PWV with VISIR

Outline

Measuring the amount of precipitable water vapour with VISIR

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Water

PWV with VISIR

Water vapour

Water

Measuring wate vapour

VISIR

Method

Results

Comparison with photometric data

Conclusions

Water:

- one of the main constituents of the atmosphere:
- mainly found in the lower layers
- can be found
 - in condensed form (important for optical and NIR):

- liquid, as droplets in most clouds
- solid, as small ice crystals in cirrus clouds
- as water vapour (important for NIR and MIR).



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Figure: Atmospheric opacity

http://www-atm.physics.ox.ac.uk/group/mipas/ac

What?

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Usual quantity is the Precipitable Water Vapour (PWV):

- amount of water vapour in the atmospheric column above the observatory
- equivalent to the amount of liquid precipitation that would result if all the water vapour in the column is condensed (in mm)

Why?

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- in the NIR (L, M, also in K!): zero-point
- in the MIR:
 - sensitivity, conversion factor
 - passband dependent
 - observations of standard stars is time expensive
- in the sub-mm: *crucial* to correct variations of the path length in the atmosphere

How and when? 1/3: from above

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• at ESO (Marc Sarazin, Andre Erasmus): http://www.eso.org/gen-fac/pubs/astclim/ forecast/meteo/ERASMUS/l_p_f0.html

based on

- the Upper Tropospheric Humidity for the middle and upper troposphere (layer between 700 mb and 300 mb (approx. 3000 to 9500 m), determined from 6.7μm images
- surface relative humidity value derived from the observations of surface relative humidity at the observatory in the 24 hours preceding the forecast
- satellite images every 3 h + European Centre for Medium-Range Weather Forecasting model

How and when? 2/3: from below

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• radiometers (APEX, ALMA):

• measurement of a water emission line (e.g.: 183 GHz)

- requires accurate atmospheric model (pressure, temperature)
- high-accuracy ($\approx 10^{-6}$ mm)
- real-time
- expensive

sky dips (e.g.: SCUBA)

- simple
- expensive in execution time
- absorption lines in stellar spectra (CRIRES SV, P24)

- some lines are T-insensitive
- expensive in execution time

How and when? 3/3: from below

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• VISIR medium-resolution spectroscopy

- 10 s exposure, 1 min execution
- near real time
- sufficient precision
- requires decent atmospheric model (pressure, temperature)

PWV determination from VISIR spectra

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- VISIR sky spectra (MR, 19.5 μ m (Q3), 0.4" slit) ;
- 10 s exposure time; \approx 1 min execution time
- χ^2 fit using RFM, modifying the H2O concentration by a constant factor in each layer

• IDL interface (H2Ocalc) to RFM

RFM: Radiative Forward Model

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 RFM: Line-by-line modeling Fortran code, developped by Anu Dudhia (Oxford) to analyse data from MIPAS on-board ENVISAT

HITRAN database:

- HITRAN'2004 Database (Version 12.0)
- 1,734,469 spectral lines for 37 different molecules, including the atom O (singlet) and the ion NO+.
- Atmospheric profile:
 - Lower and upper altitudes of (possibly arbitrarily) layers (typically 50)

- Mean pressure in each layer
- Mean temperature in each layer
- Concentration of H20, CO2, O3, etc... in each layer

Predicted spectra



Figure: Predicted MR spectrum

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Observed spectra



Figure: Examples of measured and fitted spectra

Observed spectra



Figure: Examples of measured and fitted spectra

Observed spectra



Figure: Examples of measured and fitted spectra

VISIR vs satellite



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VISIR vs satellite



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VISIR vs satellite





Figure: PAH1



Figure: PAH1

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Figure: PAH2



Figure: Q2

Conclusions

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- Measuring the amount of PWV with VISIR is now easy
- Preferable than to rely on satellite data
- Allows to easily estimate its effect on the conversion factor and sensitivity with a reasonable accuracy
- Better accuracy could be achieved by using a more appropriate atmospheric profile
- User constraint? (spectroscopy?)
- Values should be available on the web
- Method to be tested with ISAAC, SINFONI, CRIRES spectra

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