/Paranal/Paranal.html>

MAD: Multi-Conjugate Adaptive Optics Demonstrator

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MAD Installation at Paranal

The MAD dismantling has been completed on January 26th and the demonstrator has been shipped to Paranal Observatory on January 30th. The MAD boxes have been delivered in Paranal on February 6th the installation started on February 10th. In this page we will show a daily update of all the MAD installation activities lasting for about 1 month. MAD is being installed at the Nasmyth Focus A. of the VLT UT3 (Melipal) telescope. The first observing night has been scheduled for March 26th and the observing run will consist of 8 nights in total.

JANUARY

26

FEBRUARY

10	11	12	13	14
15	16	17	18	19
20	21	22	23	24
25	26	27	28	

MARCH

1	2	3	4	5
6	7	8	9	10
11	12	13	14	15

MAD: Multi-Conjugate Adaptive Optics Demonstrator

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MAD dismantling: January 26th

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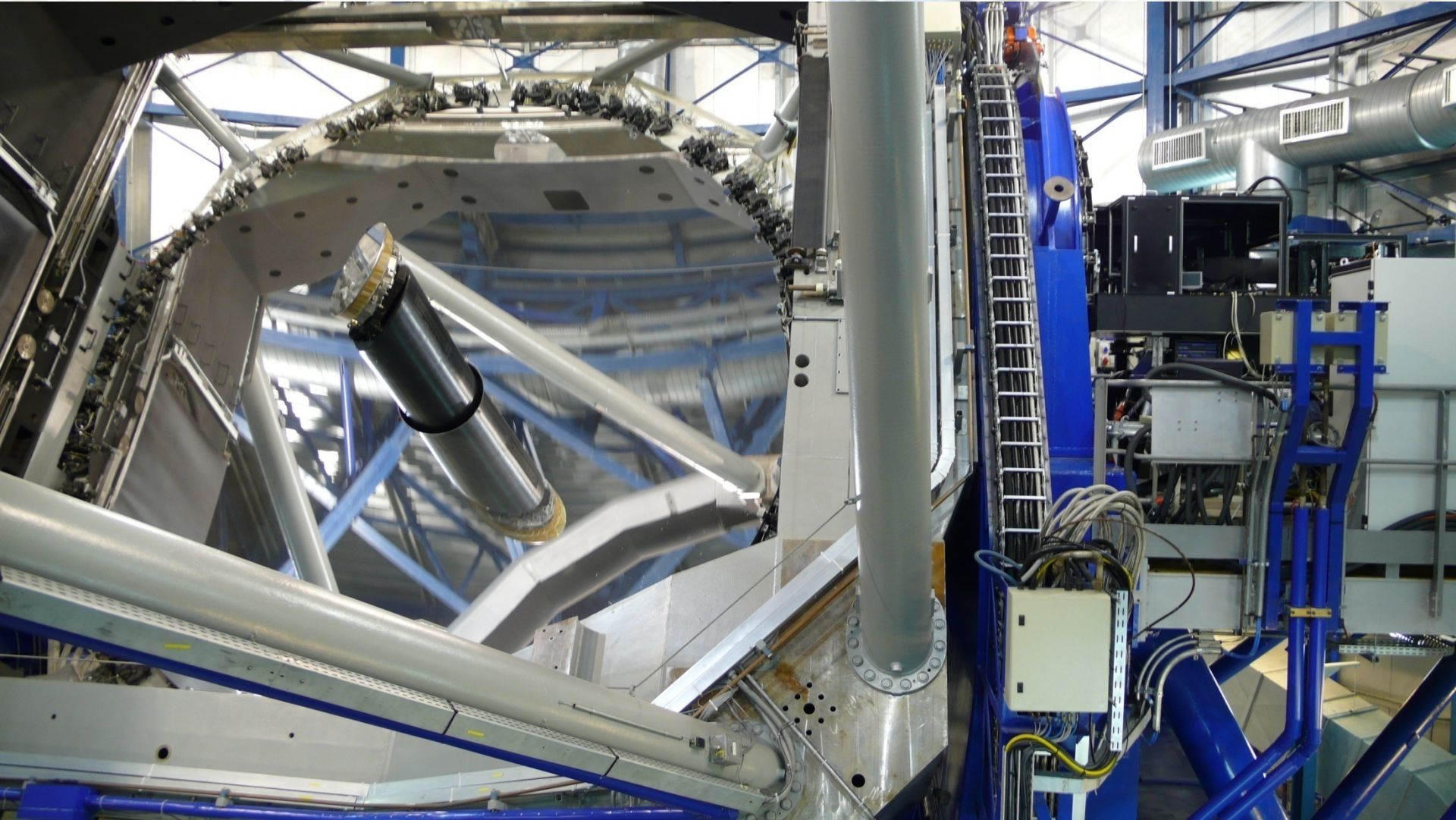
Despite the snow, the dismantling and packing of MAD has been completed well in schedule. A total of 18 boxes, 7.3 tons, are waiting for shipment to Paranal Observatory.



Done

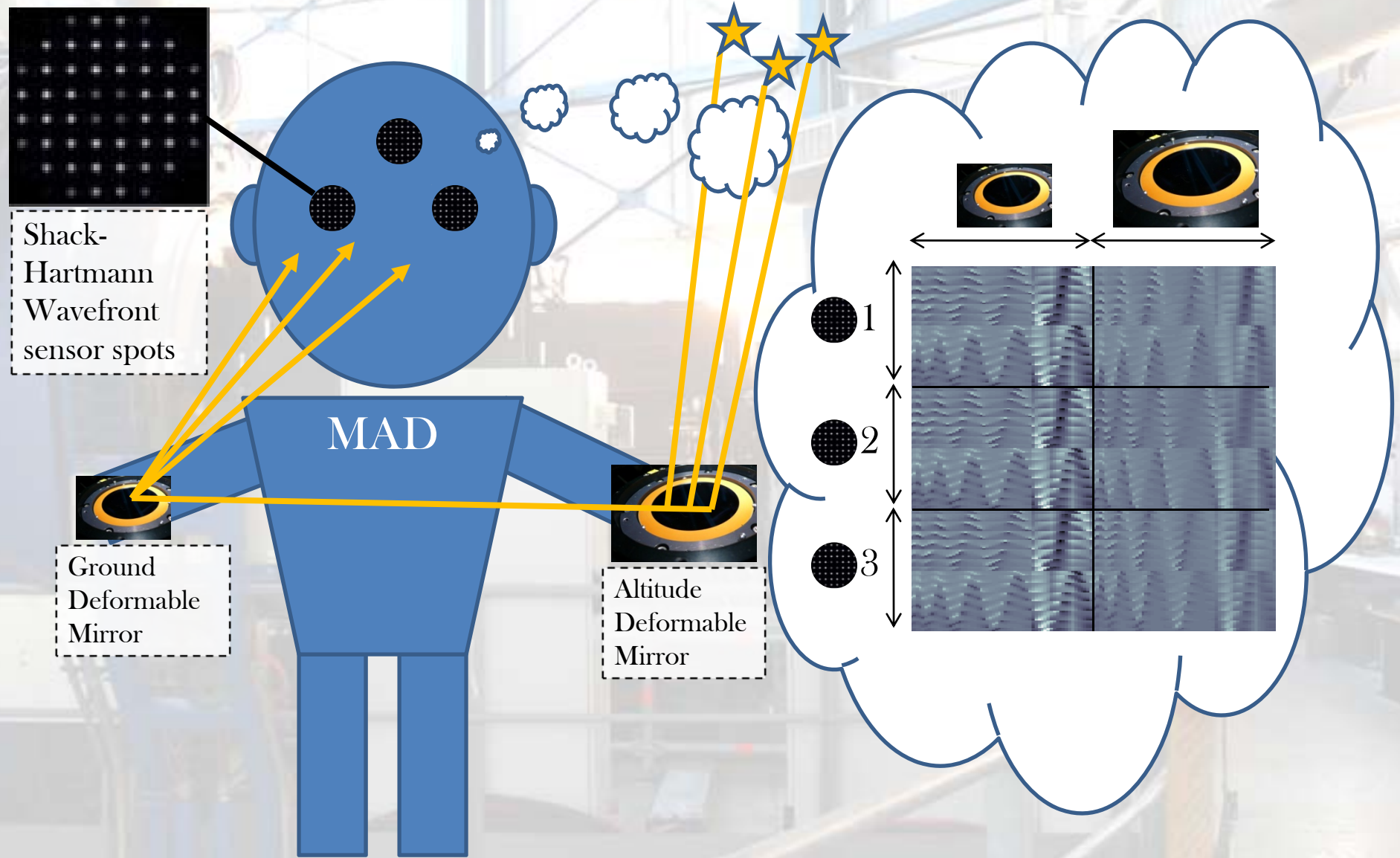


Installation at Paranal: Feb-Mar '07





Interaction Matrix, the heart of AO

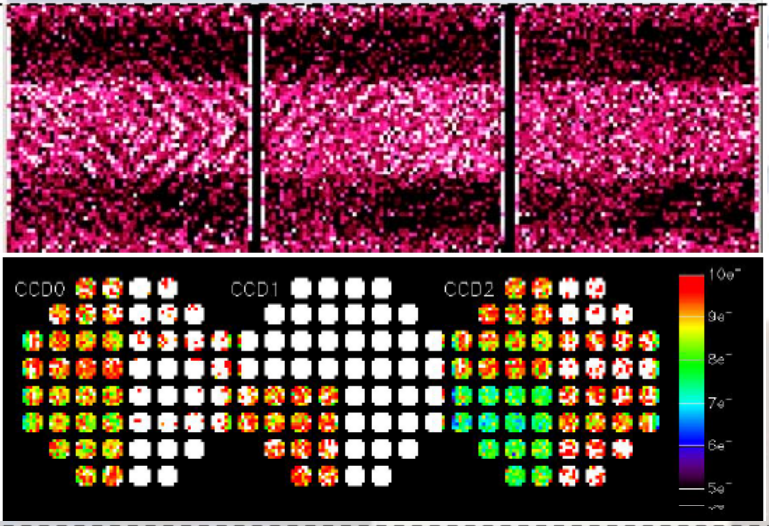




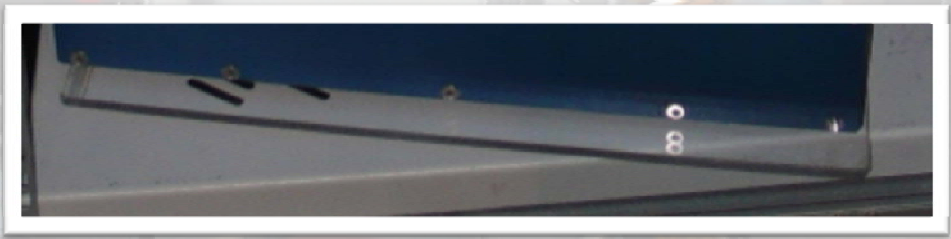
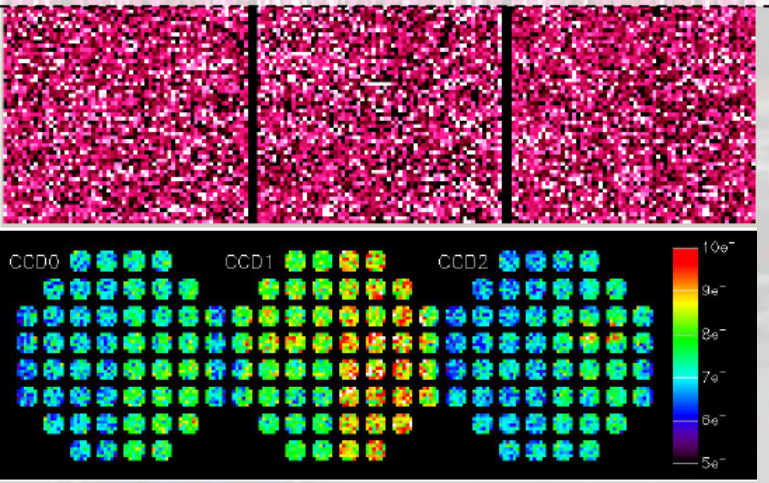
Problems... solved before
commissioning nights

Problems... solved before commissioning nights

Before: Interference on the WFS detector



After isolation of detector electronics



Problems... solved before commissioning nights

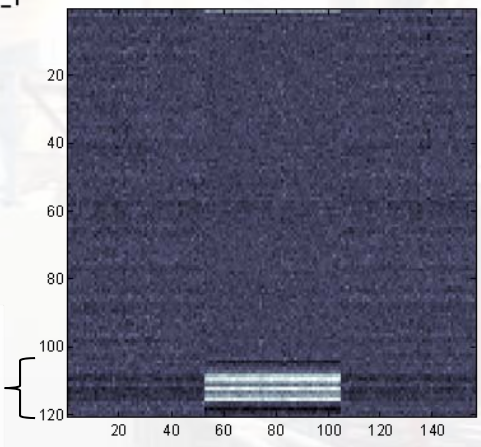
DM backplate unscrewed during transport: severe risk of damage
 → Dismounting and testing



Difference of 2 Interaction Matrices separated by a few minutes shows pattern

IM on the x axis

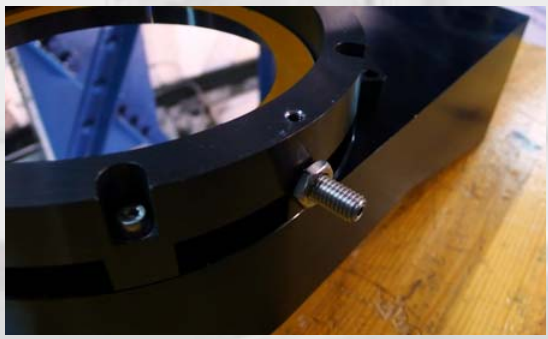
WFS #2



Deformable Mirror outer electrodes

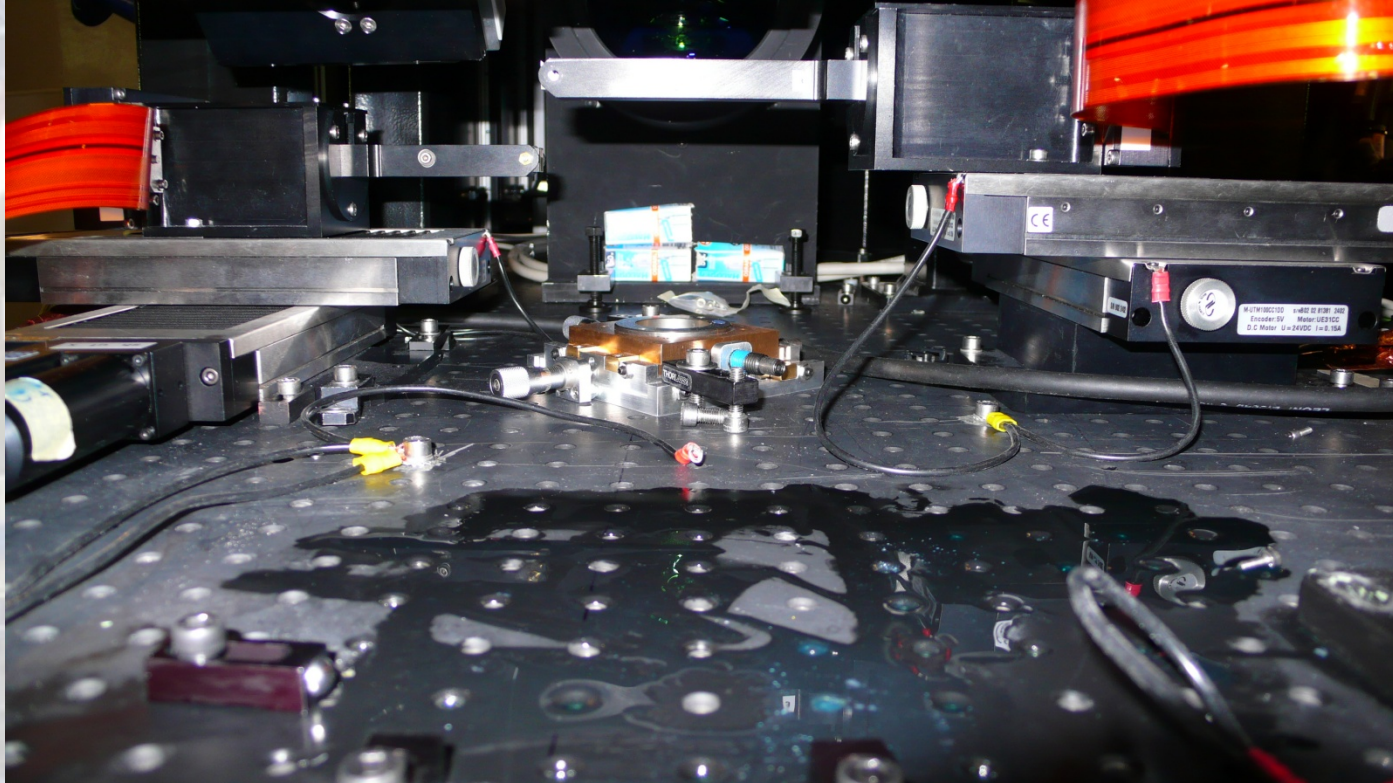


The origin was a fixation screw to be better adjusted



Problems... solved before commissioning nights

Wavefront sensor head coolant leak! 4 days before the first light

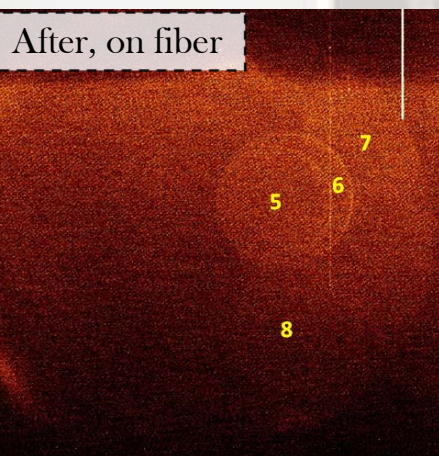
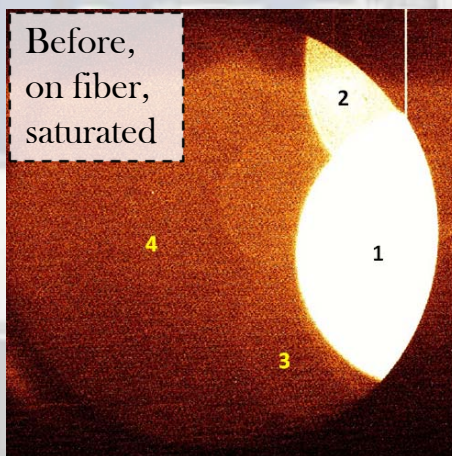
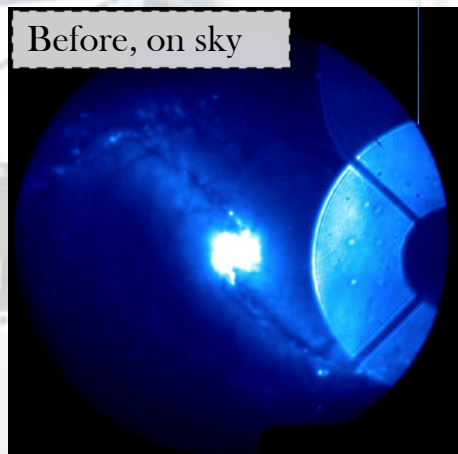
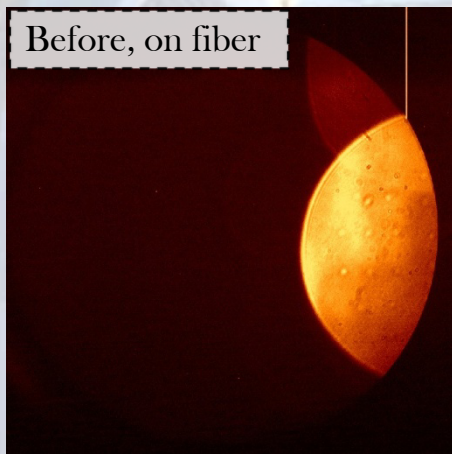


- Dismounting, replacing head, cleaning, remounting and re-alignment
- No liquid cooling used for WFS heads during all MAD observations

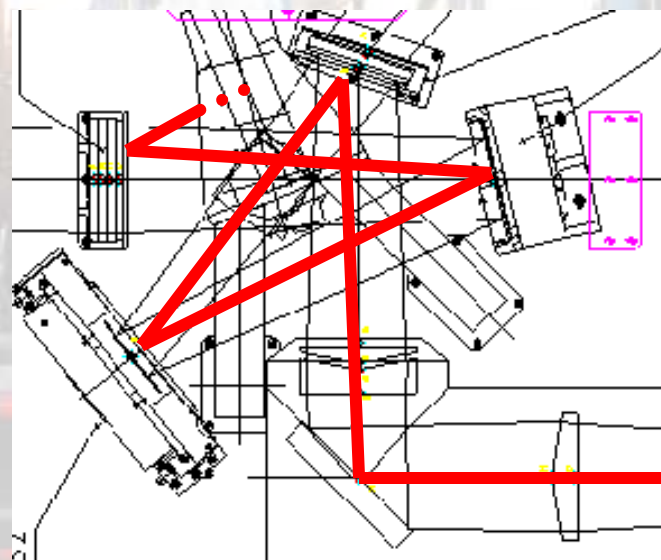


Problems... solved before commissioning nights

Static ghosts on the acquisition camera: disturb the acquisition of Guide Stars



→ Cause: Light from a linear stage limit switch LED, after multiple reflections on MAD optics

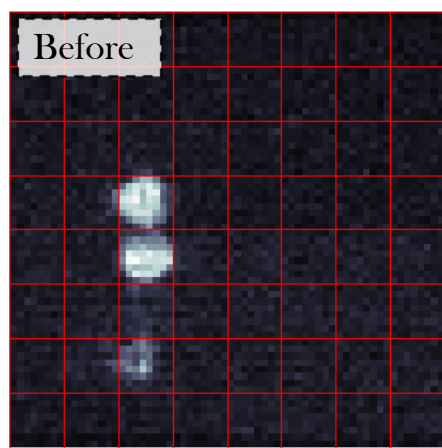


To Acq camera



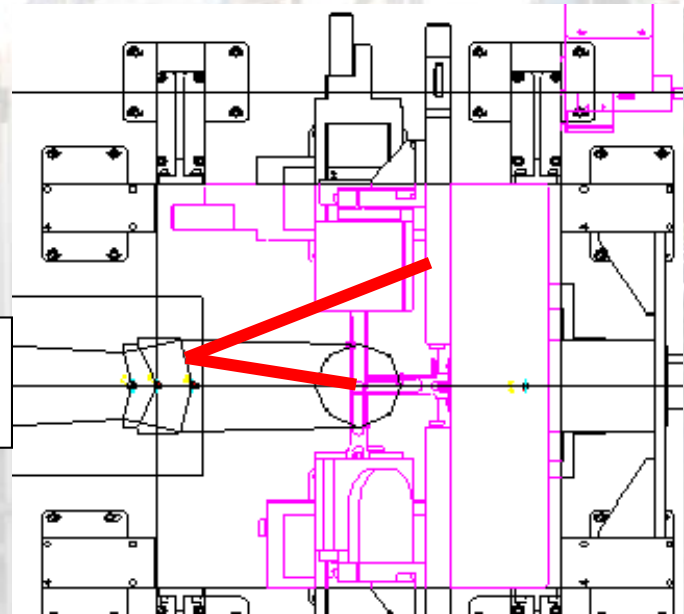
Problems... solved before commissioning nights

Ghosts on the Wavefront sensors: disturb AO close loop for certain GS positions



→ Cause: Light from another linear stage limit switch LED, after one reflections on MAD WFS objective

WFS objective



Problems... solved before commissioning nights

2 days before SD2 first night:

“a SIMM module blew up in one of the clock boards causing likely a short circuit and made the +15V power supplier module not working and the +5V one providing only 2V and this was visualized by the LED #6 on in the commboard. The error message from the selftest about a problem with the communication with the DSP was somehow misleading. We replaced the broken SIMM module and everything got back to normality”



- Solved at 17:00
- Telescope open at 19:00 for on-sky flats
- 2 hours later: MCAO closed loop with 75 mas resolution in K band over the full 2 arcmin FoV...

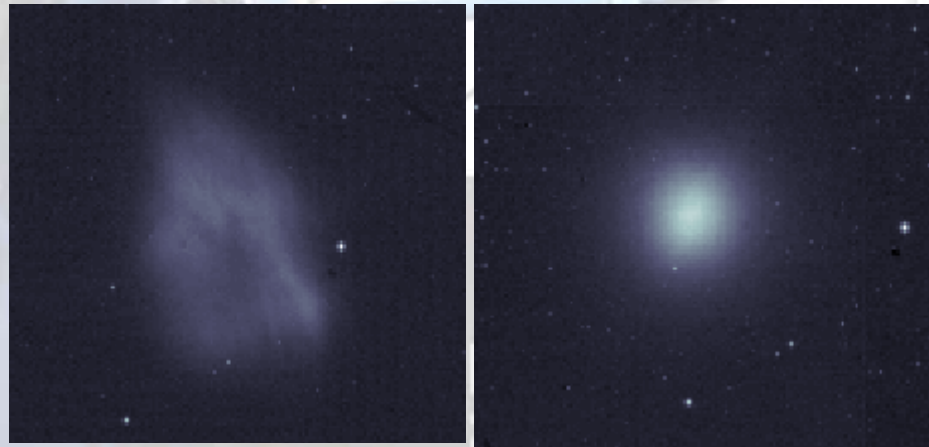


Observations



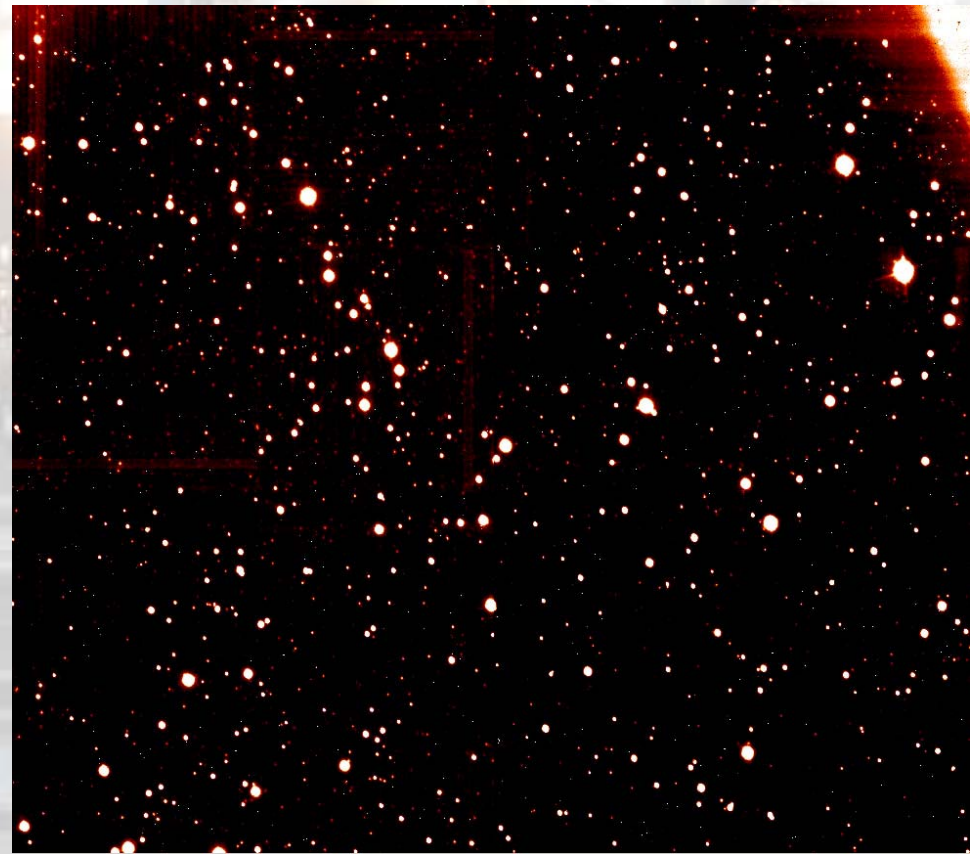
First light: without Deformable Mirrors

First blobs without and with VLT active optics



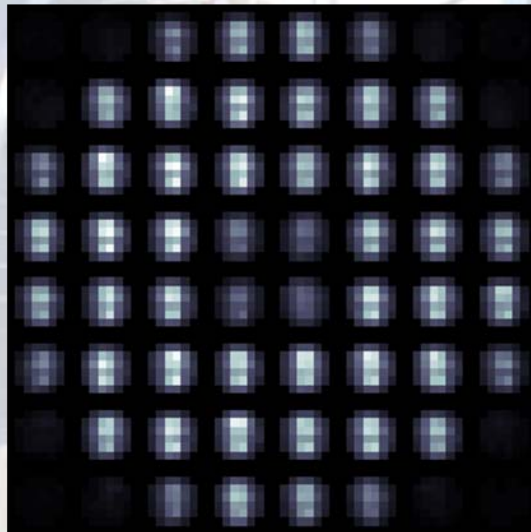
Omega Centauri - open loop

Targets for the derotator tracking test

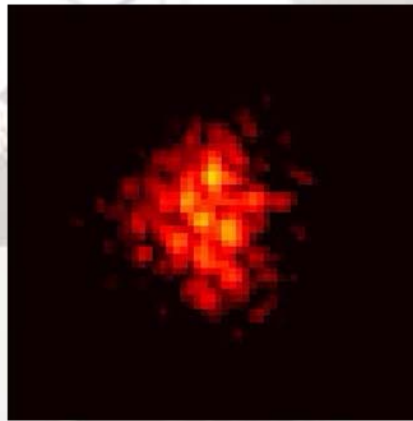


Double stars...

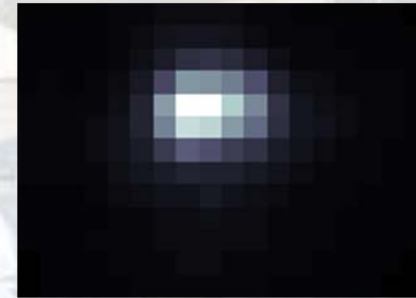
WFSing on a double star,
 $\approx 0.6''$ separation



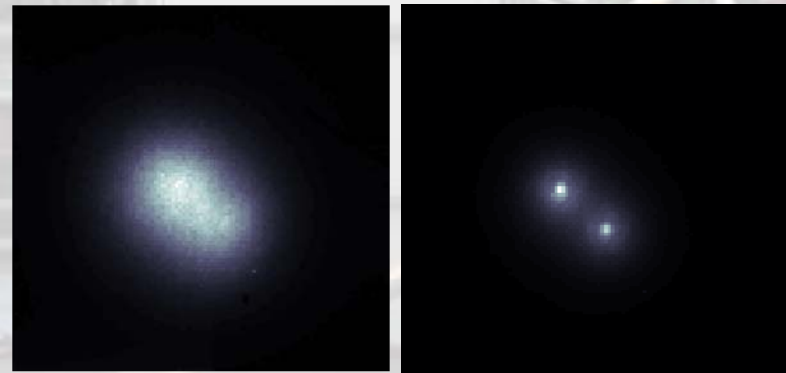
SCAO on double star
07508+0317,
173 mas separation



SCAO on a double star,
78 mas separation



SCAO on a double
star in J band,
 $\approx 0.54''$ separation

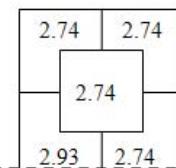




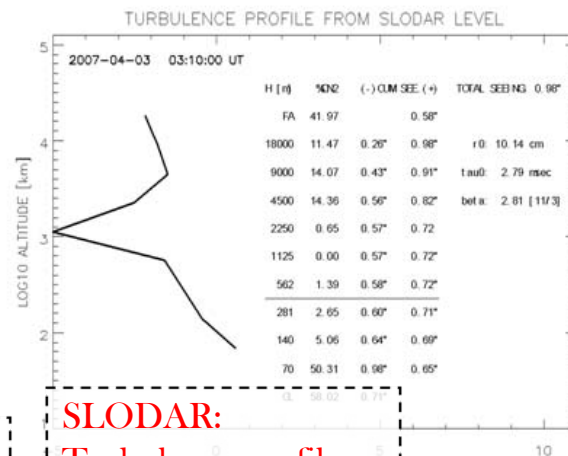
Turbulence profiling

MCAO performance dependant on turbulence strength and profile

Seeing in image mosaics



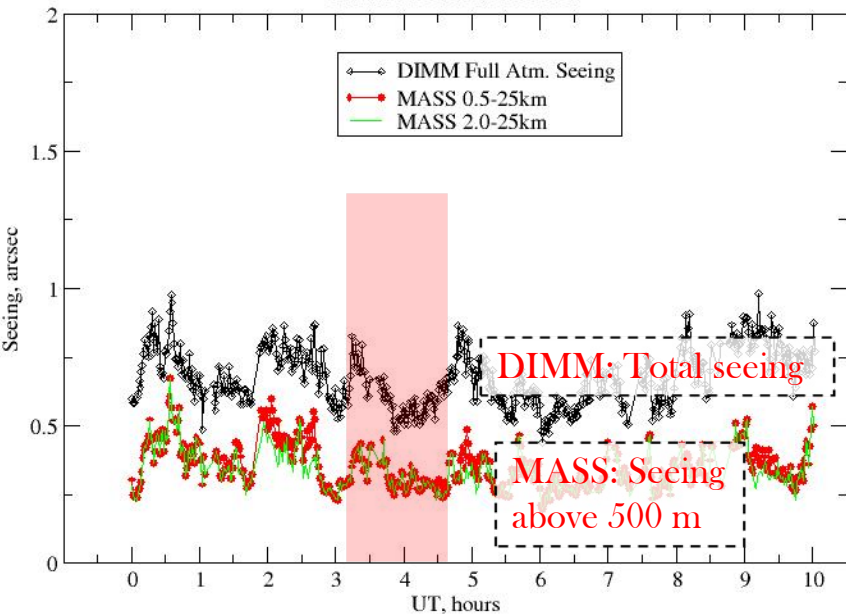
Tau₀ in image mosaics



SLODAR: Turbulence profile

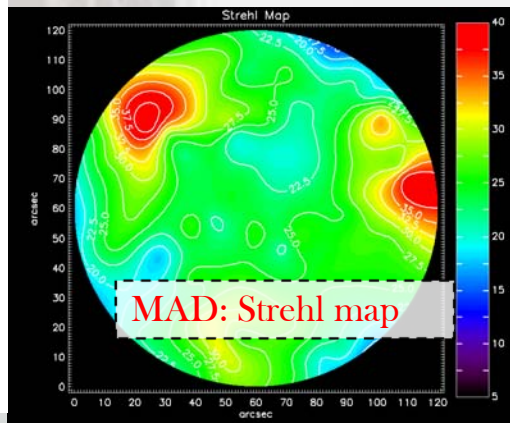
Cerro Paranal, MASS & DIMM

MASS overshoots corrected

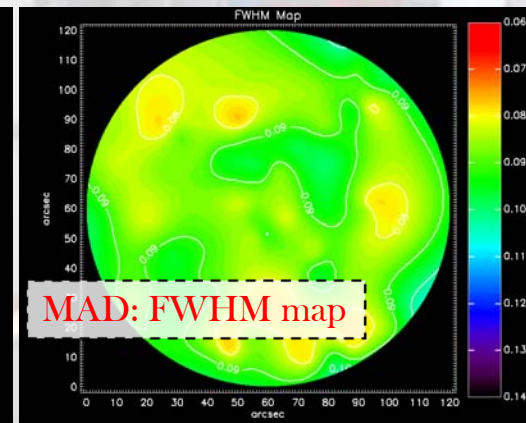


DIMM: Total seeing

MASS: Seeing above 500 m



MAD: Strehl map

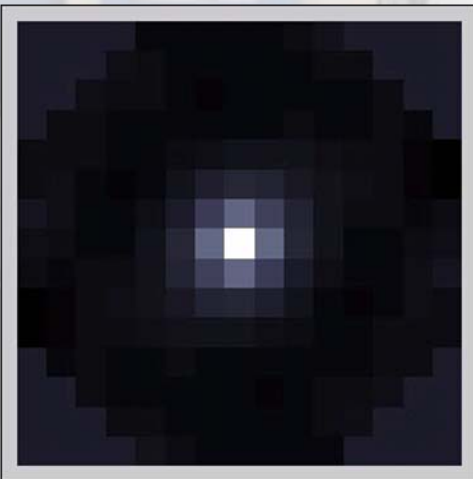
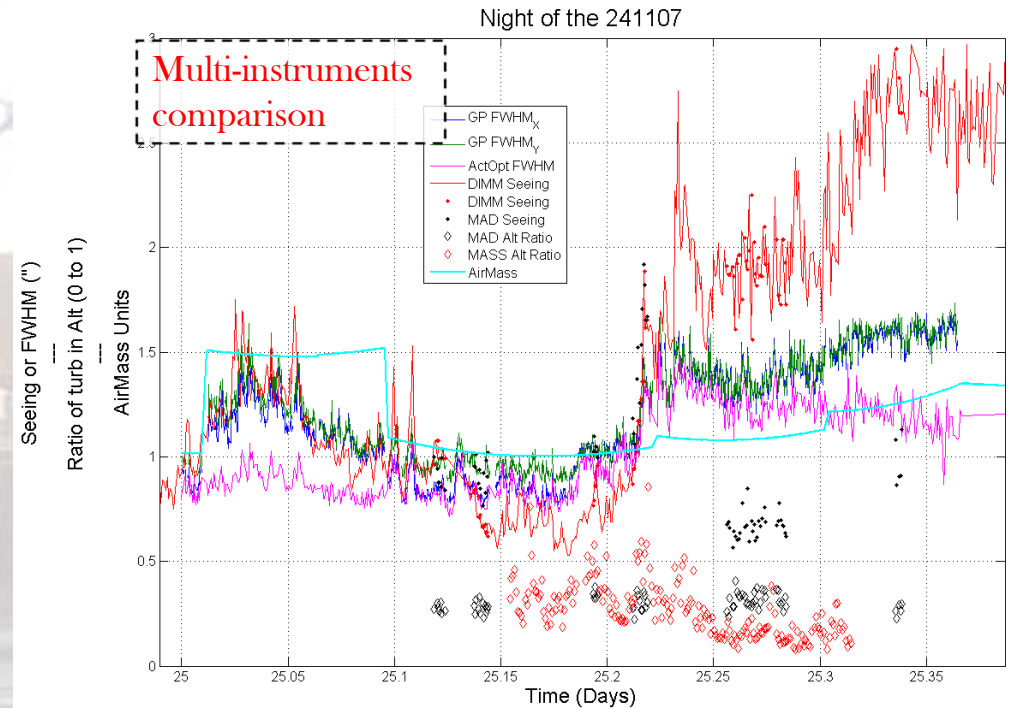
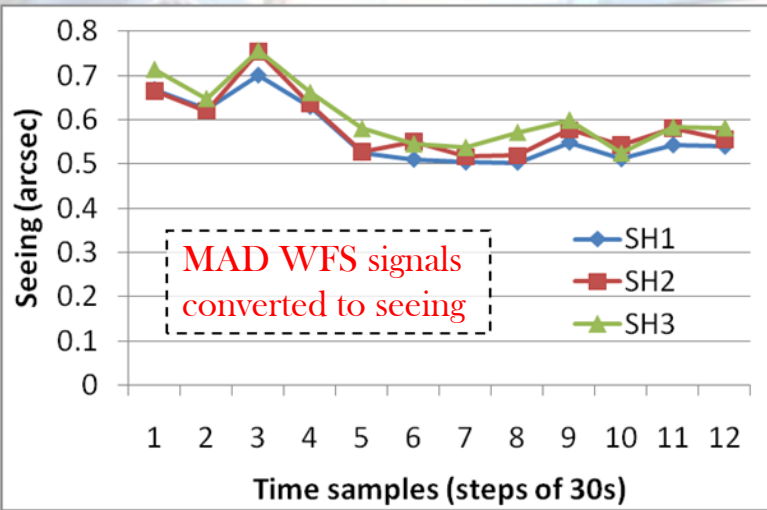


MAD: FWHM map



Turbulence study

Use MAD and other VLT tools as seeing estimators...



Temporal cross-correlation → turbulence layers strength and speed (T. Butterley, Durham)

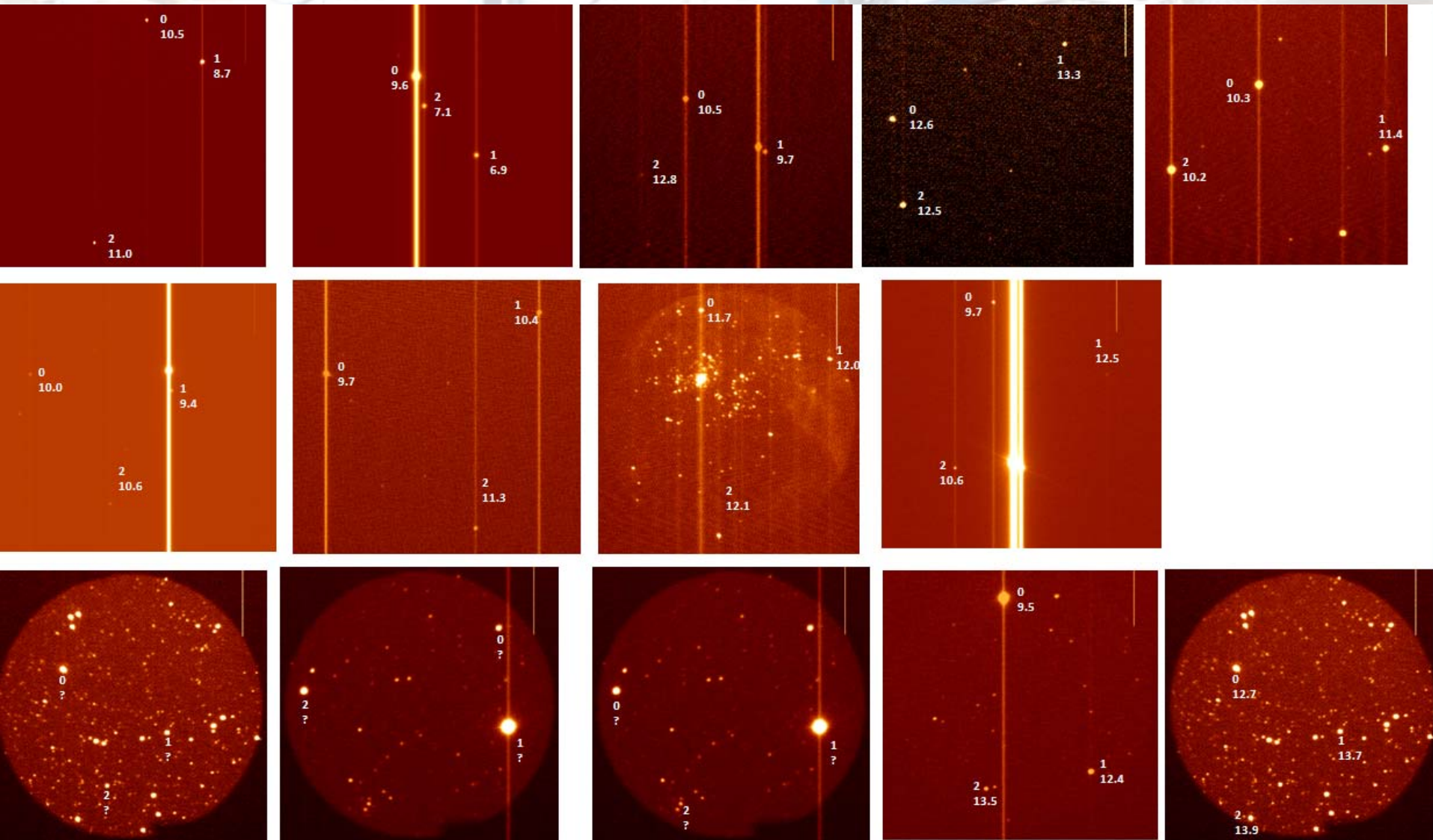


Manual filter change on CAMCAO...



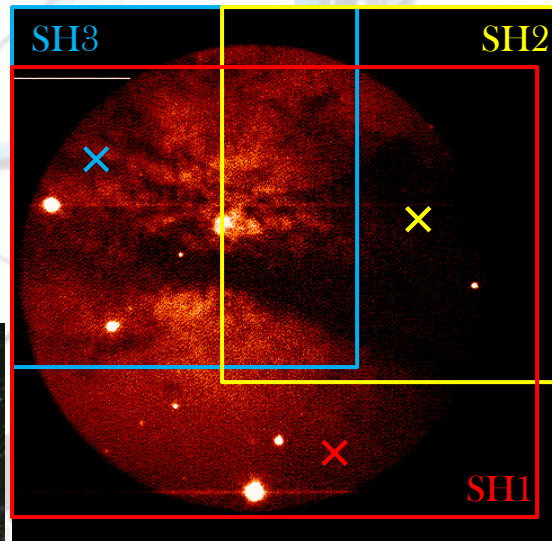


Acquisition camera



Acquisition with MAD

Acquisition of the 2 arcmin FoV with a visible CCD



Position the WFSs

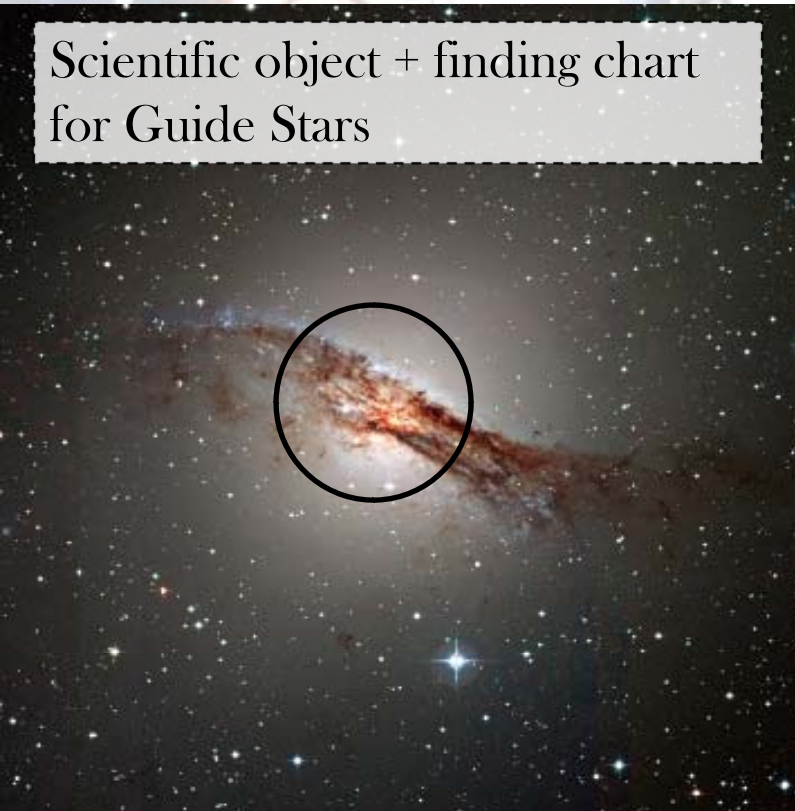
Center the WFSs

Record the IM

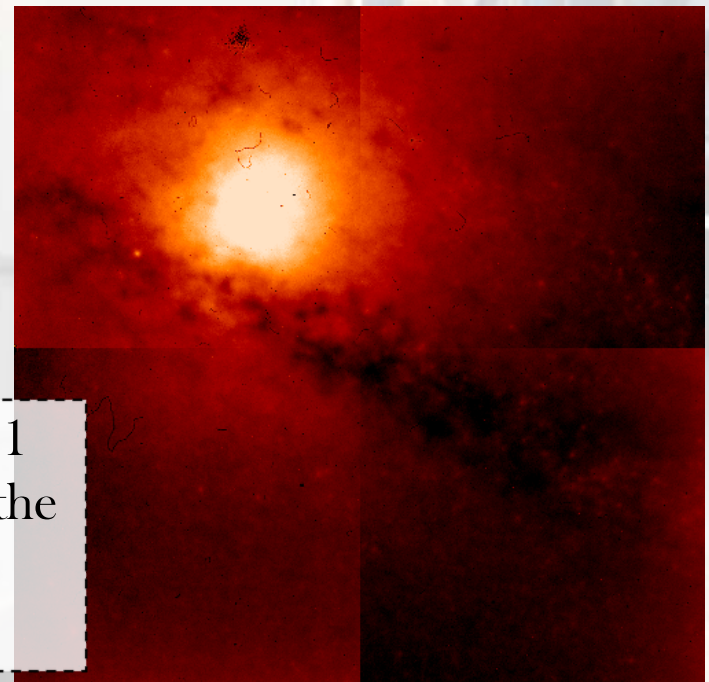
Invert the IM

CLOSE THE LOOP

Scientific object + finding chart for Guide Stars



Imaging of 1 arcmin on the IR camera CAMCAO





To be continued...

Lessons from MAD:
From the Lab to the Sky

Paola Amico