

# Interferometric observations of the nuclear dust in active galactic nuclei

Obscured AGN Across Cosmic Time

**Konrad R. W. Tristram**

Collaborators:

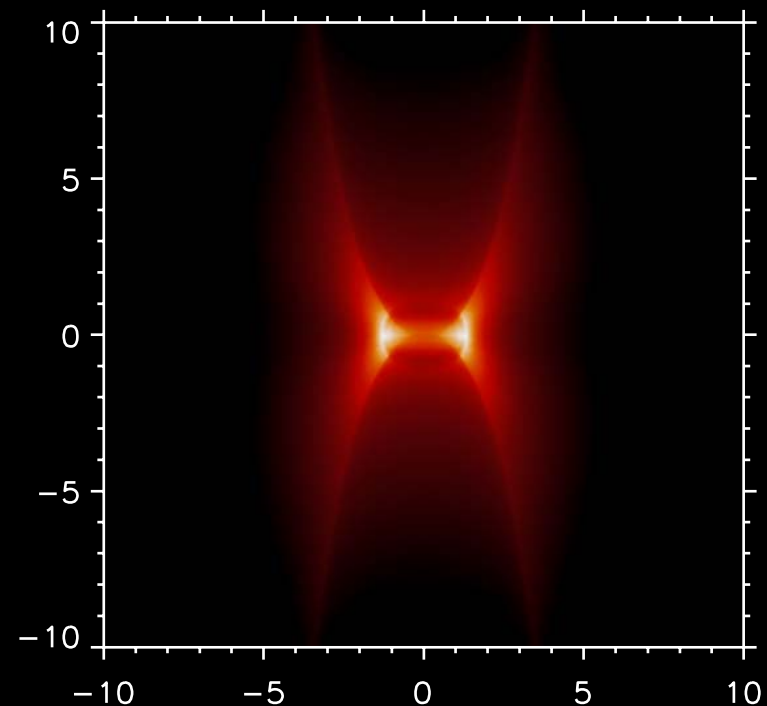
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W. Jaffe, D. Raban (Sterrewacht Leiden), Bill Cotton (NRAO)

# Introduction: Which instrument?



- Torus size:  $\sim 1$  pc
  - ↳ Angular sizes for the nearest AGN:  $< 20$  mas

continuous torus model  
 $\lambda = 10 \mu\text{m}$ ,  $i = 60^\circ$   
Schartmann (2005)



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- Need MIR instruments with highest resolution!
  - ↳ Find this only at the VLTI: MIDI



VLT with the Residencia  
February 2005

# Introduction: Which instrument?

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- MIDI
  - Dispersed fringes in the N-band (8 to 13  $\mu\text{m}$ )
  - NaCl prism ( $R \sim 30$ )
  - Grism ( $R \sim 240$ )

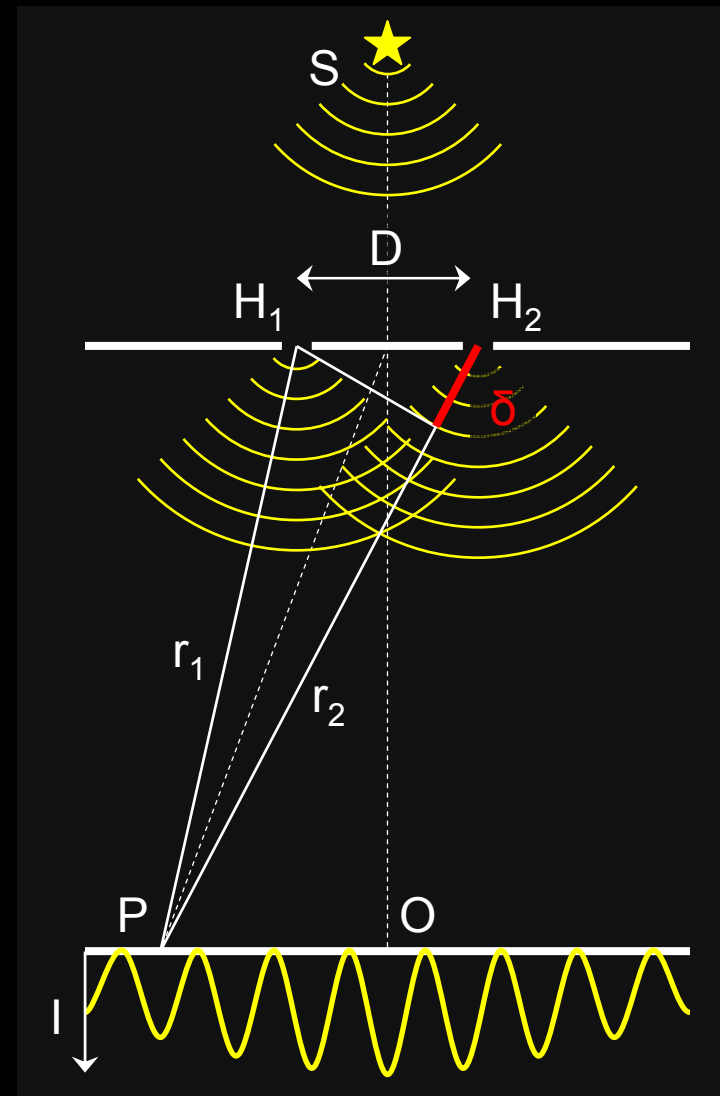


MIDI in the VLT lab

# Introduction: MIR interferometry



- Essentially Young's experiment

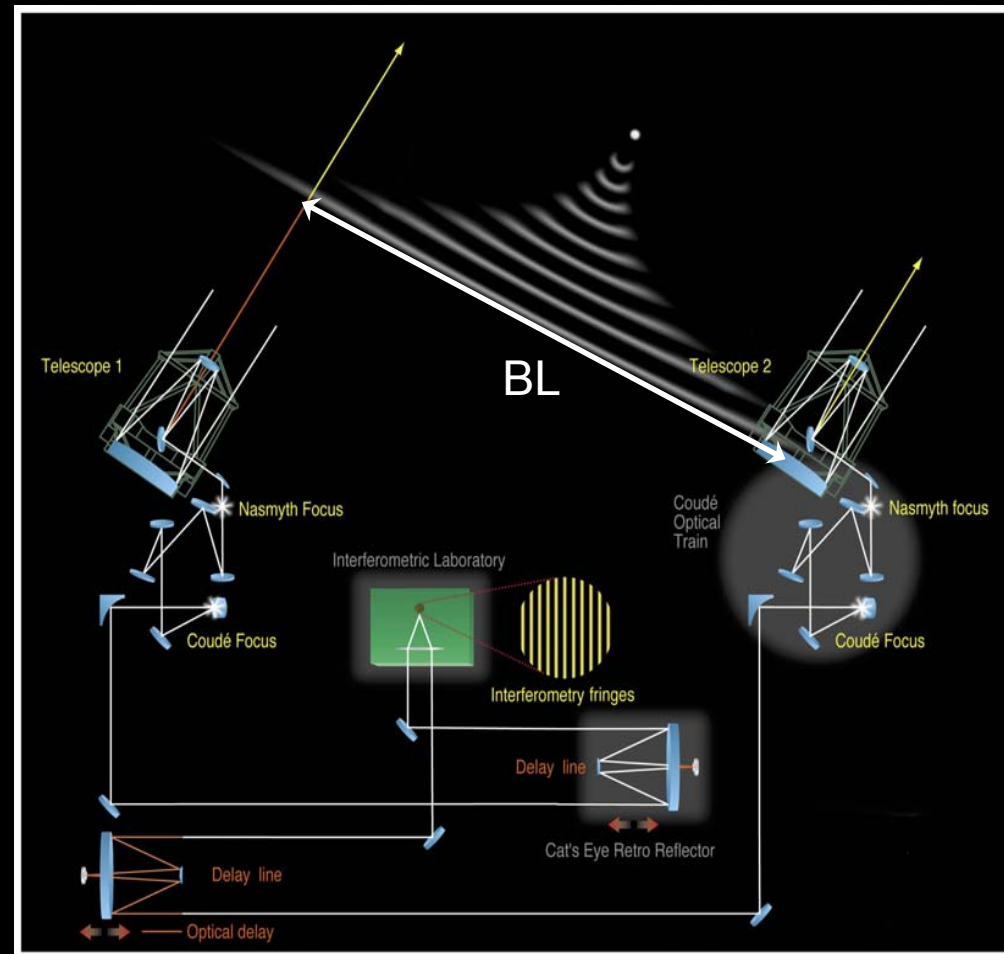


# Introduction: MIR interferometry



- Essentially Young's experiment
- Correlated flux  $F_{\text{cor}}$ : flux in fringes
- Resolution:

$$\varphi = \frac{\lambda}{2 \cdot BL}$$



The VLTI interferometer

# Introduction: Target list



List of all AGN observed to date:

<u>Name</u>	<u>Flux (10 <math>\mu\text{m}</math>)</u>	<u>Time</u>
Circinus	5.0 Jy	GTO
NGC 1068 (M77)	13.0 Jy	SDT / OT
NGC 5128 (Centaurus A)	0.6 Jy	GTO
NGC 3783	0.5 Jy	OT
Mrk 1239	0.4 Jy	GTO
MCG -05-23-016	0.3 Jy	GTO
NGC 7469	0.4 Jy	GTO
NGC 1365	0.6 Jy	GTO
IC 4329A	0.6 Jy	GTO
3C 273	0.3 Jy	GTO

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# Circinus: General properties



- Spiral galaxy SA(s)b,  $i = 65^\circ$
- Seyfert type 2
- $4 \times 10^6 M_\odot$  nucleus
- Distance  $\sim 4$  Mpc
  - ↳  $50 \text{ mas} \sim 1 \text{ pc}$
- Circumnuclear starburst

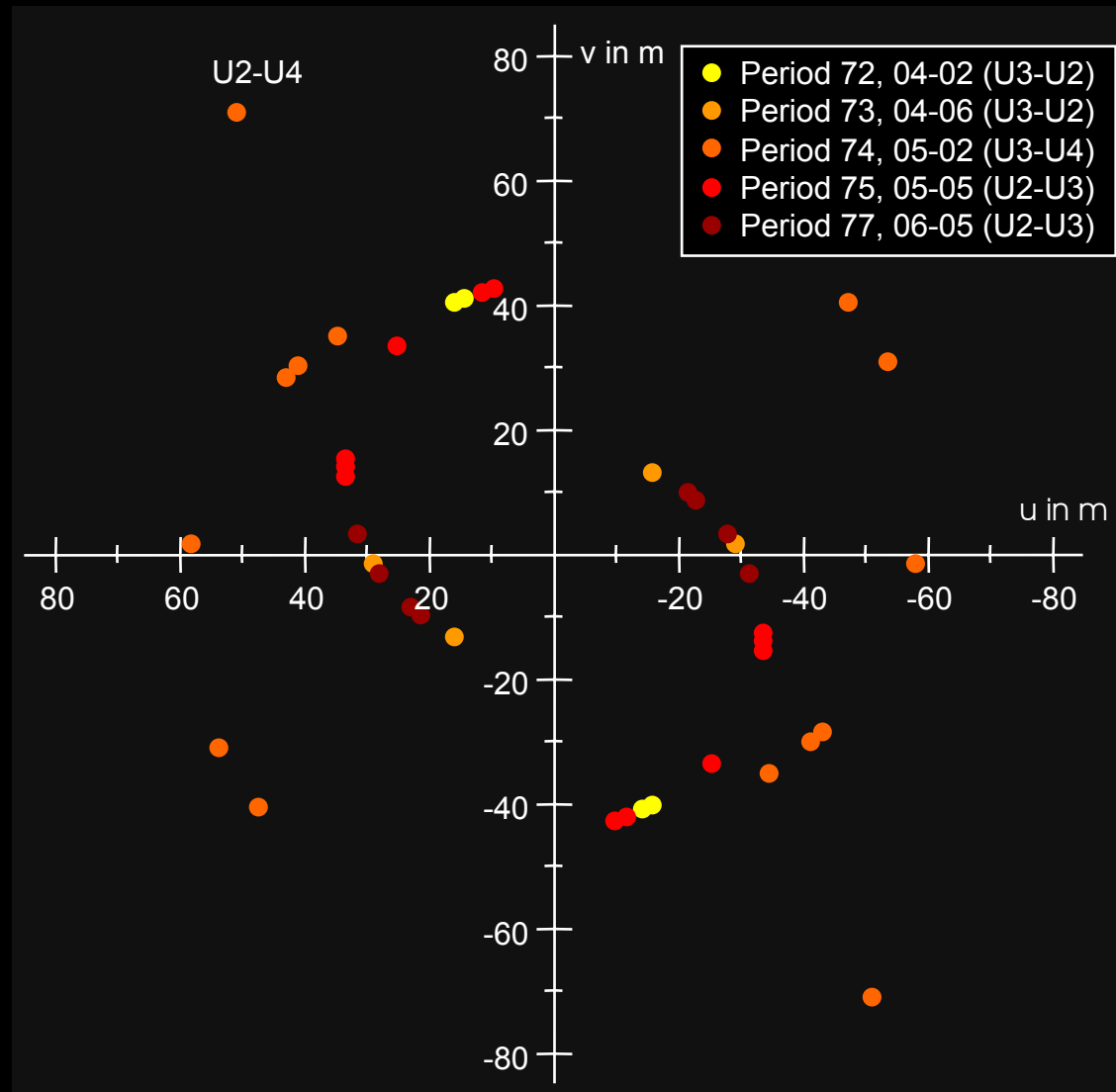
2MASS J, H,  $K_s$  colour mosaic

# Circinus: UV coverage

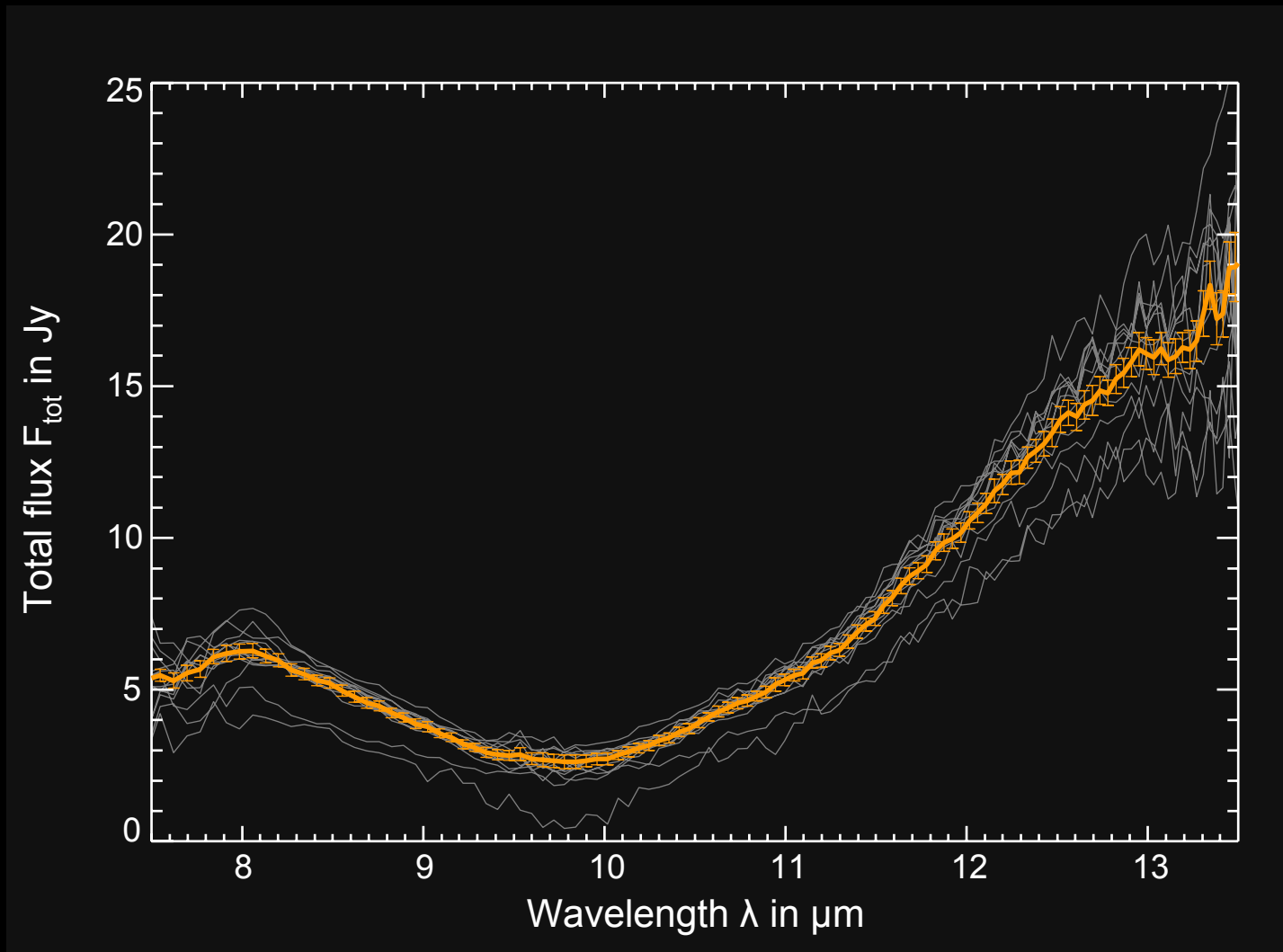


GTO observations  
from P72 to P77:

- 21 fringe tracks
- 16 photometries



# Circinus: Total flux

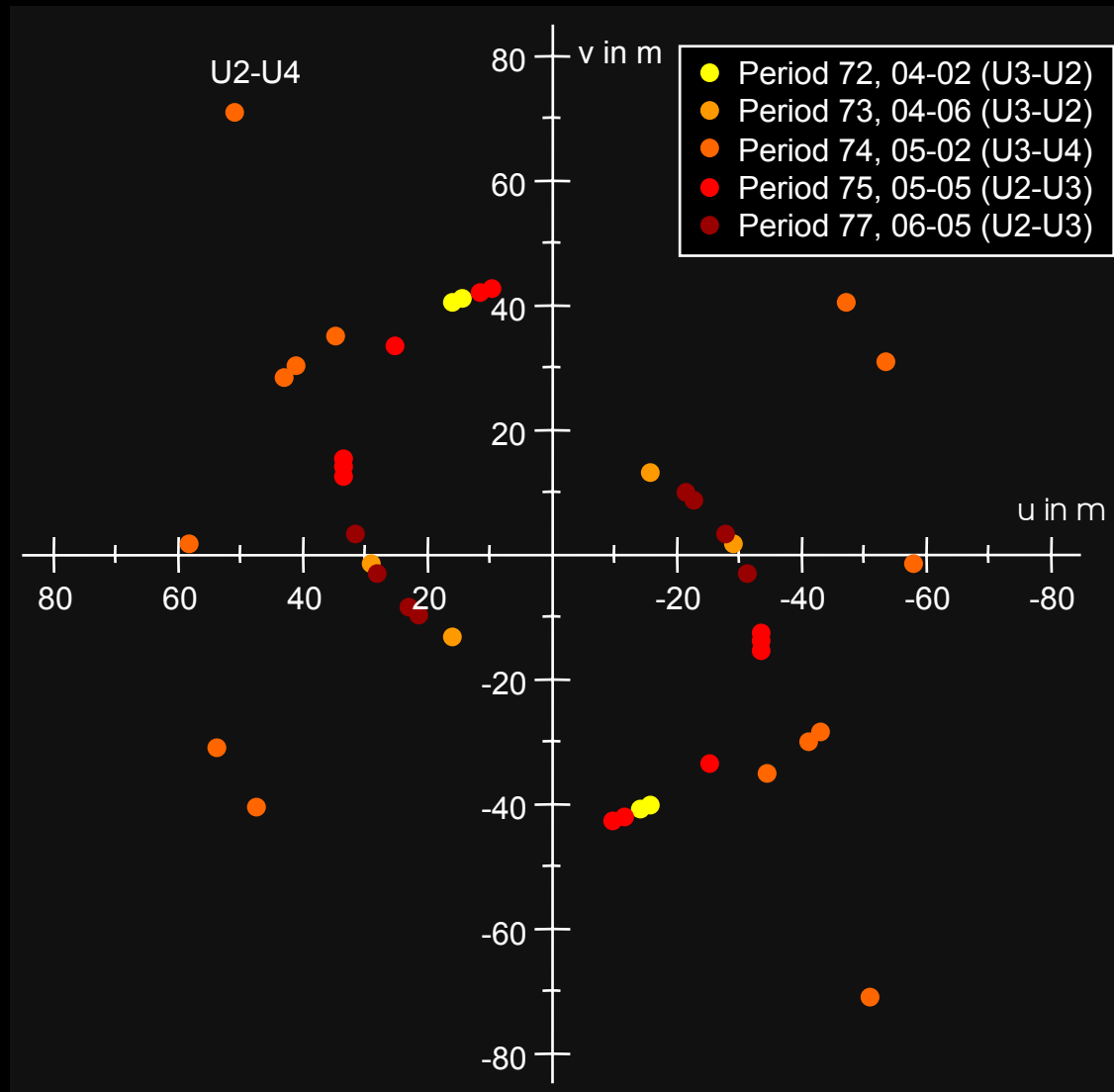


# Circinus: UV coverage

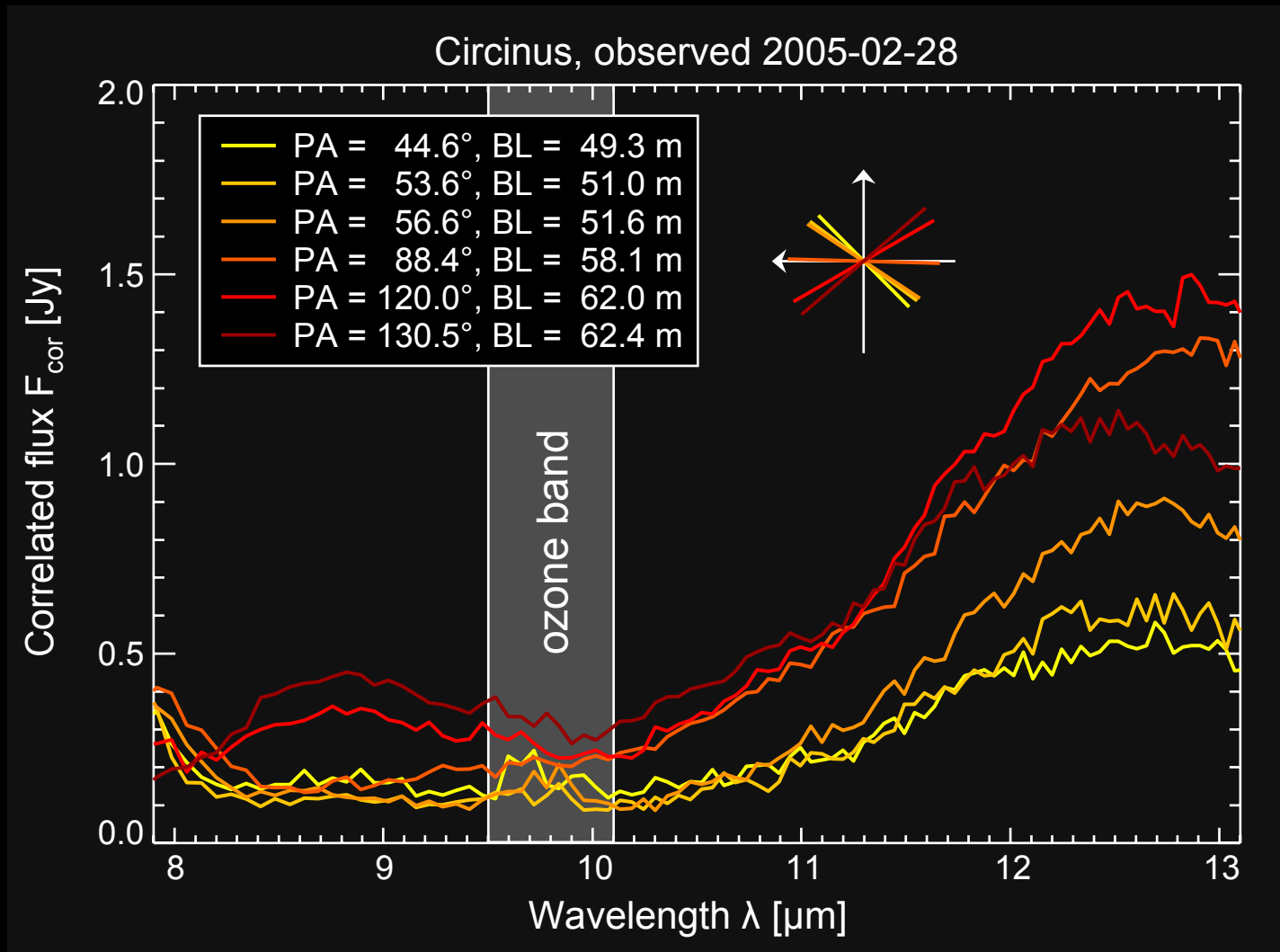


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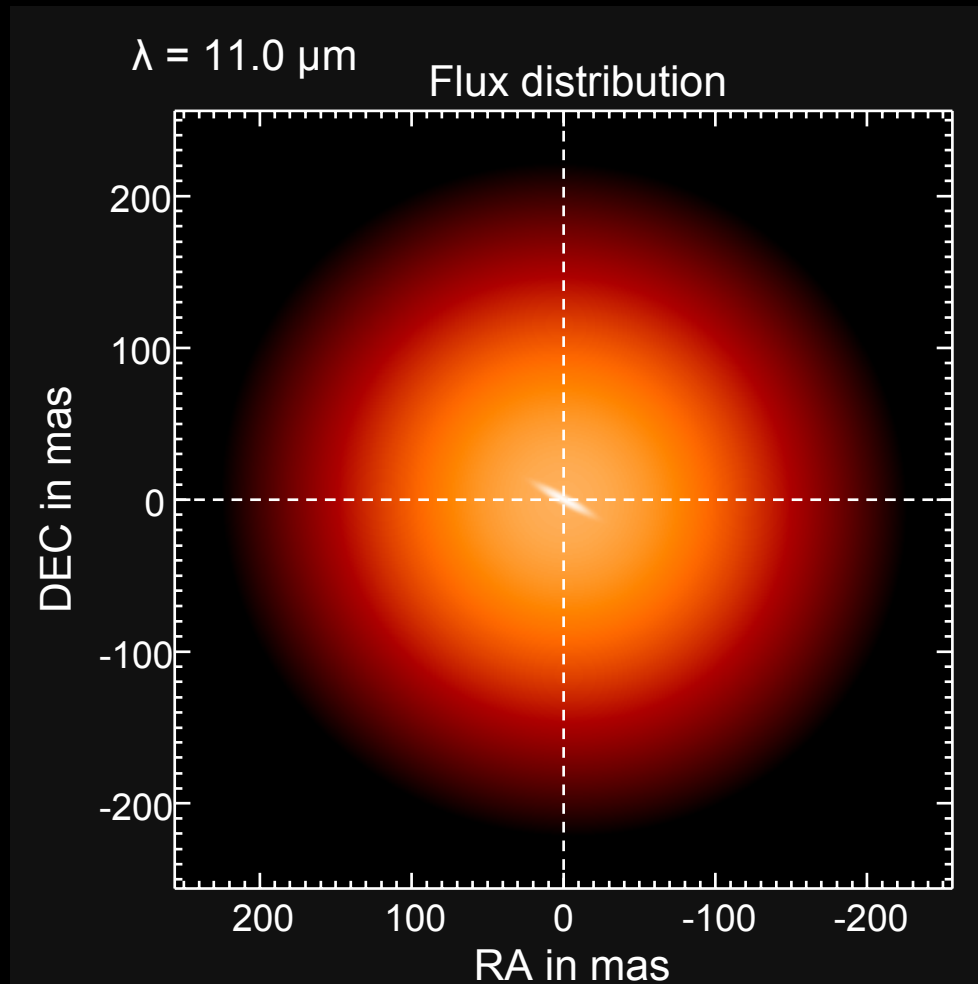
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# Circinus: Correlated flux



# Circinus: Gaussian fit



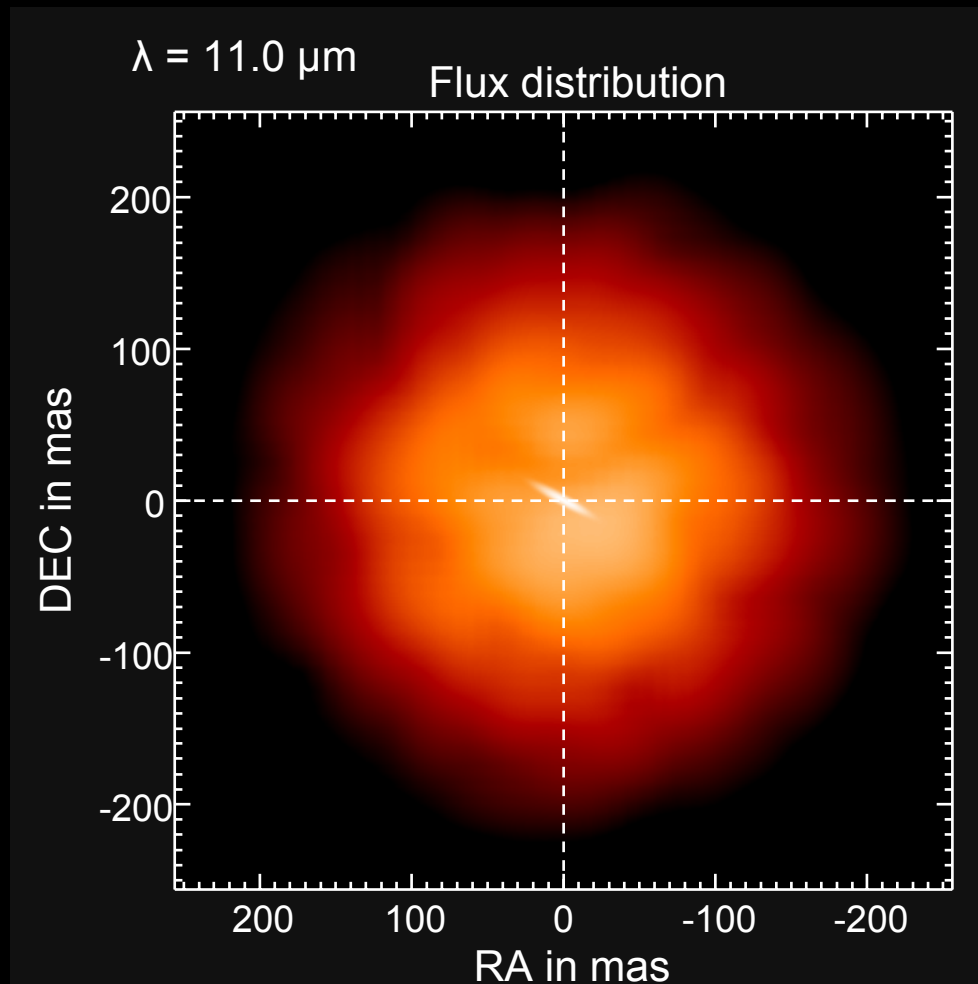
Log of model flux distribution at  $11\mu\text{m}$

## Two blackbody Gaussians:

Size:	$\Delta_1 = 21 \text{ mas}$
Axis ratio:	$r_1 = 0.21$
Silicate depth:	$\tau_1 = 1.18$
Temperature:	$T_1 = 334 \text{ K}$
Covering factor:	$f_1 = 1.0$
Size:	$\Delta_2 = 97 \text{ mas}$
Axis ratio:	$r_2 = 0.97$
Silicate depth:	$\tau_2 = 2.21$
Temperature:	$T_2 = 298 \text{ K}$
Covering factor:	$f_2 = 0.20$
Position angle:	$\alpha = 61^\circ$

Tristram et al.

# Circinus: Clumpiness



Log of model flux distribution at  $11\mu\text{m}$

## Two blackbody Gaussians:

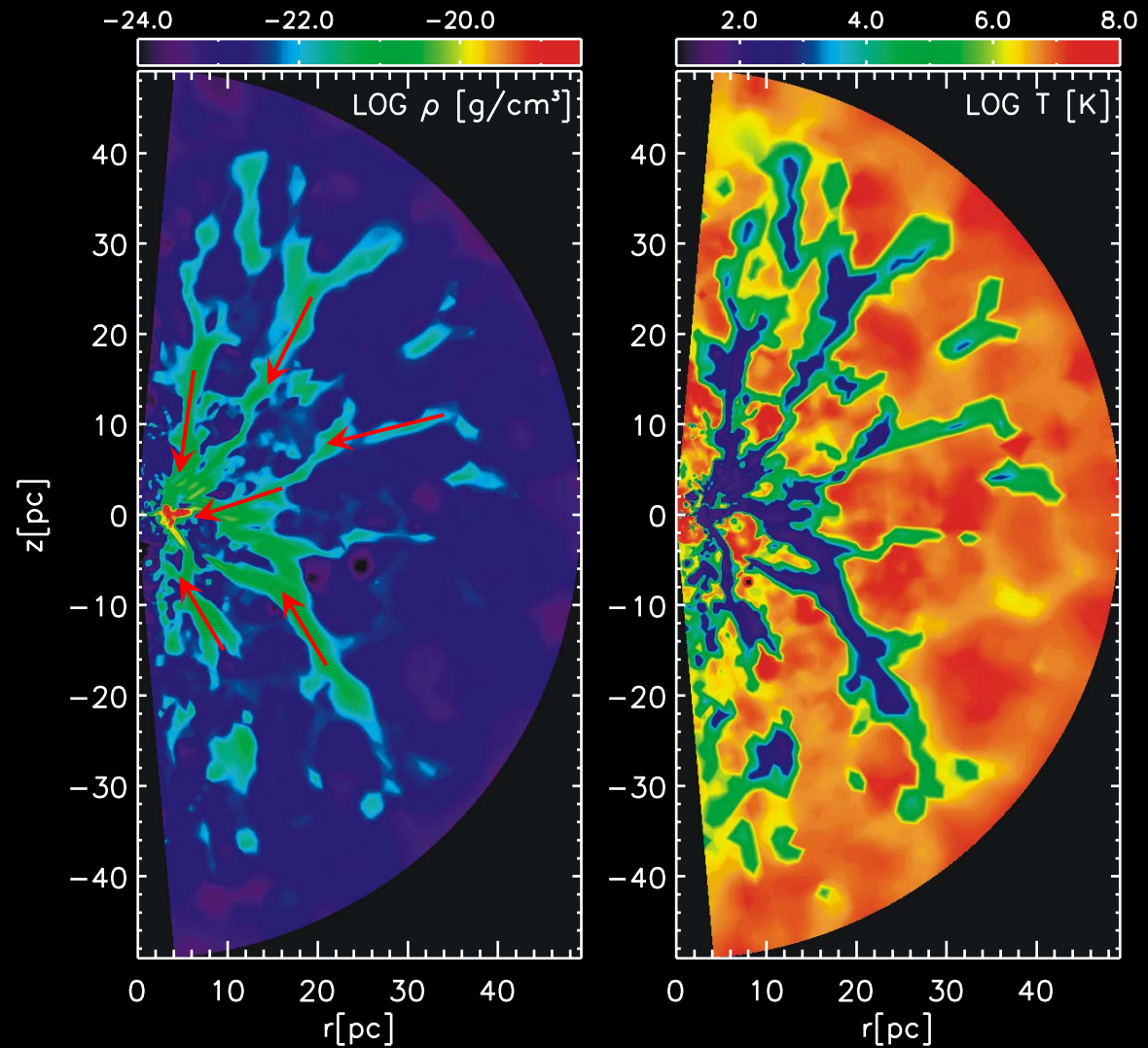
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Tristram et al.

# Simulations of AGN tori



## Hydrodynamical torus model



Schartmann et al.

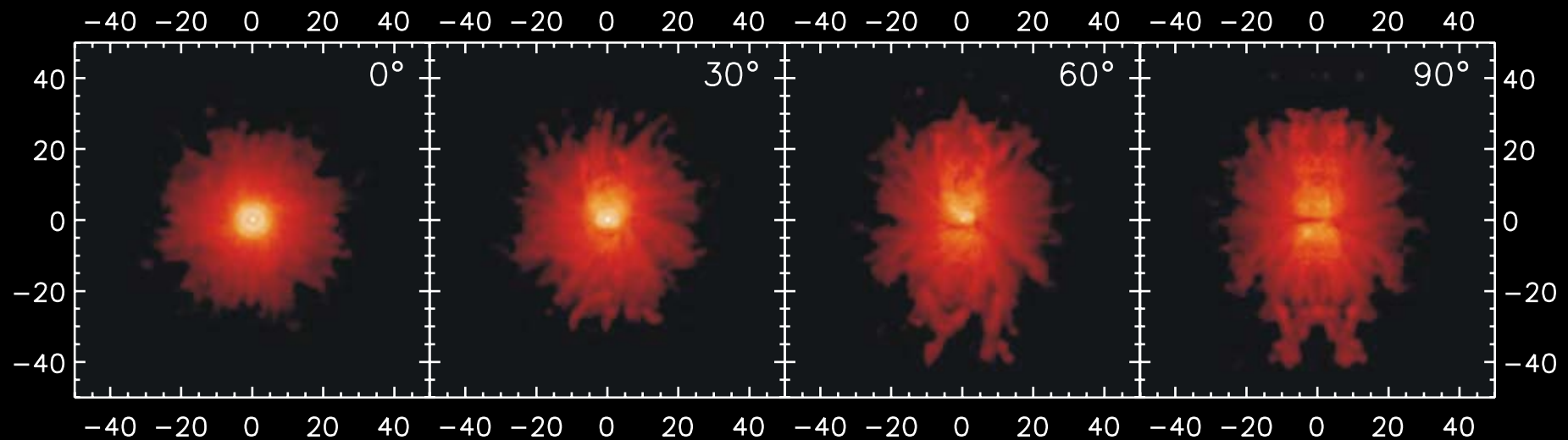


# Simulations of AGN tori



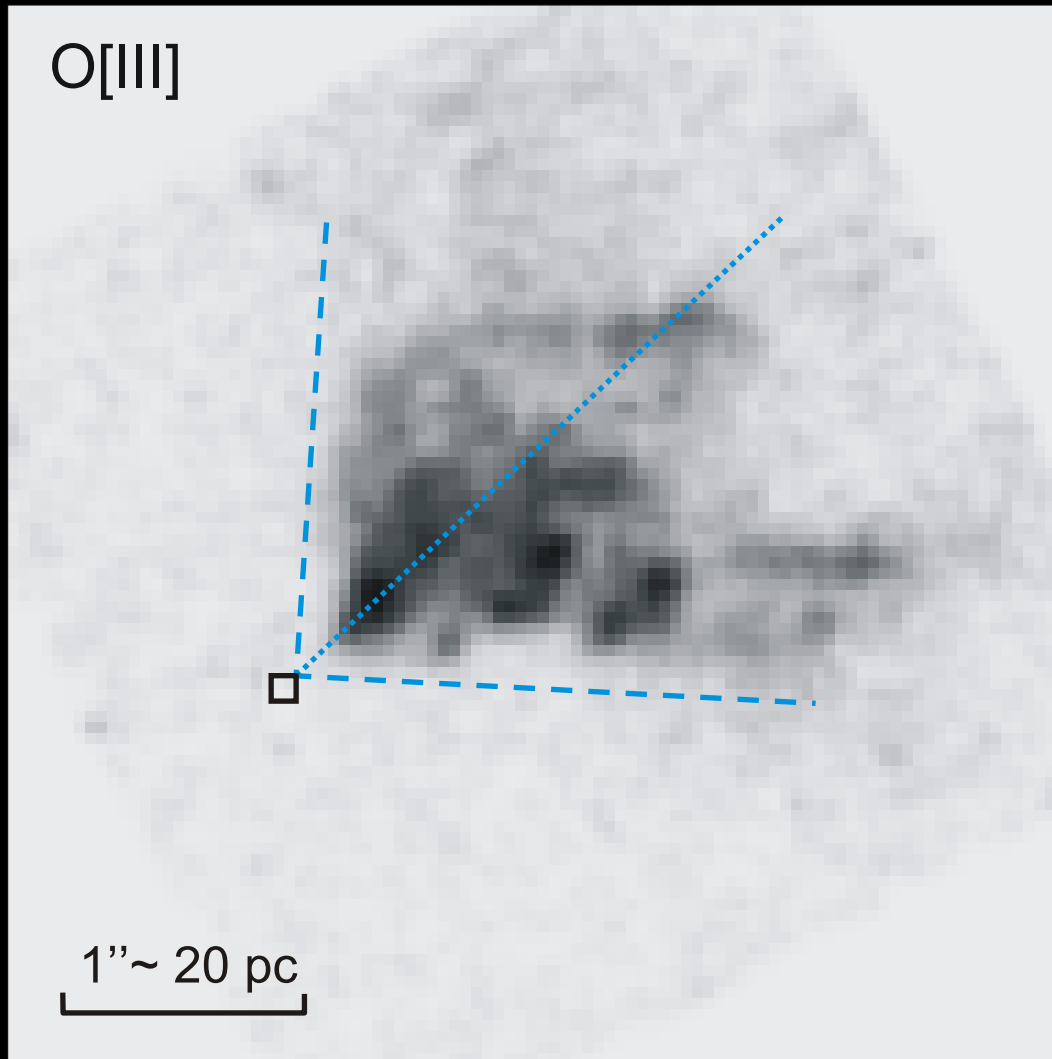
## Hydrodynamical torus model

Brightness distribution (in logarithmic scaling, axis in pc):



Schartmann et al.

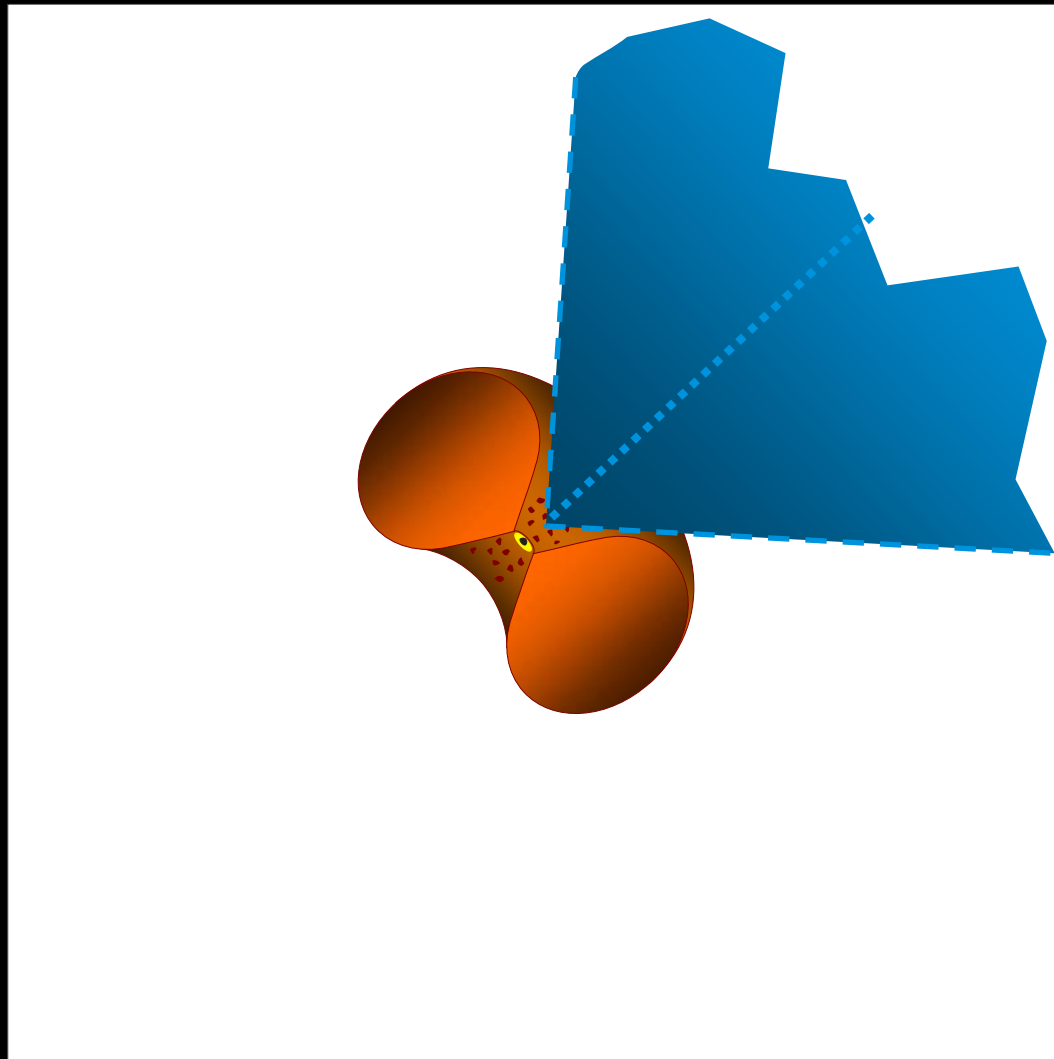
# Circinus: Geometry



Known geometry

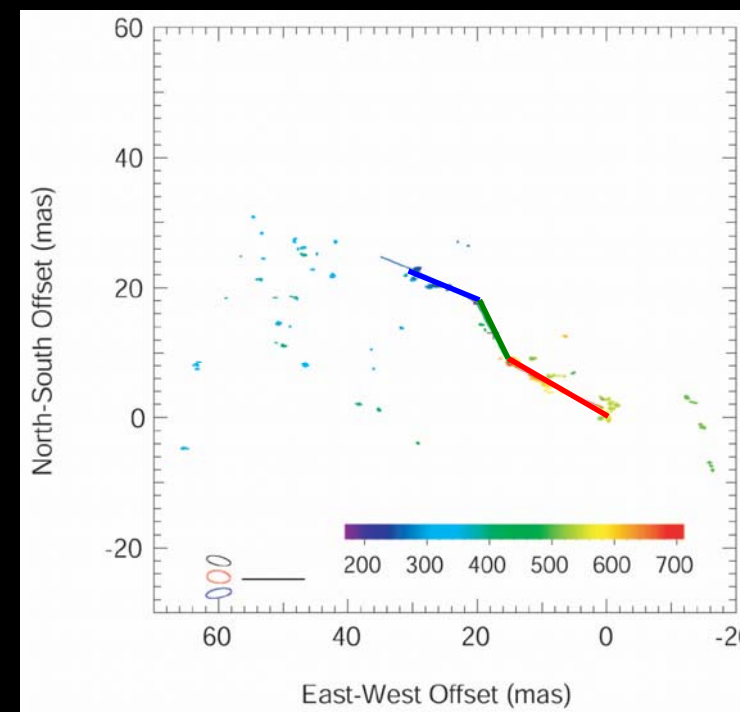
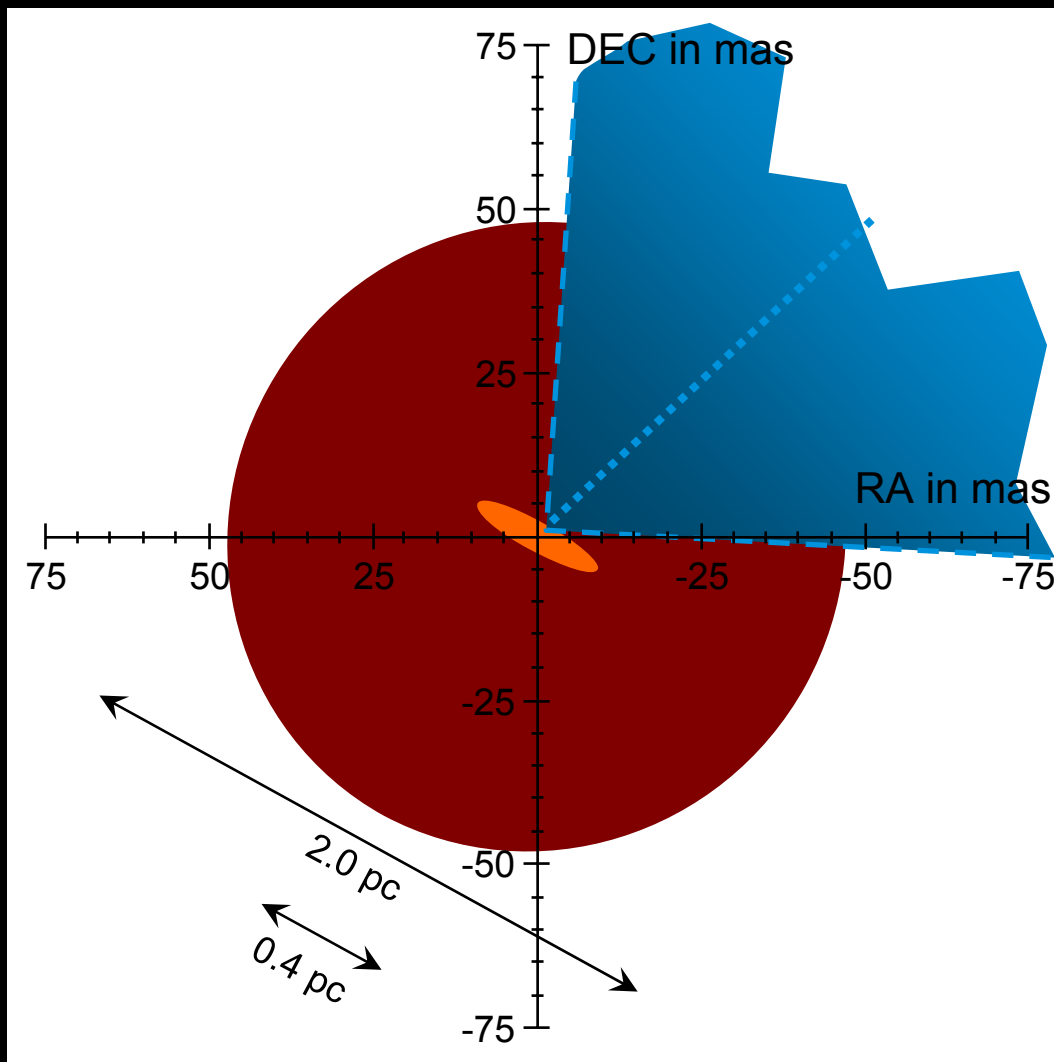
Wilson et al. 2000

# Circinus: Geometry



Expected torus configuration

# Circinus: Geometry



Greenhill et al. 2000

# NGC 1068: General properties



- Messier M77
- barred spiral galaxy  
(R)SA(rs)b,  $i = 51^\circ$
- Seyfert type 2
- $1 \times 10^7 M_\odot$  nucleus
- Distance  $\sim 14$  Mpc  
↳  $14 \text{ mas} \sim 1 \text{ pc}$



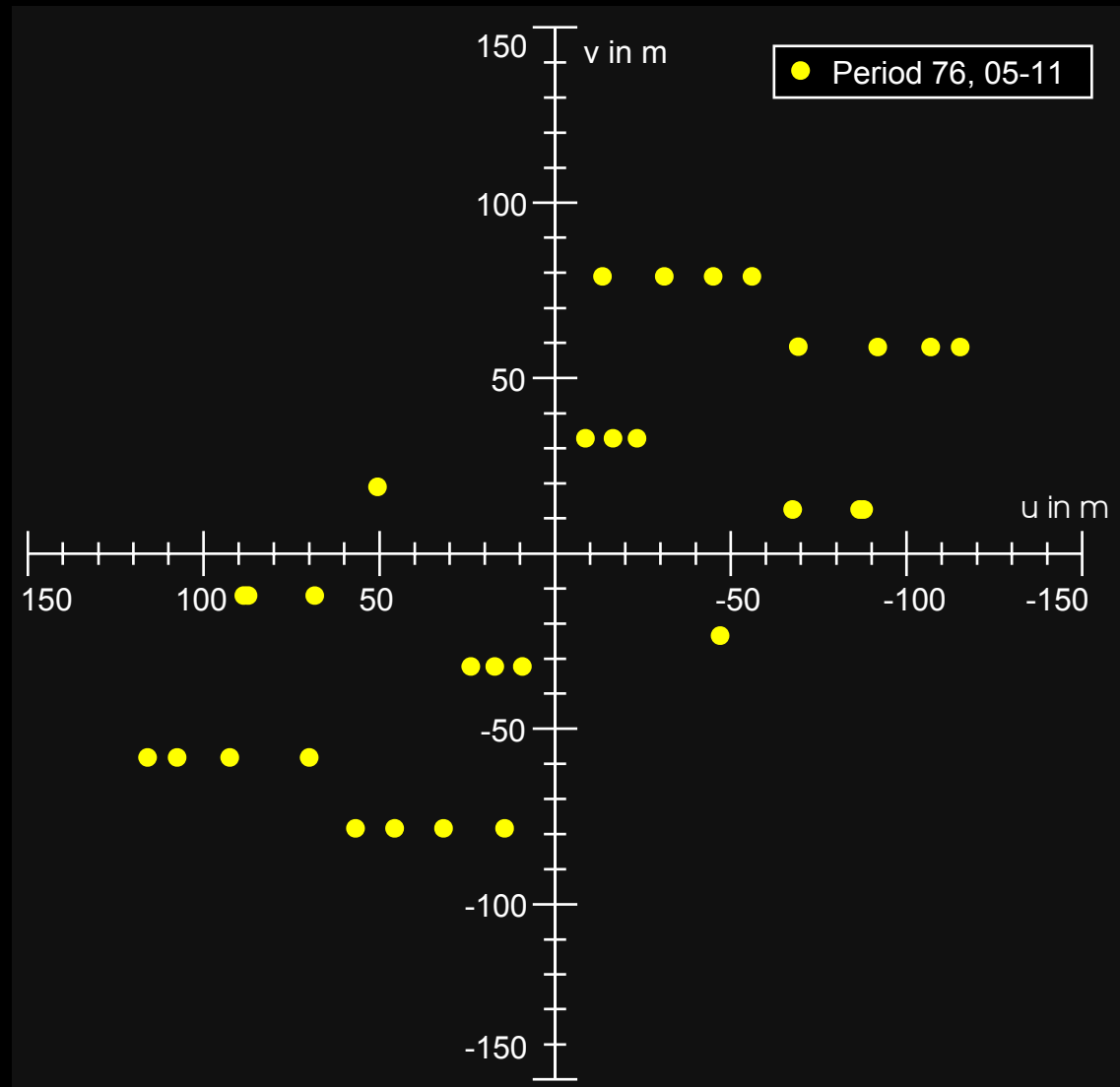
SDSS (<http://www.wikisky.org>)

# NGC 1068: UV coverage

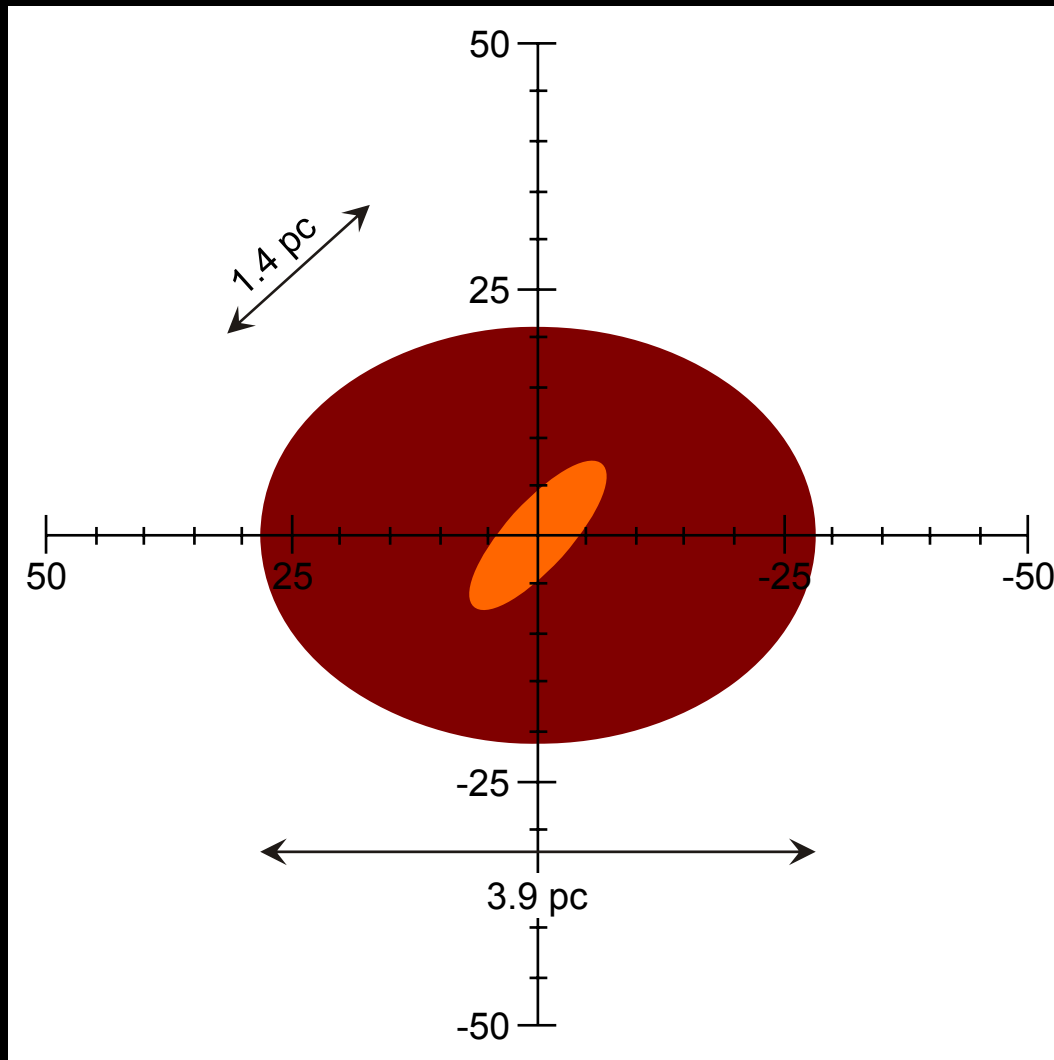


OT observations  
in period P76:

- grism
- 15 fringe tracks



# NGC 1068: Modelling

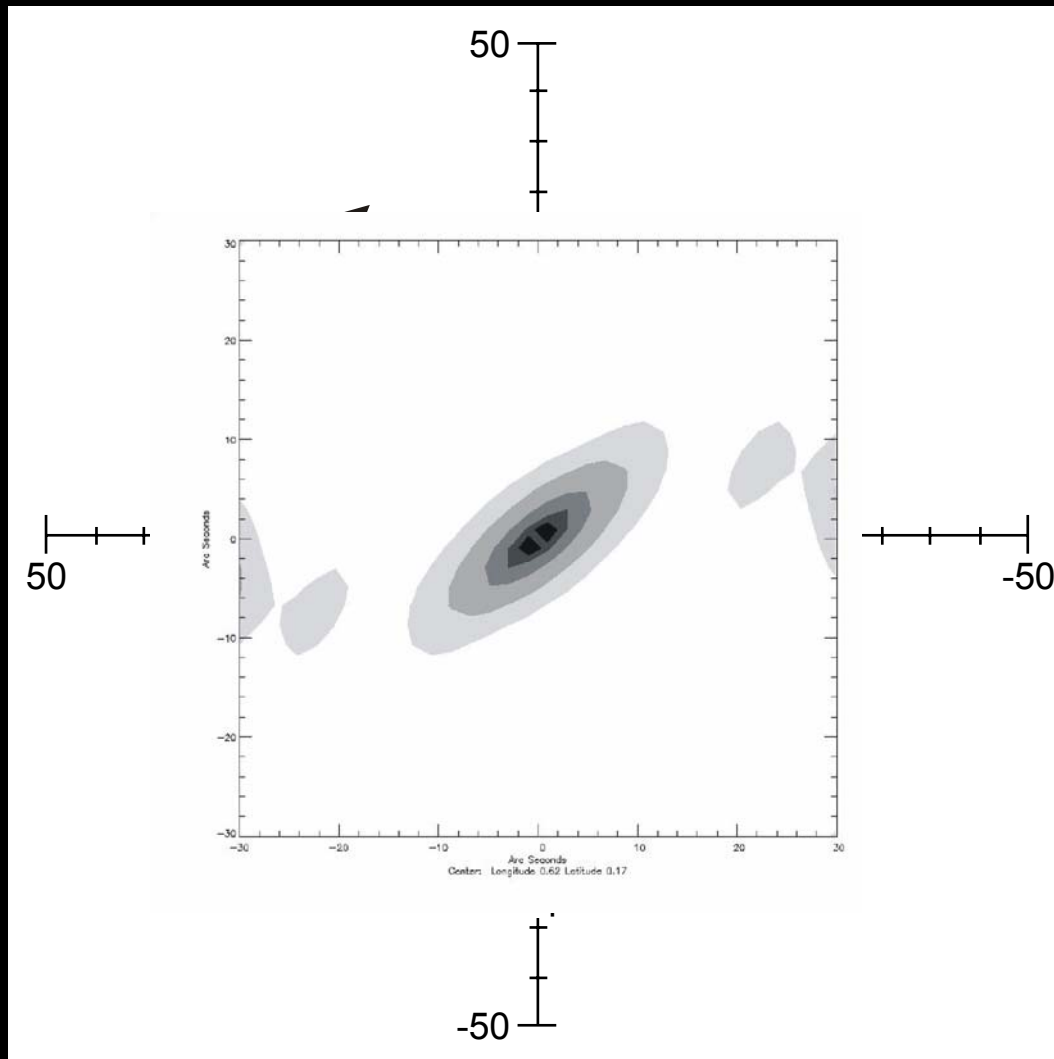


## 2 blackbody Gaussians:

Size:	$\Delta_1 = 20 \text{ mas}$
Axis ratio:	$r_1 = 0.34$
Silicate depth:	$\tau_1 \approx 5$
Temperature:	$T_1 \approx 700 \text{ K}$
Position angle:	$\alpha_1 = -42^\circ$
Size:	$\Delta_2 = 57 \text{ mas}$
Axis ratio:	$r_2 = 0.75$
Silicate depth:	$\tau_2 \approx 4$
Temperature:	$T_2 \approx 300 \text{ K}$
Position angle:	$\alpha \approx 0^\circ$

Raban et al.

# NGC 1068: Modelling



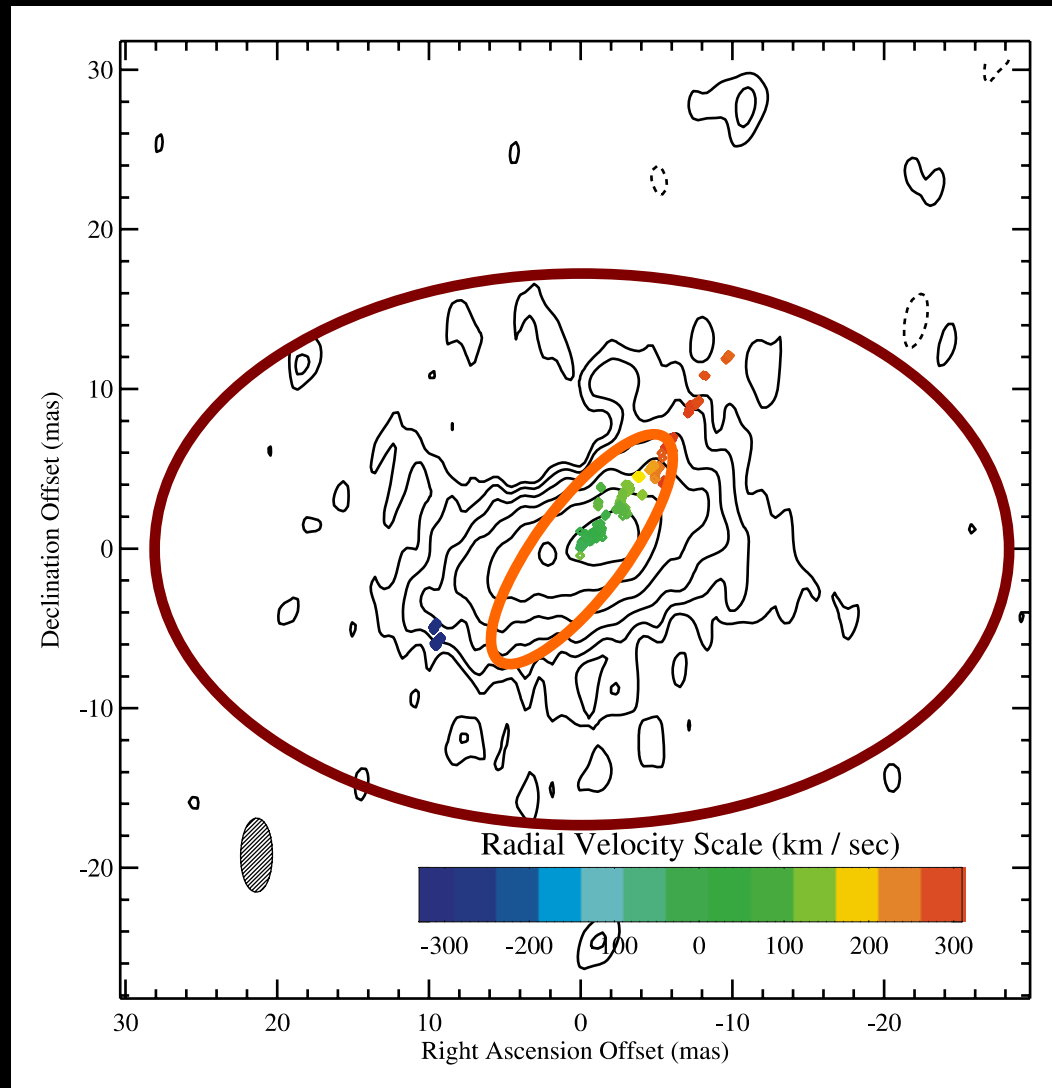
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# NGC 1068: Geometry

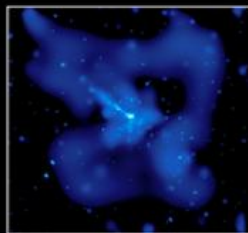


Gallimore et al.  
2004

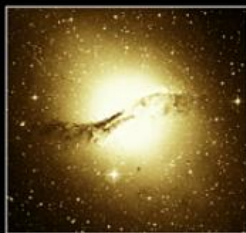
# Centaurus A: Properties



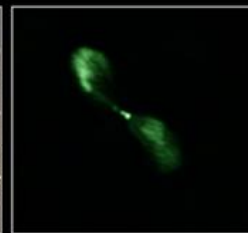
- elliptical galaxy
- recent merger
- dust lane, edge-on
- FR 1 radio galaxy
- narrow line RG
- $6 \times 10^7 M_{\odot}$  nucleus
- distance  $\sim 4$  Mpc  
↳ 50 mas  $\sim 1$  pc



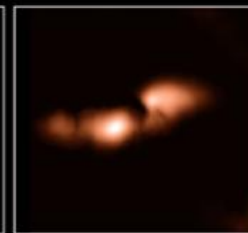
CHANDRA X-RAY



DSS OPTICAL



NRAO RADIO  
CONTINUUM



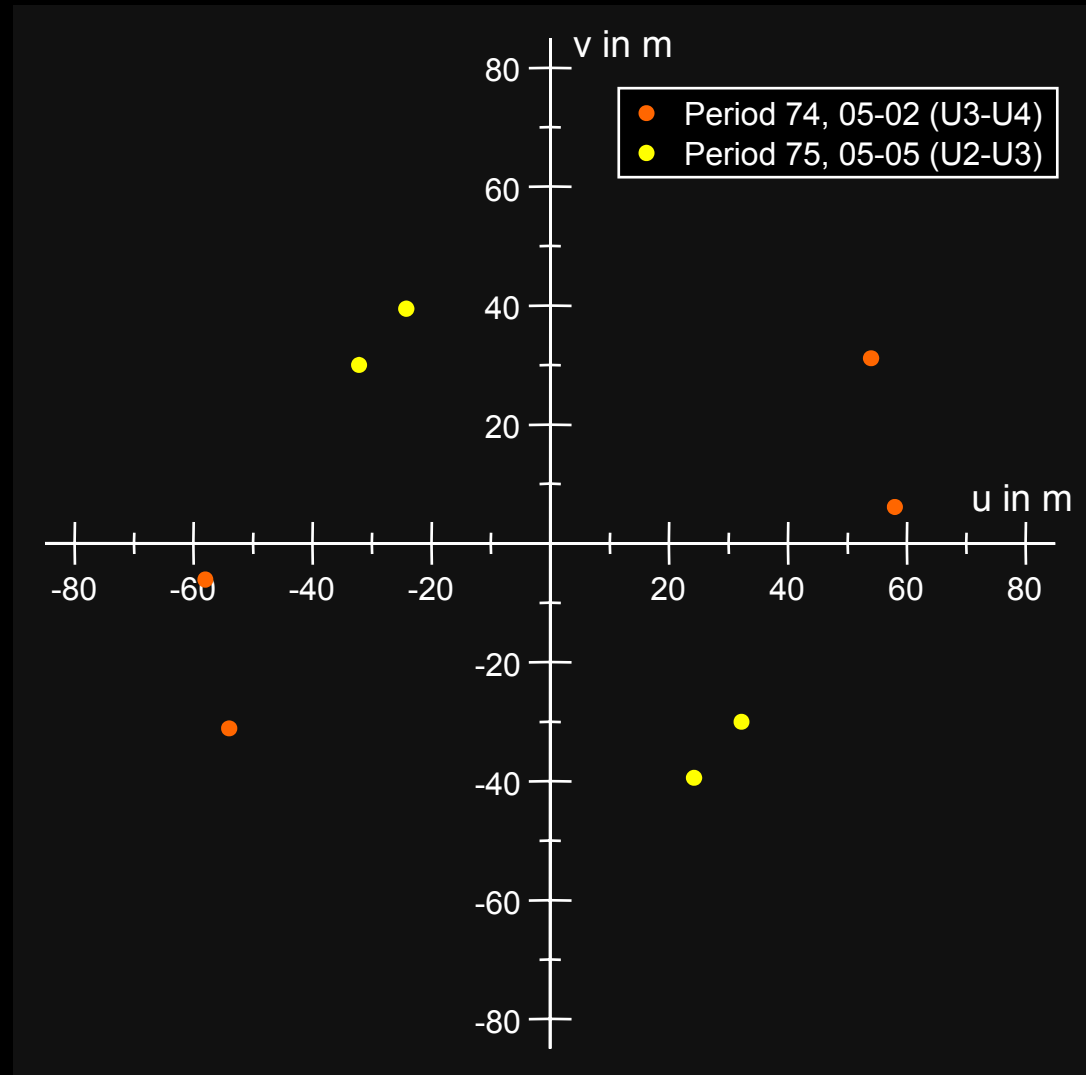
NRAO RADIO  
(21-CM)

# Centaurus A: UV coverage

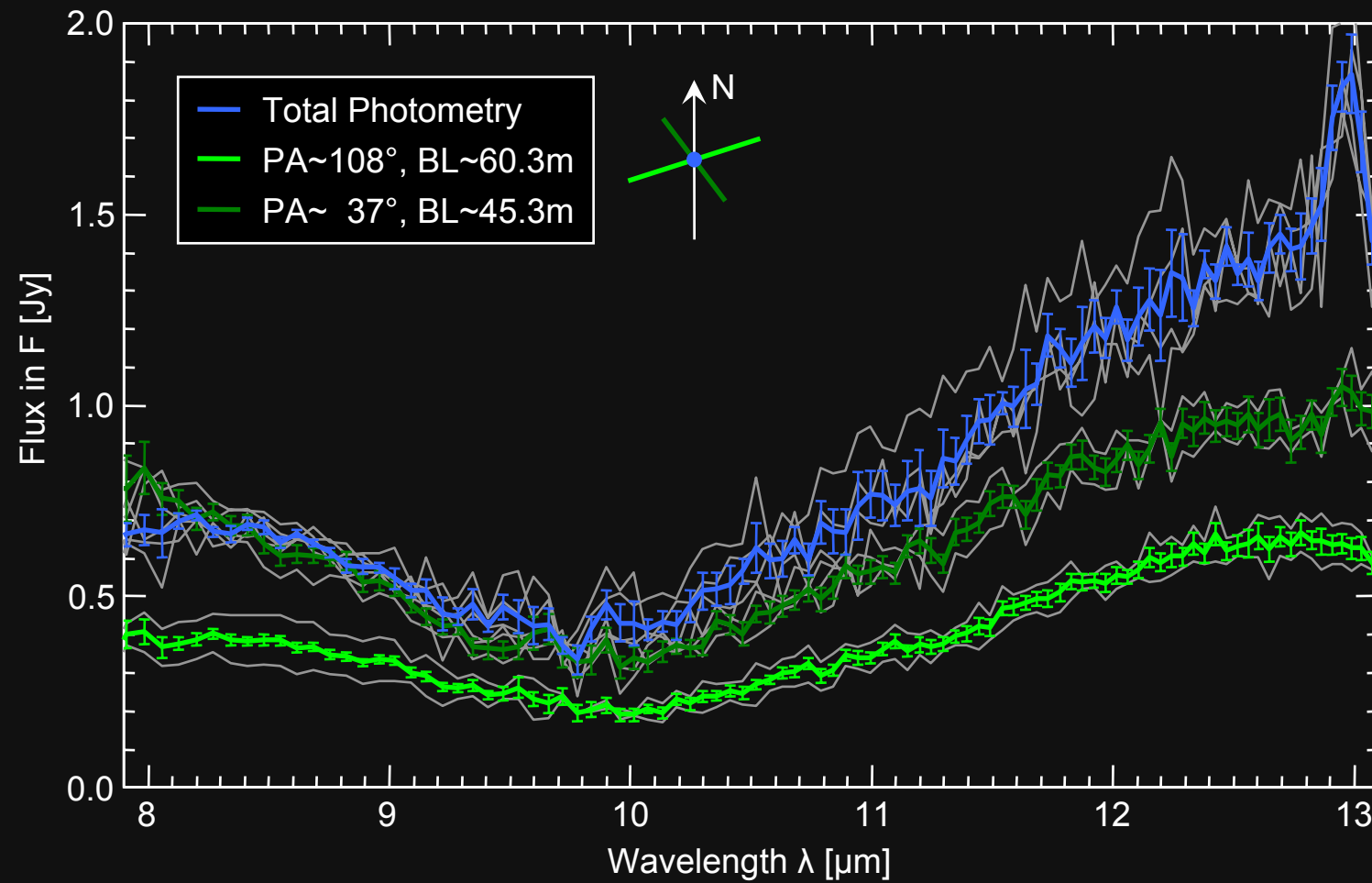


GTO observations in  
P74 and P75:

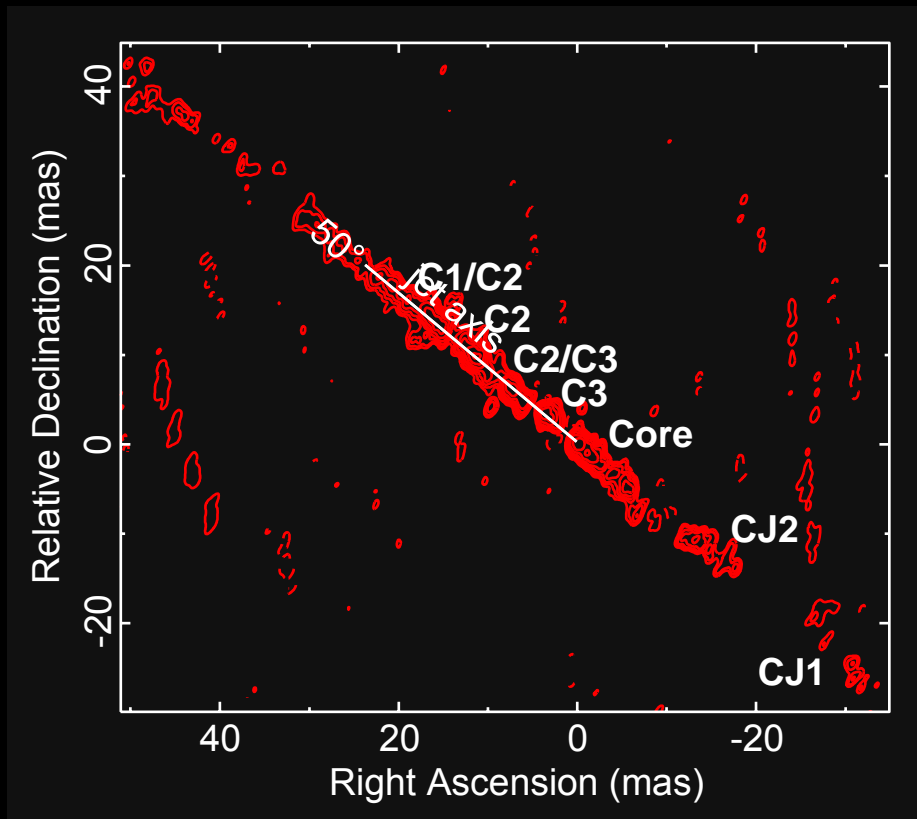
- 4 independent visibility points
- problems with photometry



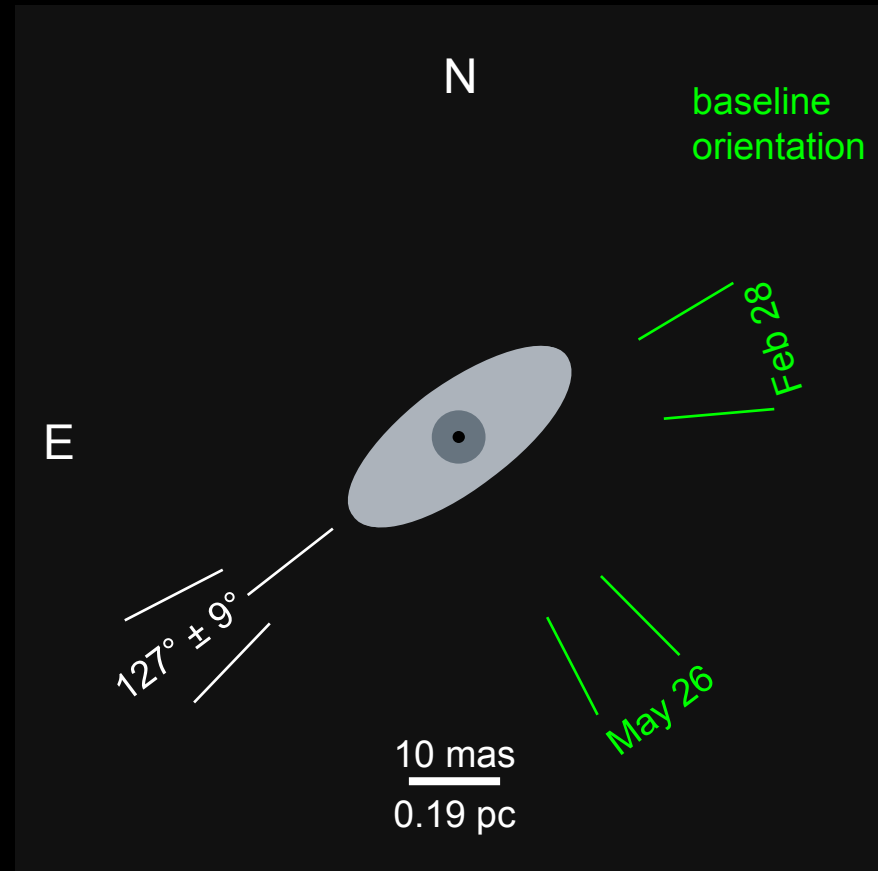
# Centaurus A: Fluxes



# Centaurus A: Results



Horiuchi et al. 2006

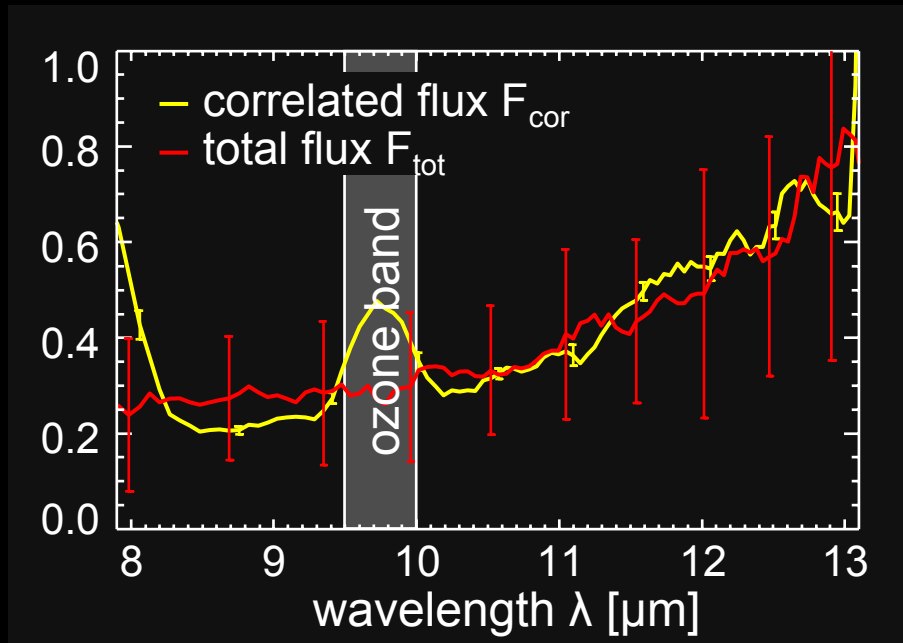


Meisenheimer et al. 2007

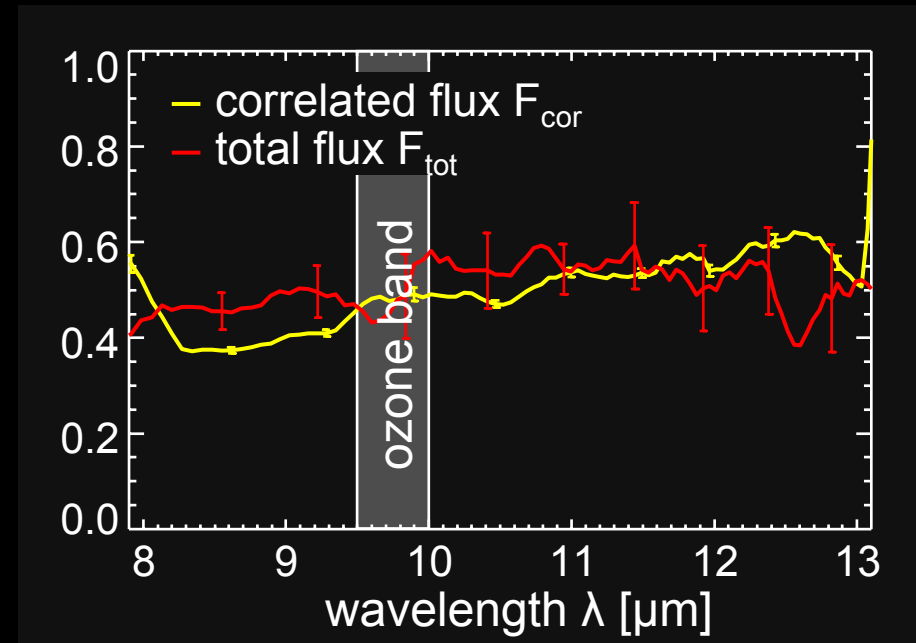
# Other targets: Mrk 1239, MCG -05-..



MCG -05-23-016 (Seyfert 2)



Mrk 1239 (Seyfert 1)



# Conclusions

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- Circinus & NGC 1068:
  - dusty tori perpendicular to ionisation cone
  - trace maser locations
- Centaurus A
  - dust disk
  - synchrotron source
- Weaker targets:
  - hardly resolved



VLT on Cerro Paranal, May 2005



# Prospects



- Long baselines for unresolved targets
- Shorter baselines for highly resolved targets
- Look at the phases
  - ↳ imaging
- Comparison to simulations
- Extension of sample

One of the galaxies to  
be observed: NGC 1365

