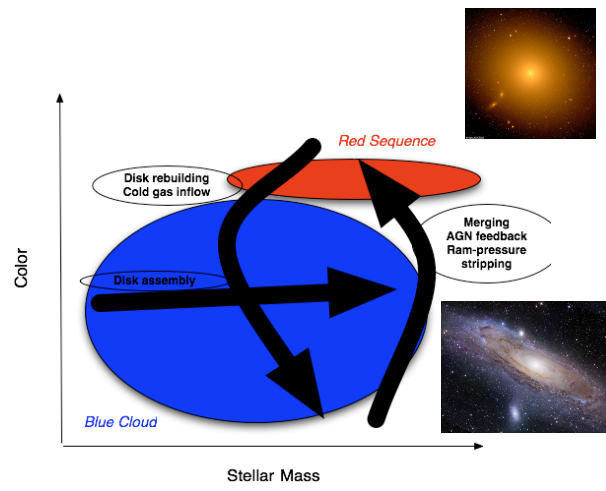


# Morphologies of galaxies in intermediate redshift clusters

*M. Huertas-Company, G. Soucail, G. Foex, R. Pelló*

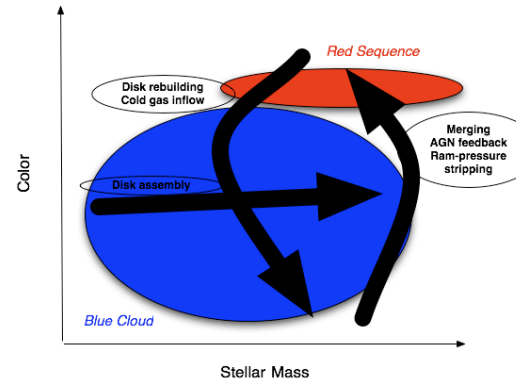
*ESO - Chile*

# Morphology/Color bimodality



# Some opened questions

- Ellipticals?
  - how and when did they form their stars? downsizing: more massive galaxies formed stars faster. AGN feedback?
  - when were they assembled? strong size evolution of passive galaxies. Dry major/minor mergers?
- Spirals?
  - inside the blue cloud: disk rebuilding?
  - from the blue cloud to the red sequence?
  - already formed at  $z \sim 1$ ?



# clusters: laboratories for testing effects of environment

- highest density in the universe
- laboratories for studying the effects of the surrounding environment in the morphological transformation of galaxies

• **when and how E-T galaxies were assembled?**

- *requirements:*
  - high angular resolution (morphology)
  - spectroscopy (cluster membership)

# Sample

**Wide imaging: ~ 12 Mpc @  
z~0.5  
(see Gael's talk)**

9 massive clusters at z~0.5

4 photometric bands (g,r,i,z)

**Completeness: r=25**

**Stellar mass completeness:  
50% at  $\text{Log}(M/M_{\text{sol}})=9.5$   
(red galaxies)**

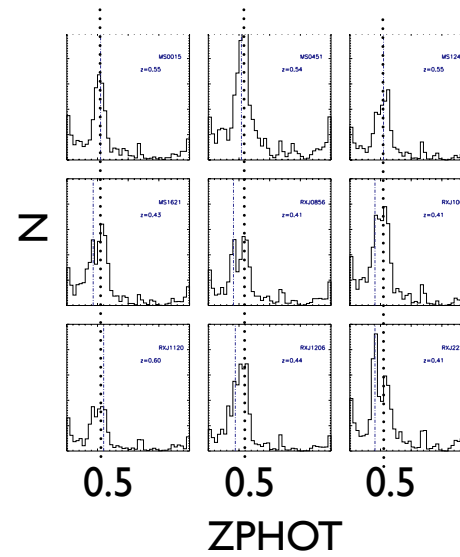


# cluster membership

- photometric redshifts, HyperZ, 4 filters: g, r, i, z
- 2 selection criteria:

$$|z_{phot} - z_{cluster}| < 0.1$$

$$P_{clust} = \int_{z_{cluster}-0.1}^{z_{cluster}+0.1} P(z) dz > 0.55$$



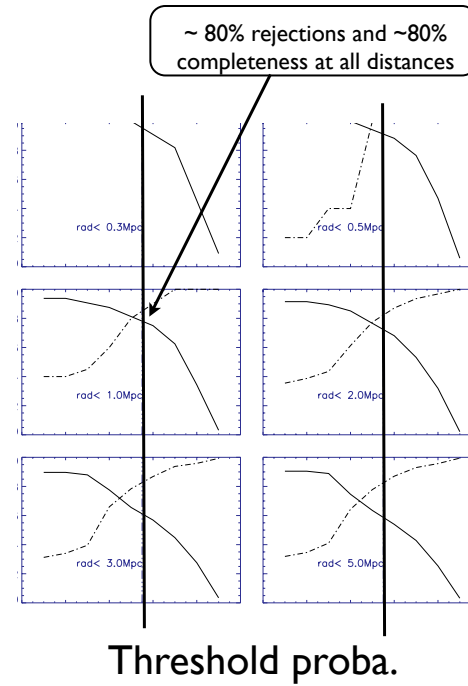
# cluster membership

- photometric redshifts, HyperZ, 4 filters: g, r, i, z
- 2 selection criteria:

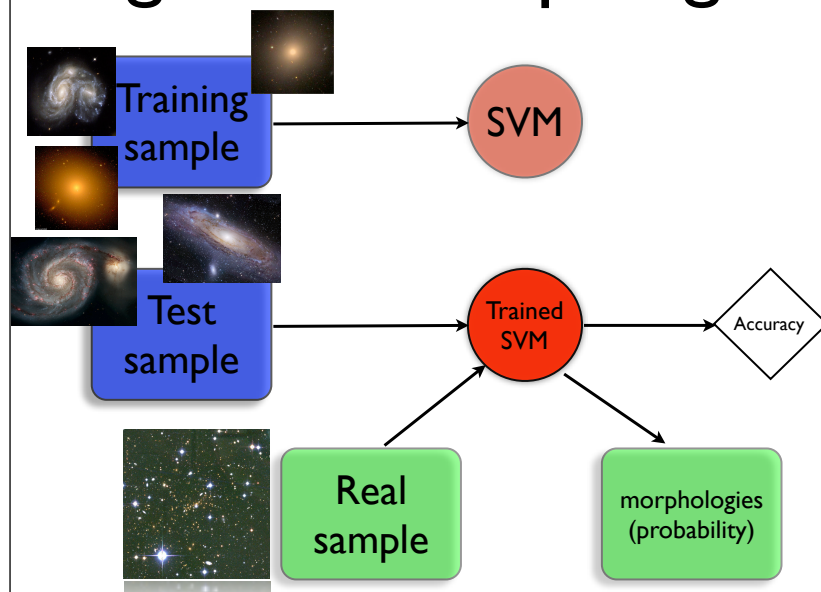
$$|z_{phot} - z_{cluster}| < 0.1$$

$$P_{clust} = \int_{z_{cluster}-0.1}^{z_{cluster}+0.1} P(z) dz > 0.55$$

Fraction

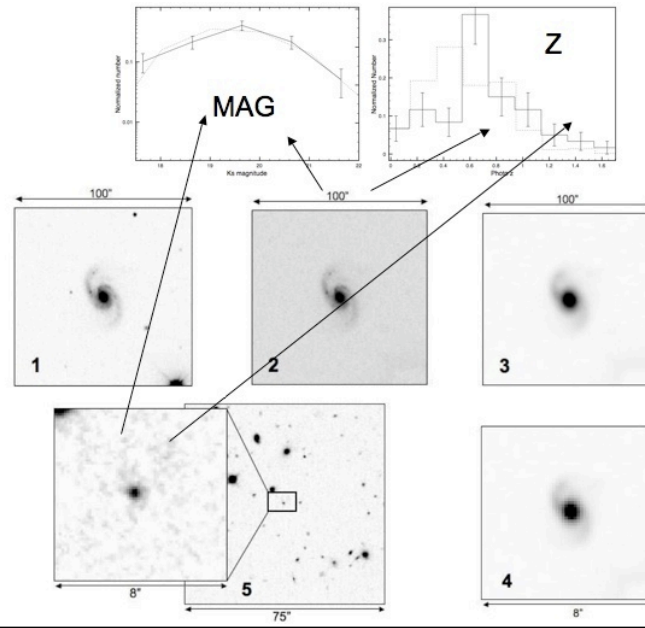


# galSVM morphologies (I)



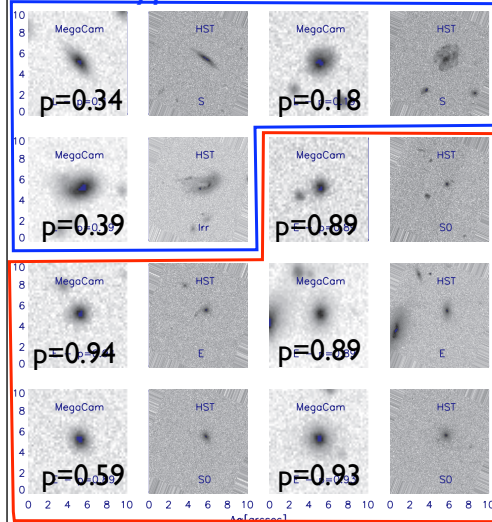


# galSVM morphologies (II)

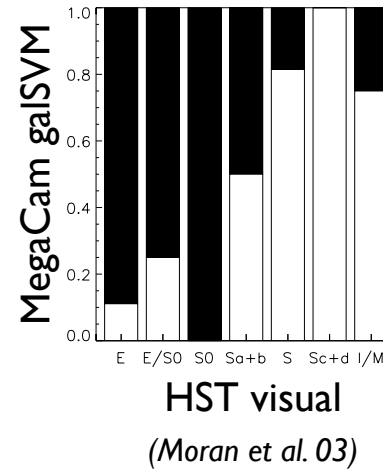


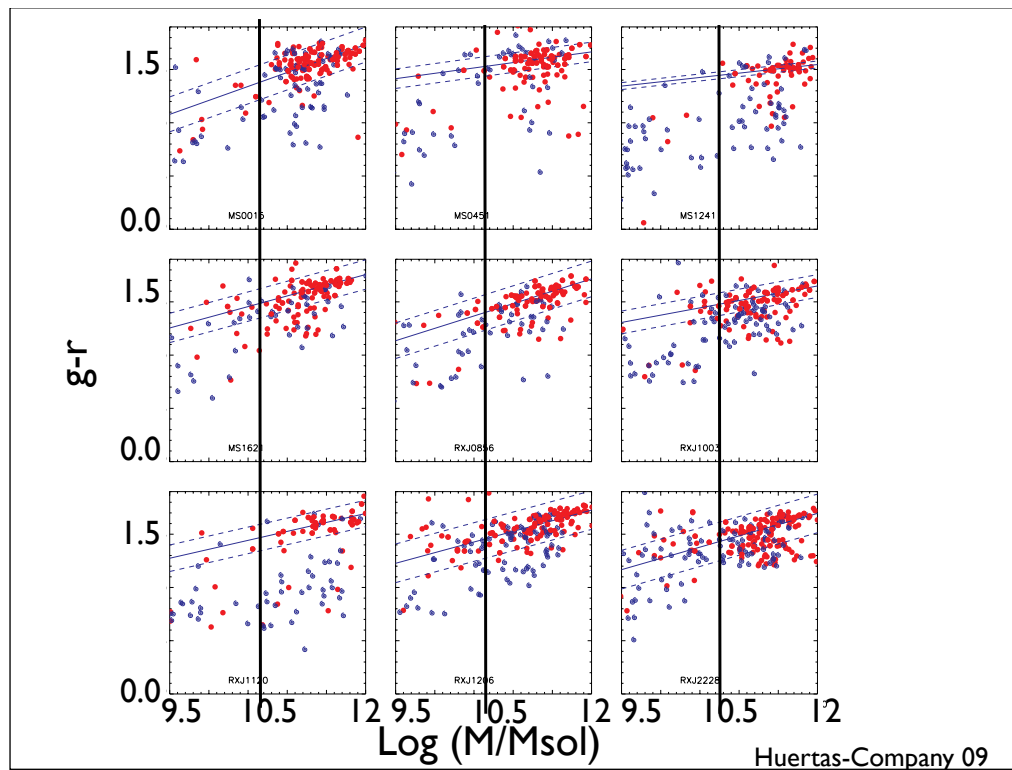
# Cross-checking with HST

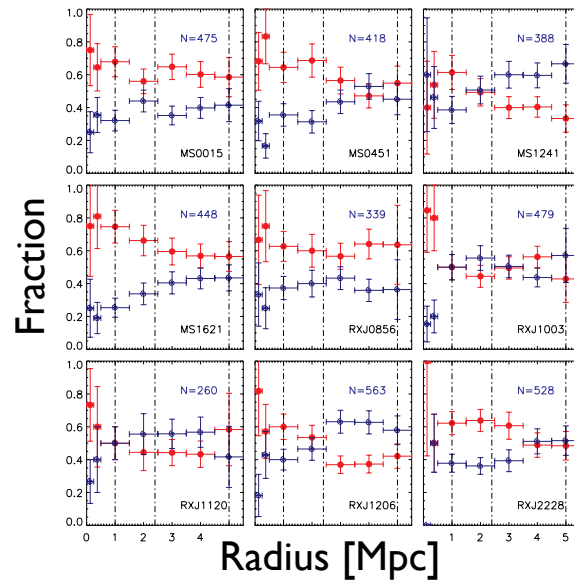
Late-Type



Early-Type



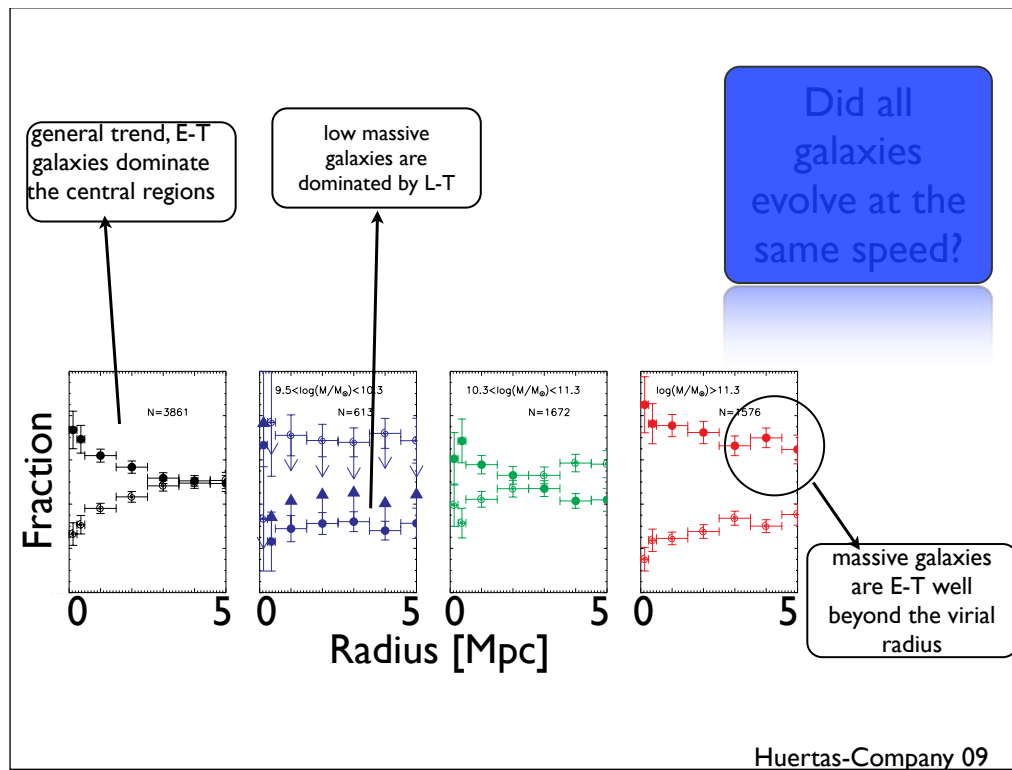


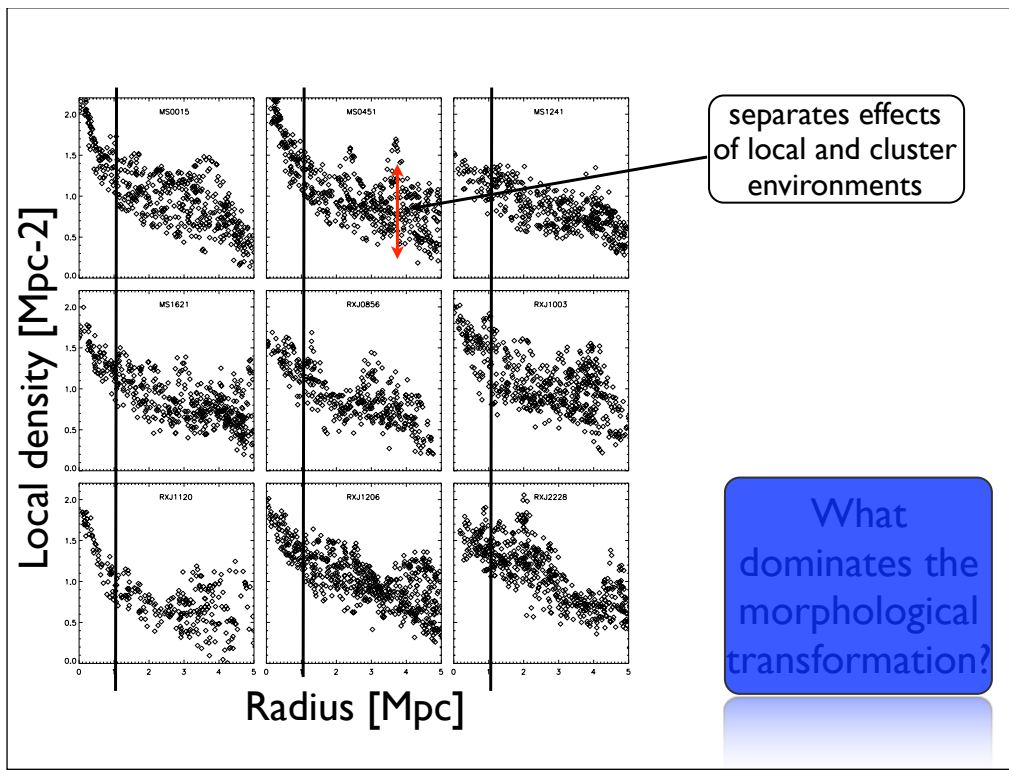


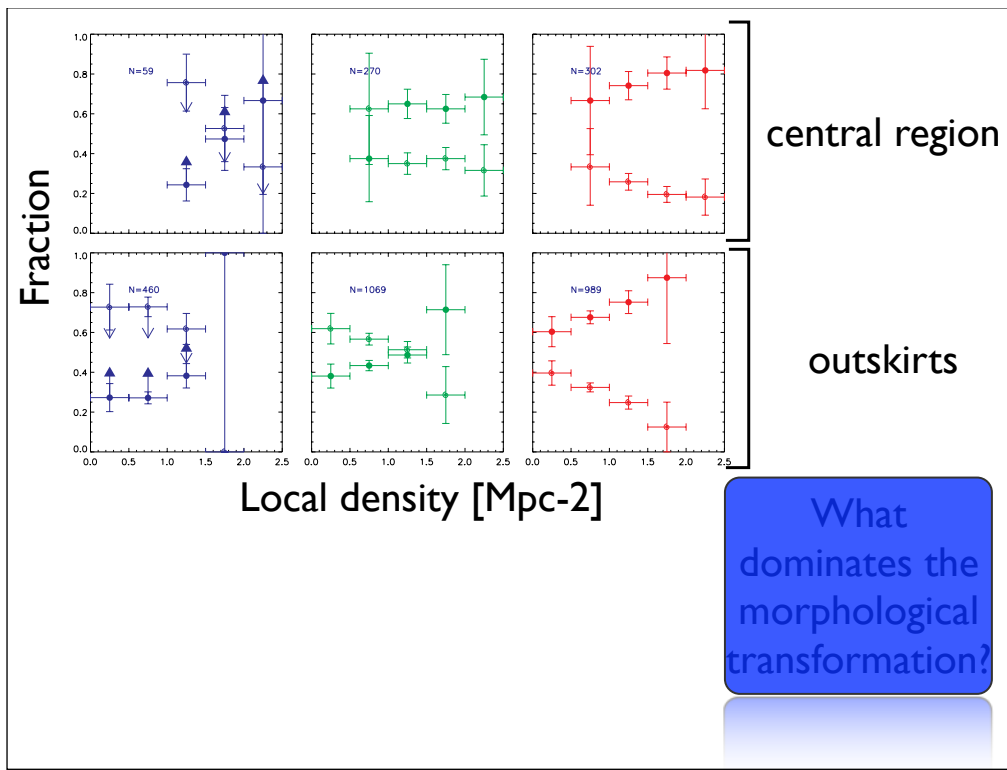
E-T galaxies  
dominate the  
central regions

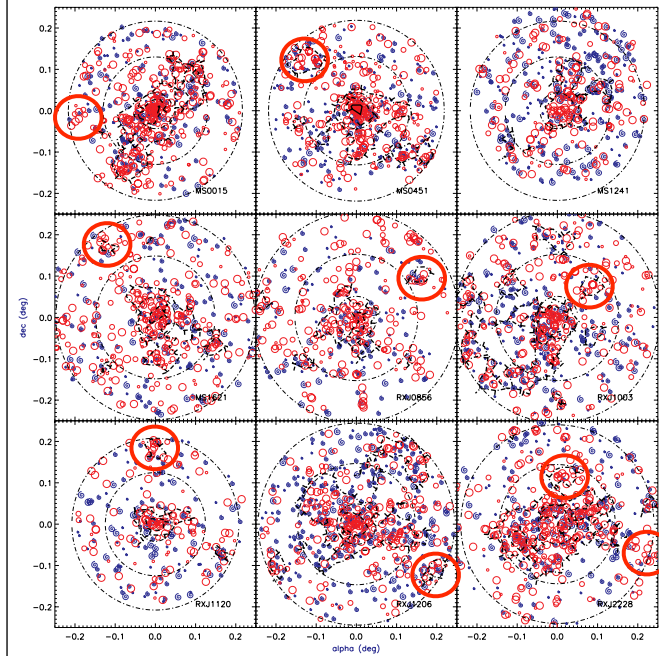
Did all  
galaxies  
evolve at the  
same speed?

What  
dominates the  
morphological  
transformation?









high density regions in the outskirts are dominated by E-T galaxies

What dominates the morphological transformation?



# Conclusions

- morphology density relations @  $z \sim 0.5$  with a 4m class telescope and no spectroscopy
  - control samples for high- $z$  studies?
- RS already in place at  $z \sim 0.5$
- Massive E-T formed earlier outside the clusters
- Most of the morphological evolution is taking place in intermediate mass galaxies
- What's next?
  - used to find clusters (over densities of E-T galaxies)