



Mass Assembly Survey with SINFONI in VVDS

Thierry Contini

LATT - Observatoire Midi-Pyrénées - Toulouse





Mass Assembly Survey with SINFONI in VVDS

People & Places

About 15 astronomers located in France, Italy, Germany and United Kingdom are involved in the MASSIV project

LATT – Toulouse

T. Contini (PI)

F. Lamareille

J. Moutaka

E. Perez-Montero

J. Queyrel

LAM – Marseille

P. Amram

B. Epinat

O. Le Fèvre

L. Tresse

INAF – Milano

B. Garilli

L. Paoro

INAF – Bologna

S. Bardelli

D. Vergani

E. Zucca

ESO – Garching

M. Kissler-Patig

Oxford Astrophysics

M. Lemoine-Busserolle

Staff – Postdoc – PhD student – Database manager

Galaxy Kinematics at High Redshift: **Why?**

Probing the **mass assembly history of galaxies**
at the crucial **$1 < z < 2$** epoch:

- peak of the cosmic star formation activity
- transition between the morphological diversity observed @ $z \sim 3-4$ and the modern-day Hubble sequence
- Complete census of the **mix of dynamical types**
mergers, spheroids, disks, perturbed kinematics, etc
- Determination of **dynamical mass, mass-to-light ratio,**
and **metallicity**
- Evolution of fundamental scaling relations:
Tully-Fisher, Mass-Metallicity, Size-Velocity, etc

Galaxy Kinematics at High Redshift: **How?**

- At $z < 1.5$: velocity fields obtained with **optical spectrographs**, [OII]3727 emission line is still visible
- At higher redshifts ($1.5 < z < 3.5$), **NIR observations** are mandatory:
[OII] in J-band, H β & [OIII] in H-band, H α in K-band
- First attempts on small samples using **long-slit NIR spectrographs** (ISAAC/VLT, NIRSPEC/Keck, ...)
Pettini et al. (2001), Lemoine-Busserolle et al. (2003), Erb et al. (2003,2004), van Starkenburg et al. (2006)

but ... long-slits have **strong limitations**
- **Integral-field spectrographs (IFU)** are better to derive **accurate velocity fields**
Flores et al. (2004) @ $z \sim 0.5$, Swinbank et al. 2005, Forster-Schreiber et al. (2006) @ $z \sim 2.0$, Law et al. (2007) @ $z \sim 2-3$



Mass Assembly Survey with SINFONI in VVDS

VVDS offers the best suited parent sample to select high- z galaxies for NIR 3D spectroscopic follow-up :

- **largest & deepest** spectroscopic survey of distant universe
- well-defined & **minimal bias**
- complete, **representative** & statistically significant
- secure & **accurate redshifts** for efficient NIR follow-up





Mass Assembly Survey with SINFONI in VVDS

ESO "Large Program" (PI: T. Contini)
started in P79 for **2 years (2007-09)**
--> **300 hours in Service mode**



Sample: ~ **140 VVDS star-forming galaxies selected**
in three fields F02, F22 & F14

- spanning a **wide range of stellar masses** $\log(M)=[9,12]$
- with "**bright**" **[OII] emission** up to $z \sim 1.4$
- selection in the **ultra-deep** ($I_{AB} < 24.75$) VVDS sample **for $z > 1.4$**

Strategy: **AO-assisted SINFONI** observations

- *JH* bands to observe redshifted **bright emission lines** ($H\alpha$, [OIII])
- **high spatial resolution** ($\sim 0.2''$) to probe the dynamics
from EL velocity fields



Mass Assembly Survey with SINFONI in VVDS

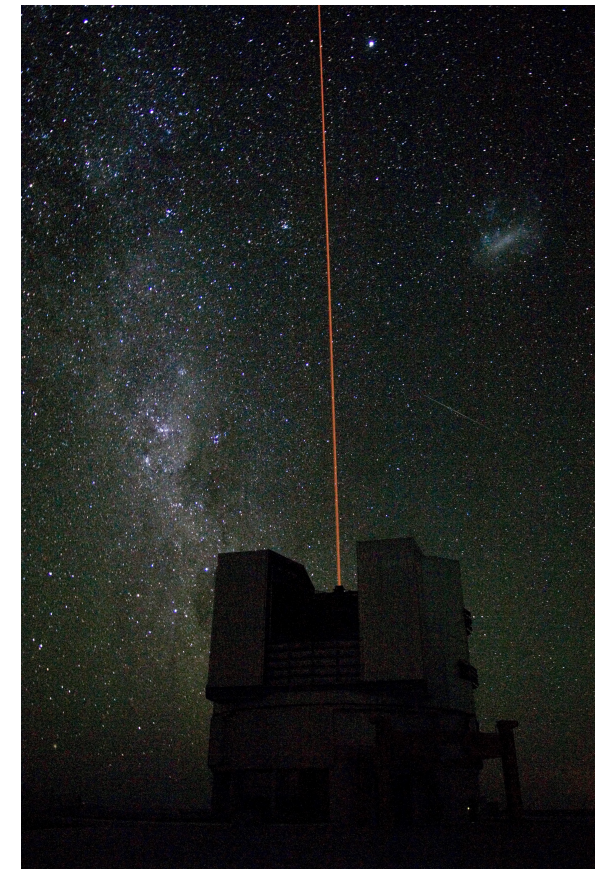
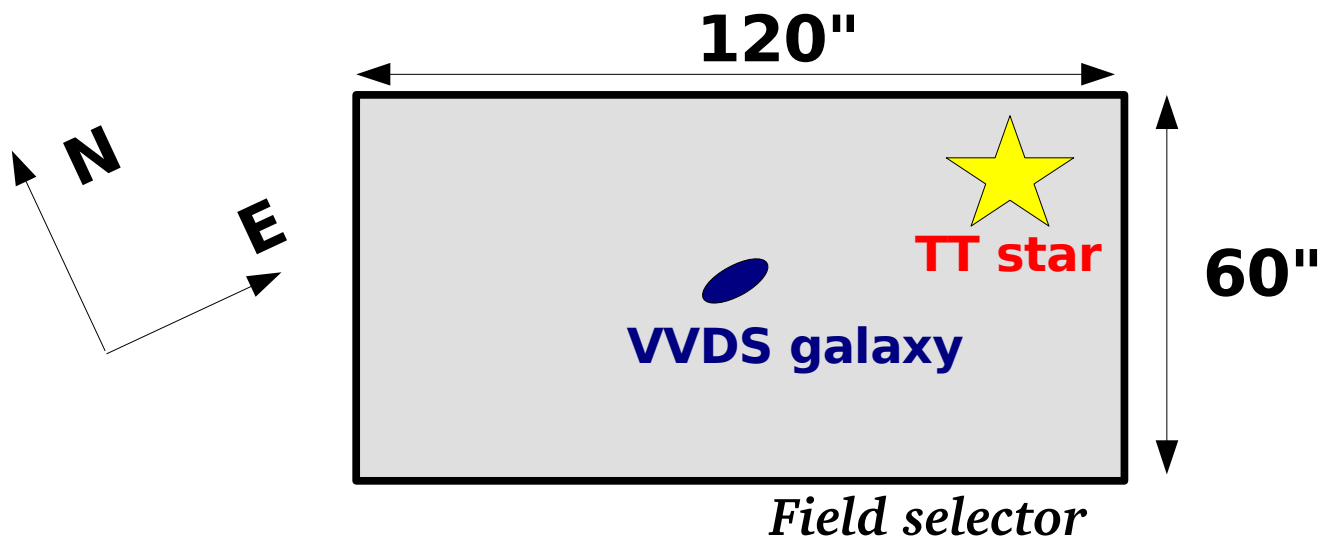
Sample selection criteria

- **Secure redshift:** $z_{\text{flag}}=2-9$ (or 22-29) + visual (+ z_{phot}) check
 - $z=0.900-1.049$ -> $H\alpha$ in J band
 - $z=1.193-1.240$ -> [OIII] in J band
 - $z=1.240-1.814$ -> $H\alpha$ in H band
- **Avoid the $z=1.049-1.193$ range** ($H\alpha$ between the J & H bands)
- **Avoid bright sky lines** ($\Delta\lambda \geq 9\text{\AA}$ for $f_{\text{OH}} > 100$)
- **Lower limits on [OII]3727 EW:**
 - $\text{EW}([\text{OII}]) < -40$ and $\text{S/N} > 6$ **or**
 - $\text{EW}([\text{OII}]) < -25$ and $\text{S/N} > 10$

Mass Assembly Survey with SINFONI in VVDS

AO/LGS assisted observations

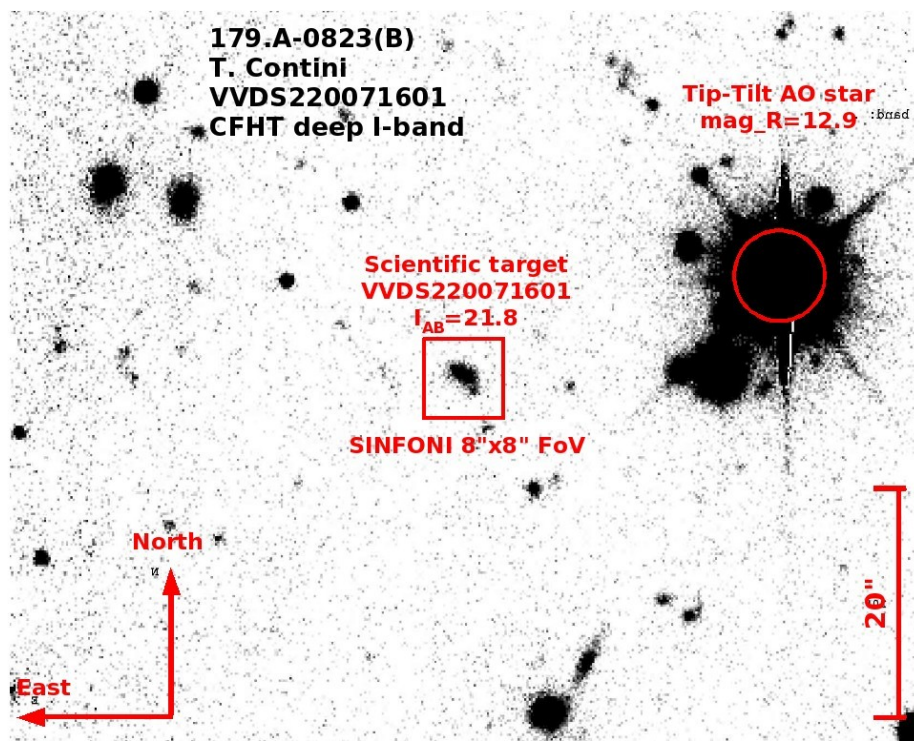
--> **constraints on distance/mag of TT star**



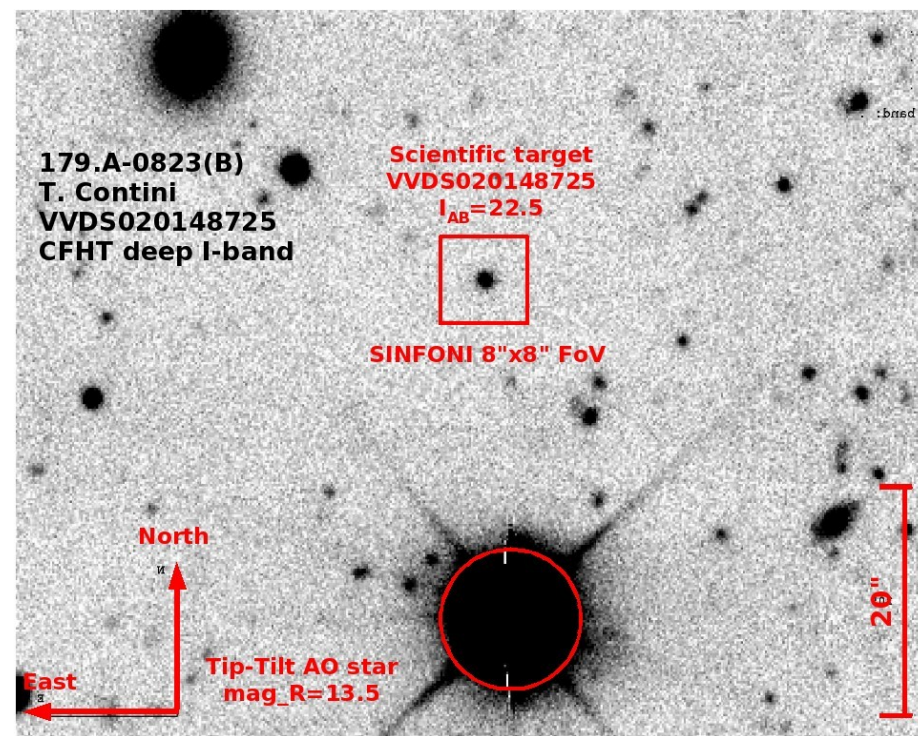
- **J band:** $\text{magR} < 16.5$ (< 15.5 if $\text{dist} < 30''$)
- **H band:** $\text{magR} < 17.5$

Mass Assembly Survey with SINFONI in VVDS

Examples of $1.0 < z < 1.4$ VVDS targets

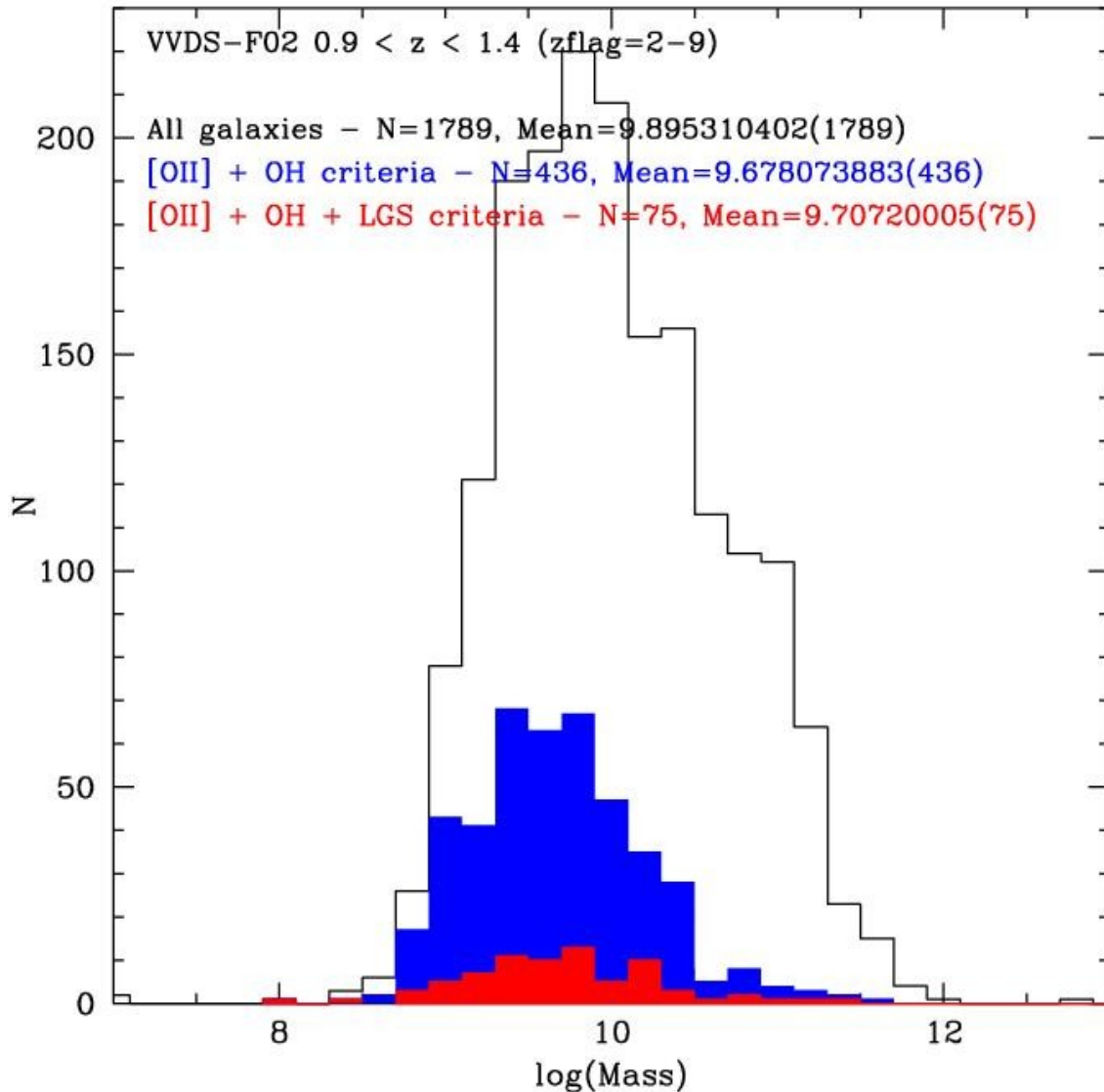


VVDS220071601 - $z=1.3538$



VVDS020148725 - $z=1.3266$

Mass Assembly Survey with SINFONI in VVDS

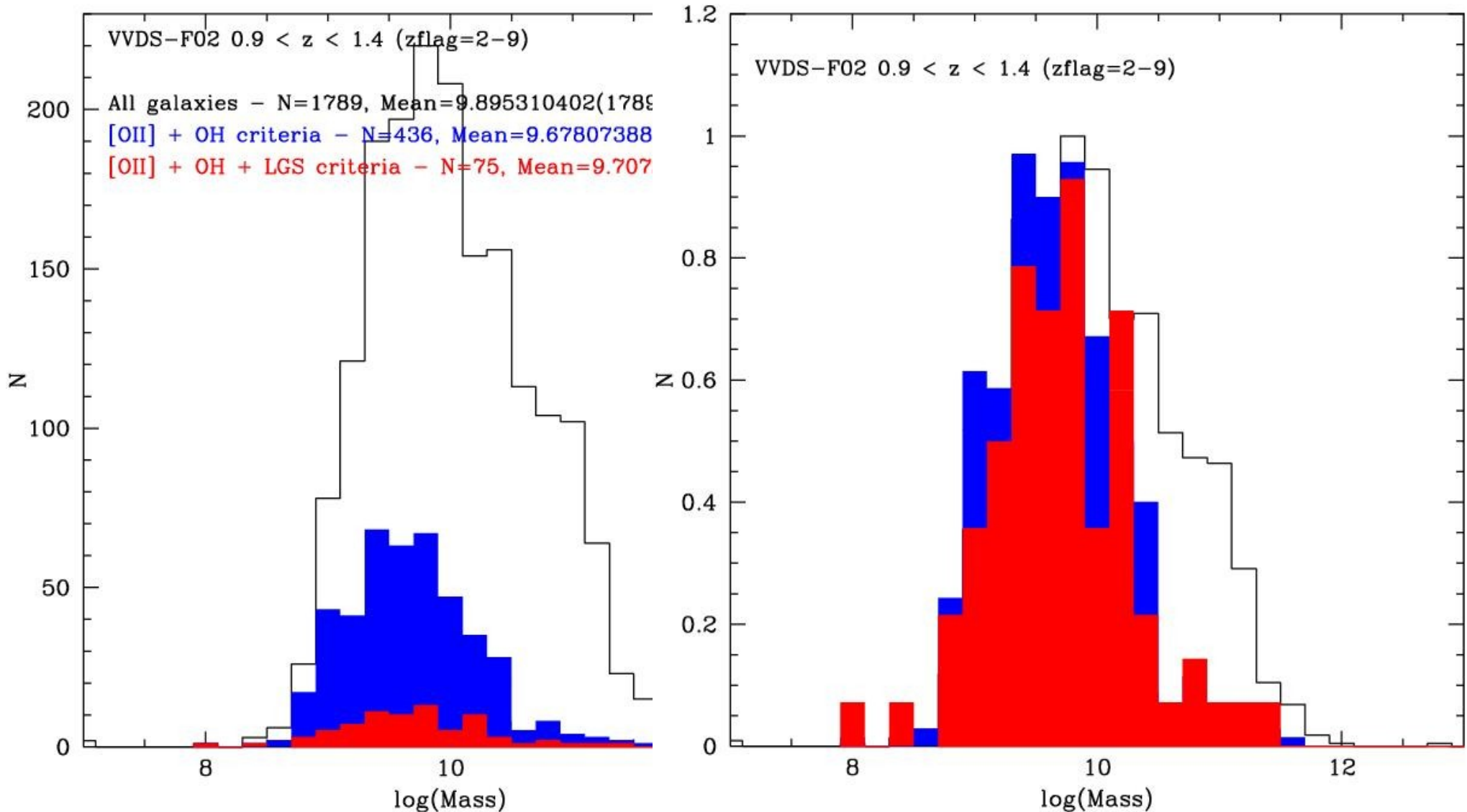


**Full F02 $0.9 < z < 1.4$ sample
 $N=1789$**

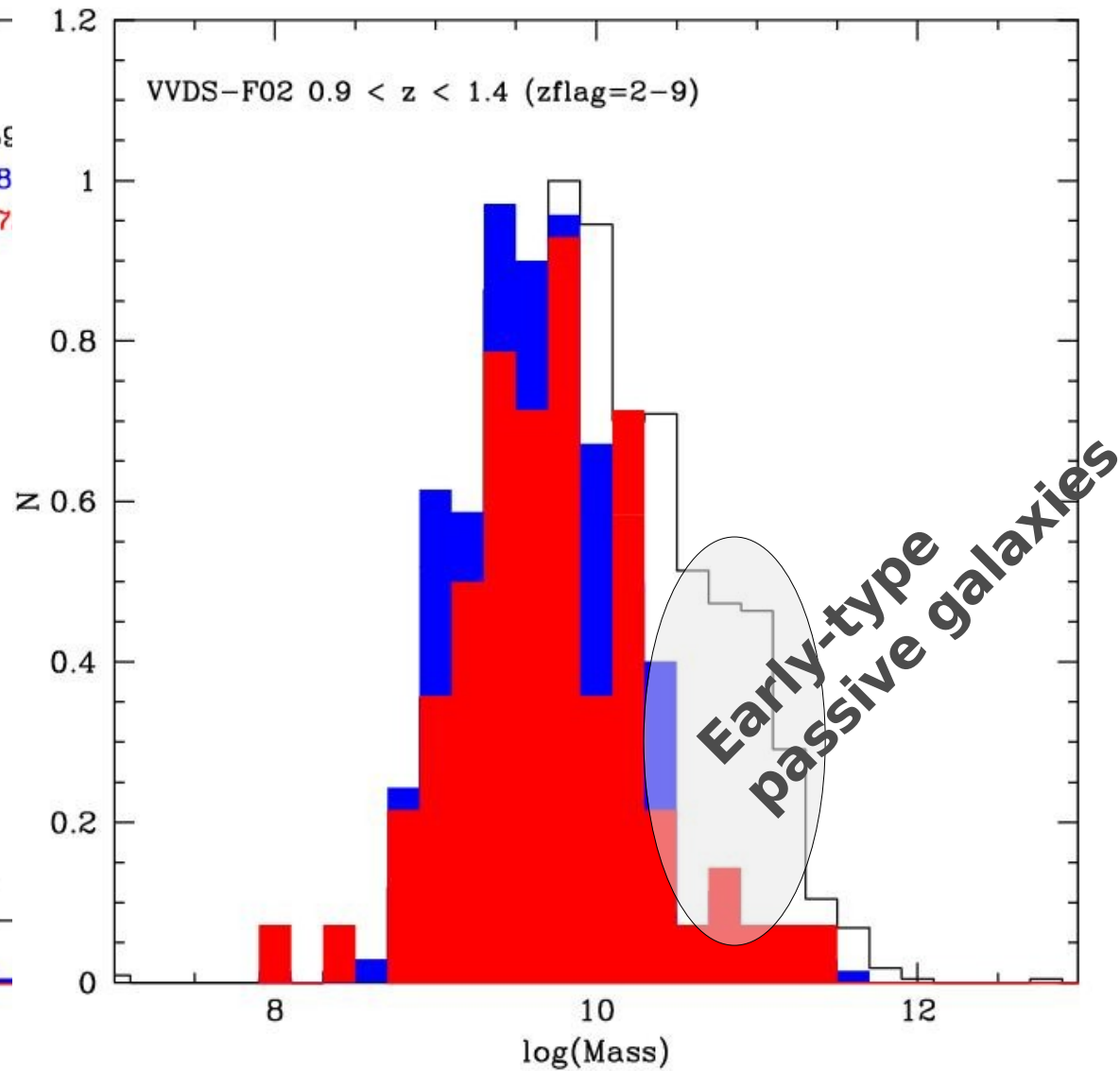
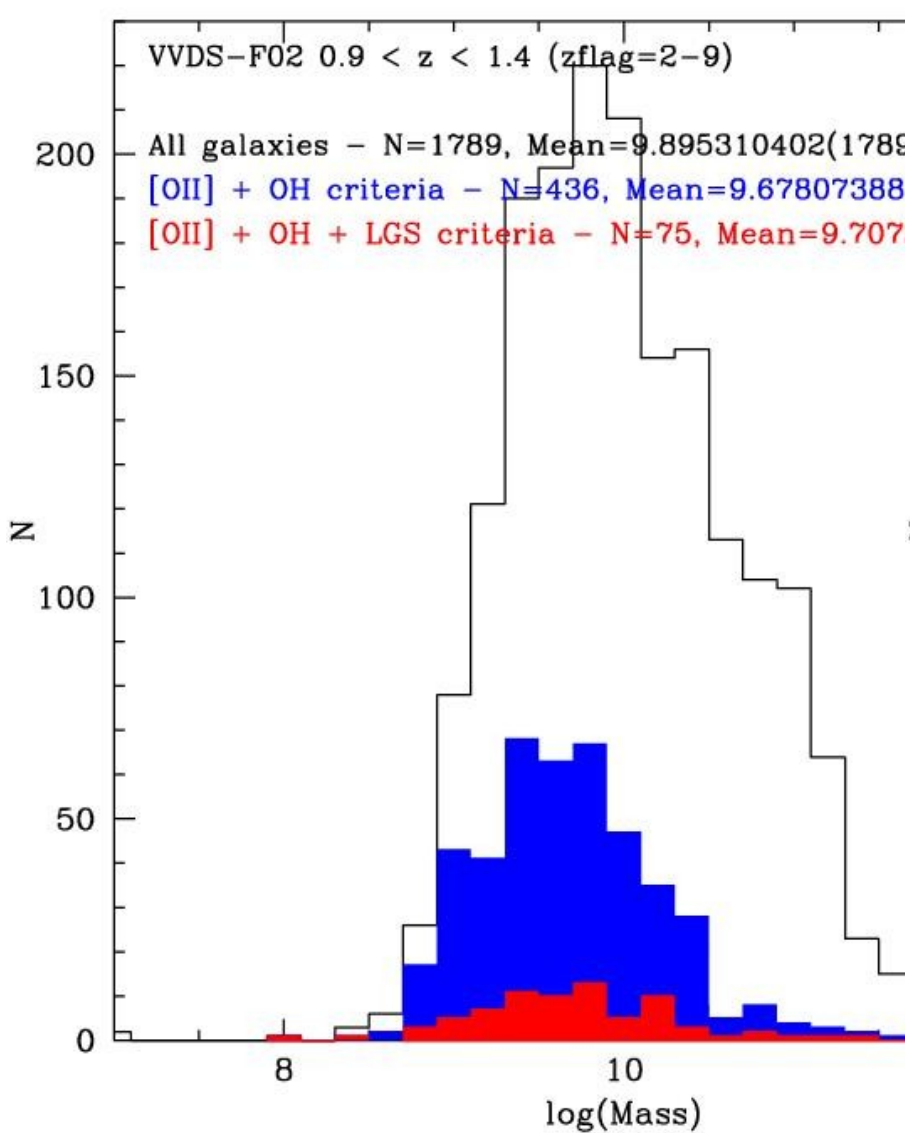
**[OII]+OH criteria
 $N=436$ (24%)**

**[OII]+OH+LGS criteria
 $N=75$ (4%)**

Mass Assembly Survey with SINFONI in VVDS



Mass Assembly Survey with SINFONI in VVDS





Mass Assembly Survey with SINFONI in VVDS

Observations scheduling

2007

P79A spring '07	9.5h F14	$z < 1.4$ no AO	completed
P79B sum/aut '07	61.5h F22/F02	$z < 1.4$ AO-LGS	open(10% comp)
P80C autumn '07	76.0h F02	$z < 1.8$ AO-LGS	open(0%)

2008

P81D spring '08	10.0h F14	$z < 1.4$ no AO	completed
P81E sum/aut '08	64.0h F22,F02	$z < 1.8$ AO-LGS	open
P82F autumn '08	82.0h F02	$z < 1.8$ AO-LGS	



Mass Assembly Survey with SINFONI in VVDS

Observations scheduling

2007

P79A spring '07	9.5h F14	z<1.4 no AO	completed
P79B sum/aut '07	61.5h F22/F02	z<1.4 AO-LGS	open(10% comp)
P80C autumn '07	76.0h F02	z<1.8 AO-LGS	open(0%)

2008

P81D spring '08	10.0h F14	z<1.4 no AO	completed
P81E sum/aut '08	64.0h F22,F02	z<1.8 AO-LGS	open
P82F autumn '08	82.0h F02	z<1.8 AO-LGS	

Important delay (~ 1 year), mainly due to major problems with Laser Guide Star Facility
 --> **Carry-over status for 2007 runs**



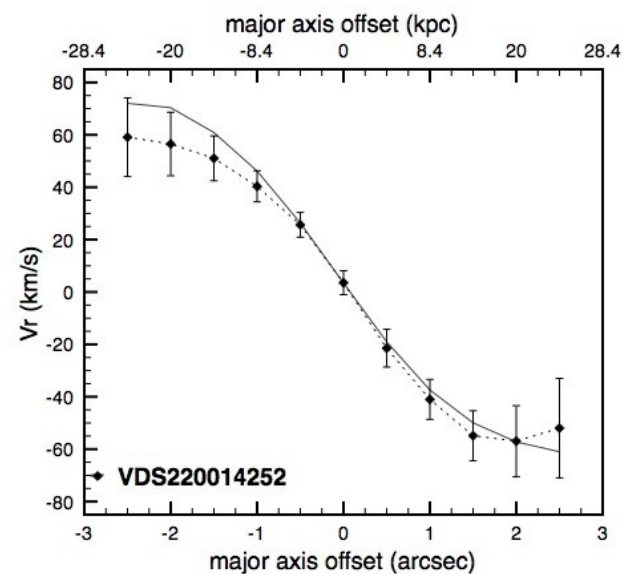
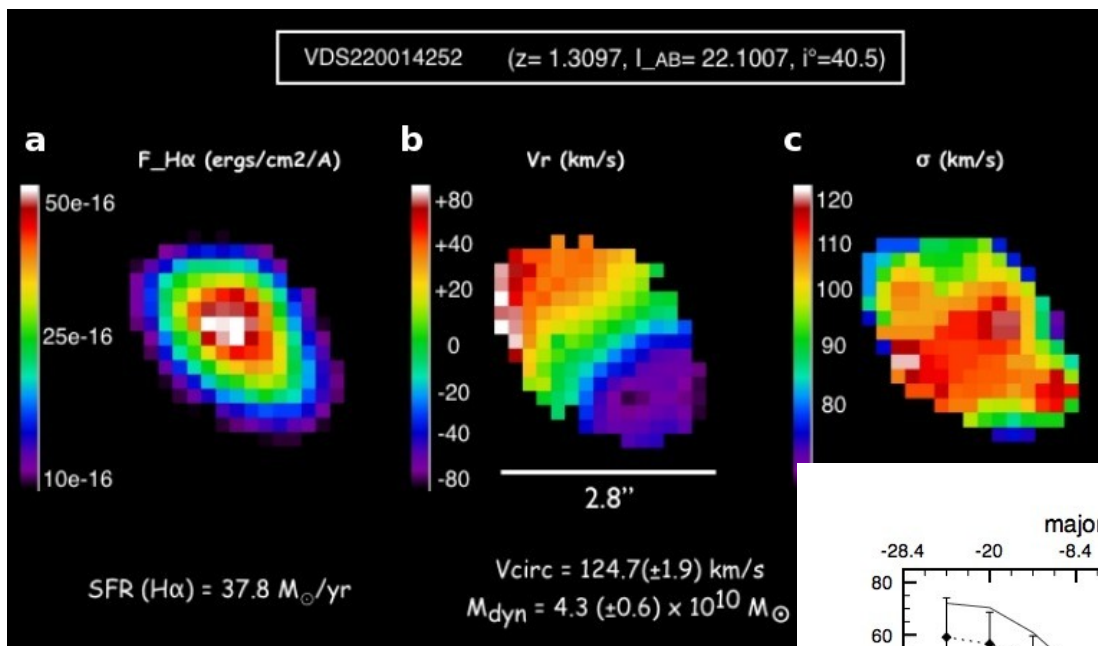
Mass Assembly Survey with SINFONI in VVDS

First results from pilot runs

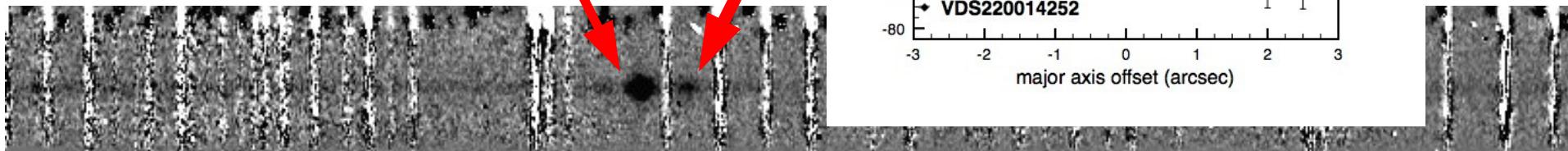
- ✓ **PI: M. Lemoine-Busserolle**
- ✓ **8 nights in Nov 05/06**
- ✓ **13 VVDS galaxies in F02 ($I_{AB} < 24$)
and F22 ($I_{AB} < 22.5$) fields**
 - 4 galaxies @ **$z \sim 3.3-3.7$** (5h exp. time)
 - 9 galaxies @ **$z \sim 1.3-1.5$** (2h exp. time)
- ✓ **no AO – FoV = 8"x8" – 0.25 spaxel**

Mass Assembly Survey with SINFONI in VVDS

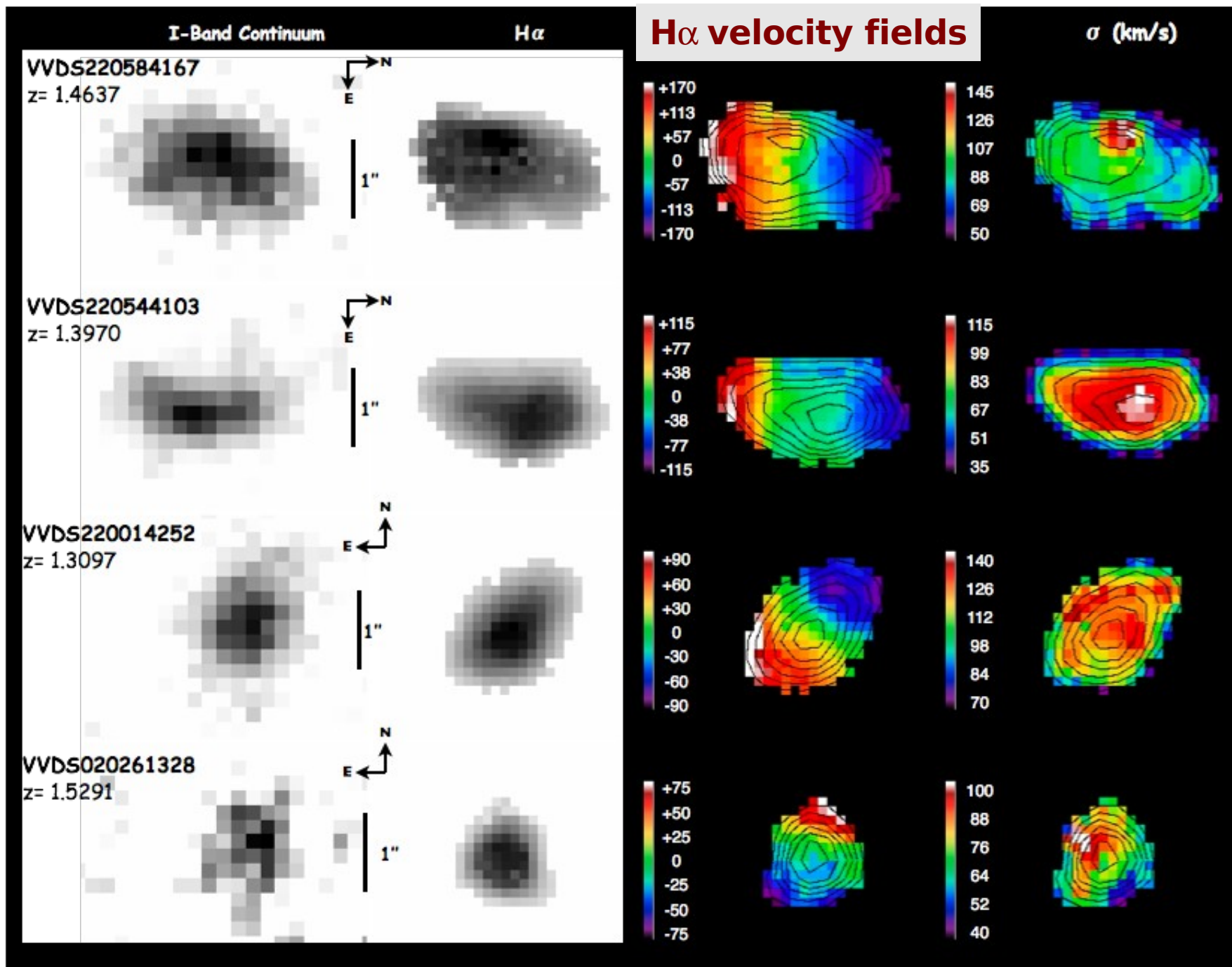
VVDS220014252
z=1.3097



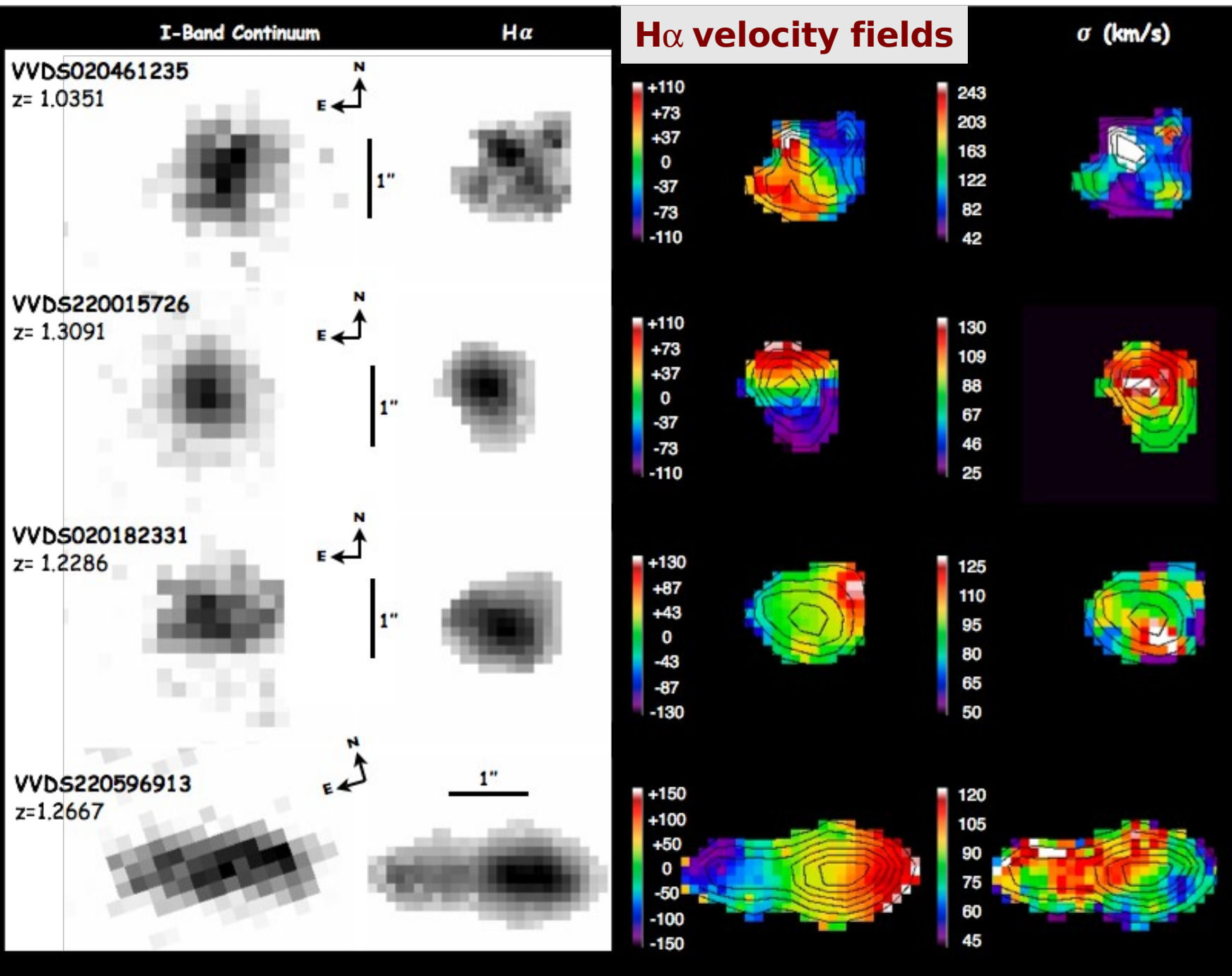
H α **[NIII]**



First results from Pilot runs - Galaxies @ $z \sim 1.2-1.5$

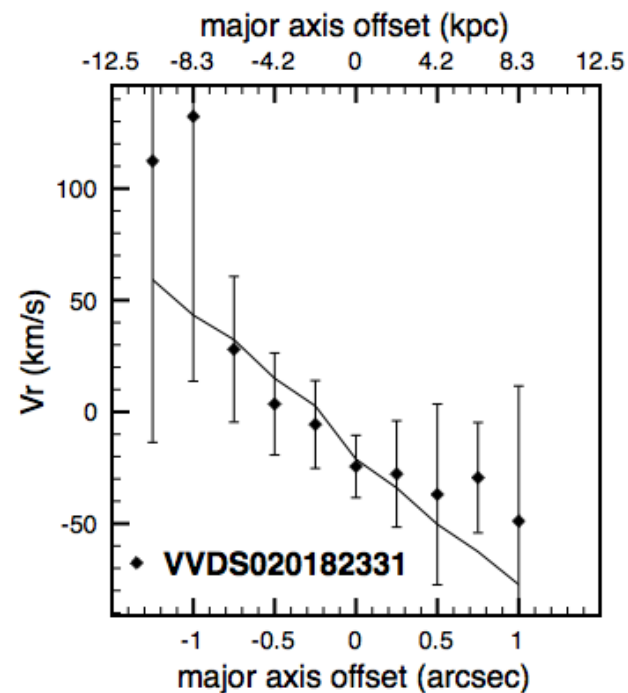
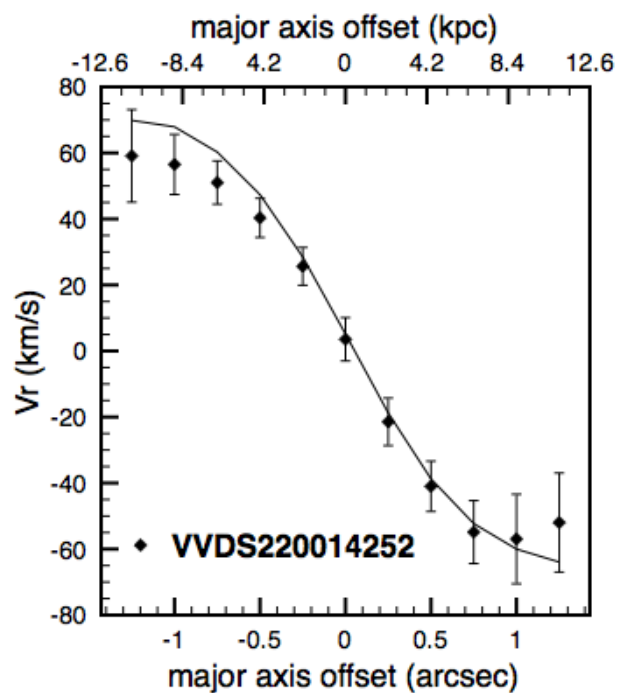
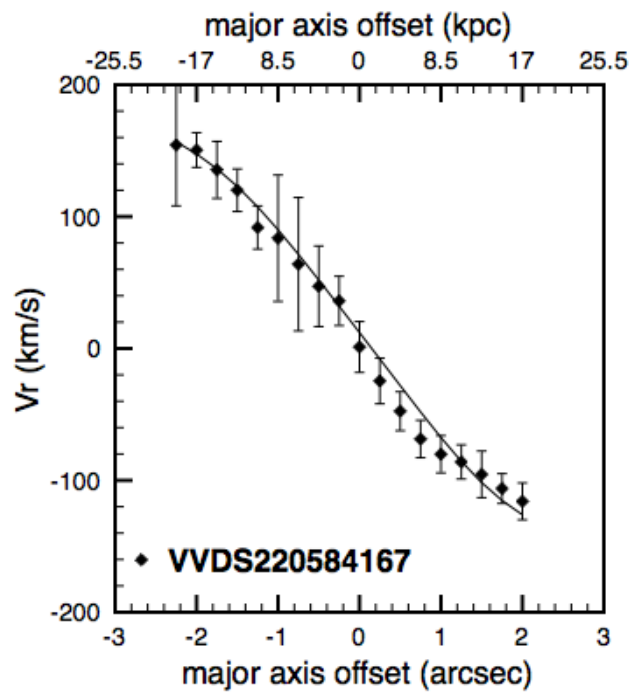
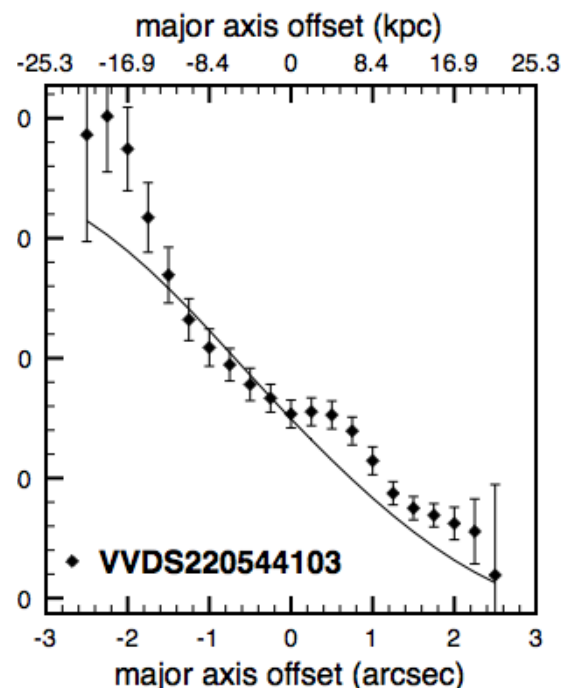
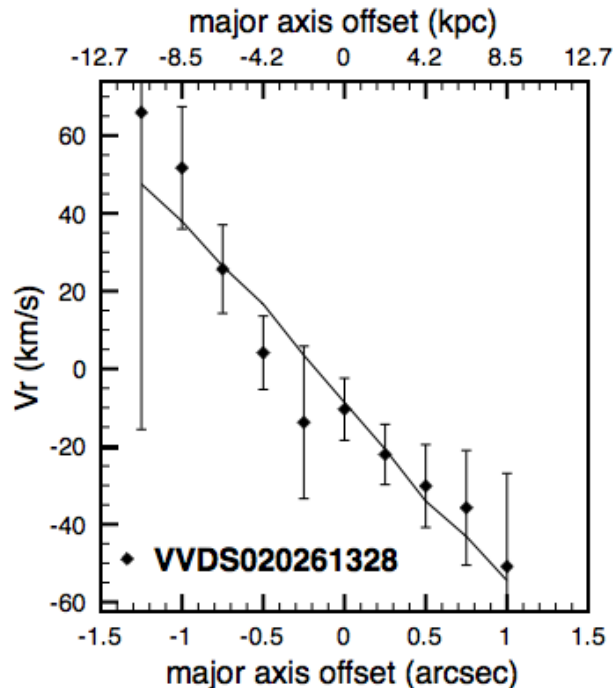
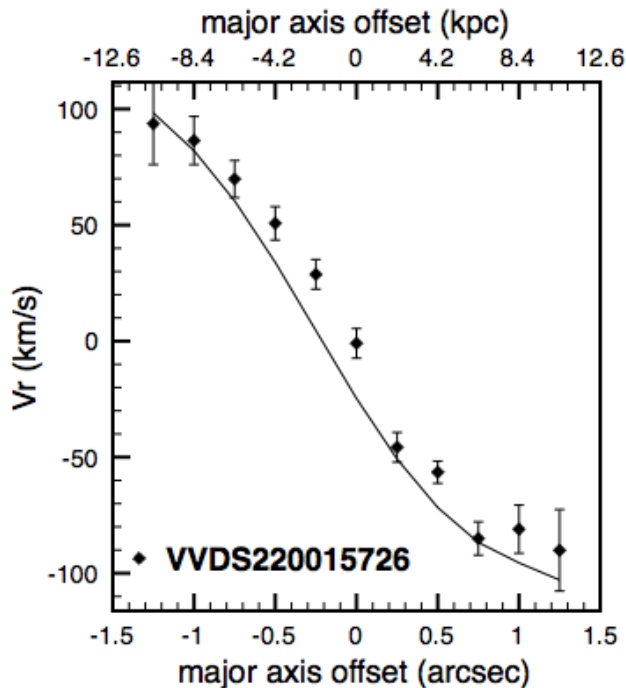


First results from Pilot runs - Galaxies @ $z \sim 1.2-1.5$



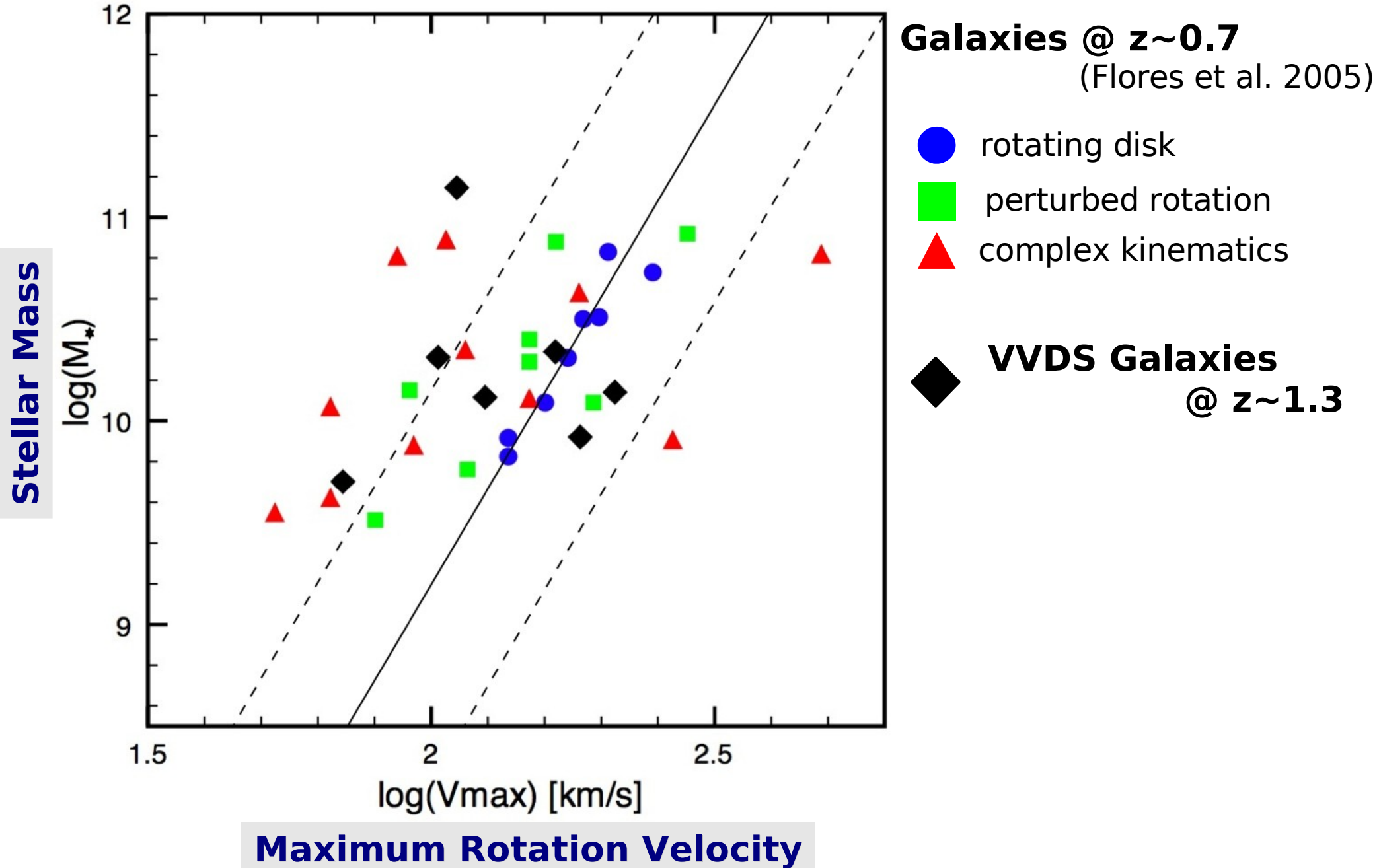
First results from Pilot runs - Rotating disks @ $z \sim 1.2-1.5$

Lemoine-Busserolle et al. (2008a)



First results from Pilot runs - Rotating disks @ $z \sim 1.2-1.5$

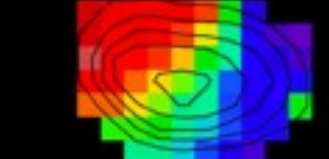
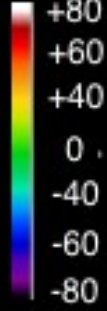
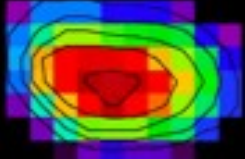
Lemoine-Busserolle et al. (2008a)



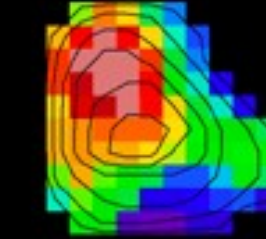
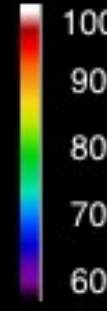
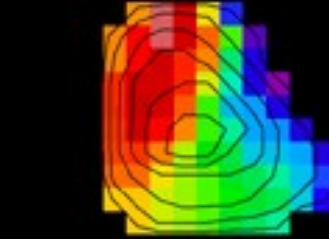
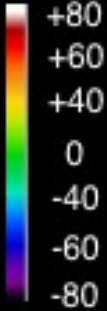
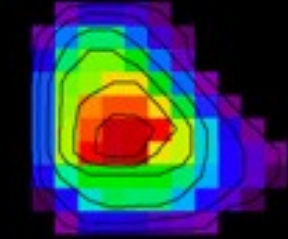
First results from Pilot runs - Galaxies @ $z \sim 3.3-3.7$

**[OIII]5007
velocity fields**

VDS020298666
 $z = 3.2911$,
 $I_{AB} = 23.671$,
 $i = 0^\circ$



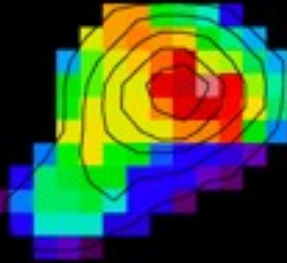
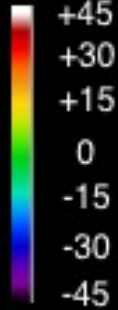
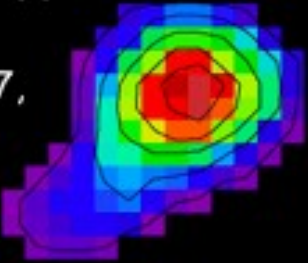
VDS020463884
 $z = 3.2780$,
 $I_{AB} = 23.388$,
 $i = 88.3^\circ$



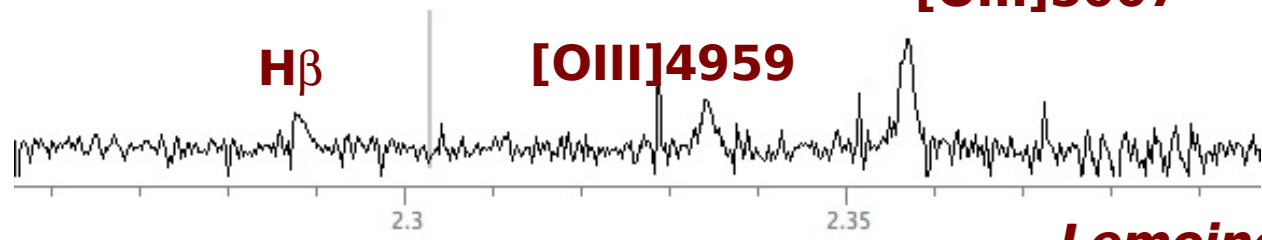
1,6'

with AO/NGS

VDS020335183
 $z = 3.6993$,
 $I_{AB} = 23.687$,
 $i = 83.9^\circ$

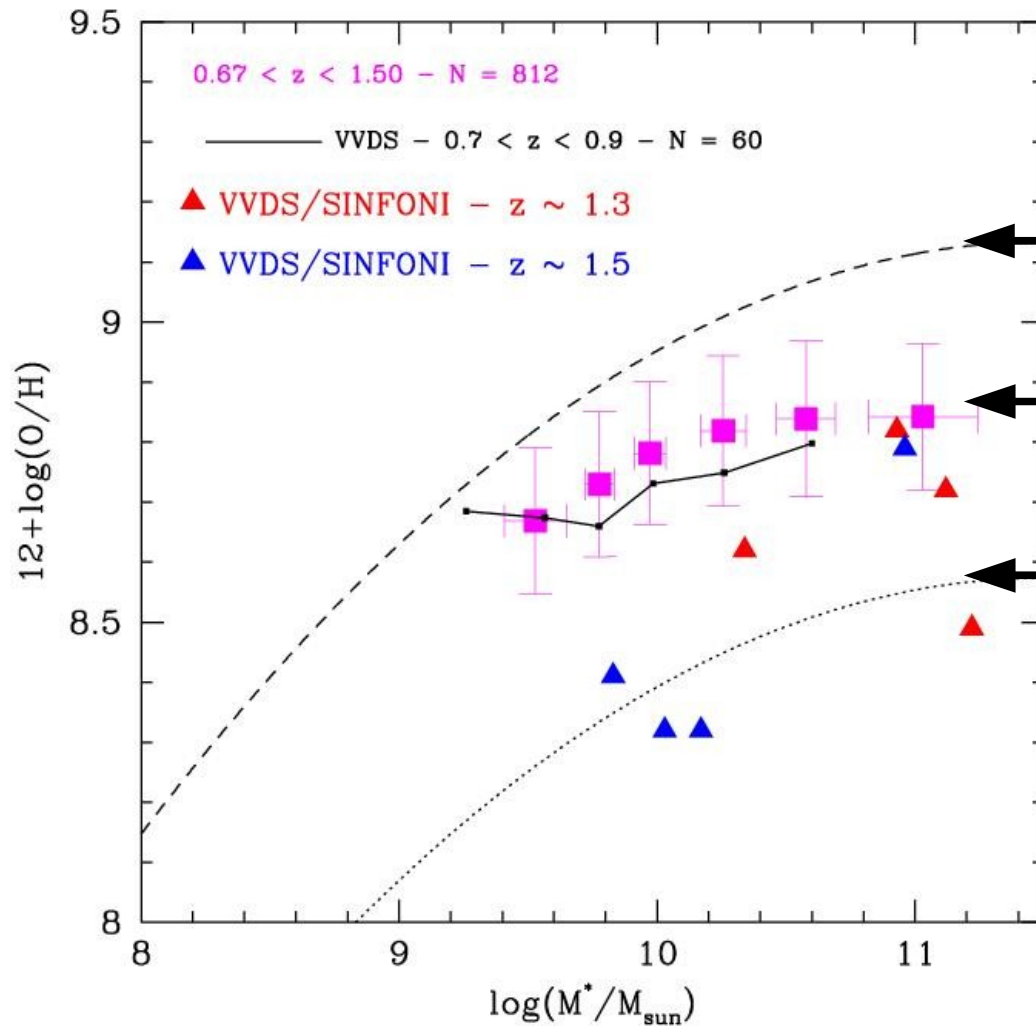


[OIII]5007



Lemoine-Busserolle et al. (2008b)

The MZ relation @ $z \sim 1.4$ with

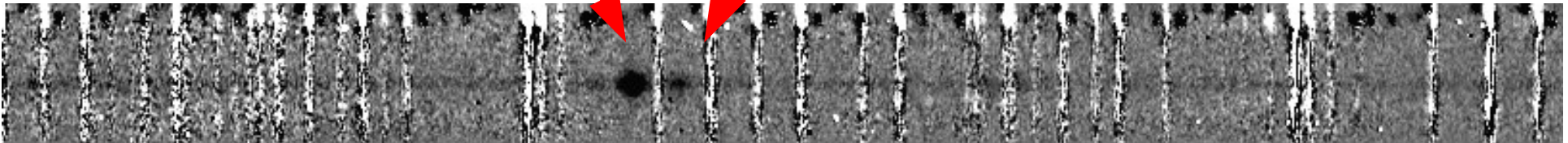


z~0 SDSS
Tremonti et al. (2004)

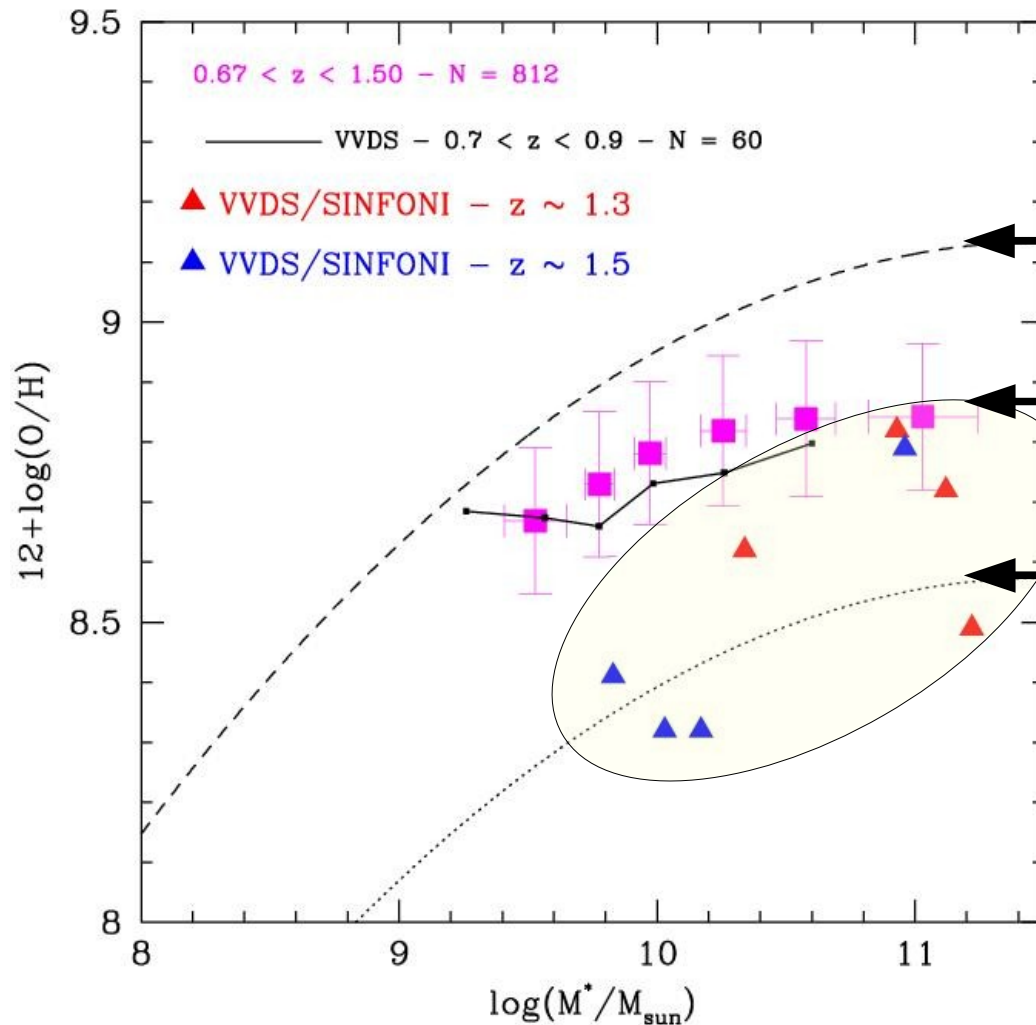
z~0.9
VVDS - Lamareille et al. (2008)
zCOSMOS - Contini et al. (2008)

z~2 Erb et al. (2004)

H α [NII]



The MZ relation @ $z \sim 1.4$ with



$z \sim 0$ SDSS
Tremonti et al. (2004)

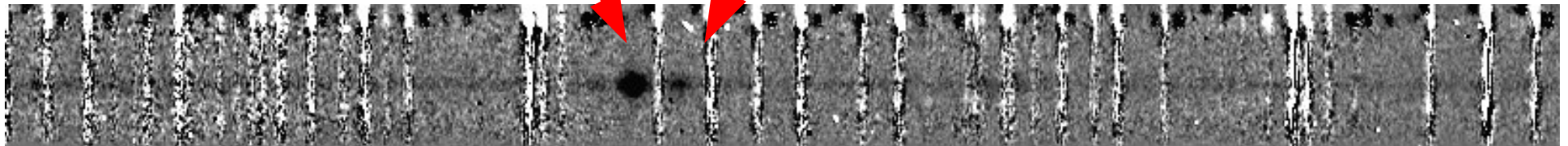
$z \sim 0.9$
VVDS - Lamareille et al. (2008)
zCOSMOS - Contini et al. (2008)

$z \sim 2$ Erb et al. (2004)

$z \sim 1.3-1.5$
MASSIV survey
Queyrel et al. (2008)

O/H from [NII]/H α ratio

H α [NII]



See **Benoit Epinat** poster + talk
for preliminary  results

