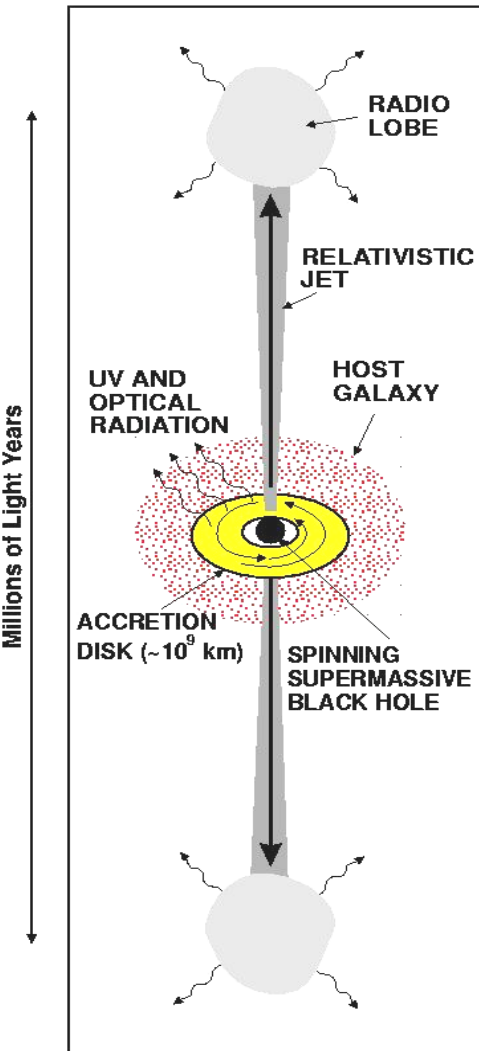
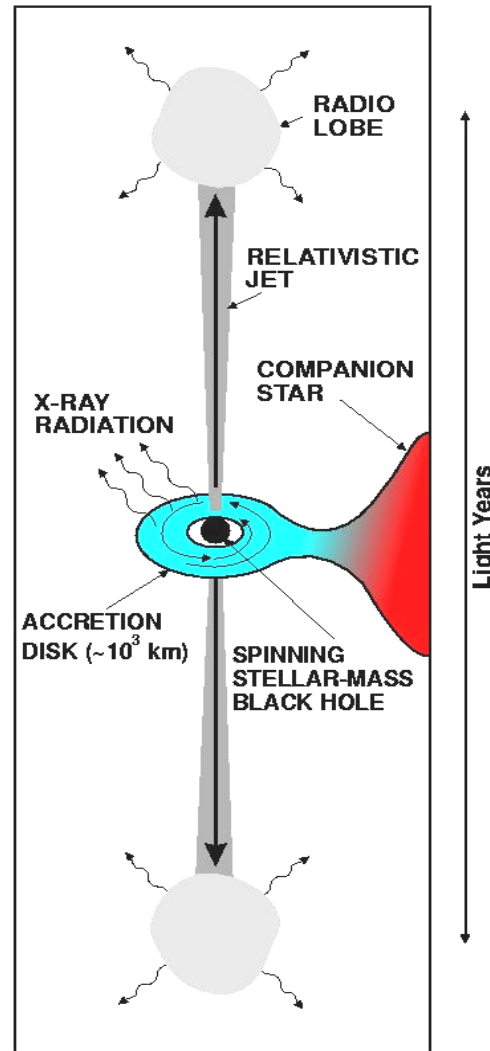


# QUASAR-MICROQUASAR ANALOGY

## QUASAR



## MICROQUASAR



M. & Rodriguez; Nature 1998

The scales of length and time are proportional to  $M_{BH}$

$$R_{sh} = 2GM_{BH}/c^2 ; \Delta T \propto M_{BH}$$

Unique system of equations:

The maximum color temperature of the accretion disk is:

$$T_{col} \propto (M/10M_{\odot})^{-1/4}$$

(Shakura & Sunyaev, 1976)

For a given accretion rate:

$$L_{Bol} \propto M_{BH} ; l_{jet} \propto M_{BH} ;$$

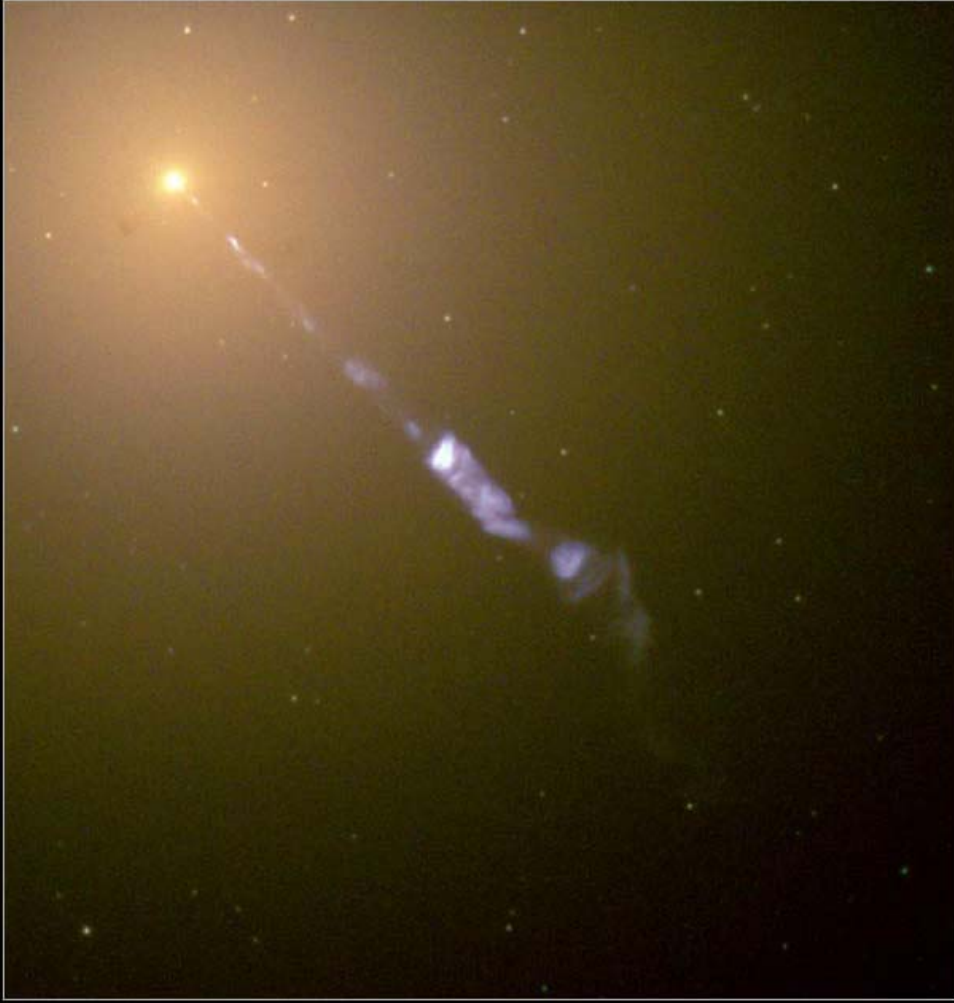
$$\phi \propto M_{BH}^{-1} ; B \propto M_{BH}^{-1/2}$$

(Sams, Eckart, Sunyaev, 96; Rees 04)

APPARENT SUPERLUMINAL MOTIONS IN  $\mu$ QSOs AS IN QSOs ?

# SUPERLUMINAL MOTIONS IN QSOs & AGN

The M87 Jet



- OBSERVED IN  $> 30$  QSOs & AGN
- IN RADIO & OPTICAL WAVES
- PROPER MOTION SEEN IN YEARS
- $V_{\text{app}}$  UP TO  $30c$  in blazars
- One sided because of Doppler boosting
- WHAT IS THE NATURE ?

“PLASMONS” OR SHOCKS ?

# SUPERLUMINAL EJECTION IN A $\mu$ QSO

M. & Luis Rodriguez, 1994

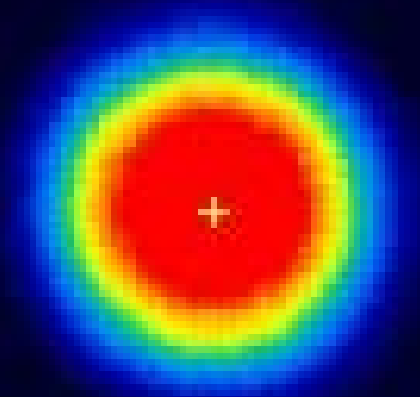
THE PLASMA THAT RADIATES IN THE HARD  
X-RAYS IS BLOWN IN SUPERLUMINAL JETS

$\lambda 3.6$  cm

1 arcsec



GRS 1915+105



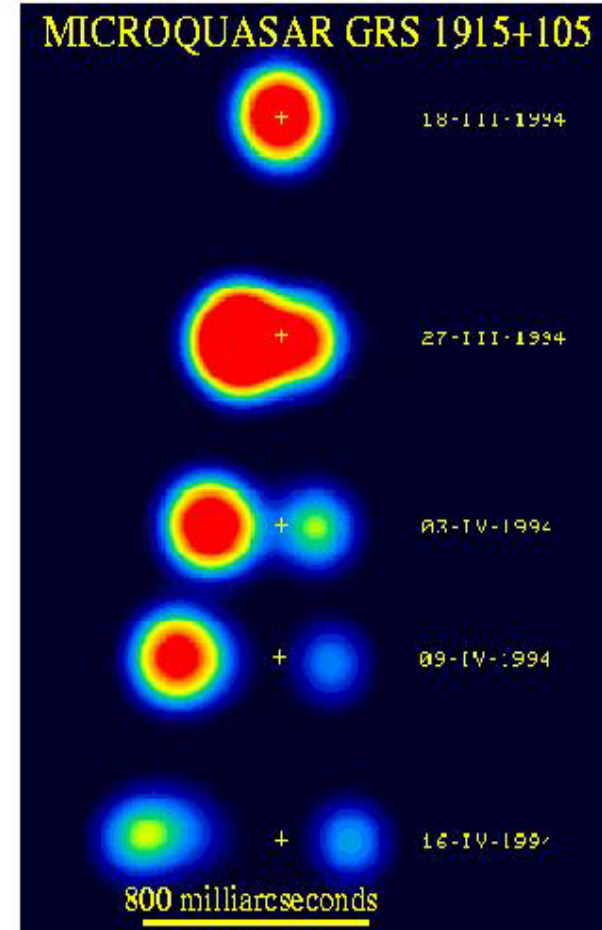
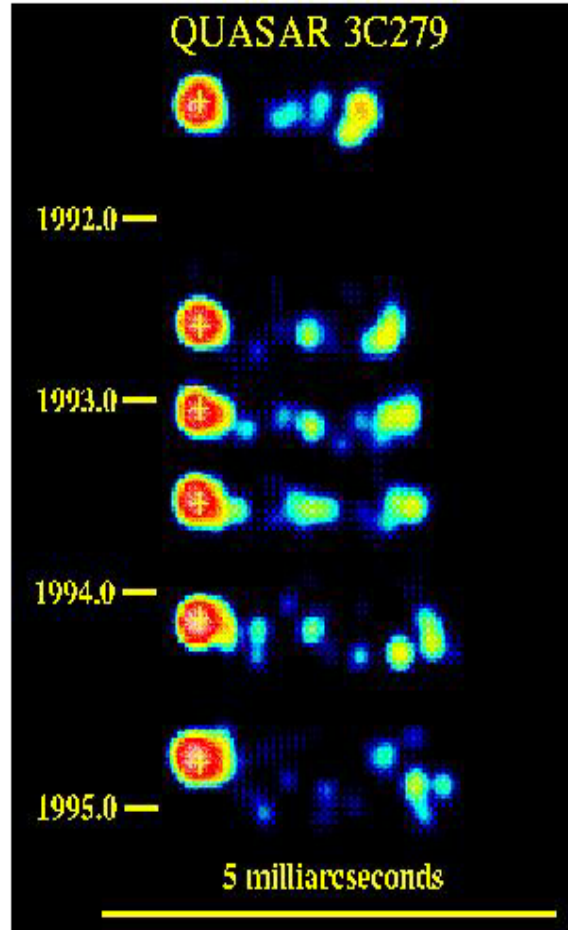
18-III-1994

$V_{\text{app}} > C$  for a DISTANCE  $> 8$  Kpc

# SUPERLUMINAL MOTIONS IN THE GALAXY

M. & Rodriguez, 1994

WITH SAME BULK LORENTZ FACTORS AS IN QSOs

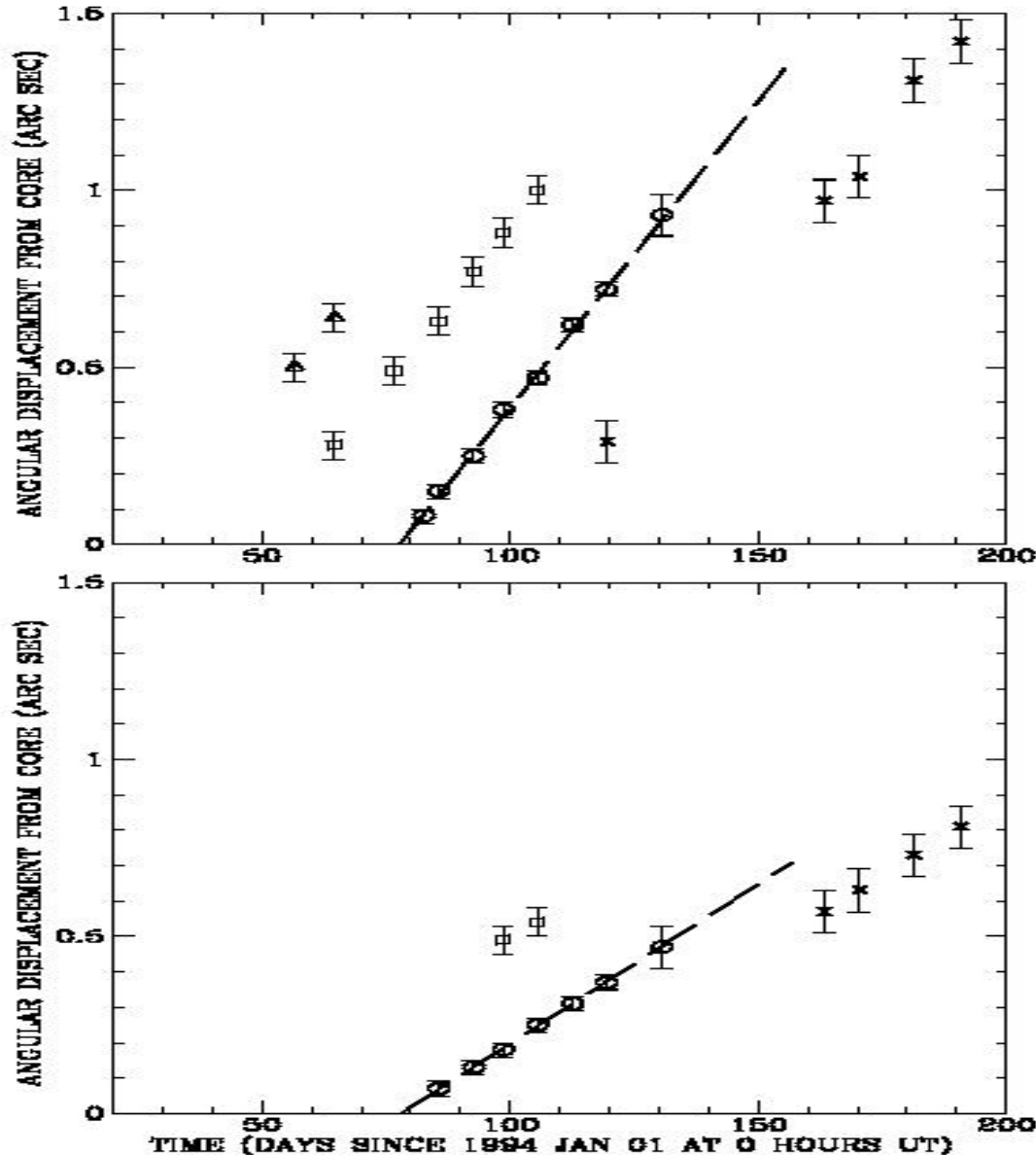


- 1) MOVE ON THE PLANE OF THE SKY  $\sim 10^3$  TIMES FASTER
- 2) JETS ARE TWO-SIDED WHICH ALLOWS TO SOLVE EQUATIONS
- 3) BETTER AGN AT  $< 100$  Mpc. e.g. Collimation @  $30-100 R_{sh}$  in M87 (Biretta)

# GRS 1915+105: SEVERAL EJECTION EVENTS

VLA (Rodriguez & Mirabel, 1999)

Also with MERLIN (Fender et al. 01)



$$\text{Bulk } \Gamma = (1 - \beta^2)^{-1/2} \sim 5$$

for  $\beta = v/c$

Are these jets

**Fireballs**

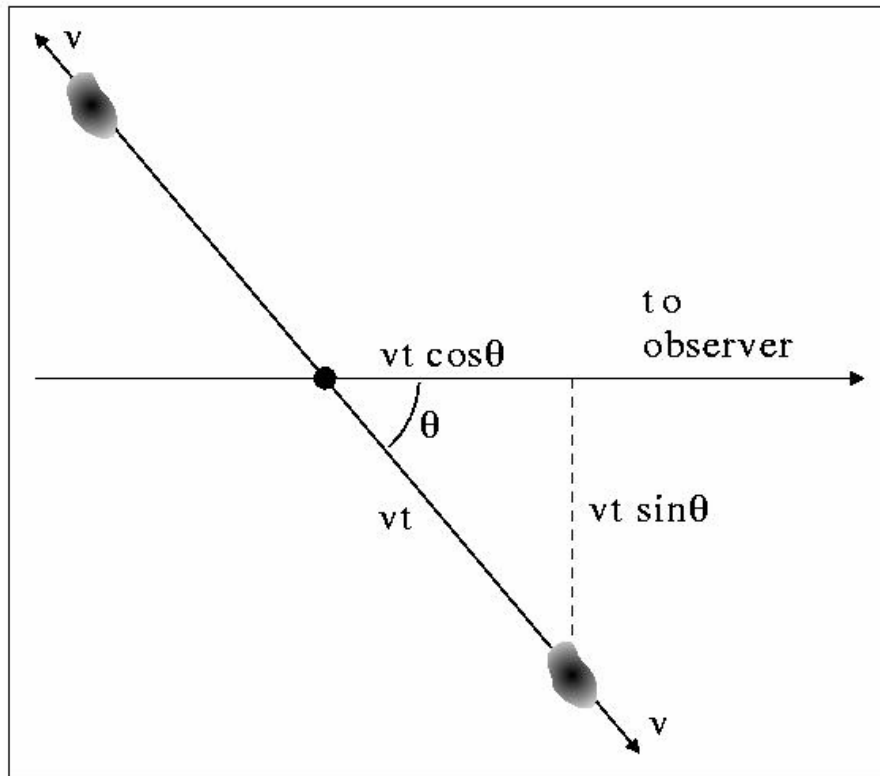
(Mezaros & Rees)

or

**Cannonballs ?**

(Daar & De Rujula)

# RELATIVISTIC ABERRATION IN ANTISYMMETRIC TWIN JETS



$$\frac{S_a}{S_r} = \left( \frac{1 + \beta \cos \theta}{1 - \beta \cos \theta} \right)^{k-\alpha},$$

$$\mu_a = \frac{\beta \sin \theta}{(1 - \beta \cos \theta)} \frac{c}{D},$$

$$\mu_r = \frac{\beta \sin \theta}{(1 + \beta \cos \theta)} \frac{c}{D},$$

$$\beta \cos \theta = \frac{\mu_a - \mu_r}{\mu_a + \mu_r},$$

$$D = \frac{c \tan \theta}{2} \frac{(\mu_a - \mu_r)}{\mu_a \mu_r}.$$

Same bulk Lorentz factors as in QSOs: 2-10

$$D \leq \frac{c}{\sqrt{\mu_a \mu_r}}.$$

Relativistic upper limit:  $D < 14$  kpc

# RELATIVISTIC METHOD FOR DISTANCES IN ASTRONOMY

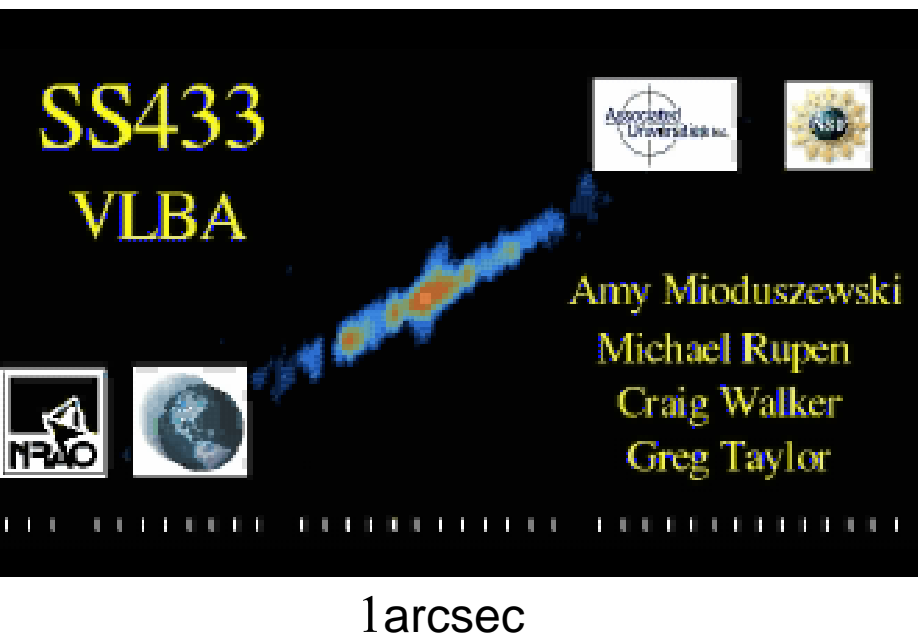
$$\left[ \frac{\mu_{a,r}}{\text{radians sec}^{-1}} \right] = \frac{v \sin \theta}{(1 \mp (v/c) \cos \theta) D} \quad \text{FROM PROPER MOTIONS}$$

IF THE JETS ARE DOMINATED BY ELECTRON-PROTON PLASMA:

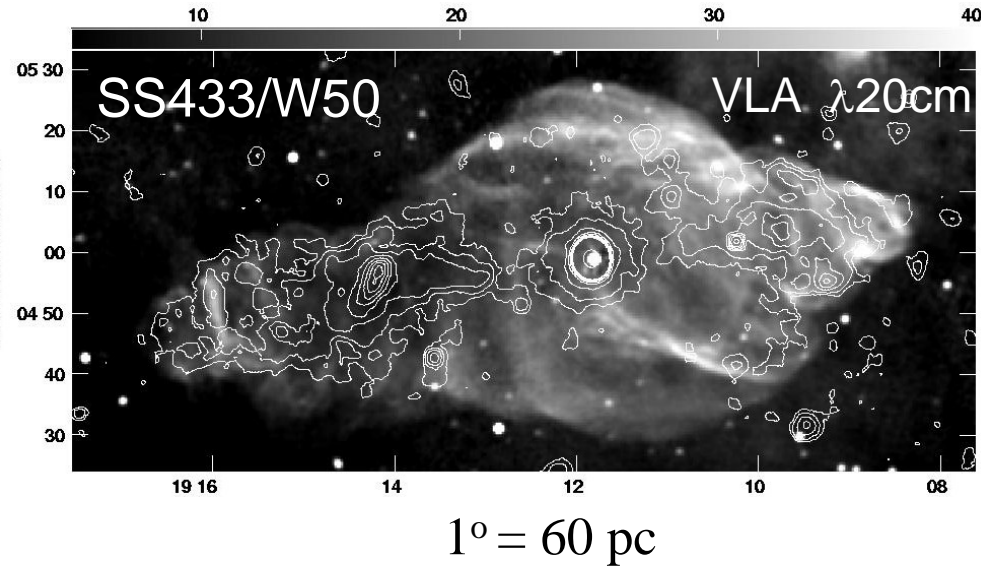
$$\frac{\lambda_{a,r}}{\lambda_{rest}} = \frac{(1 \mp (v/c) \cos \theta)}{[1 - (v/c)^2]^{1/2}} \quad \text{FROM DOPPLER FACTOR OF ION LINES}$$

- IF ANTISYMMETRIC, THE VELOCITY, DIRECTION OF THE JETS, & DISTANCE CAN BE FOUND (M. & Rodriguez, 1994)
- BUT SPECTRAL LINES (H, He, Fe) HAVE BEEN SO FAR DETECTED ONLY IN SS 433. **WHY ?**

# DARK JETS FROM BLACK HOLES



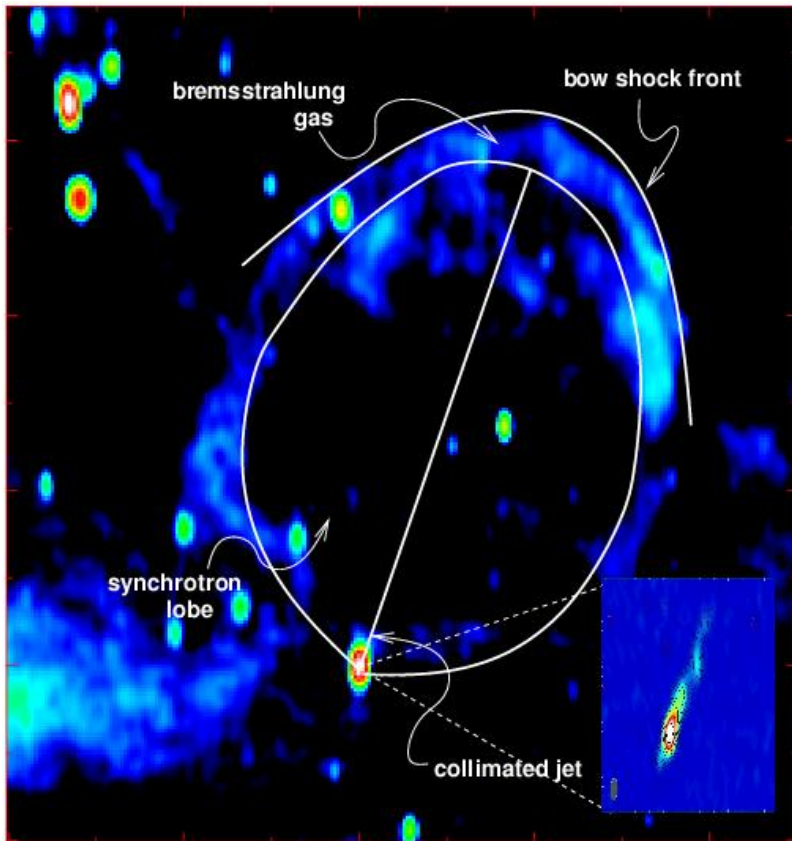
Radio (Dubner et al); X-rays: (Brinkmann et al)



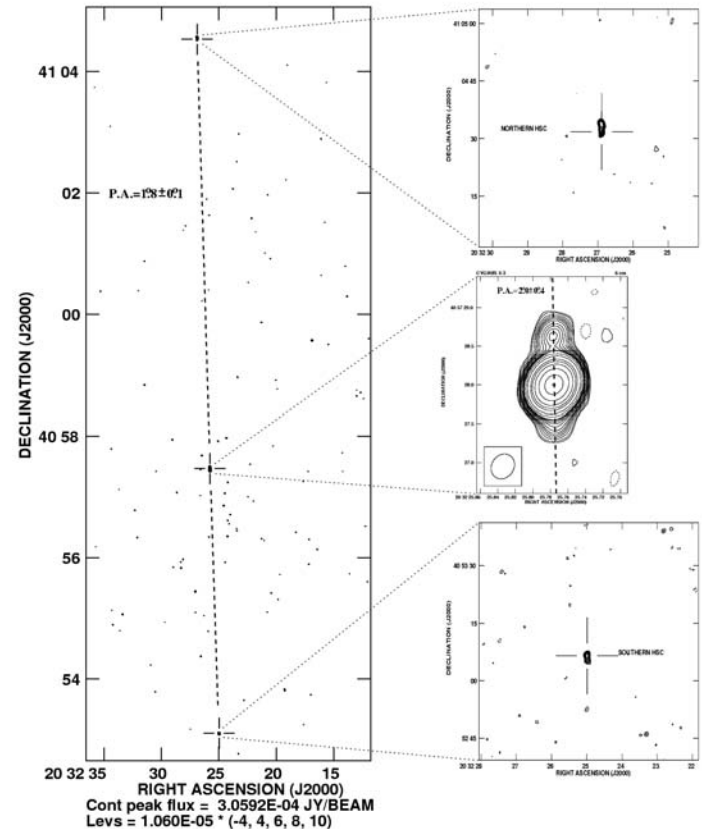
- ATOMIC NUCLEI MOVING AT  $0.26c \Rightarrow$
- MECHANICAL LUMINOSITY  $> 10^{40}$  erg/sec
- NON RADIATIVE JETS = "DARK" JETS
- $>50\%$  OF THE ENERGY IS NOT RADIATED

# LARGE-SCALE JETS BECOMING COMMON

Cyg X-1 (Gallo, Fender et al. Nature 05)



Cyg X-3 (Marti et al. 2005)



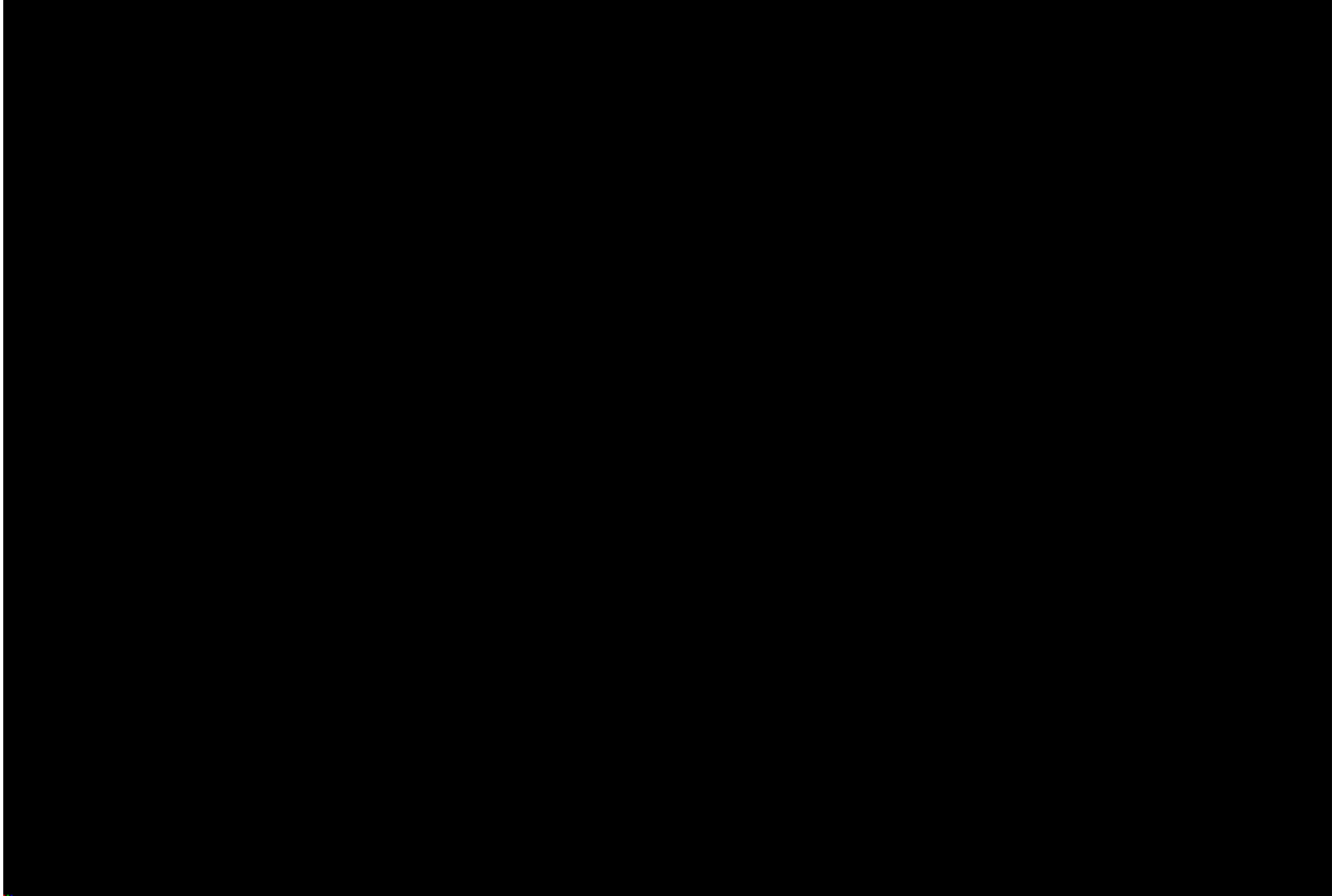
ALSO IN **GRS 1915+105** (Luis Rodriguez & M. 2006)

>50% OF THE RELEASED ENERGY IS NOT RADIATED

# MOVING X-RAY JETS IN A $\mu$ QSO

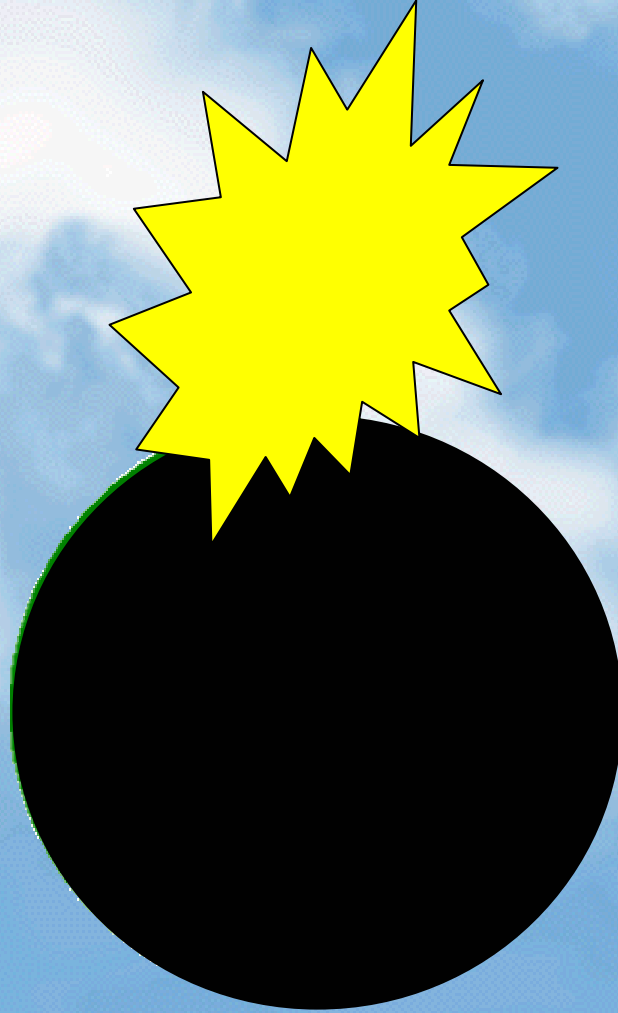
$\mu$ QSOs XTE J1550-564 & H1743-322

Corbel et al. (2002, 05)



X-rays are produced by synchrotron  $\Rightarrow$  electrons accelerated to TeV energies

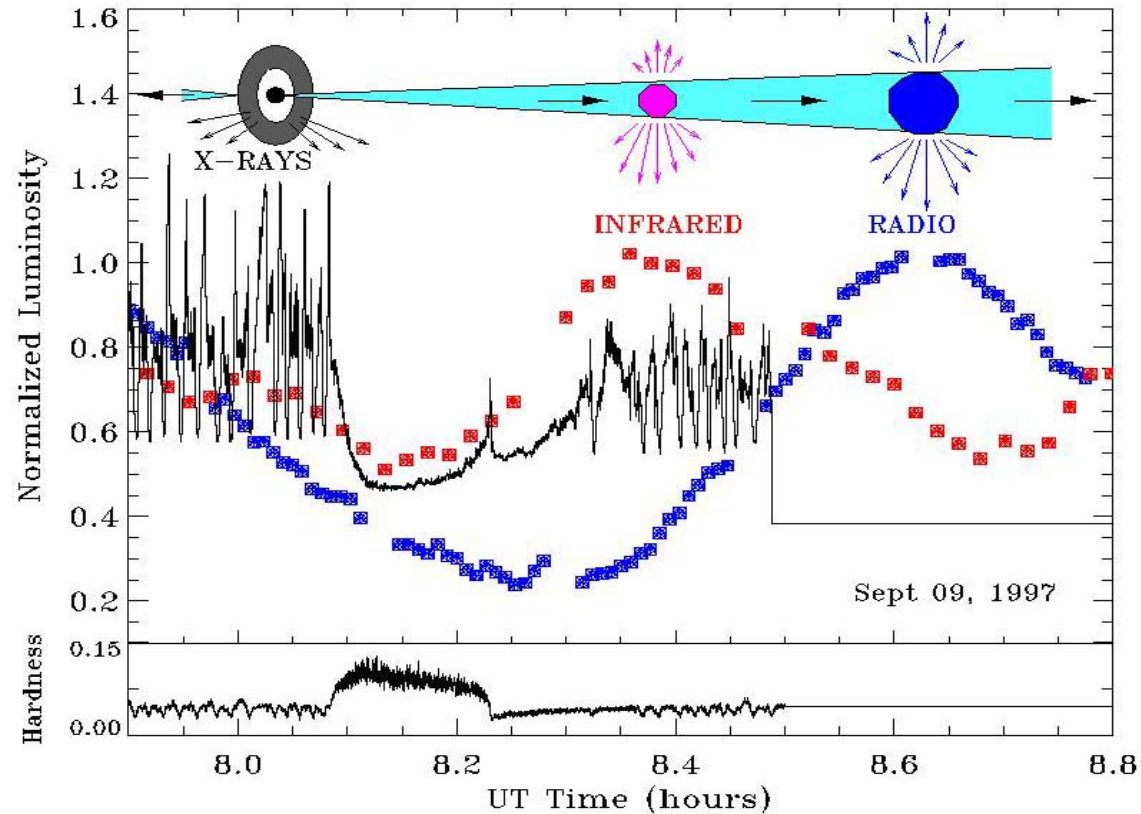
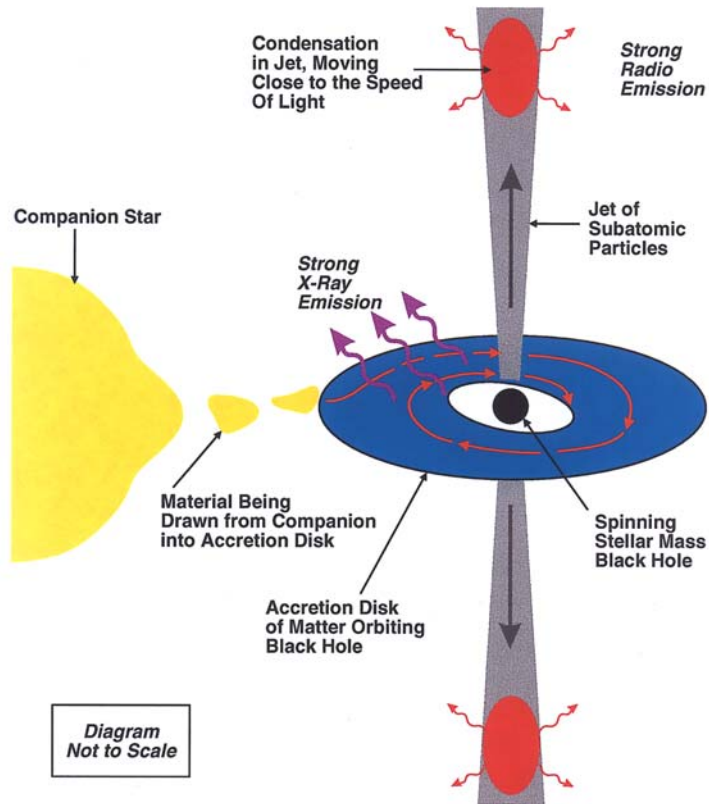
**HORIZON IS THE BASIC CONCEPT THAT DEFINES BLACK HOLES AND  
IT IS DIFFICULT TO OBTAIN DIRECT EVIDENCE OF THEIR EXISTENCE**



# ACCRETION-JET CONNECTION

$$\Delta T \propto M_{\text{BH}}$$

1 hr = 30 yr in SgrA\* Multiwavelength approach (M. et al.1998)



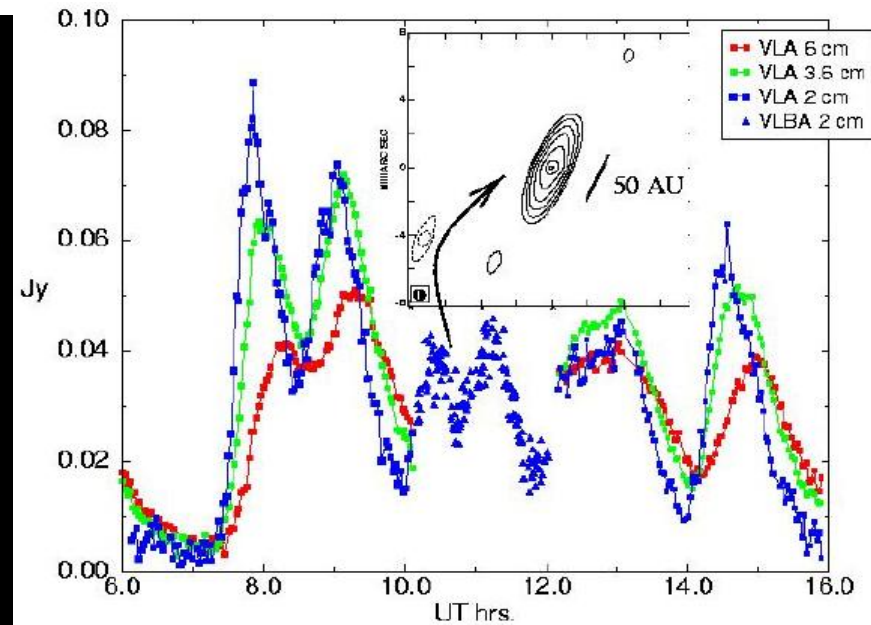
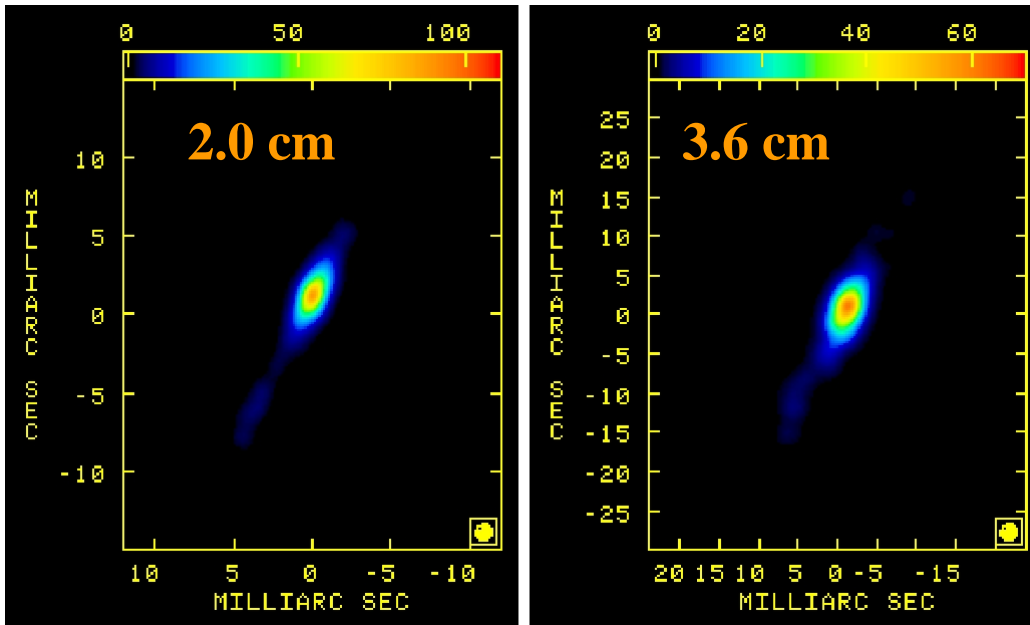
**ABSCENCE OF EVIDENCE FOR A MATERIAL SURFACE AROUND  $14 M_{\text{sun}}$**

**THE ONSET OF THE JET IS AT THE TIME OF A X-RAY “SPIKE”:  
SUDDEN REFILL OF THE DISK & SHOCK THROUGH COMPACT JET**

# COMPACT STEADY JETS

Ribo, M. & Dhawan 2005

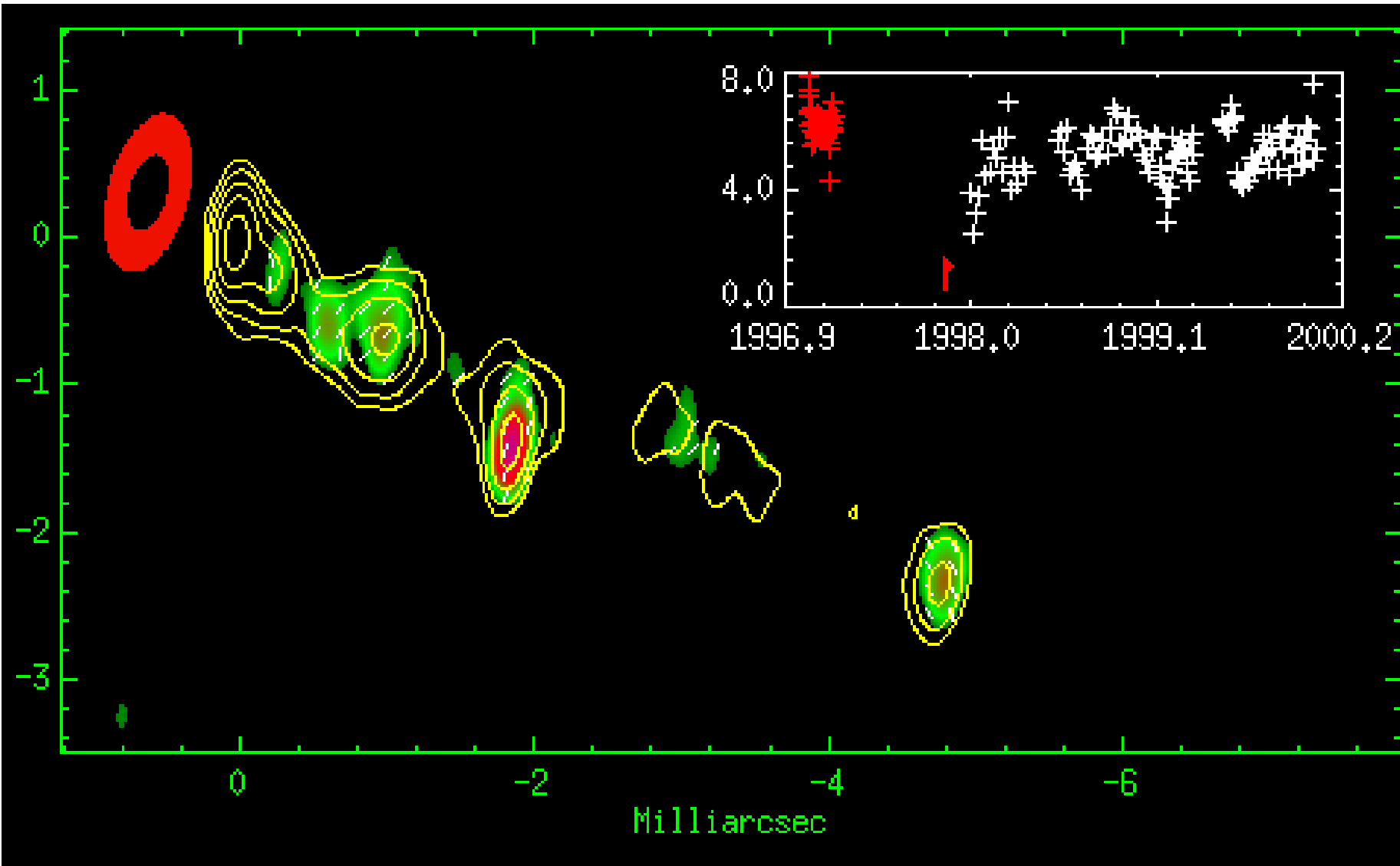
M. et. al, 1998



- ~100 AU IN LENGTH PRESENT DURING PLATEAU STATE
- SPEED OF THE FLOW  $< 0.4c$  (Ribo, M. & Dhawan 2005)
- LARGE SCALE JETS ARE SHOCKS PROPAGATING AT  $\sim 0.98c$  THROUGH THE SLOWER MOVING COMPACT JET
- THE COMPACT JET REAPPEARS HOURS AFTER FLARES

# ANALOGOUS ACCRETION-EJECTION IN 3C 120

Marscher et al. Nature 2002



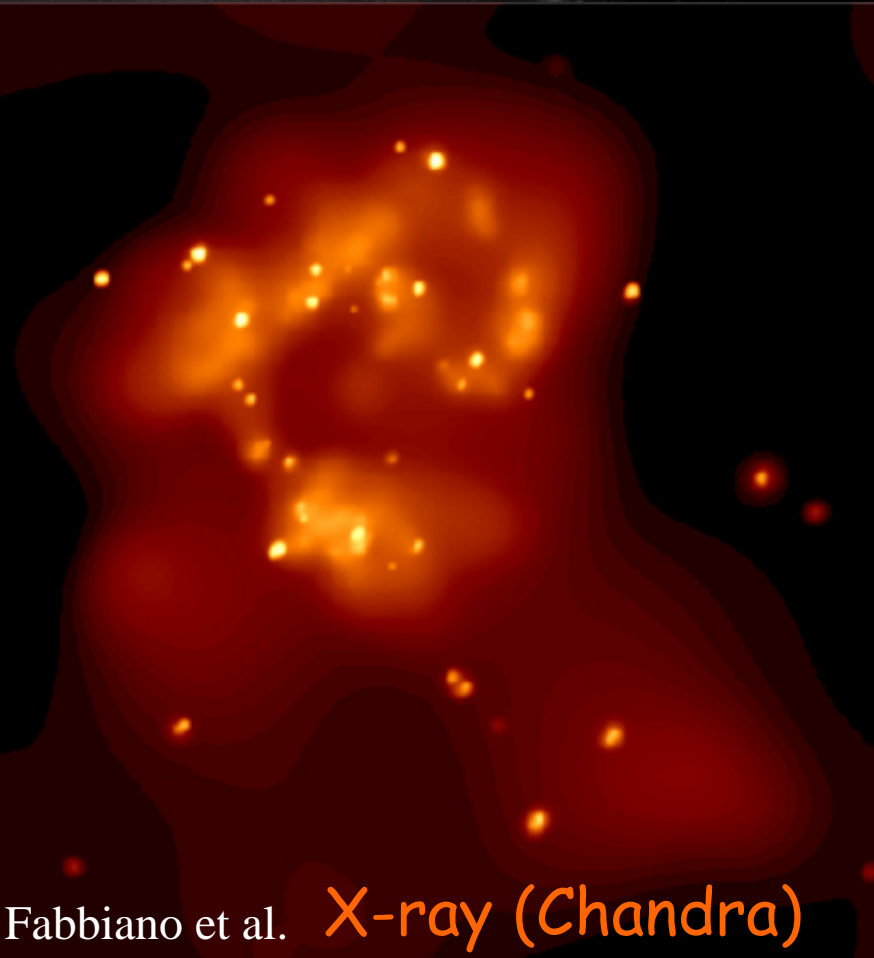
# $\mu$ QSOs & ULTRALUMINOUS X-RAY SOURCES

## Microquasars in external galaxies

ISOTROPIC BUT  $M_{\text{BH}} > 30 M_{\odot}$  (Pakull et al. 2002)

ANISOTROPIC BUT NOT BEAMED (King et al 2001)

ANISOTROPIC AND BEAMED (Mirabel & Rodriguez, 1999)



Fabbiano et al. X-ray (Chandra)



Antennae

# MICROBLAZARS

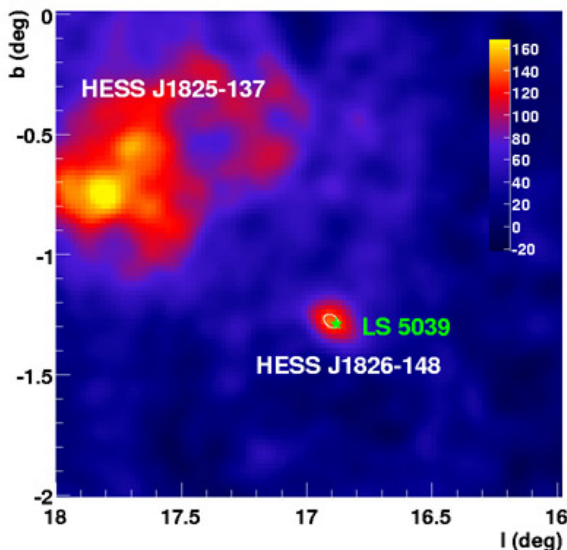
(M. & L.F. Rodriguez, ARAA 1999)

**Due to relativistic beaming:  $\Delta t \propto 1/2\gamma^2$ ;  $I \propto 8\gamma^3$**

**e.g. If  $\gamma = 5$ ,  $\Theta < 10^\circ \Rightarrow \Delta t < 1/50$  and  $\Delta I > 10^3$**

**-Should appear as variable and luminous gamma-ray sources**

**-By inverse Compton of jet particles on UV photons some HMXBs may be gamma-ray sources (Romero et al.; Paredes et al. 2002)**



**TWO HMXBs detected at  $> 100$  Gev**

- **LS 5039 by HESS** (Aharonian et al 2005)
- **LSI +61 303 by EGRET** (Paredes et al. 04)