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# Future Developments of Infrared Controllers at ESO

Gert Finger

# Next Generation IR Detectors

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Detector	Format	Outputs	Pixel Time [ $\mu$ s]	Frame Time [ms]	Baud Rate Gbaud
Hawaii-2RG MBE HgCdTe	2Kx2K	32	0.2	26	2.6
Si:As BIB	640x480	32	0.3	2.8	1.7
ORION 2Kx2K InSb	2Kx2K	64	1.5	100	0.7
Adaptive optics sensors	256x256	32	0.2	0.41	2.6

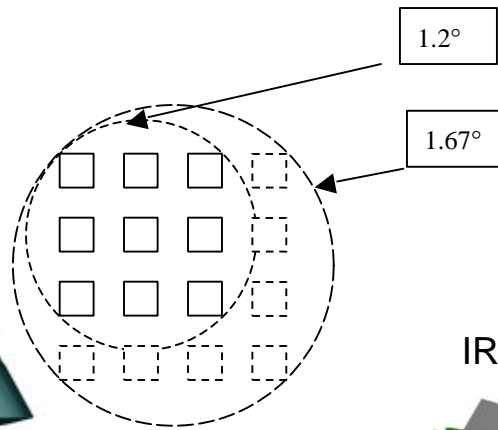
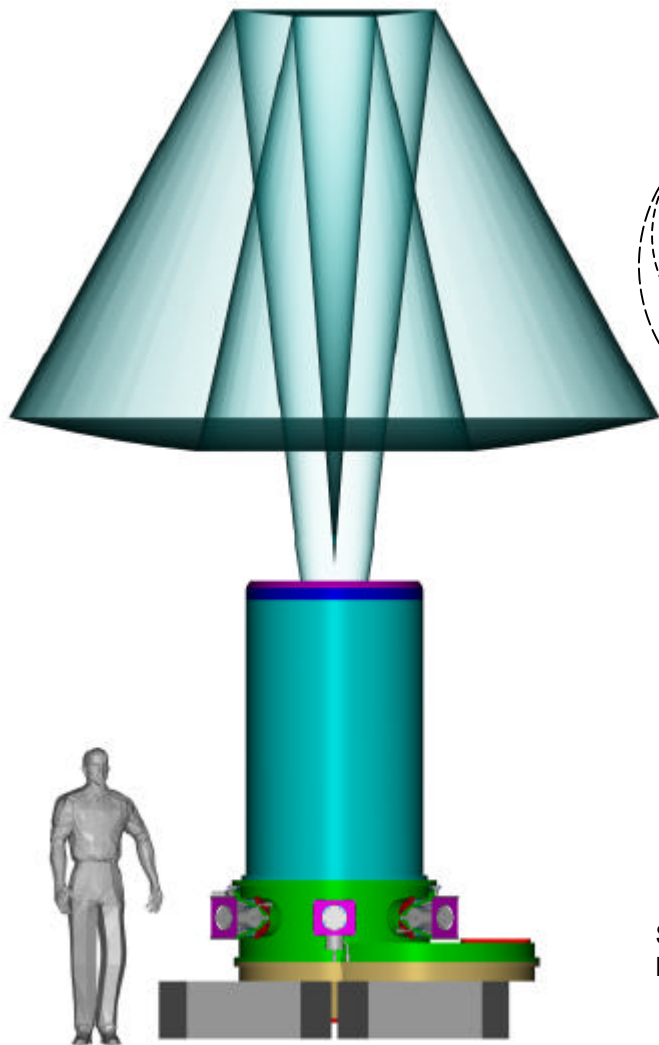
- need 5 MHz 16 bit ADC's
- need 3 Gbaud fiber link

# Future Requirements for IRACE

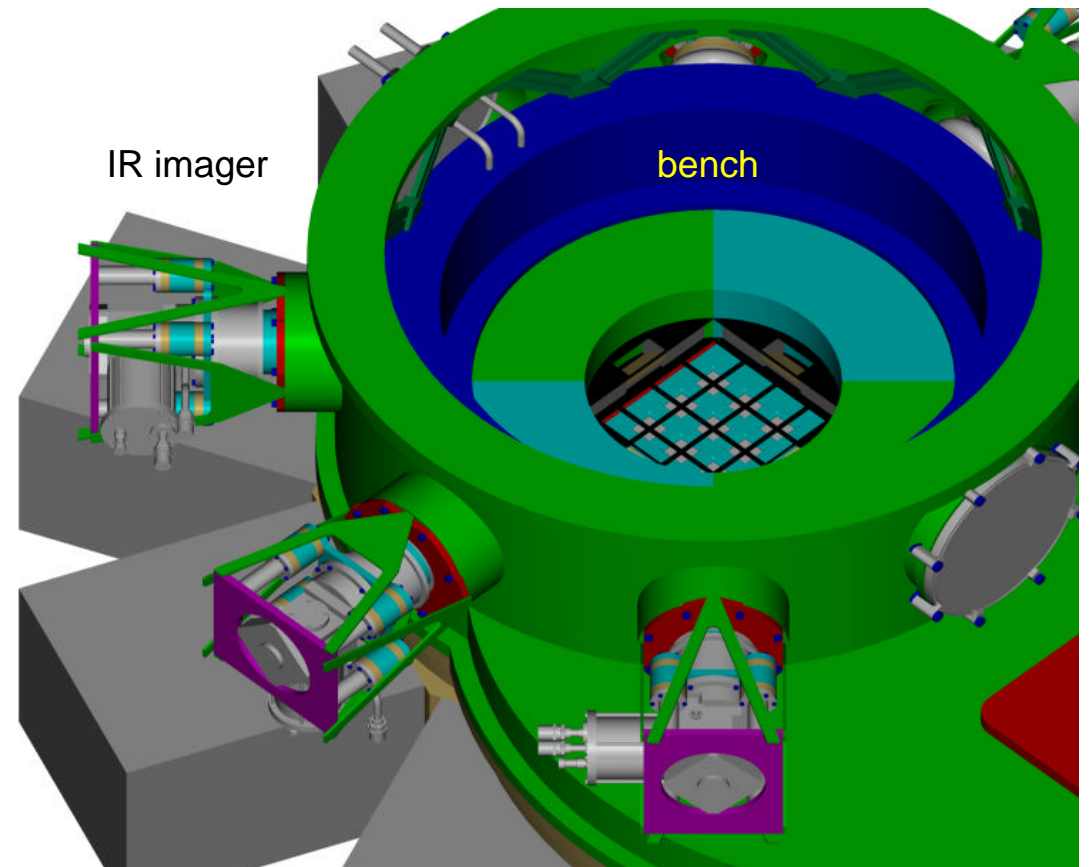
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- Large format mosaics of 4x4x2Kx2K IR arrays each having  $\geq 32$  parallel video outputs (VISTA)
- Subpixel sampling by digital filter in EPLD
- Reference unit cells for true differential signal chain
- Interpolated digital clamp of reference unit cell
- Embedded real time pre-processing in high level programming language such as IDL ( cosmic ray rejection)
- Guiding on science frame
- diagnostics , self test of acquisition system, on line help
- Audio DAC's to control clock slopes ?
- Put detector front end in ASIC close to focal plane

# Large Format : VISTA



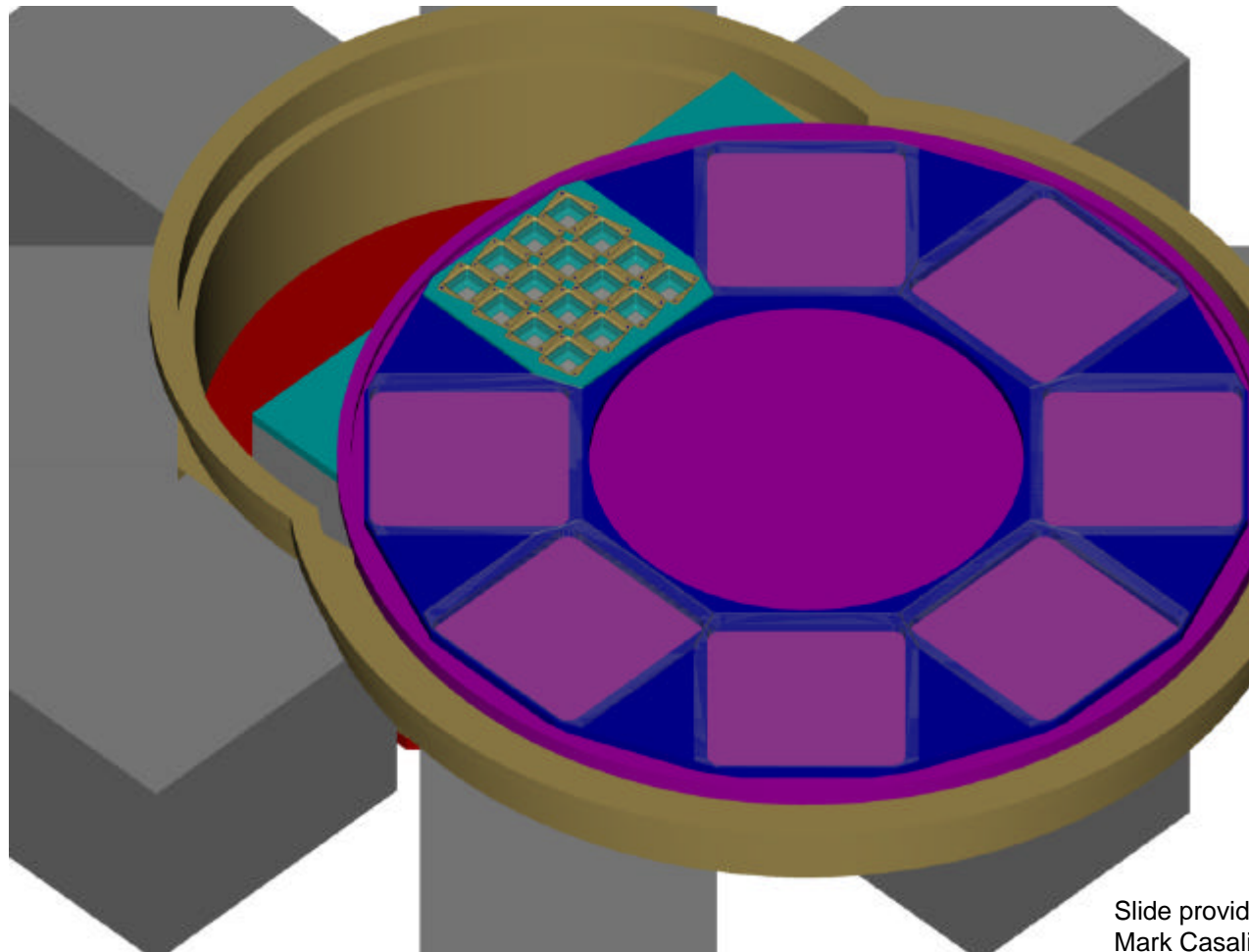
- 4 meter telescope, field 1.67 deg
- mosaics of 4x4x2Kx2K IR arrays



Slide provided by  
Mark Casali / UKATC

# Large Format: VISTA

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- eight filter units, each containing sixteen filters, each 54mm square
- with HAWAII2 mux use four quadrants for tracking, guiding and wave-front sensing
- with HAWAII2-RG use guiding feature built into detector

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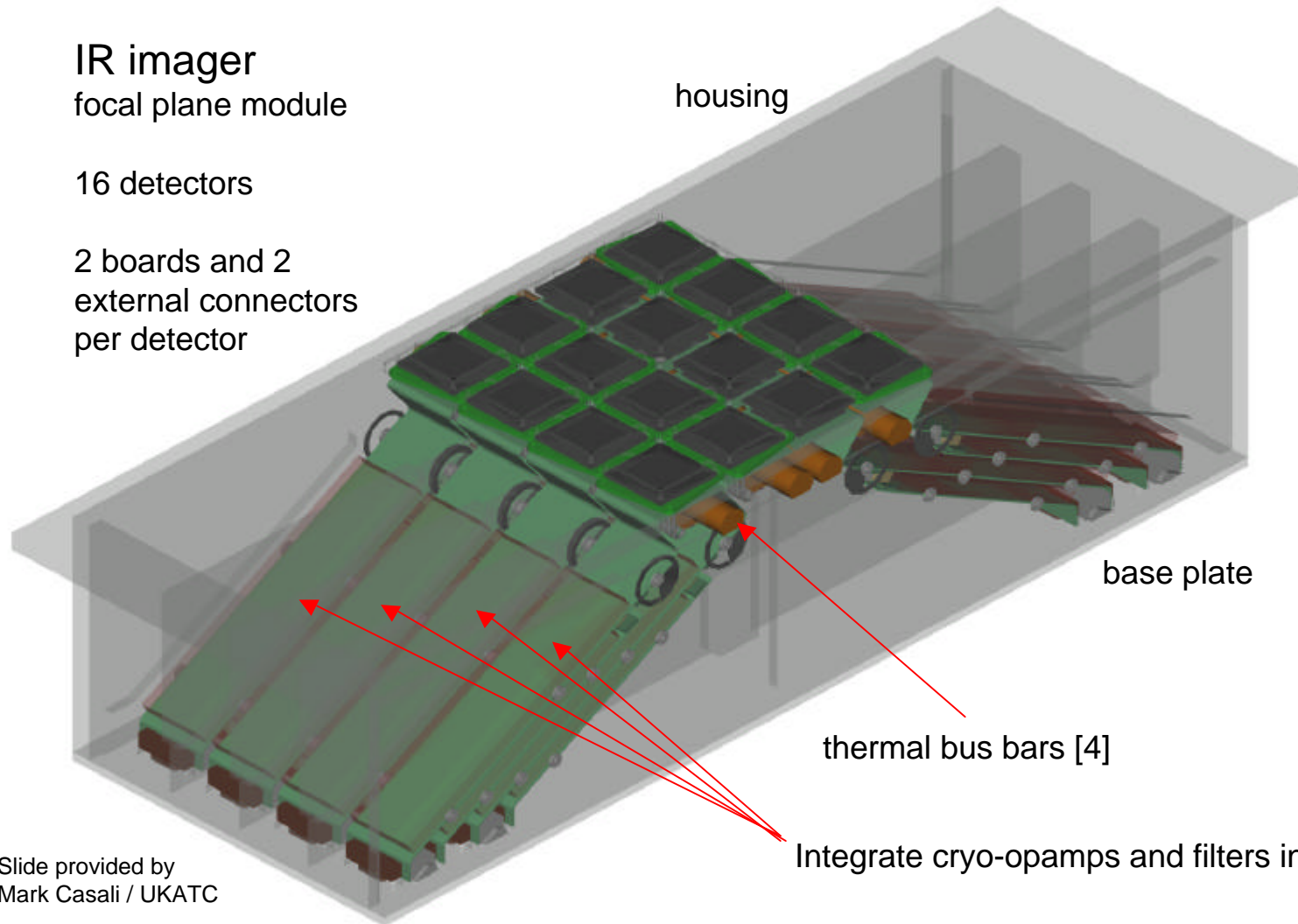
# VISTA focal plane

IR imager  
focal plane module

16 detectors

2 boards and 2  
external connectors  
per detector

housing



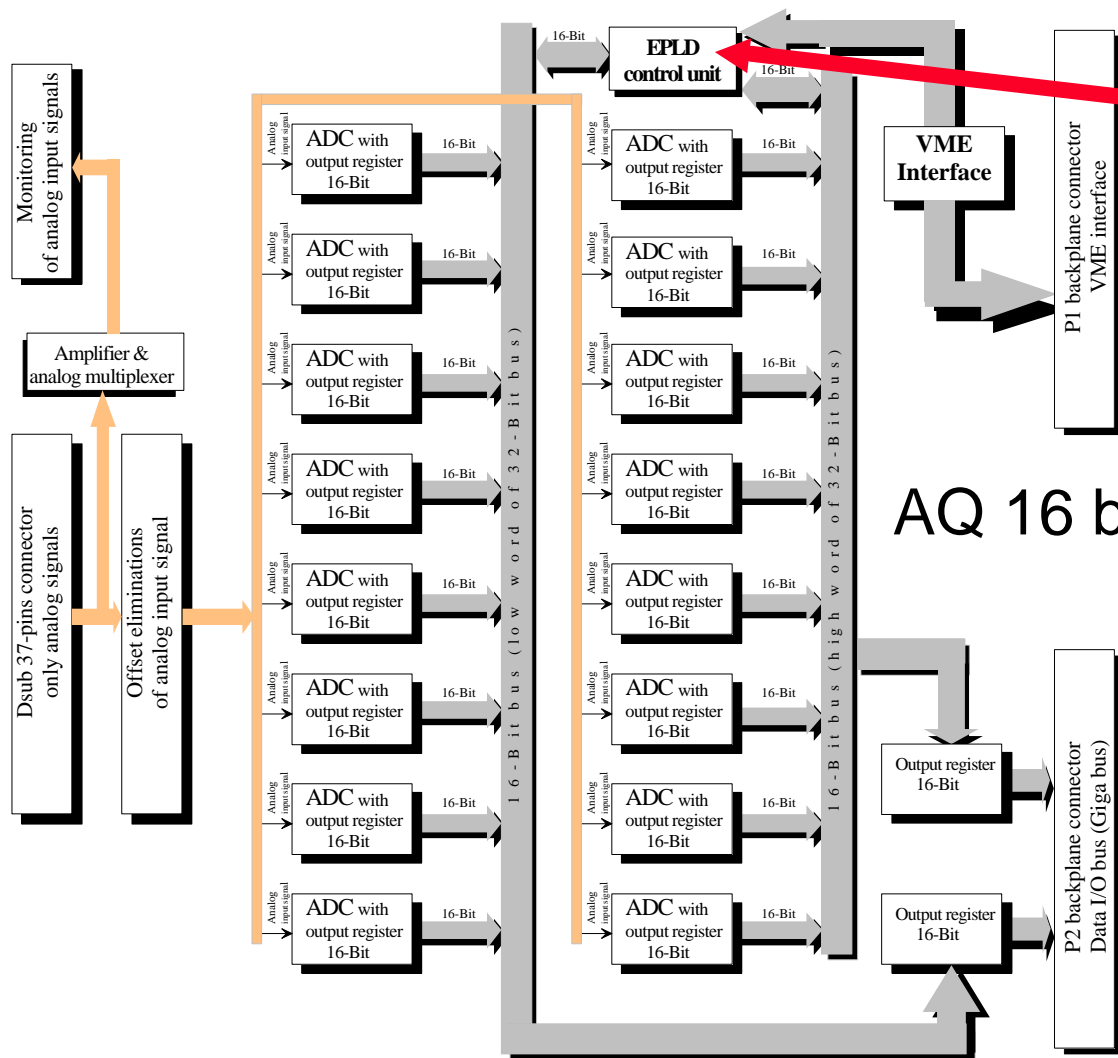
base plate

thermal bus bars [4]

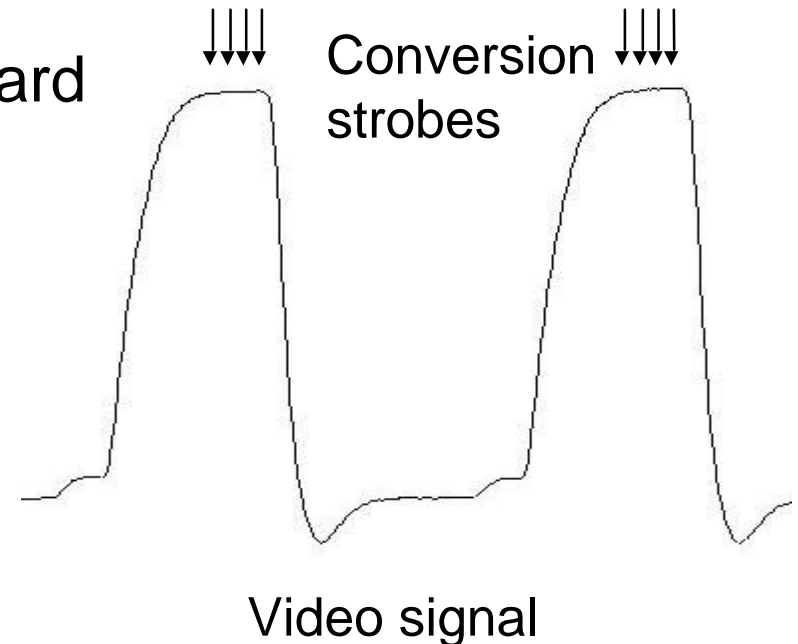
Integrate cryo-opamps and filters in flex board

- 16 2Kx2K arrays each having 32 parallel outputs
- 512 channel IRACE system
- readout time 1.3 s

# Digital filter in EPLD of AQ16

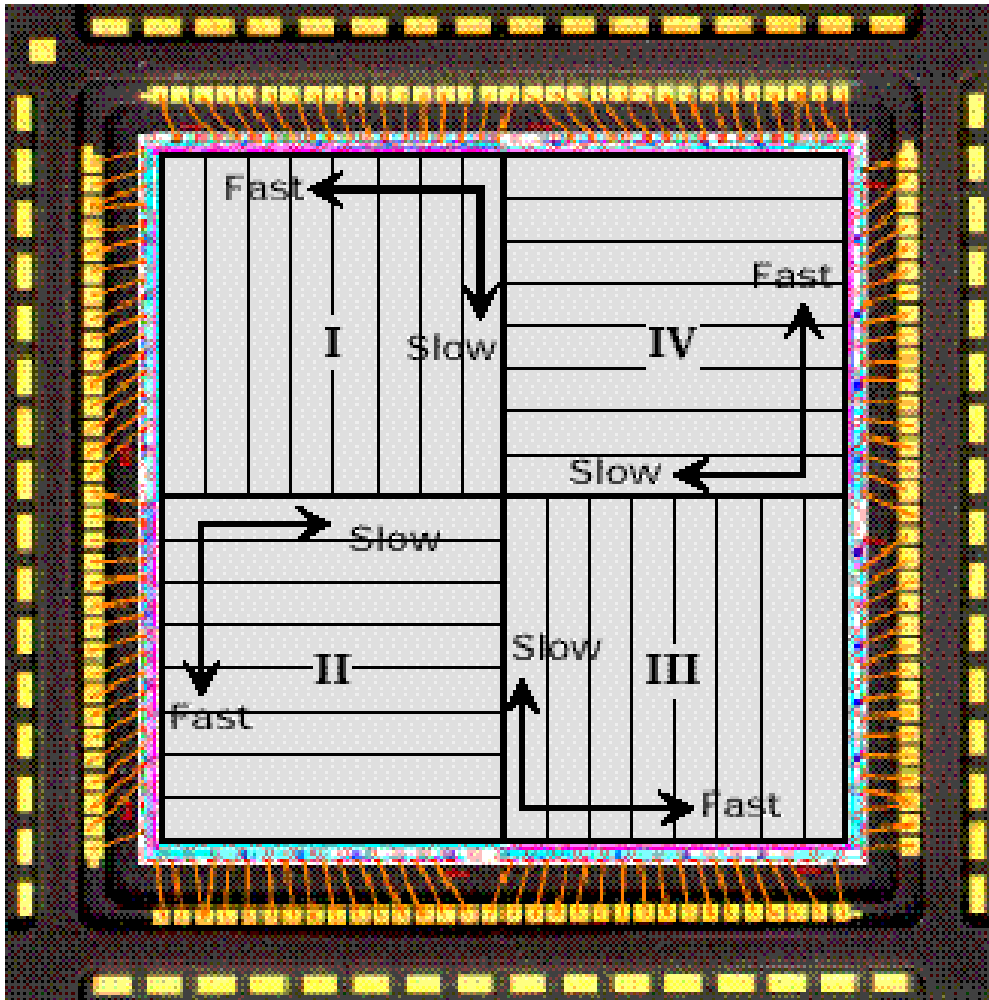


- process subpixel samples in EPLD
- parallel processing 1 processor /AQ board
- digital filter of video signal

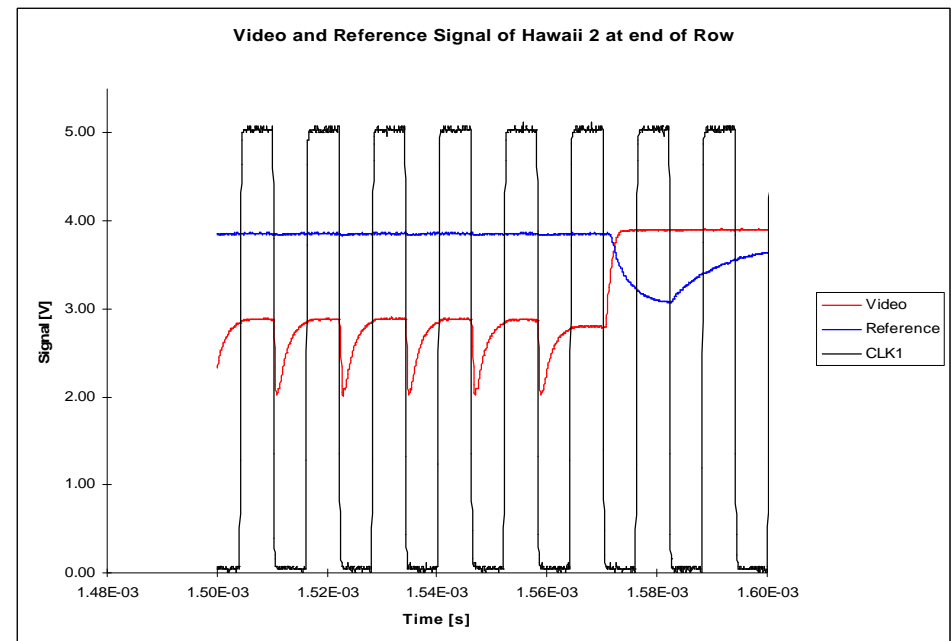


AQ 16 board

# Reference unit cell (Hawaii2)

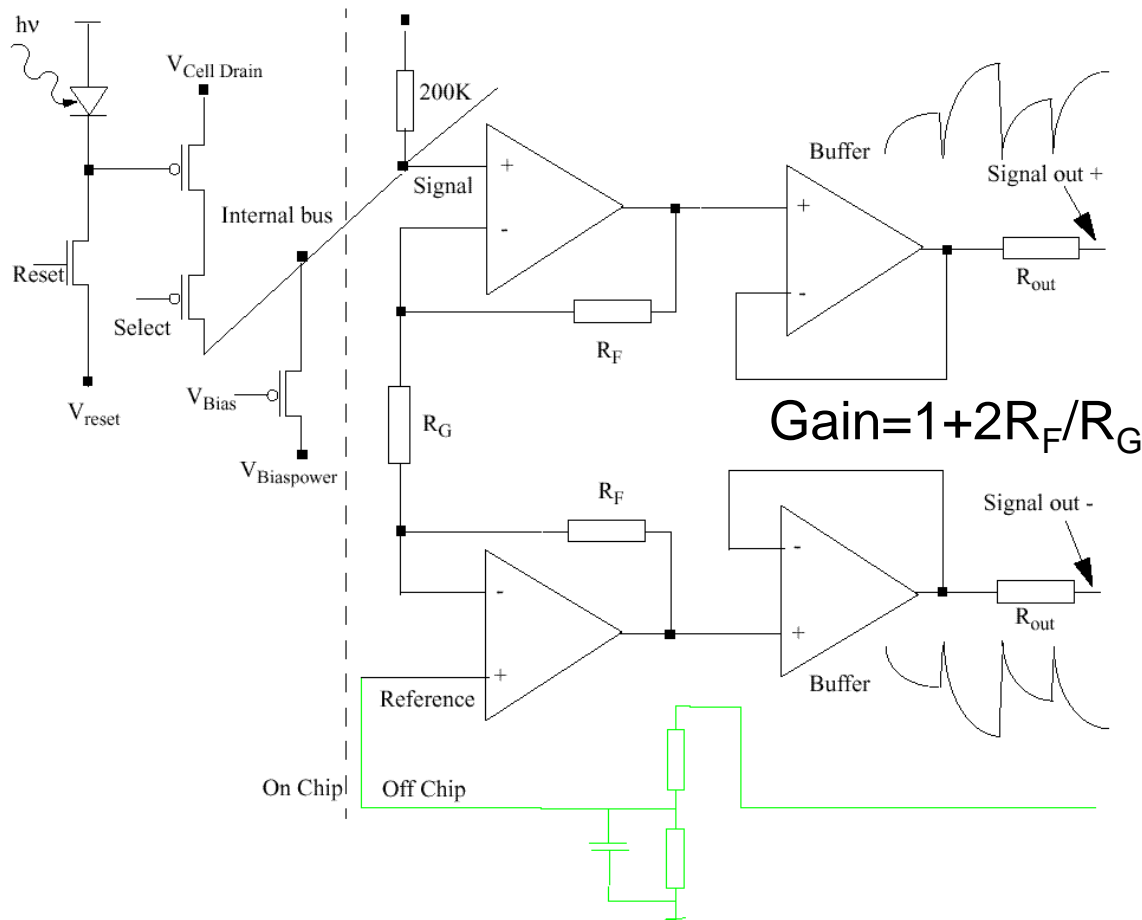


- 8 video outputs / quadrant
- 128 pixels in fast direction
- reference pixel for thermal drift is 129<sup>th</sup> pixel on 9<sup>th</sup> output
- 36 channels / detector required



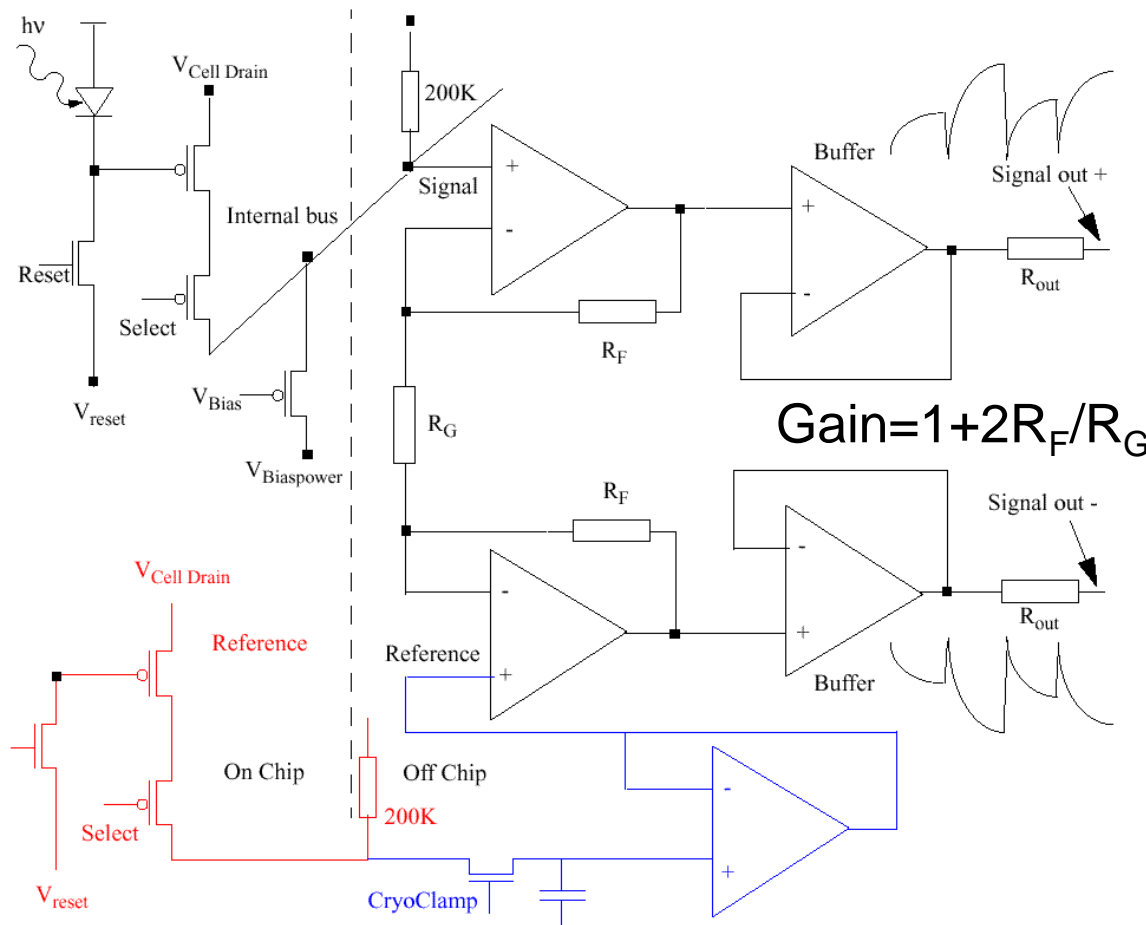


# External reference with Hawaii1



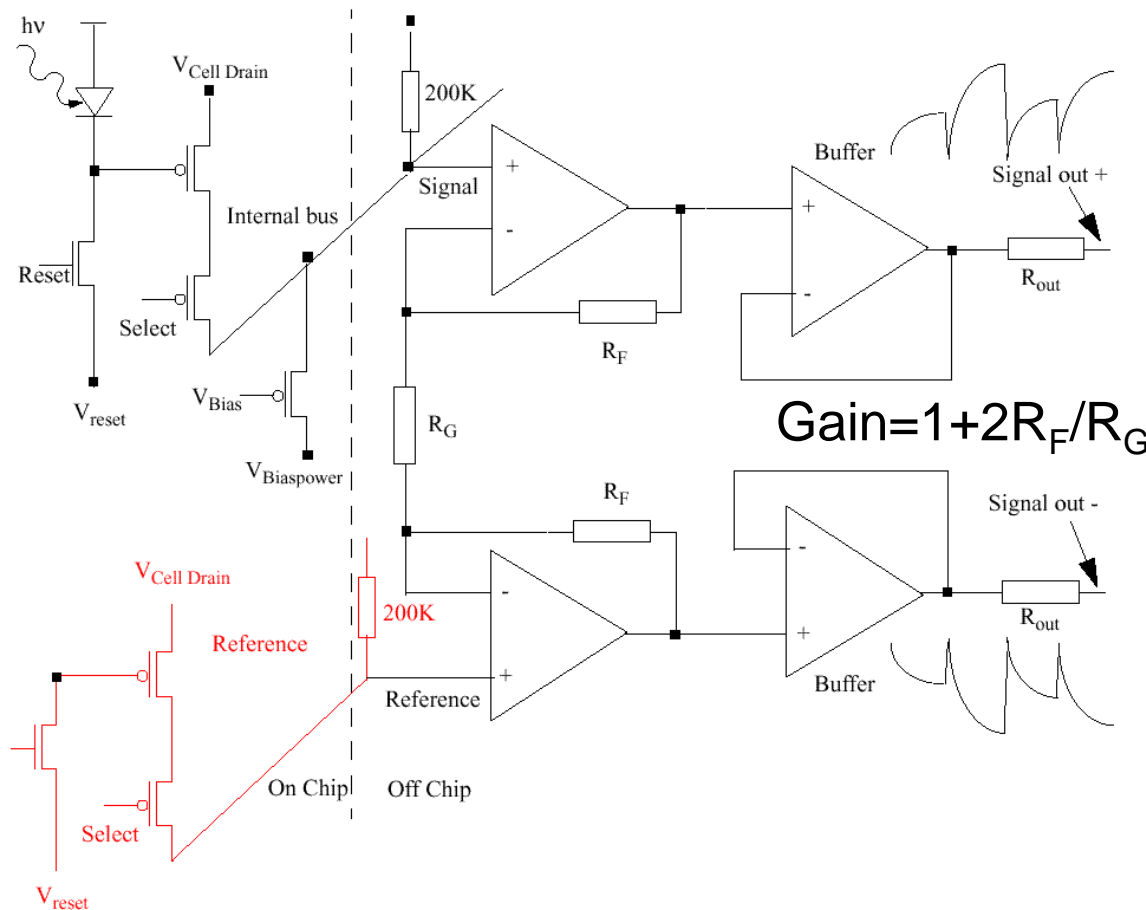
- No reference output provided on chip
- external reference has to be used

# On chip reference cell of Hawaii2



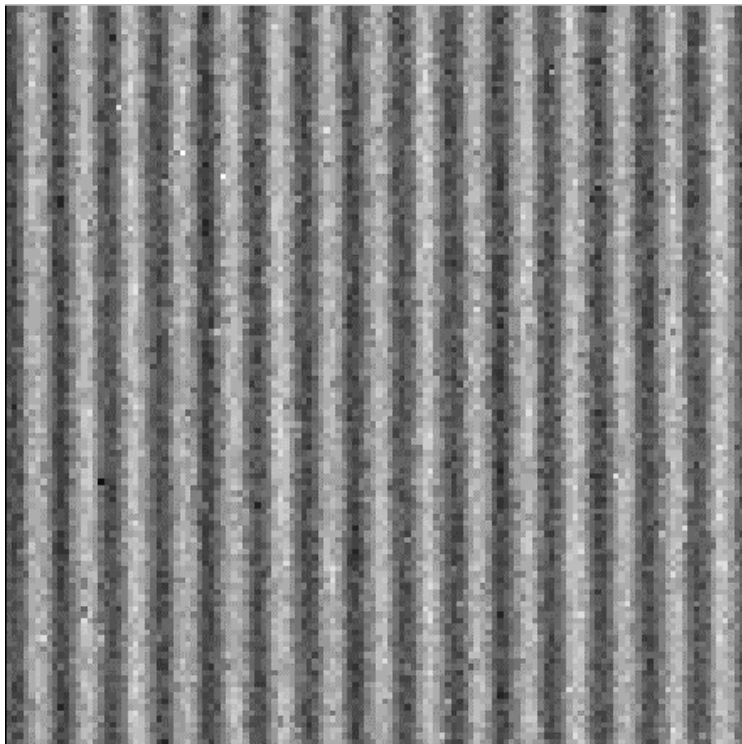
- reference pixel for thermal drift compensation and suppression of pickup is 129<sup>th</sup> pixel on 9<sup>th</sup> output
- reference pixel not available while reading pixels
- clock first to reference pixel
- clamp reference pixel with **cryogenic sample and hold**
- read pixels of row with clamped voltage at reference input of cryo-opamp

# On chip reference cell of Hawaii2-RG

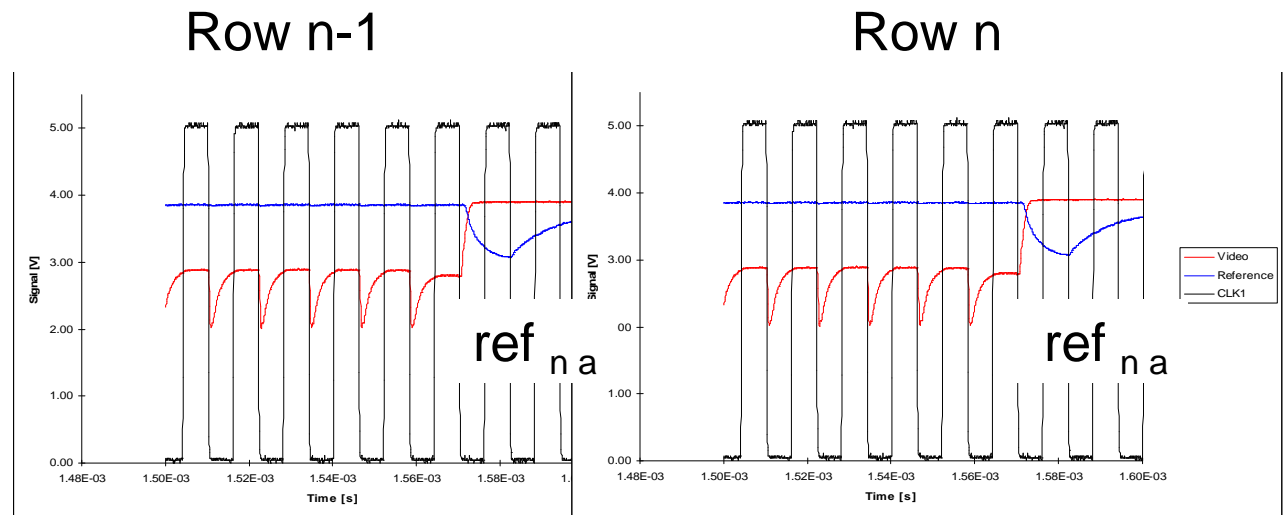


- reference pixel for thermal drift compensation and suppression of pickup on 9<sup>th</sup> output, but
- **reference pixel available while reading pixels**
- read pixels of row with voltage of reference cell at reference input of cryo-opamp
- best suppression of noise pickup

# 50 Hz pickup of Hawaii2 without clamp of reference

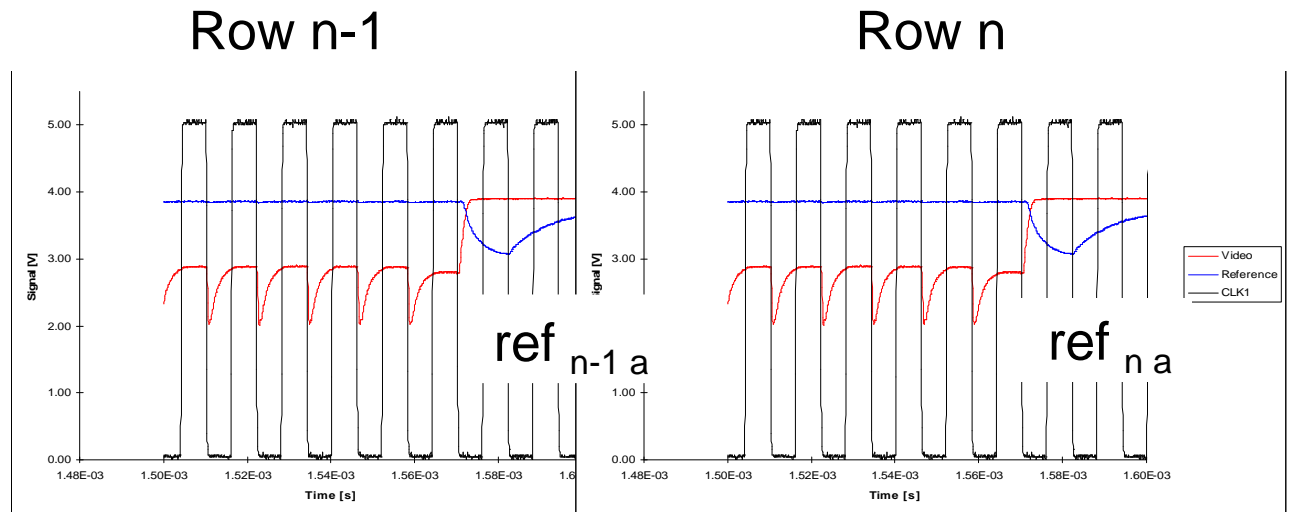
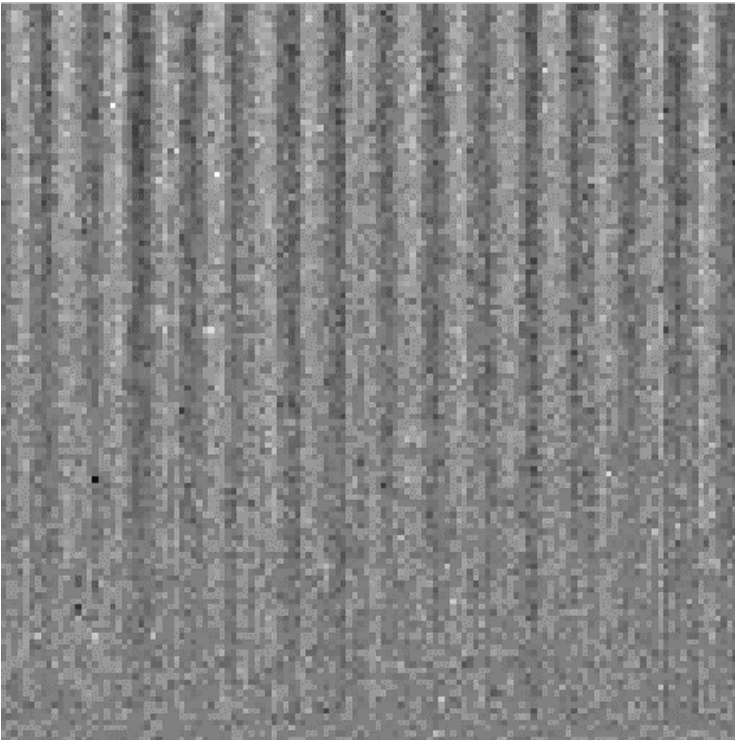


- multiple sampling reduces readout noise and requires lower pickup noise
- Clock each row once and read once
- reference output ( 129<sup>th</sup> pixel on 9<sup>th</sup> output ) is not used
- 32 channels / detector required

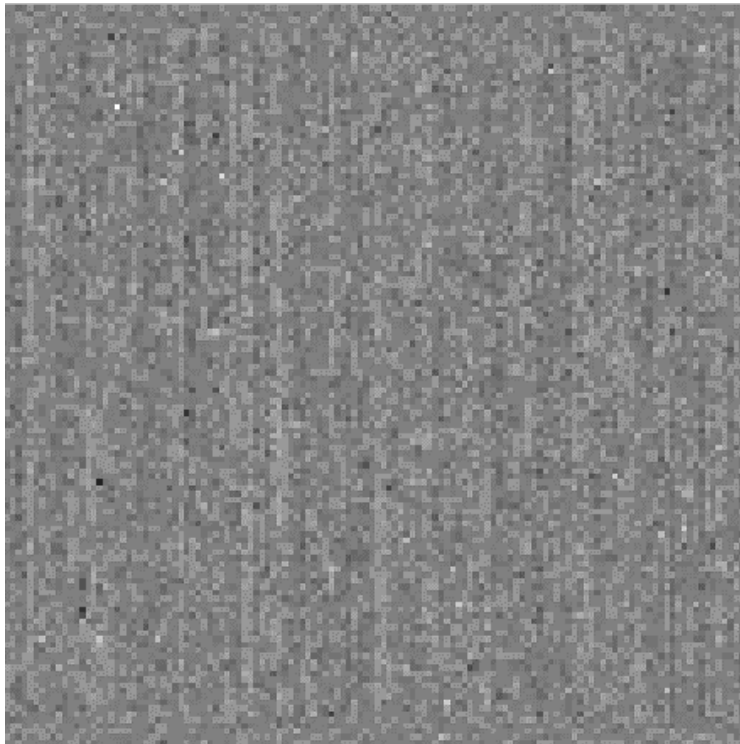


# 50 Hz pickup of Hawaii2 with digital clamp

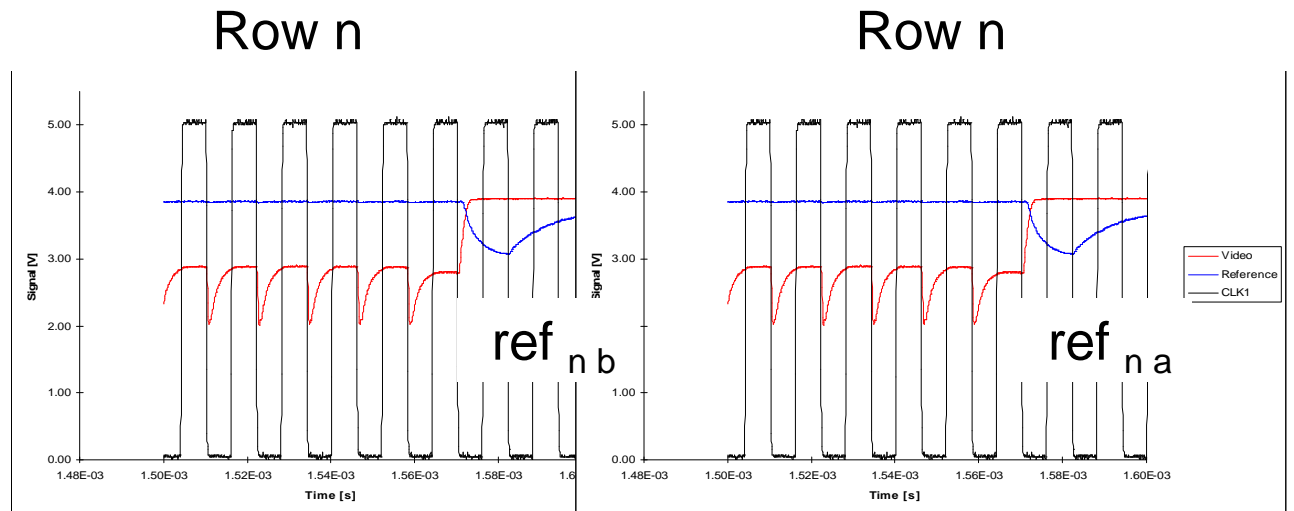
- Clock each row once and read once
  - read reference after reading row and subtract reference signal for each pixel
- $$\text{sig}_{nk} = \text{pix}_{nk} - \text{ref}_{na}$$
- 36 channels / detector required



# 50 Hz pickup of Hawaii2 with interpolated digital clamp



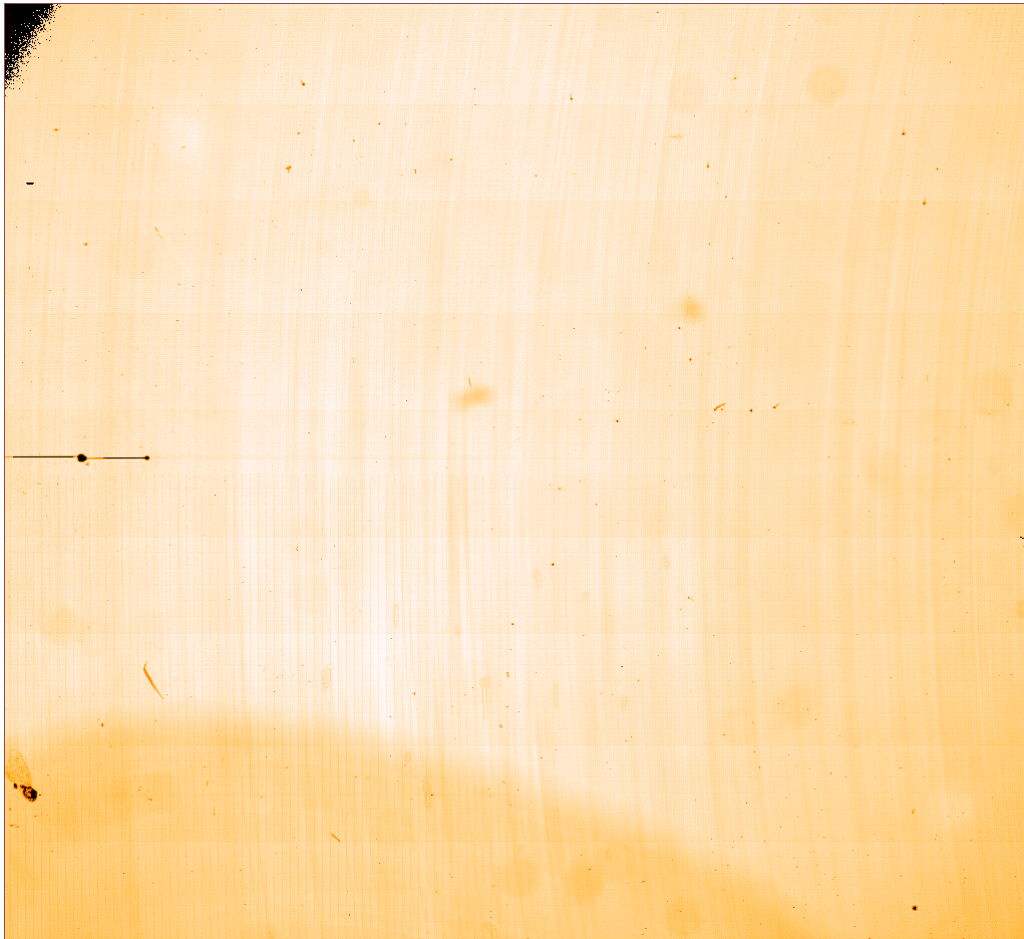
- Clock each row twice and read once
- read reference before and after reading row and interpolate reference signal for each pixel
$$\text{sig}_{nk} = \text{pix}_{nk} - (\text{ref}_{nb} - (\text{ref}_{nb} - \text{ref}_{na}) * (k-1)/127)$$
- 36 channels / detector required
- reference allows rejection of low frequency pickup and compensation of thermal drifts



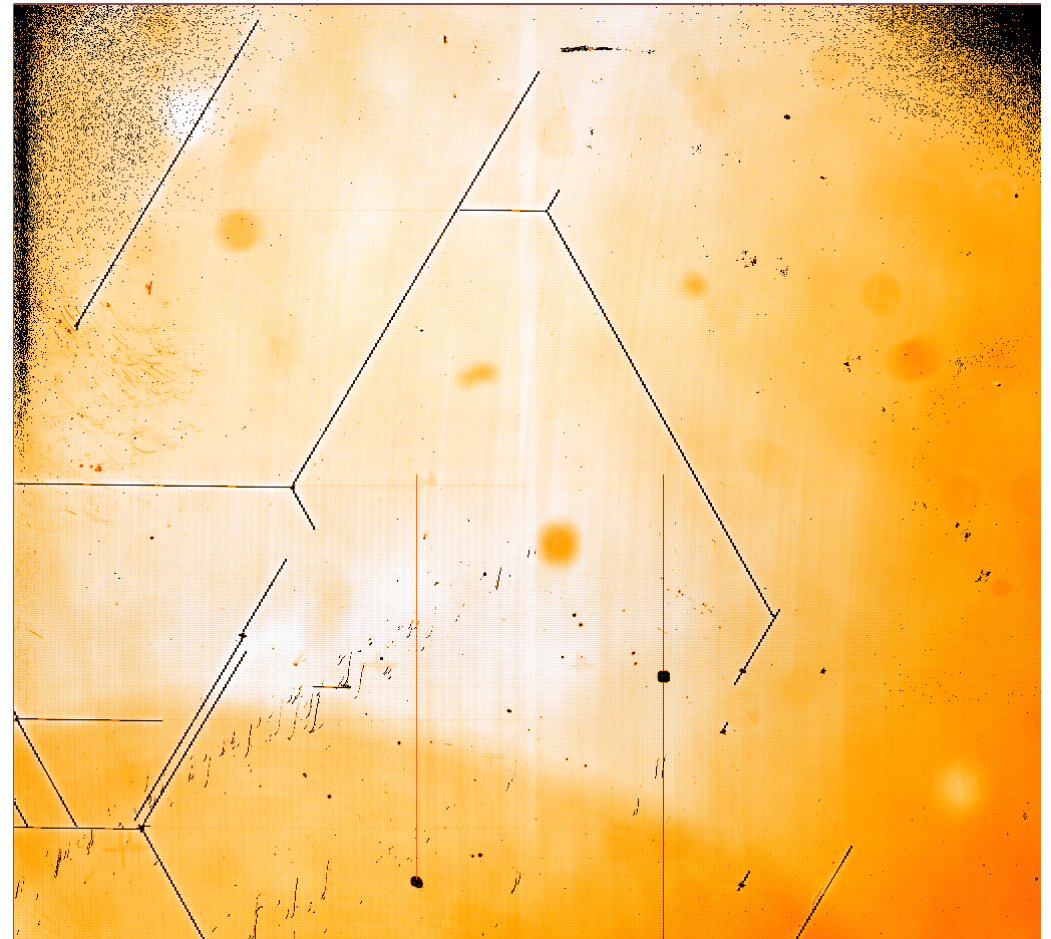


# Aladdin #3 & #4

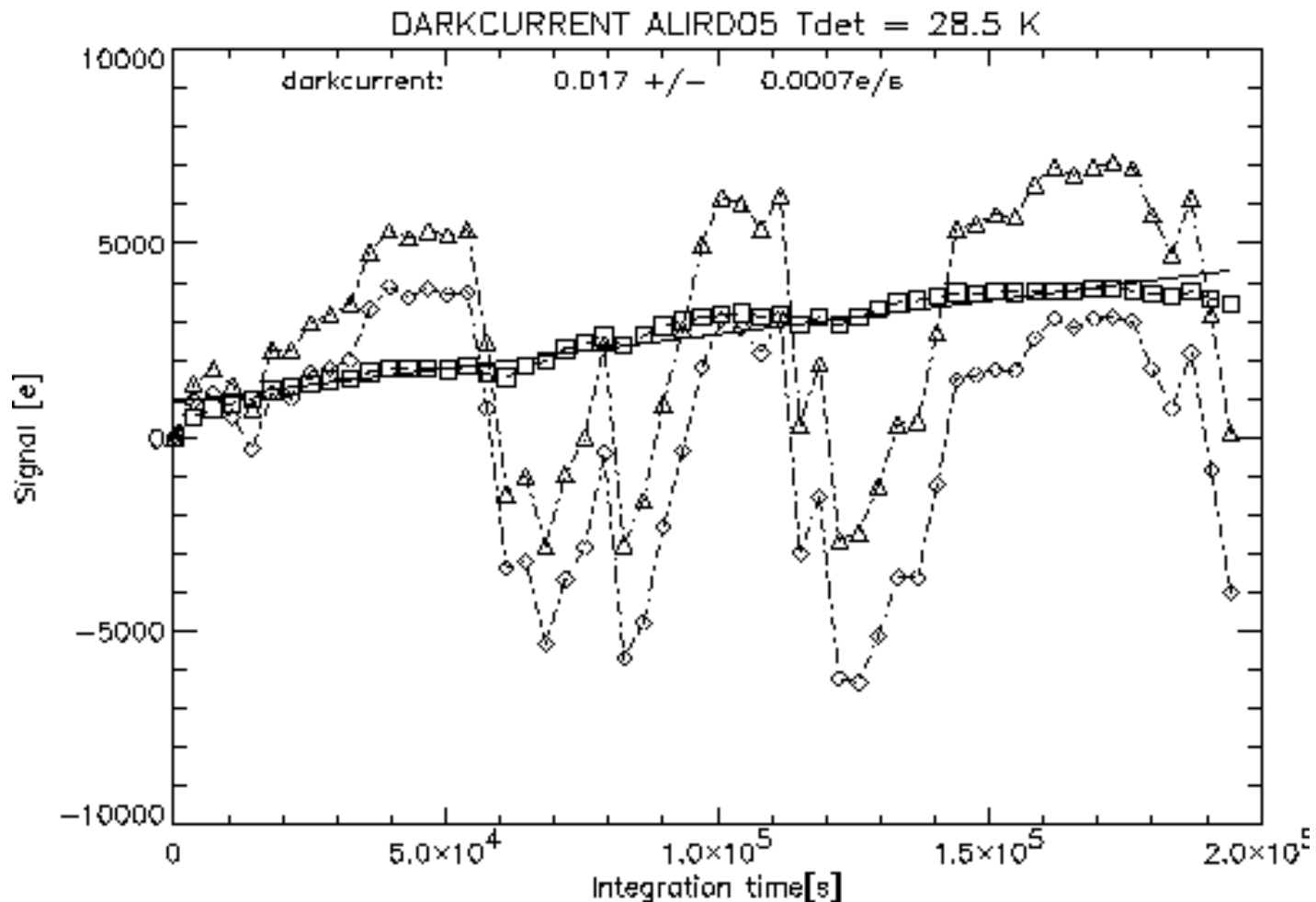
#5: QE H = 0.82 (ISAAC)  
darkcurrent = 0.017 e/s



#4: QE H = 0.86  
darkcurrent = 0.004 e/s



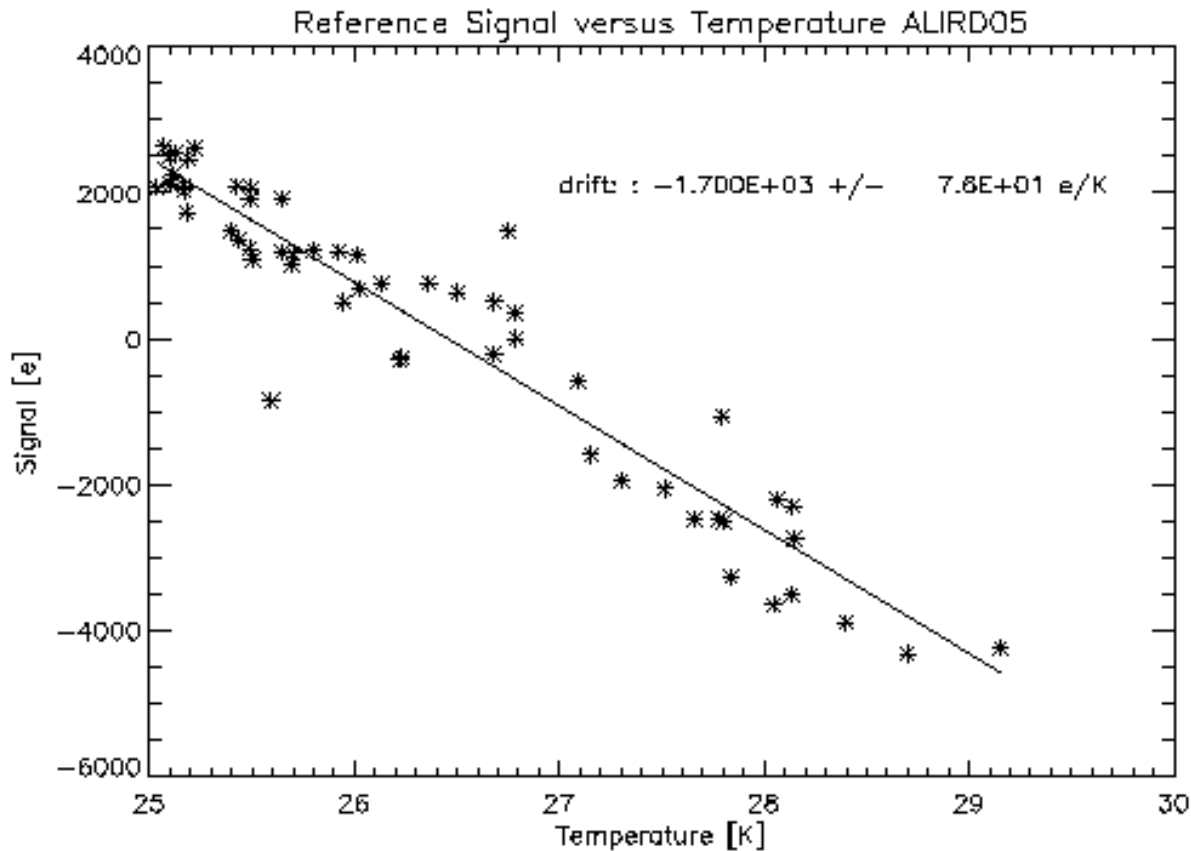
# Monitor temperature drift using dead pixels



- Triangles: measured integration ramp
- Diamonds: dead pixels
- Open In bump bonds are used to monitor drifts
- Squares: drift corrected integration ramp
- darkcurrent at 28.5 K:  
**0.017 e/s/pixel**



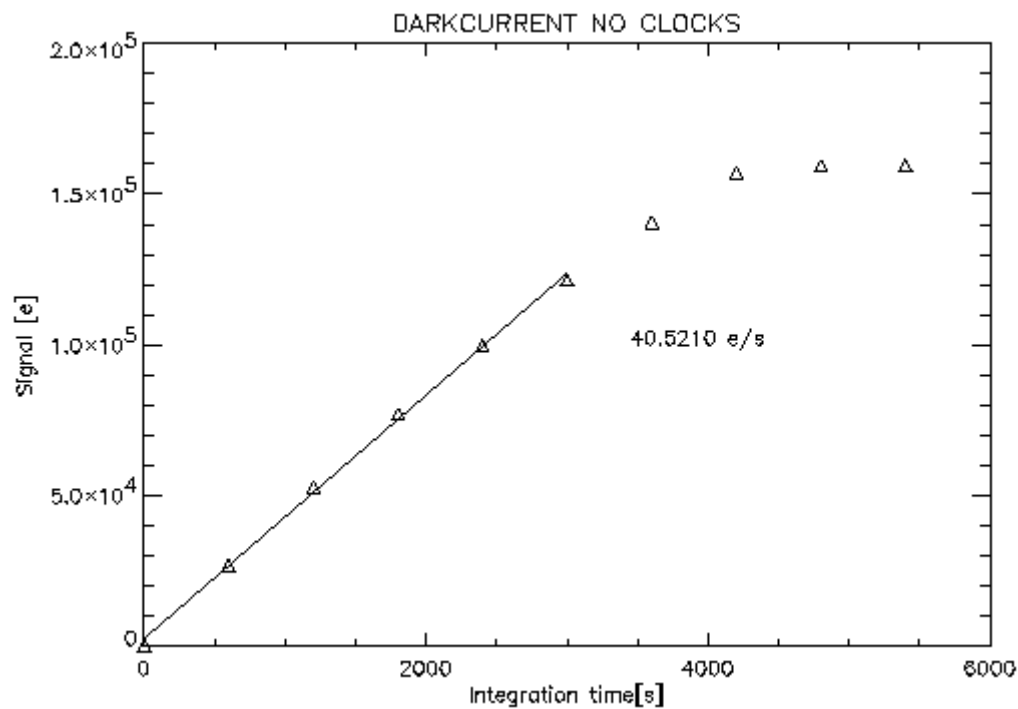
# Temperature drift of dead pixels in Aladdin array



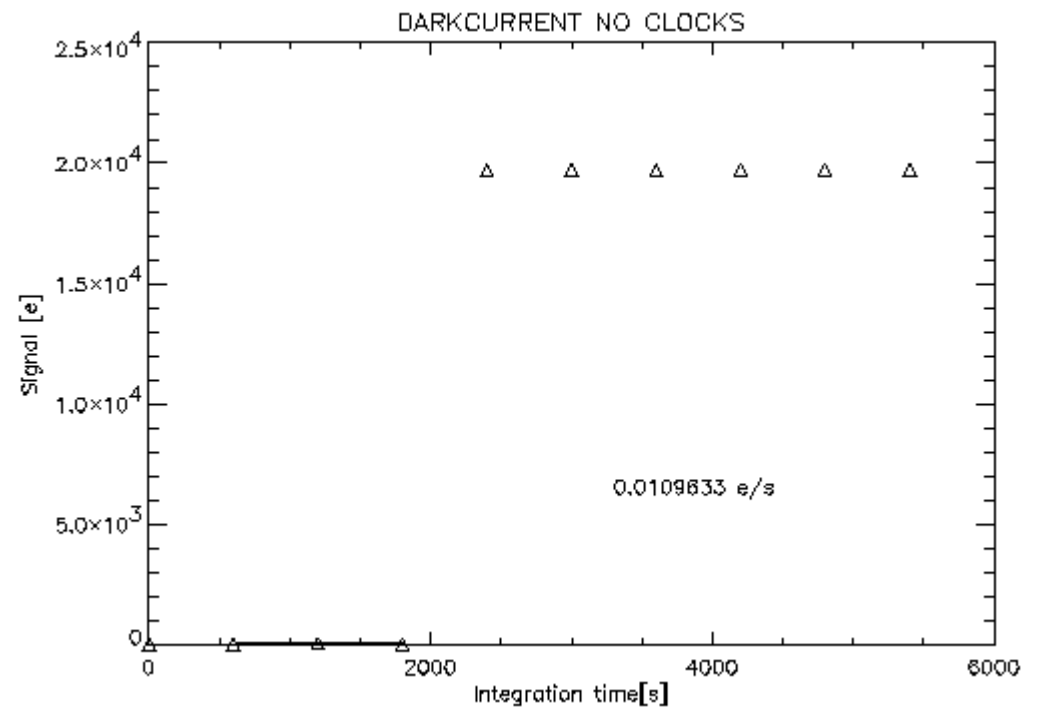
- Temperature drift: 1700 electrons / K
- required temperature stability of array: 6 mK
- temperature drift will be monitored with reference unit cell of Hawaii-2RG

# Elimination of cosmic rays by multiple sampling

Warm pixels  
continuous integration ramp



Cosmic rays create charge burst  
and step in integration ramp

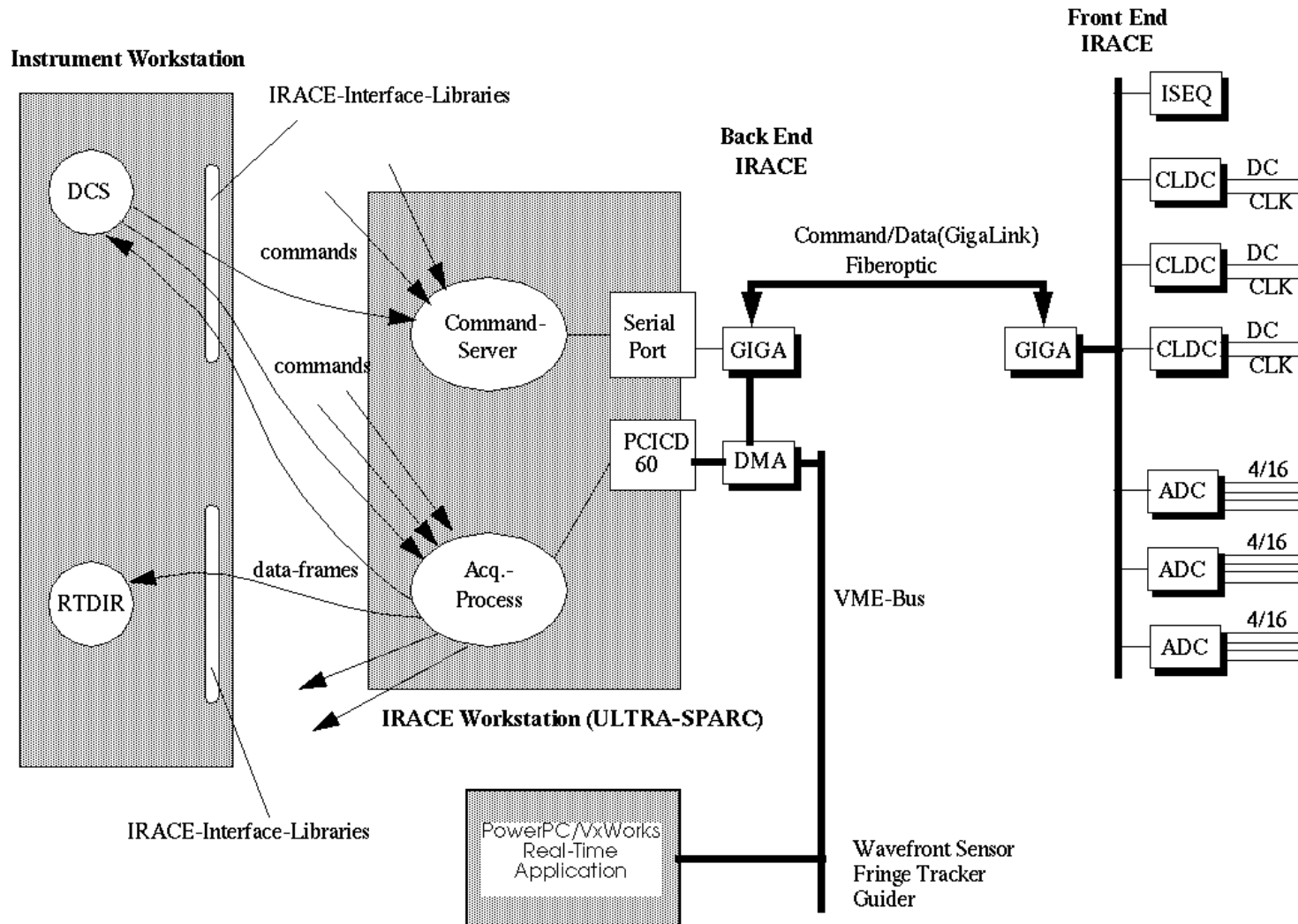


# Cosmic Ray Correction

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- At Paranal (altitude 2600 m)  
2038 pixels / hour are hit
- maximum charge injected by cosmic ray  
~ 2E4 electrons < full well (1E5 electrons)
- correction by multiple nondestructive  
sampling and real time processing in IRACE  
number cruncher using idl routines
- at present more warm pixels than cosmics,  
but arrays may improve (MBE)

# Parallel Processing of Detector Data in IRACE



- mark digitized video signal with different headers
- process video signal in IRACE number cruncher (ULTRA-SPARC) for science frame
- process subwindow in PowerPC on VME bus for real time applications

# Conclusion

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- Mosaics of 2Kx2K arrays are coming
- need 5 MHz 16 bit ADC's and 3 Gbaud fiber link
- use of unit cell reference for true differential signal chain
- more real time image processing with faster processors
- General purpose ASIC needed for ground based applications